

# **$V_p/V_s$ structure and $P_n$ anisotropy across the Louisville Ridge, seaward of the Tonga-Kermadec Trench**

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## **Abstract**

The Pacific Plate within the collision zone between the Louisville Ridge and the Tonga-Kermadec Trench was formed at the Osborn Trough, a paleo spreading center that became inactive during the Cretaceous. In this region, the trench shallows from a depth of 8-11 km to ~6 km below sea surface, while the outer rise topography is obscured by Louisville seamounts that rise 4-5 km above the adjacent seafloor. We derive 2-D P-wave ( $V_p$ ) and S-wave ( $V_s$ ) velocity-depth models along a wide-angle seismic profile oriented sub-parallel to the trench axis, intersecting the 27.6°S seamount. The seismic profile is located in the down-going Pacific Plate eastwards from the trench axis (~100 km distant at the south end and ~150 km at the north end), where bending-related faulting is limited or absent. Using the derived P- and S-wave velocity-depth models we calculate the corresponding  $V_p/V_s$  ratio model which shows values of 1.7-1.85 throughout the oceanic crust either side of the Louisville Ridge where it is unaffected by magmatism associated with its formation. This range of observations lies within those documented by laboratory measurements on basalt, diabase, and gabbro. Conversely, in the vicinity of the summit of 27.6°S seamount, the relatively elevated  $V_p/V_s$  (~1.9) ratio observed can be attributed to water-saturated cracks within the shallow sub-seabed section of the intrusive core. Beneath the seamount the uppermost mantle has a  $V_p$  ranging from 8.0 to 8.9 km/s. Comparing our P-wave model with a pre-existing model running sub-perpendicularly along the Louisville Ridge axis, we observe an anisotropy of up to ~6% at a depth of 3-4 km below the Moho. The predominant

30 orientation of the faster axis follows the direction of paleo spreading flow when the plate was  
31 formed at the Osbourn Trough.

32

### 33 **Key Points:**

- 34 • We obtain a  $V_p/V_s$  ratio of  $\sim 1.9$  near the summit of the 27.6°S seamount.
- 35 • We find mean  $V_p/V_s$  ratios of 1.7-1.85 for the oceanic crust adjacent to the Louisville  
36 Ridge.
- 37 • Seismic anisotropy of up to 6% is observed in the upper 15 km of the mantle.

### 38 **1. Introduction**

39 Oceanic hotspot magmatism results from the dynamic interaction between an overlying  
40 oceanic tectonic plate and an ascending, thermally buoyant mantle plume originating in the  
41 Earth's deeper mantle. Ridges that form along oceanic hotspot tracks far from a spreading  
42 center (off-ridge setting) are generally characterized by a rough topography composed of  
43 isolated volcanic edifices typically constructed on old ( $>20$  Ma) and rigid oceanic lithosphere  
44 (Orellana-Rovirosa and Richards, 2017). In contrast, hotspot track ridges that form proximal  
45 to a spreading center (near-ridge setting) are generally characterized by smooth topography,  
46 with intrusive magmatism in the young ( $<10$  Ma), weak oceanic lithosphere (Pollack et al.,  
47 1981; Orellana-Rovirosa and Richards, 2017; Contreras-Reyes et al., 2022).

48 Velocity-depth models obtained from wide-angle seismic data show how the structure  
49 of the crust and upper mantle is affected by hotspot magmatism (e.g., Richards et al., 2013).  
50 Examples of off-ridge hotspot trails such as the Juan Fernández Ridge (Kopp et al., 2004)  
51 and Hawaii (McGregor et al., 2023; Dunn et al., 2024) show a generally normal-thickness  
52 oceanic crust deformed (deflected) by the weight of the volcanic edifice (Manríquez et al.,  
53 2014; Watts et al., 2021; McGregor et al., 2023). In contrast, on-ridge hotspot trails such as  
54 the Cocos (Walther 2003), Carnegie (Sallares et al., 2003), and Nazca Ridges (Hampel et al.,  
55 2004; Contreras-Reyes et al., 2022) have an anomalously thick crust ( $>10$  km) but a similar  
56 velocity-depth structure to that of standard oceanic crust (Christeson et al., 2019; Grevenmeyer

57 et al., 2018a; White et al., 1992). In addition, some off-ridge hotspot trails are characterized  
58 by oceanic crust that is highly intruded, with intrusions forming a dense, high P-wave  
59 velocity ( $>6.0$  km/s) core; for example, the Great Meteor seamount (Weigel and Grevemeyer,  
60 1999), Hawaii (McGregor et al., 2023; Dunn et al., 2024), Tenerife (Canales et al., 2000),  
61 and the Louisville Ridge (Contreras-Reyes et al., 2010; Robinson et al., 2018).

62 In this paper, we investigate the lithological structure of the Louisville Ridge (LR),  
63 which formed in an off-ridge setting more than 500 km away from a spreading center axis.  
64 Two-dimensional P-wave ( $V_p$ ) velocity-depth models reveal that LR seamounts contain  
65 intrusive cores characterized by a  $V_p \geq 6.0$  km/s (Contreras-Reyes et al., 2010; Stratford et  
66 al., 2015; Funnell et al., 2017; Robinson et al., 2018), in contrast to the adjoining Pacific  
67 oceanic crust which is unaffected by hotspot-related magmatic activity. Contreras-Reyes et  
68 al. (2010) interpreted these high  $V_p$  regions as comprising basaltic and intrusive rocks that  
69 form a symmetrical, semi-conical core, implying that the seamount has undergone growth  
70 both vertically and laterally. To gain further insight into the seismic structure of the  $27.6^\circ\text{S}$   
71 seamount, and what its lithological implications are, we model the S-wave velocity ( $V_s$ )  
72 structure in the vicinity of the summit and surroundings areas. We derive both 2-D  $V_p$  and  $V_s$   
73 tomographic inversion models along one wide-angle seismic profile that intersects the  $27.6^\circ\text{S}$   
74 seamount and follows a sub-perpendicular trend across the LR (profile P03 shown in Fig. 1).  
75 The sediment thickness along-profile is constrained by a coincident multichannel seismic  
76 profile (Fig. 2). We further use these models to calculate a  $V_p/V_s$  ratio model, which provides  
77 valuable constraint for deciphering the lithological composition of the oceanic crust and  
78 upper mantle (Christensen 1996; Carlson & Miller, 2003; Grevemeyer et al., 2018b; Li et al.,  
79 2022; Contreras-Reyes et al., 2022, 2023).

80 Hotspot magmatism can also result in upper mantle anisotropy. Initially, seismic  
81 anisotropy is governed by the lattice preferred orientation (LPO) of olivine acquired during  
82 the formation of the oceanic lithosphere at mid-ocean ridges (Nicolas and Christensen, 1987).  
83 Compressional waves exhibit their highest speed (fast direction) along the  $a$ -axis of olivine,  
84 typically aligning with the direction of plate motion/mantle flow (Hess, 1964; Skemer and  
85 Hansen, 2016). However, some refraction experiments that have measured upper mantle P-

86 wave velocity away from the spreading center indicate a rotation relative to a paleo spreading  
87 direction. This has been attributed to various processes, including mantle serpentinization  
88 induced by plate bending (Contreras-Reyes et al., 2008; Mishra and Gordon, 2016), the  
89 presence of 3-D mantle flow patterns instead of exclusively 2-D corner flow in the plate  
90 direction (VanderBeek et al., 2016; Toomey et al., 2007), upwelling of a hotspot mantle  
91 plume (Fontaine et al., 2005), and a complex configuration of cracks (Mark et al., 2019).

92 We present 2-D  $V_p$  and  $V_s$  models for the oceanic crust formed >90 Ma at the Osbourn  
93 Trough, an extinct spreading center (Worthington et al., 2006). Here, we also model  
94 lithospheric mantle refractions ( $P_n$ ) with offsets of up to 150 km, generated with an active  
95 seismic source, which sample up to a depth of ~30 km below seafloor (*bsf*). We compare our  
96 results with previous, and orthogonal, 2-D  $V_p$  models obtained along the LR (Stratford et al.,  
97 2015; Funnell et al., 2017; Robinson et al., 2018) to further study the magnitude of seismic  
98 anisotropy within the uppermost 3-4 km of the mantle. Finally, we discuss our results in  
99 terms of hotspot magmatism and upper mantle seismic anisotropy.

100

## 101 **2. Geological Setting**

102 The Pacific Plate undergoes oblique subduction beneath the overriding Indo-  
103 Australian Plate with an approximate azimuth of N60°W (e.g., Bird, 2003) and with a high  
104 convergence rate of ~240 mm/yr (DeMets et al., 2010; Seton et al., 2012). This, combined  
105 with the LR oblique alignment relative to the trench axis, results in the southward migration  
106 of the LR-trench intersection. Lonsdale (1988) and Balance et al. (1989) estimate the rate of  
107 southward migration to be 120-180 mm/yr. This migration leads to lateral variations in the  
108 structural composition of the trench (e.g., Clift et al., 1998), forearc (Clift and MacLeod,  
109 1999; Contreras-Reyes et al., 2011), arc (England et al., 2004), and backarc (Bevis et al.,  
110 1995). In particular, the collision between the LR and the Indo-Australian Plate has  
111 accelerated subduction erosion, causing the arc-ward migration of the trench axis and the  
112 regional subsidence of the easternmost portion of the overriding plate (Ballance et al., 1989;  
113 Clift et al., 1998).

114           The LR marks a major hotspot track that stretches for over ~4,300 km, that was  
115 formed by the interplay between the Pacific Plate and the Louisville hotspot mantle plume  
116 located >500 km from the SW Pacific-Antarctic spreading center axis (Fig. 1A; Craig and  
117 Sandwell, 1988; Lonsdale, 1986; Downey et al., 2007). The LR comprises a succession of  
118 seamounts (Fig. 1C), with Osbourn being the oldest (prior to subduction) based on the dating  
119 of rock samples dredged along its length (Koppers et al., 2004, 2011). Along the Tonga-  
120 Kermadec Trench, the subducting old, cold Pacific Plate (>90 Ma) has a prominent outer  
121 rise-forebulge region, that is characterized by extensively fractured oceanic basement as a  
122 result of extensional bending-related faulting (Ballance et al., 1989; Contreras-Reyes et al.,  
123 2011; Stratford et al., 2015; Funnell et al., 2017; Robinson et al., 2018). At the point of LR-  
124 Tonga-Kermadec Trench collision, however, the trench-outer rise is further deformed by the  
125 significant elevation of Osbourn seamount; the next seamount to subduct (Ballance et al.,  
126 1989; Robinson et al., 2018). Here, the seafloor shallows by approximately 4,000-5,000 m  
127 (Fig. 1B).

128           The LR has volcanic edifices of 4,000-5,000 m in height (Fig. 1C) and is flanked by a  
129 flexural moat (Fig. 2) possibly filled by a mixture of volcanoclastic material, eroded island  
130 volcanic debris, and pelagic sediments. Contreras-Reyes et al. (2010) show a  $V_p$  of 4.0-4.5  
131 km/s at depths of 1.0-1.5 *bsf* adjacent to the 27.6°S seamount, consistent with volcanoclastic  
132 material as interpreted from the seismic models of Great Meteor seamount (Weigel and  
133 Grevemeyer, 1999), Marquesas (Caress et al., 1995; Wolfe et al., 1998), Tenerife (Watts et  
134 al., 1997), and Hawaii (Dunn et al., 2024). Most of the volcanoclastic material and eroded  
135 island volcanic debris that form part of the infill material were likely deposited during and/or  
136 a few years after seamount formation when the volcano was active (68-69 Ma). Subsequently,  
137 sedimentation is expected to be continuous, forming a thin pelagic drape (Fig. 2).

138

### 139 **3. Seismic Data and Tomographic Inversion**

#### 140 **3.1. Seismic data**

141 Seismic profile P03 is located ~150 km to the east of the Tonga-Kermadec Trench axis  
142 (Fig. 1), crossing the LR axis obliquely and running through the 26.7°S seamount. It is  
143 aligned sub-parallel to the trench axis, an orientation chosen to mitigate the structural  
144 inheritance of plate-bending in the outer rise region (Grevemeyer and Flueh, 2008). Both  
145 multichannel seismic (MCS) reflection and wide-angle (WA) seismic refraction data were  
146 acquired along profile P03.

147 For the MCS survey, a 100 m-long, 16-channel streamer was used for acquisition  
148 which aimed to image the top of the oceanic basement. Fig. 2 shows the sediment layer which  
149 includes a prominent intra-sediment reflector. This reflector is interpreted as an unconformity  
150 that records subsidence of the crust under seamount loading. For the WA survey, a total of  
151 33 GEOMAR Ocean-Bottom Hydrophones/Seismographs (OBH/S) were deployed along the  
152 ~368 km profile, with a spacing of ~9 km between each instrument (Fig. 1). The seismic  
153 source, used for both WA and MCS contemporaneous acquisition, comprised two sub-arrays,  
154 each of six G airguns, which collectively had a total volume of 84 l. This source was fired  
155 every 60 s at a pressure of 3,000 psi. The WA seismic processing procedure included clock  
156 drift correction, location on seabed determination, band pass filtering at 1/4/40/50 Hz, and  
157 deconvolution. P-wave arrivals are observed to greater than 200 km receiver-shot distance.  
158 Example record sections are shown in Figs. 3-5, with observed phases annotated.

159 Most OBSs also record prominent P-to-S mode-converted phases (Fig. 6) with high  
160 signal-to-noise ratio adjacent to the 27.6°S seamount, such as converted crustal refractions  
161 ( $PS_gP$ ), Moho reflections ( $PS_mSP$ ), and mantle refractions ( $PS_nP$ ). Unfortunately, the  $PS_gP$   
162 and  $PS_mSP$  phases beneath the 27.6°S seamount (>2 km depth below the summit) are  
163 obscured by the  $P_g$  and  $P_mP$  phases, and the P-wave seabed-sea surface multiples.

164

### 165 **3.2 $V_p$ and $V_s$ tomographic inversion approach**

166 Tomographic models were obtained using the 2-D method developed by Korenaga  
167 et al. (2000). This approach enables the combined inversion of seismic reflection and  
168 refraction phases for a single two-dimensional layer. For modelling, the thickness of the

169 water layer was determined from the swath bathymetry data, while the sediment layer  
170 thickness was determined by converting the two-way travel (TWT) time measurements of  
171 the corresponding reflectors in the MCS data into depth (Fig. 2). The time-to-depth  
172 conversion used a fixed velocity of 1700 m/s, the average P-wave velocity for pelagic  
173 sediment (Tenzer and Gladkikh, 2014). While lateral variations in this velocity might cause  
174 errors in the modelled P-S-P travel times, the range of sediment velocities expected (1.6-2.0  
175 km/s; Hamilton, 1976; Fulthorpe et al., 1989; Tenzer and Gladkikh, 2014) around the  
176 Louisville Ridge results in a maximum error of 20 ms. This is well within the observed P-S-  
177 P phase picking error of 70-90 ms. Therefore, using a fixed velocity for depth determination  
178 is appropriate given the overall travel time pick error.

179 In the  $V_p$  inversion, we used the  $P_g$  picks and initial model of Contreras-Reyes  
180 et al. (2010). Here, we further incorporate  $P_n$  picks (>100 km receiver-shot offset; Figs. 3-5).  
181 Fig. 7A shows the resulting 2-D  $V_p$  inversion model in which the upper mantle is constrained  
182 to depths of ~30 km.

183 For the  $V_s$  inversion, we used the PSP phase picks, where that arrival is interpreted to  
184 originate at the interface between the sedimentary layer and crystalline basement away from  
185 the 27.6°S seamount, and at the bottom of the pelagic sediment layer near the seamount (Figs.  
186 2 and 6), where the acoustic impedance ( $Z = \rho V$ ) contrast is most likely largest (Zoeppritz  
187 1919; Trummer, 2002). The latter is supported by many other studies such as Au and Clowes  
188 (1984), Contreras-Reyes et al. (2008, 2022, 2023), Latta and Dunn (2020), Li et al. (2021),  
189 Grevemeyer et al. (2018b), Guo et al. (2023), Spudich and Orcutt (1980), Trummer (2002)  
190 and White and Stephen (1980); see reference seismic velocity and density values in Fig. 6.  
191 Near the seamount's flanks (where the moat-depression is filled by volcanoclastic material),  
192 the most likely seismic interface for P-to S-wave conversion is the volcanic clastic  
193 material/pelagic sediment interface (Fig.6).

194 Seismic studies conducted in settings where the oceanic crust has not been affected  
195 by hotspot magmatism (without a flexural moat filled with volcanoclastic material) show a  
196 relatively thin pelagic sediment layer (<500 m; Au and Clowes, 1984; Contreras-Reyes et al.,  
197 2008, 2022, 2023; Latta and Dunn, 2020; Grevemeyer et al., 2018b; Spudich and Orcutt,

198 1980; Trummer, 2002; White and Stephen, 1980 among others). For these studies, the high  
199 acoustic impedance contrast between pelagic sediment and oceanic igneous crust (Fig. 2A)  
200 at the sediment/oceanic crust interface promotes P-to-S-wave conversion.

201 Near the 27.6°S seamount, the infill material within the flexural moat has a  $V_p$  of 4.0-  
202 4.5 km/s at depths of 1.0-1.5 *bsf* (Contreras-Reyes et al., 2010). These values are similar to  
203 those measured in the uppermost oceanic crust (Grevemeyer et al., 2018a; Christeson et al.,  
204 2019) and result in a weak velocity and density contrast at the infill material/basement  
205 interface. In contrast, pelagic carbonate sediments usually have  $V_p$  and density values of 1.6-  
206 2.0 km/s and 1,700 kg/m<sup>3</sup>, respectively (Hamilton 1976; Fulthorpe et al., 1989; Tenzer and  
207 Gladkikh, 2014), which results in higher acoustic impedance contrasts at the pelagic  
208 sediment/infill material interface than at the infill material/basement interface (Fig. 6).

209 For the  $V_s$  inversion, we used converted P-S-P crustal phases with picking uncertainties  
210 smaller than 100 ms. Due to our conservative selection of converted S-waves arrivals for  
211 picking, the  $V_s$  model has lower spatial coverage than the  $V_p$  model. We applied damping  
212 constraints to maintain a consistent  $V_p$  within the pelagic sedimentary layer. The initial 2-D  
213  $V_s$  model was constructed by dividing the final inverted 2-D  $V_p$  model by a constant  $V_p/V_s$   
214 ratio of 1.8 (see Supporting Information (SI)). Thus, the  $V_s$  inversion is independent of the  
215  $V_p$  inversion, but the initial  $V_s$  model is dependent on the final  $V_p$  model. Despite this  
216 limitation, the high-density of OBH/S receivers and shots used in controlled seismic source  
217 experiments (the case here), provides better resolution for crust and uppermost mantle than,  
218 for example, that provided by 3-D earthquake tomography.

219  $V_s$  tomographic inversions all converge to a similar outcome, regardless of which  $V_p/V_s$   
220 ratio in the range of 1.65 and 1.85 is used to construct the initial model. The final 2-D  $V_s$  and  
221  $V_p/V_s$  models are presented in Figs. 7B and 7C, respectively. See Sections 3.3 and 3.4 and SI  
222 for details of the resolution and uncertainty of the  $V_p$  and  $V_s$  models.

223

### 224 **3.3 Resolution of the $V_p$ and $V_s$ models**

225 To assess the resolvability of the obtained velocity models, we constructed synthetic  
226 velocity models by applying Gaussian anomalies (Figs. 8A and 8B for  $V_p$  and  $V_s$  inversion,  
227 respectively) to the final inversion velocity models shown in Figs. 7A and 7B for  $V_p$  and  $V_s$ ,  
228 respectively. Each anomaly has a maximum amplitude of  $\pm 5\%$  in velocity. Synthetic travel  
229 time data with the same source-receiver geometry as the actual dataset, were generated using  
230 the perturbed models. Subsequently, these data were inverted using an initial unperturbed  
231 model, and the velocity anomalies recovered appraised. The results show that most of the  
232 velocity anomalies are reasonably well reproduced in position, shape, and amplitude. Some  
233 of the recovered velocity anomalies show certain shape deterioration in the upper mantle  
234 (Figs. 8C and 8D) and for the  $V_s$  inversion in particular at the NE edge of the model (Fig.  
235 8D). Nevertheless, the results show that the geometry and instrument spacing provide  
236 sufficient resolution to discern most of the velocity anomalies, and are capable of  
237 distinguishing between positive and negative variations.

238 We also compute the Derivative Weight Sum (DWS), a proxy for the ray density  
239 (Korenaga et al., 2000), for the  $V_p$  and  $V_s$  models (Figs. 8E and 8F). Most of the oceanic crust  
240 and uppermost mantle have good coverage for the  $V_p$  model (except the edges of the models  
241 and the mantle at greater than 20 km depth). For the  $V_s$  model, DWS values are high to the  
242 NE of 27.6°S seamount, but negligible beneath it.

243

### 244 **3.4 Uncertainties in the $V_p$ and $V_s$ models**

245 To assess the accuracy of the final model, we adopted a Monte Carlo approach (Korenaga  
246 et al., 2000) which consisted of randomly perturbing velocities in our reference model.  
247 Specifically, we generated 10 distinct initial velocity models by adding randomly distributed  
248 perturbations. For the P- and S-wave models, respectively. For the 2-D initial P-wave  
249 velocity models, we applied smooth perturbations distributed randomly, with maximum  
250 velocity perturbations reaching  $\pm 5\%$  of the original amplitude. Additionally, we introduced  
251 phase errors ( $\pm 50$  ms) and common receiver errors ( $\pm 50$  ms) to the original dataset to  
252 establish corresponding perturbed travel times, following the approach outlined by Korenaga  
253 et al. (2000). A similar procedure was followed for the S-wave velocity model.

254

255 Subsequently, we conducted a tomographic inversion for each velocity model using  
256 a single noisy dataset. This allowed us to evaluate not only the solution's reliance on the  
257 reference model but also the impact of phase arrival time picking errors. The termination  
258 criterion for each inversion was set at  $\chi^2 < 1$  (when the travel time fit becomes lower than the  
259 travel time uncertainties).

260

261 Figs. 8G and 8H show the calculated percentage standard deviation ( $\Delta V$ ) for the average  
262 velocity-depth model from the ten final models for both P- and S-wave velocity. The root-  
263 mean-square travel time misfit ( $T_{\text{RMS}}$ ) and  $\chi^2$  parameter for the final model are summarized  
264 in the SI. The  $\Delta V$  is generally lower than 0.2 km/s ( $\leq 4\%$ ) for both models with average values  
265 of 1.5%.

266

267 Figs. 9A and 9B show the calculated  $V_p/V_s$  model and its standard deviation  $\Delta \left(\frac{V_p}{V_s}\right)$ ,  
268 computed as

269

$$(1) \quad \Delta \left(\frac{V_p}{V_s}\right) = \pm \left(\frac{V_p}{V_s}\right) \sqrt{\left(\frac{\Delta V_p}{V_p}\right)^2 + \left(\frac{\Delta V_s}{V_s}\right)^2}$$

270

271 where  $\Delta V_p$  and  $\Delta V_s$  correspond to the uncertainties of the  $V_p$  and  $V_s$  models,  
272 respectively. Overall,  $\Delta \left(\frac{V_p}{V_s}\right)$  magnitudes are lower than 0.06 and reach maximum values of  
273 0.1 in the summit region of the 27.5°S seamount (Fig. 9B).

274

#### 275 276 **4. Results and Interpretation**

277 We summarize the seismic findings in three contexts: (1) sediment/volcaniclastic  
278 deposits; (2) seismic structure of the Pacific crust and upper mantle adjacent to the 27.6°S  
279 seamount (NE and SW domains along profile P03) unaffected by hotspot magmatism; and

280 (3) seismic structure of the crust beneath the 27.6°S seamount that results from hotspot  
281 magmatism.

282

#### 283 **4.1. Sediment/volcaniclastic deposits layer**

284 The sediment layer has a thickness of between 100 m at the NE and SW end of the  
285 profile to 1,200-1,500 m on the flanks adjacent to the 27.6°S seamount (Fig. 2). The P-wave  
286 velocity within this layer ranges from 1.6 km/s at the seabed to between 1.7 and 4.0-4.5 km/s  
287 at 5.5-6.0 km depth, depending on seamount proximity (Fig. 7). The P-wave model shows a  
288 moat-like depression around the seamount, up to 200 km across, and having a maximum  
289 depth of 1,000-1,500 m at 160 km and 240 km along-profile (Fig. 7). This moat is formed by  
290 the flexure of the oceanic lithosphere under the load of the LR, and the seismic velocity-  
291 depth structure suggests that its infill is composed mainly of volcaniclastic material  
292 (Contreras-Reyes et al., 2010; Watts et al., 2021; McGregor et al., 2023; Dunn et al., 2024;  
293 Xu et al., 2023).

294 The infill material within the flexural moat adjacent to the 27.6°S seamount was  
295 interpreted as an accumulation of volcaniclastic and debris flow deposits based on the range  
296 in P-wave velocity of 1.9-4.0 km/s observed by Contreras-Reyes et al. (2010). A few hundred  
297 meters below seafloor,  $V_p$  ranges from 1.7 to 1.9 km/s in the vicinity of the seamount,  
298 indicating the likely presence of a pelagic drape, hyaloclastics, or sub-areal volcanic rocks  
299 (Grevemeyer and Flueh, 2008; Contreras-Reyes et al., 2010). The base of the pelagic  
300 sedimentary layer is highlighted by the intra-sediment reflector shown in Fig. 2. Specifically,  
301 around the moat a discontinuity likely separates pelagic sediment and eroded oceanic island  
302 material from volcaniclastic deposits formed during the creation of the 26.7°S seamount. Our  
303 2-D  $V_s$  inversion model reveals velocities within the volcaniclastic layer ranging from 1.8 to  
304 2.8 km/s (Fig. 7B), which results in a  $V_p/V_s$  ratio of between 1.7-1.75, which is similar to that  
305 obtained by laboratory measurements of pillow basalts (Fig. 7B; Christensen, 1996; Carlson  
306 and Miller, 2003).

307

308 **4.2 Seismic structure of the Pacific crust and upper mantle adjacent to the 27.6°S**  
309 **seamount**

310 To the NE and SW of 27.6°S seamount, the Pacific upper crust (oceanic crustal seismic  
311 layer 2) is approximately 1.5-2.0 km thick and has a P-wave velocity ranging between 4.0  
312 and ~6.7 km/s (Fig. 7A). The corresponding S-wave velocity ranges between 2.5 and 3.0  
313 km/s and is similar to that of the volcanoclastic deposits sited in the deeper portion of the  
314 moat. The  $V_p/V_s$  ratio ranges between 1.7 and 1.8. At the upper-to-lower crust transition  
315 (boundary between seismic layer 2 and 3),  $V_p$  and  $V_s$  velocities are 6.4-6.7 km/s and 3.3-3.7  
316 km/s, respectively, corresponding to a  $V_p/V_s$  ratio of 1.7-1.85. The  $V_p/V_s$  values observed at  
317 the top and bottom of the upper crust of ~1.7-1.85 are consistent with laboratory  
318 measurements of basalts and diabase reported by Christensen (1996) and Carlson and Miller  
319 (2003).

320 At a distance of more than 50 km to both the NE and SW from the 27.6°S seamount, the  
321 lower crust (seismic layer 3) has a thickness of ~4 km within an overall crustal thickness of  
322 5.5-6.0 km. Lowermost crustal P-wave velocities range between 7.0-7.2 km/s, whereas S-  
323 wave velocity ranges between 3.7-4.0 km/s, with corresponding  $V_p/V_s$  ratios of 1.7-1.85  
324 (Figs. 7 and 9). These  $V_p/V_s$  values are slightly lower but still in agreement with laboratory  
325 measurement of gabbroic samples reported by Christensen (1996) and Carlson and Miller  
326 (2003).

327 The P-wave velocity within the uppermost mantle increases significantly from 8.1 to  
328 ~8.9 km/s from the seismic Moho to a depth of ~30 km, except to the SW, furthest from the  
329 Osbourn Trough. Here, the P-wave velocity varies from 8.1 to 8.4 km/s at equivalent depths.  
330  $V_s$  ranges from 4.5 to 4.7 km/s, with a corresponding  $V_p/V_s$  ratio of 1.8-1.9 between 0 km and  
331 300 km along-profile (Figs. 7, 10, 11). At the SW end of the profile, the  $V_p/V_s$  ratio is lower  
332 at 1.65-1.8, which we interpret as reflecting a region of dry mantle peridotite (e.g.,  
333 Christensen, 1996). We attribute the region of relatively elevated  $V_p/V_s$  of 1.8-1.9 to reflect  
334 anhydrous mantle. This interpretation contrasts with studies that associate a  $V_p/V_s$  range of  
335 1.8-1.9 with serpentinized mantle (e.g., Christensen, 1996; Carlson and Miller, 1997; 2003).  
336 This alternative interpretation will be discussed in Section 5.

337

### 338 **4.3 Seismic structure of the crust beneath the 27.6°S seamount**

339       Near the summit of 27.6°S seamount, the P-wave model shows velocities in the range  
340 6.0-6.5 km/s (Fig. 7A) overlying an apparent intrusive core ( $V_p = 6.5-7.5$  km/s). Similar  
341 features are also observed at Osbourn and 28.5°S seamounts (Robinson et al., 2018). The S-  
342 wave velocity-depth model exhibits velocities in the range 3.2-3.4 km/s in the near sub-  
343 seabed, resulting in a  $V_p/V_s$  ratio of 1.85-1.95 (Figs. 7B and 7C). Such  $V_p/V_s$  values are higher  
344 than those observed in the upper crust, particularly when compared to similar subseabed  
345 depths away from the seamount (Figs. 7-9). Such elevated  $V_p/V_s$  ratios are indicative of  
346 basaltic rocks with a high degree of porosity and fracturing (Christensen, 1996). Lower  $V_p/V_s$   
347 ratios are observed adjacent to the seamount, which are interpreted as reflecting  
348 volcanoclastic and debris flow deposits (Fig. 9).

349

## 350 **5. Discussion**

### 351 **5.1. Volcanoclastic and sediment deposition around the 27.6°S seamount**

352       Seamount formation induced by plate-hotspot interaction involves episodic  
353 eruptions and the deposition of volcanoclastic material. Consequently, significant clastic  
354 volumes amass in the summit and upper flanks. The susceptibility to mass wasting events  
355 and subsequent re-sedimentation in these regions contributes to the development of debris  
356 flows along the flanks and aprons (Moore et al., 1978; Wolfe et al., 1994). Furthermore,  
357 large-scale landslides and avalanches may occur, that can span hundreds of kilometers, as  
358 has been reported for Hawaii (Moore et al., 1978), La Reunion (Lenat et al., 1978), Tenerife  
359 (Watts and Masson, 1984), Great Meteor (Weigel and Grevemeyer, 1999), and Marquesas  
360 (Wolfe et al., 1994).

361

362       Following the passage of the Pacific Plate over the LR hotspot mantle plume,  
363 volcanoclastic deposition ceases, and sedimentation is primarily attributed to pelagic

364 deposition and sub-aerial island erosion. According to  $^{40}\text{Ar}/^{39}\text{Ar}$  age progression along the  
365 Louisville seamount trail (Koppers et al., 2004, 2011), the 27.6°S seamount was active some  
366 68-69 Ma, with a plate age of ~20 Myr at that time (Mueller et al., 2008). Pelagic sediment  
367 accumulates at remarkably slow rates of 2-10 m/Myr (Straume et al., 2019; Rotzien et al.,  
368 2022), which would result in a pelagic sediment thickness of 130-700 m in our study area.  
369 This thickness range is consistent with our observations from the MCS data away from the  
370 seamount regions (between 100 km and 300 km along-profile distance; Fig. 2). On the flanks  
371 of the seamount, the material lying above the volcanoclastic deposits is most likely to be a  
372 mixture of sub-aerial erosion products and pelagic drape.

373

374 We interpret the velocity-depth structure ( $V_p = 2.0\text{-}4.0$  km/s,  $V_s = 1.8\text{-}2.4$  km/s and  
375  $V_p/V_s = 1.7\text{-}1.8$ ; Fig. 7) observed above the oceanic basement as a mixture of volcanoclastic  
376 deposits such as breccias, tephra, slump blocks, and lava flows (Staudigel and Schmincke,  
377 1983). Weigel and Grevemeyer (1999) note that these deposits exhibit a significantly higher  
378 porosity when compared to basaltic lavas extruded during crustal accretion at deeper water  
379 spreading centers. Our model shows a P-wave velocity greater than 4.0 km/s in the uppermost  
380 oceanic crust adjacent to the LR (*i.e.* unaffected by hotspot magmatism; Fig. 7), indicating a  
381 porosity lower than the volcanoclastic deposits. The upper crust, with a thickness of 1.0-1.5  
382 km (see Section 3.2), has a velocity-depth structure characteristic of pillow lavas overlying  
383 a sheeted dike complex (Vera et al., 1990; Carlson, 2001; Christeson et al., 2019).

384

385 At the summit of the 27.6°S seamount, the upper part exhibits a layer characterized by  
386 a P-wave velocity in the range 4.5-6.0 km/s and thickness of <1.0 km. This layer undergoes  
387 a transition to higher velocities (>6.4 km/s; Figs 7 and 10) consistent with the contact between  
388 alkali basaltic pillow lavas and intrusive rocks. However, the  $V_p/V_s$  ratio range of 1.85-1.95  
389 in this region is remarkably higher than those observed at the flanks of the 27.6°S seamount  
390 ( $V_p/V_s = 1.7\text{-}1.8$ ; Figs. 7 and 9) but is also observed in laboratory measurements of pillow  
391 basalts and diabase complexes (Fig. 9; Christensen, 1996). This suggests, in the summit  
392 region, the presence of pillow basalts and intrusive rocks characterized by water-saturated

393 cracks and fractures (Popp and Kern, 1994; Wang et al., 2012). These geological features are  
394 likely the result of eruptions taking place at shallower depth (Peterson and Moore, 1976).

395

396 Most likely, the now shallow crustal depth of the intrusive core in the summit region  
397 results from many tens of millions of years of erosion. Although converted P-S-P-waves are  
398 confined to depths of 1-2 km beneath the top of the seamount, these seismic phases are  
399 prominent (see Fig. 5) and constrain (to a limited depth) the relatively high  $V_p/V_s$  ratio beneath  
400 the summit.

401

## 402 **5.2. Intrusion beneath the 27.6°S seamount**

403 Unfortunately, the  $V_s$  model does not constrain the middle and lower crustal sections  
404 of the 27.6°S seamount due to the lack of recorded mode-converted S-waves that would have  
405 propagated within the volcanic edifice (Fig. 7). Nevertheless, the P-wave model reveals a  
406 region of high velocity (6.5-7.5 km/s) that can be interpreted as a symmetrical, semi-conical-  
407 shaped core (Fig. 7; Contreras-Reyes et al., 2010). Similar observations have been reported  
408 for Great Meteor seamount (Weigel and Grevenmeyer, 1999), and the volcanic oceanic islands  
409 of Hawaii (Zucca et al., 1982; McGregor et al., 2023; Dunn et al., 2024) and Tenerife  
410 (Canales et al., 2000). Conversely, other seamounts exhibit a conical structure with lower  
411 relative P-wave velocity (<5.5 km/s), which has been interpreted to reflect construction by  
412 extrusive processes. Some examples include the Juan Fernández Ridge (Kopp et al., 2004),  
413 Musicians Seamount Province (Kopp et al., 2003), Jasper seamount (Hammer et al., 1994),  
414 and the Marquesas islands (Caress et al., 1995).

415 Along the LR, Robinson et al. (2018) also report the presence of discrete cores of high  
416  $V_p$  ( $\geq 6.0$  km/s) for the Osbourn and 28.5°S seamounts (Figs. 1 and 11). In each case, the high  
417  $V_p$  ( $\geq 6.0$  km/s) region extends upwards to depths of 1.0-1.5 km beneath each summit.  
418 Canopus seamount (the widest) is similarly interpreted to have a predominantly intrusive  
419 composition, albeit the higher velocity core only extends upward to 2.0-3.0 km beneath its  
420 summit (Robinson et al., 2018).

421 Within the lowermost crust beneath 27.6°S seamount, P-wave velocities exceed 7.2-7.5  
422 km/s (Fig. 7), lying in the range observed for gabbroic (6.9-7.2 km/s) and peridotite/dunite  
423 (7.9-8.1 km/s) lithologies, as determined in laboratory experiments (e.g., Carlson and Miller,  
424 2003; Christensen, 1996). Similar values are observed in seismic experiments for the lower  
425 crust and upper mantle, respectively (Grevemeyer et al., 2018a; Christeson et al., 2019).  
426 Contreras-Reyes et al. (2010) and Richards et al. (2013) have interpreted the high P-wave  
427 velocity regions in the lower crust beneath the LR to result from the intrusion of relatively  
428 buoyant intrusive bodies at just above Moho depth within the crust, in contrast to findings at  
429 the Marquesas (Caress et al., 1995) and Reunion (Charvis et al., 1999) islands, where such  
430 bodies are located beneath the Moho.

431

### 432 **5.3. Upper mantle structure**

433 Our seismic results reveal a notably elevated  $V_p$ , ranging from  $\sim 8.1$  to  $\sim 8.9$  km/s, within  
434 the uppermost  $\sim 15$  km of the oceanic mantle in the NE segment and beneath the 27.6°S  
435 seamount of profile P03 (Fig. 7). Towards the SW, there is a localized reduction in the upper  
436 mantle P-wave velocity, ranging from 8.1 to 8.3 km/s, such that the upper mantle velocity-  
437 depth structure appears asymmetric about 27.6°S seamount (even considering the maximum  
438  $V_p$  uncertainty of 0.2 km/s; Fig. 8G). There is also no evidence to support the presence of  
439 sub-crustal bodies/magmatic underplating centered beneath the LR at this location,  
440 equivalent to that reported by Watts et al. (1985), Caress et al. (1995), and Charvis et al.  
441 (1999) beneath the Hawaiian, Marquesas, and Reunion islands, respectively.

442 In the NE segment, where a  $V_p \geq 8.0$  km/s and  $V_s \geq 4.5$  km/s are observed, there is a  
443 notable region characterized by relatively high  $V_p/V_s$  ratios of 1.8-1.9. These ratios exhibit a  
444 decreasing trend to 1.7-1.8, predominantly in the SW sector of the seamount, at depths of 1-  
445 3 km below the Moho (Figs. 7 and 9).  $V_p/V_s$  ratios higher than 1.8 are usually interpreted in  
446 terms of some degree of hydration of the mantle. For example, Carlson and Miller (2003)  
447 estimated that  $V_p$  and  $V_s$  values of  $\sim 7.7$  and  $\sim 3.9$  km/s ( $V_p/V_s \sim 1.97$ ) reflect a serpentine  
448 content of 20% (by volume) or 3% H<sub>2</sub>O (by wt%). In the context of their linear relationship,

449 the endmember composition of dry mantle peridotite is attained with  $V_p$  and  $V_s$  values of  $\sim 8.1$   
450 and  $\sim 4.3$  km/s ( $V_p/V_s \sim 1.88$ ). These values are somewhat higher when compared to  
451 Christensen (1996) who reported  $V_p/V_s$  of  $\sim 1.75$  for mantle dunite at 200 MPa.

452 Contreras-Reyes et al. (2008) interpreted  $V_p/V_s$  ratios exceeding 1.8 as indicative of  
453 partial hydration resulting from bending-related faulting off southern Chile. This faulting has  
454 the potential to breach the entire oceanic crust, establishing pathways for water migration to  
455 upper mantle depths (Kopp et al., 2004; Contreras-Reyes et al., 2008; Grevemeyer et al.,  
456 2018). The down-going Pacific Plate adjacent to the Tonga-Kermadec Trench has well-  
457 developed bending-related faulting as shown by bathymetric maps, seismic reflection data  
458 (Funnell et al., 2014, 2017) and low upper mantle velocities (7.4-7.9 km/s) in the trench-  
459 outer rise region (Contreras-Reyes et al., 2011; Stratford et al., 2015; Funnell et al., 2017;  
460 Robinson et al., 2018; Fig. 1). However, strong plate bending stresses prior to plate  
461 subduction are confined to near the trench-outer rise region, where the plate curvature is  
462 higher. In contrast, bending-related faulting becomes weaker or absent away from the trench  
463 axis (typically  $>120$  km seaward of the trench axis; Contreras-Reyes and Osses, 2010; Watts  
464 and Hunter, 2016).

465 The NE segment of profile P03 lies  $\sim 150$  km from the trench axis (Fig. 1). Notably, the  
466 2-D  $V_p$  model does not show evidence of reduced velocity in the crust or in the upper mantle  
467 (Fig. 7) relative to what might be expected for normal oceanic crust and mantle (e.g.,  
468 Grevemeyer et al., 2018a). This observation implies that the seismic structure of the upper  
469 oceanic lithosphere remains unaffected by plate bending in this region. Hence, the elevated  
470  $V_p/V_s$  ratio of 1.8-1.9 in the upper mantle is likely due to other factors.

471 The SW segment of profile P03 is located  $\sim 100$  km from the trench axis, (Fig. 1). In this  
472 region, the 2-D  $V_p$  model does not show any evidence of reduced velocity in the crust, but  
473 the upper mantle exhibits a distinct zone of reduced P-wave velocity compared to the NE  
474 portion (Fig. 7). Nevertheless, the asymmetrical variation in upper mantle velocity along P03  
475 may result from a complex interplay of processes, including mantle plume/lithosphere  
476 magmatic interaction, plate-bending related faulting, and the high anisotropy of mantle  
477 olivine. In this paper we investigate the likelihood of mantle anisotropy being the cause, by

478 comparing our results with Robinson et al.'s (2018) P-wave model of profile PC which runs  
479 along the LR axis.

480

#### 481 **5.4. Upper mantle anisotropy of the Pacific Plate**

482 Geophysical evidence supporting upper mantle anisotropy includes seismic tomography  
483 (Dunn et al., 2000), shear wave splitting (Russo et al., 2010), surface wave (Rayleigh and  
484 Love waves) dispersion (Regan and Anderson, 1984; Eddy et al., 2022), receiver function  
485 studies (Levin et al., 2002; Hu et al., 2015), and attenuation anisotropy (Nishimura and  
486 Forsyth, 1989; Becker and Lebedev, 2021). Generally, the two primary mechanisms invoked  
487 for the generation of upper mantle seismic anisotropy are Lattice Preferred Orientation (LPO)  
488 and directional mantle flow (Hess 1964; Zhang and Karato, 1995; Mark et al. 2019).  
489 Additionally, shape preferred orientation (SPO; Kern et al., 2008), partial melt and fluid  
490 alignment (Waff and Faul, 1992), and chemical variations (Khan et al., 2009) are also thought  
491 to be significant contributors to seismic anisotropy.

492 As olivine establishes a preferred orientation direction during flow or viscous shear  
493 deformation, it is not unreasonable to suppose that the  $P_n$  velocity perpendicular to the  
494 spreading center axis will be faster than spreading direction-parallel (Hess, 1964; Skemer  
495 and Hansen, 2016). For example, Mark et al. (2019) estimated a seismic anisotropy of  $6.0 \pm$   
496  $0.3\%$  at the Moho within 70 Ma lithosphere in the central Pacific. In addition, Shearer and  
497 Orcutt (1985) and Dunn et al. (2000) measured  $P_n$  anisotropy of 5.5% and 6-7% in the NW  
498 Pacific and at the East Pacific Rise, respectively, aligned parallel to the spreading direction.  
499 In contrast, 3-4% (Gaherty et al., 2004) or neglectable (de Melo et al., 2021)  $P_n$  anisotropy  
500 has been observed in the north and equatorial Atlantic, respectively, aligned parallel to the  
501 spreading direction. The reduced  $P_n$  anisotropy observed in the Atlantic implies increased  
502 conductive cooling at the corresponding portion of the Mid-Atlantic Ridge when the  
503 lithosphere formed and, consequently, increased localized (brittle) deformation within the  
504 mantle lithosphere and a limited extent of viscous deformation (Sleep, 1975; Gaherty et al.,  
505 2004).

506 Figs. 11A-11D show a comparison of the upper mantle P-wave structure between  
 507 profiles P03 and PC (Robinson et al., 2018) in the proximity of the 27.6°S seamount. To  
 508 assess seismic anisotropy in this area, we utilize the anisotropic percentage, defined as

$$509 \quad (2) \text{ Anisotropic percentage} = \frac{(V_f - V_s)}{V_f} \times 100 (\%)$$

510 where  $V_f$  and  $V_s$  represent the faster and slower seismic velocities, respectively.

511 Mantle velocities modelled along profile P03 are faster than those along the southern  
 512 portion of PC, except at the Moho (Fig. 11C) and we use these to calculate the minimum and  
 513 maximum seismic anisotropic percentage, as shown in Fig. 11D. In both instances,  $P_n$   
 514 anisotropy increases with depth, reaching a maximum of ~6% at about 4 km depth beneath  
 515 the Moho. The increase in  $P_n$  anisotropy with depth has also been measured by Mark et al.  
 516 (2019), who reported a vertical velocity gradient of ~0.02 s<sup>-1</sup> in the fast direction and 0 s<sup>-1</sup> in  
 517 the slow direction.

518 Our study area lacks a seismic profile oriented perfectly along the paleo spreading  
 519 direction, as such, the upper mantle velocity is unconstrained in that direction. Nevertheless,  
 520 we can extrapolate our results by noting that the azimuth ( $\varphi$ ) between profile P03 and the  
 521 paleo spreading direction is ~20° ( $\varphi_1$ ), while that for profile PC is ~50° ( $\varphi_2$ ) (Fig. 10A). We  
 522 use a simplified expression to extrapolate the mantle P-wave velocity,  $V$ , as a function of the  
 523 azimuth  $\varphi$  (in map view) in the form:

$$524 \quad (3) \quad V(\varphi) = V_{max} - (V_{max} - V_{min})|\sin(\varphi)|$$

525 Here, we have assumed that the maximum velocity ( $V_{max}$ ) is attained along the paleo  
 526 spreading direction, while the minimum velocity ( $V_{min}$ ) is perpendicular to this. This  
 527 assumption is consistent with findings in the central Pacific (Hess, 1964; Mark et al., 2019).  
 528 Thus,  $V_{max}$  and  $V_{min}$  can be written as:

$$529 \quad (4) \quad V_{max} = V_1 - \left( \frac{V_1 - V_2}{|\sin(\varphi_1)| - |\sin(\varphi_2)|} \right) |\sin(\varphi_1)| \quad \text{and}$$

530 
$$(5) V_{min} = V_1 + \left( \frac{V_1 - V_2}{|\sin(\varphi_1)| - |\sin(\varphi_2)|} \right) [1 - |\sin(\varphi_1)|]$$

531 Here,  $V_1$  and  $V_2$  denote the upper mantle seismic velocities at the intersection of measured  
 532 seismic profiles 1 and 2, respectively, and it is assumed that  $\varphi_1 \neq \varphi_2$ . For example, Figs. 11E  
 533 and 11F illustrate in polar coordinates the minimum and maximum values for  $V_1$  (P03) and  
 534  $V_2$  (PC) at a specific depth, corresponding to azimuths of  $\varphi_1 \sim 20^\circ$  and  $\varphi_2 \sim 50^\circ$ , respectively.  
 535 Thus,  $V_{max}$  and  $V_{min}$  are computed from Eqs. 4 and 5, enabling the determination of the semi-  
 536 axes of the ellipses presented in Figs. 11E and 11F.

537 The comparison of upper mantle  $V_p$  between profiles P03 and PC reveals an increase  
 538 in seismic anisotropy with depth within the uppermost 3-4 km of the mantle. The gradient of  
 539 anisotropic percentage ranges between  $\sim 0.2$  and  $\sim 2\%/km$  (Figs. 11C and 11D), which is  
 540 notably higher than that reported by Mark et al. (2019) in the central Pacific. Despite the  
 541 simplicity of our  $P_n$  anisotropy estimation in the vicinity of the P03/PC intersection, which  
 542 lacks azimuthal constraints, the resolution achieved by our active seismic study enhances the  
 543 confidence in measuring a relatively high-velocity zone extending up to  $\sim 15$  km below the  
 544 Moho along P03.

545

## 546 6. Summary

547 A re-examination of wide-angle and reflection seismic data along a profile sub-  
 548 perpendicular to the Louisville Ridge axis has revealed significant variations in P- and S-  
 549 wave velocity when compared to typical oceanic crust at equivalent sub-basement depths.  
 550 Adjacent to the 27.6°S seamount, our results indicate a  $V_p/V_s$  ratio ranging from 1.7 to 1.8 in  
 551 both the upper and lower oceanic crust, unaffected by hotspot magmatism. These results are  
 552 consistent with documented  $V_p/V_s$  ratios obtained from laboratory measurements of gabbro  
 553 and diabase. Conversely, in the vicinity of the seamount summit, there is a relatively higher  
 554  $V_p/V_s$  ( $\sim 1.9$ ), that is attributed to intrusive cores. At depths  $>1$  km beneath the summit of the  
 555 27.6°S seamount, our results suggest a predominance of intrusive features with a  $V_p$  ranging  
 556 from 6.5 to 7.5 km/s.

557  $V_p$  and  $V_p/V_s$  values of 2.0-4.0 km/s and 1.7-1.8, respectively, observed between the  
558 pelagic sediment layer and oceanic crust, indicate the possible presence of debris flows and  
559 eroded extrusive lavas. Notably, in the summit region,  $V_p$  (>6.0 km/s) and  $V_p/V_s$  (1.85-1.95)  
560 are remarkably higher than those observed in adjacent the upper oceanic crust ( $V_p = 4.0-6.0$   
561 km/s and  $V_p/V_s = 1.7-1.8$ ) at equivalent sub-basement depths. The observed relatively high  $V_p$   
562 in the summit region reflects an intrusive core, extending upwards almost to the seafloor.

563 The average  $V_p/V_s$  ratio in the upper mantle NE of 27.6°S seamount lies within the range  
564 of 1.8 to 1.9. To the SE of the seamount, there is a decrease in  $V_p/V_s$  ratio to 1.7-1.8 within  
565 the uppermost 4-5 km of the upper mantle. Additionally, an elevated  $V_p$  (~8.9 km/s) is  
566 observed at depths of 10-15 km below the Moho, both beneath the seamount and to the SW  
567 of it. In conjunction with a  $V_p$  model obtained along the Louisville Ridge axis (Robinson et  
568 al., 2018), our results indicate seismic upper mantle anisotropy of up to 6% at ~4 km below  
569 the Moho. Specifically, the fast axis is rotated by approximately 20° relative to the direction  
570 of the paleo spreading flow associated with the Osbourn Trough.

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579

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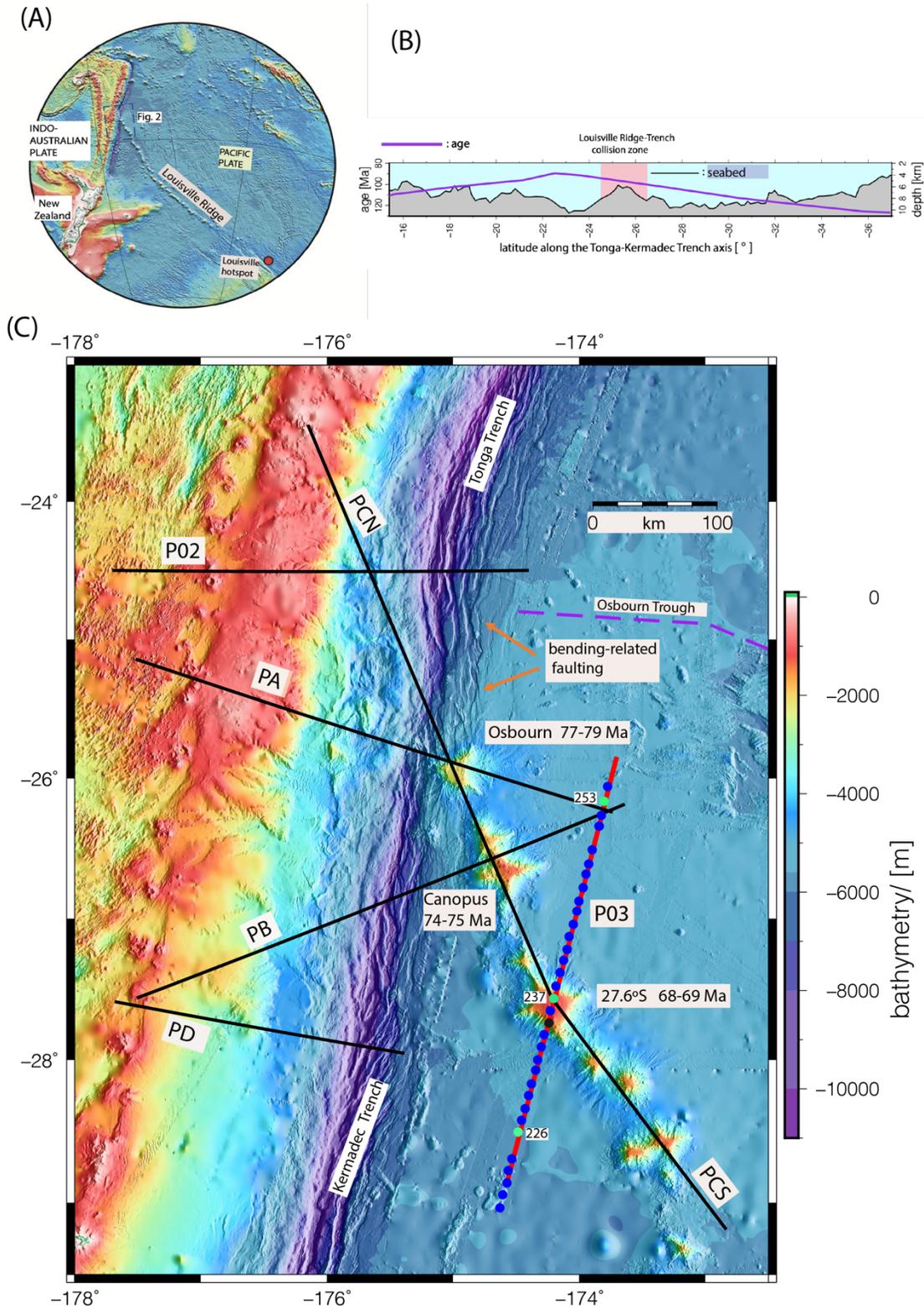
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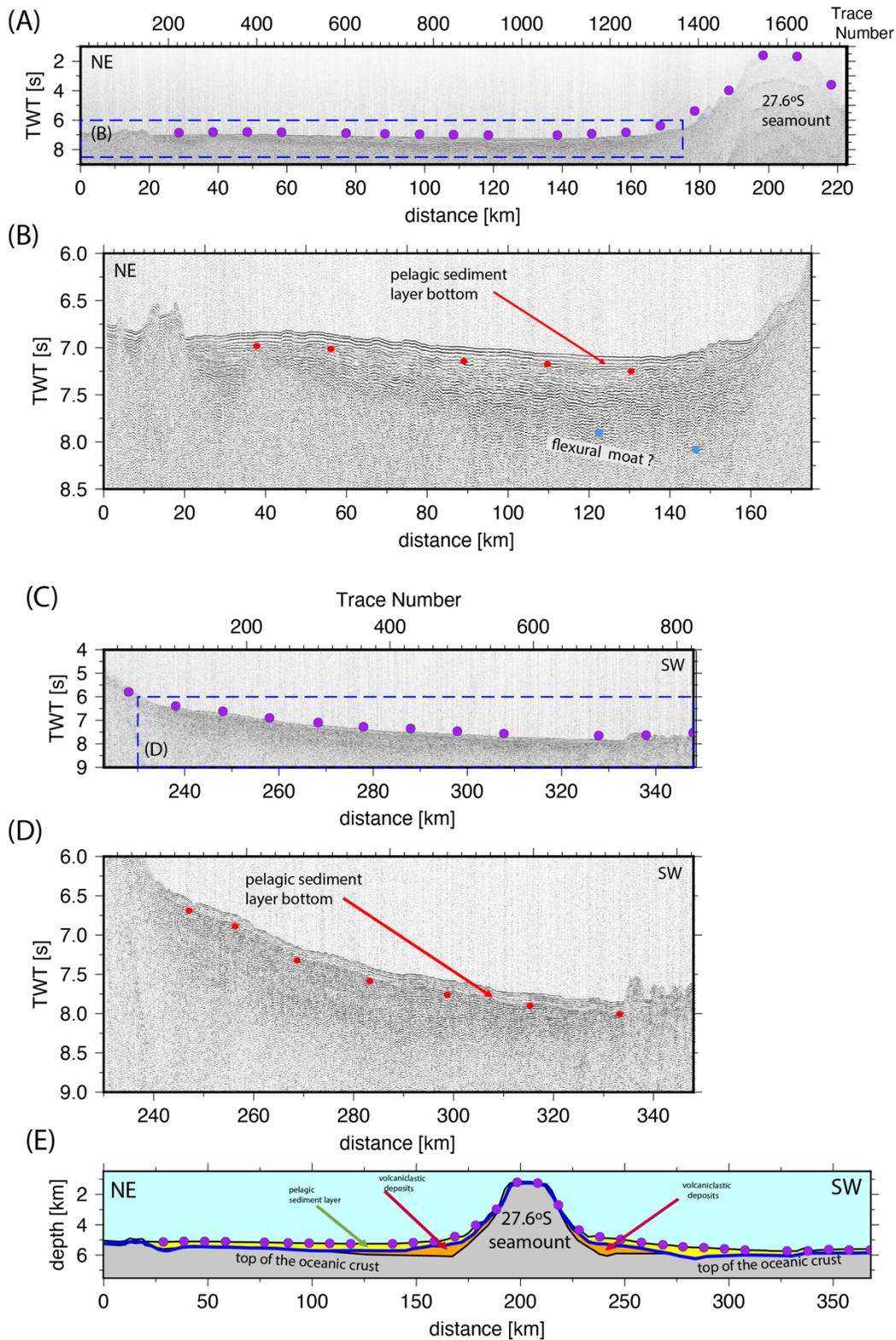


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882

883 **Fig. 1.** (A) Geodynamic setting of the Pacific, Antarctic, and Indo-Australian Plates. The Louisville Ridge is  
 884 an ~4,300 km-long hotspot track that formed by the interaction of the Pacific Plate with the Louisville hotspot

885 (red dot) located near the intersection of the Eltanin Fracture Zone with the SW Pacific–Antarctica spreading  
886 center (e.g., Watts et al., 1988). The Pacific Plate approaches the Tonga-Kermadec Trench at a convergence  
887 rate of 150 km/Myr (Lonsdale, 1986). **(B)** Tonga-Kermadec Trench axis depth taken from the global  
888 bathymetric grid of Ryan et al. (2009). The age of the oceanic Pacific Plate at the trench axis was taken from  
889 the global grid of Mueller et al. (2008). **(C)** Global satellite observations and swath bathymetry seabed  
890 topography of the Tonga-Kermadec Trench-Louisville Ridge collision zone (Ryan et al., 2009). The Osbourn  
891 seamount is the oldest remaining seamount prior to collision with the Indo-Australian Plate. Approximate  
892 basement ages for the seamounts are based on radiometric age dating of dredged rock samples (Koppers et al.,  
893 2004; 2011). The solid black lines represent the wide-angle seismic profiles reported by Stratford et al. (2015;  
894 profile PA), Funnell et al. (2017; profile PB), Robinson et al. (2018; profile PC, split into north and south parts  
895 annotated PCN and PCS), and Contreras-Reyes et al. (2011; profile P02). Profile PD was a multichannel seismic  
896 reflection profile only, reported by Funnell et al. (2017). The red line depicts the wide-angle seismic profile  
897 P03 modelled in this study, while the blue dots indicate the locations of the 33 OBH/S stations. Green dots with  
898 station numbers indicate the three OBH's (253, 237, and 226) shown in Figs. 3-5.  
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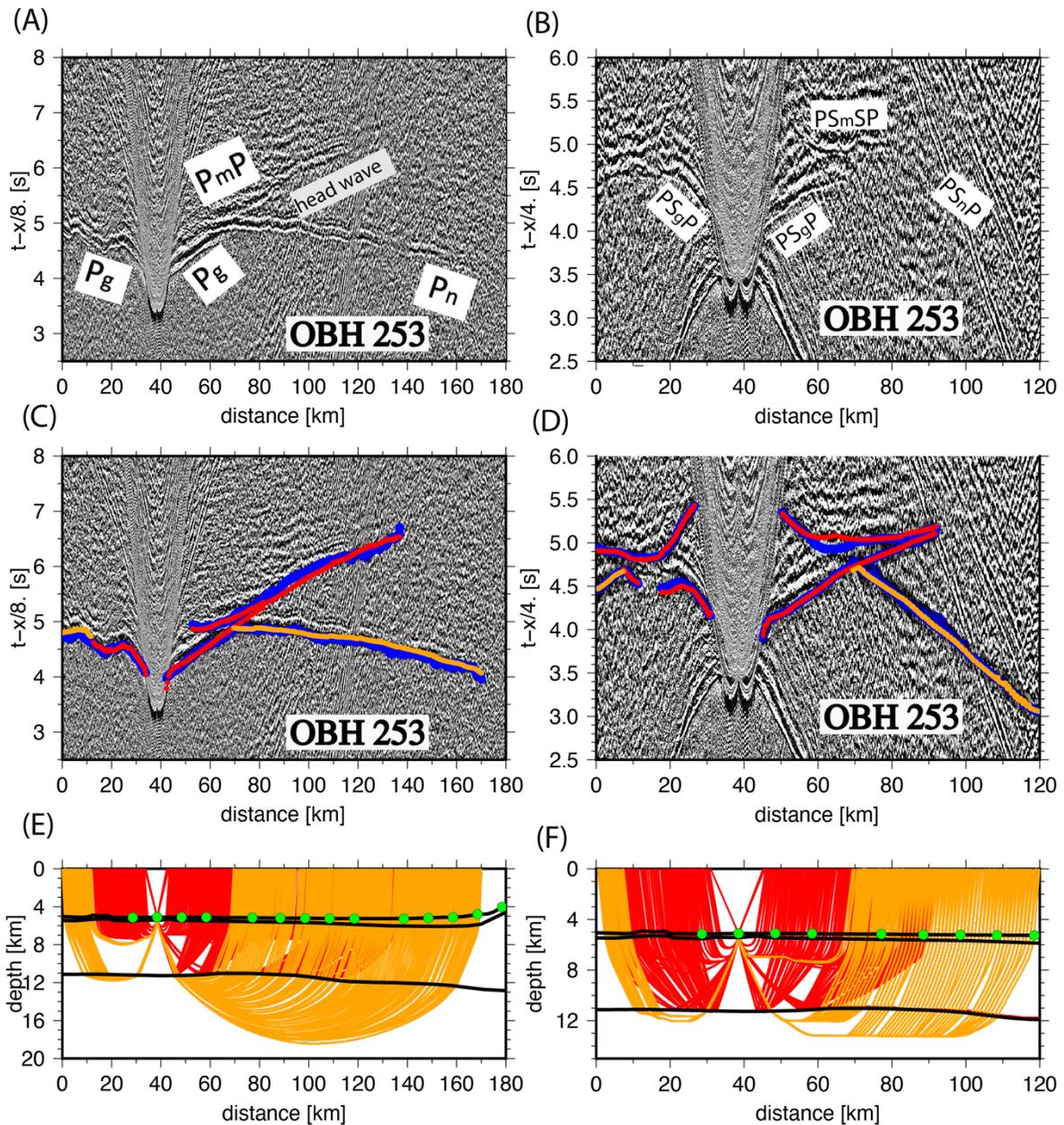


901

902 **Fig. 2.** Seismic reflection data along profile P03. **(A)** Section of profile to the NE of the 27.6°S seamount. **(B)**

903 Zoom-in as shown by the blue dashed box in (A). **(C)** Section of profile to the SW of the 27.6°S seamount. **(D)**

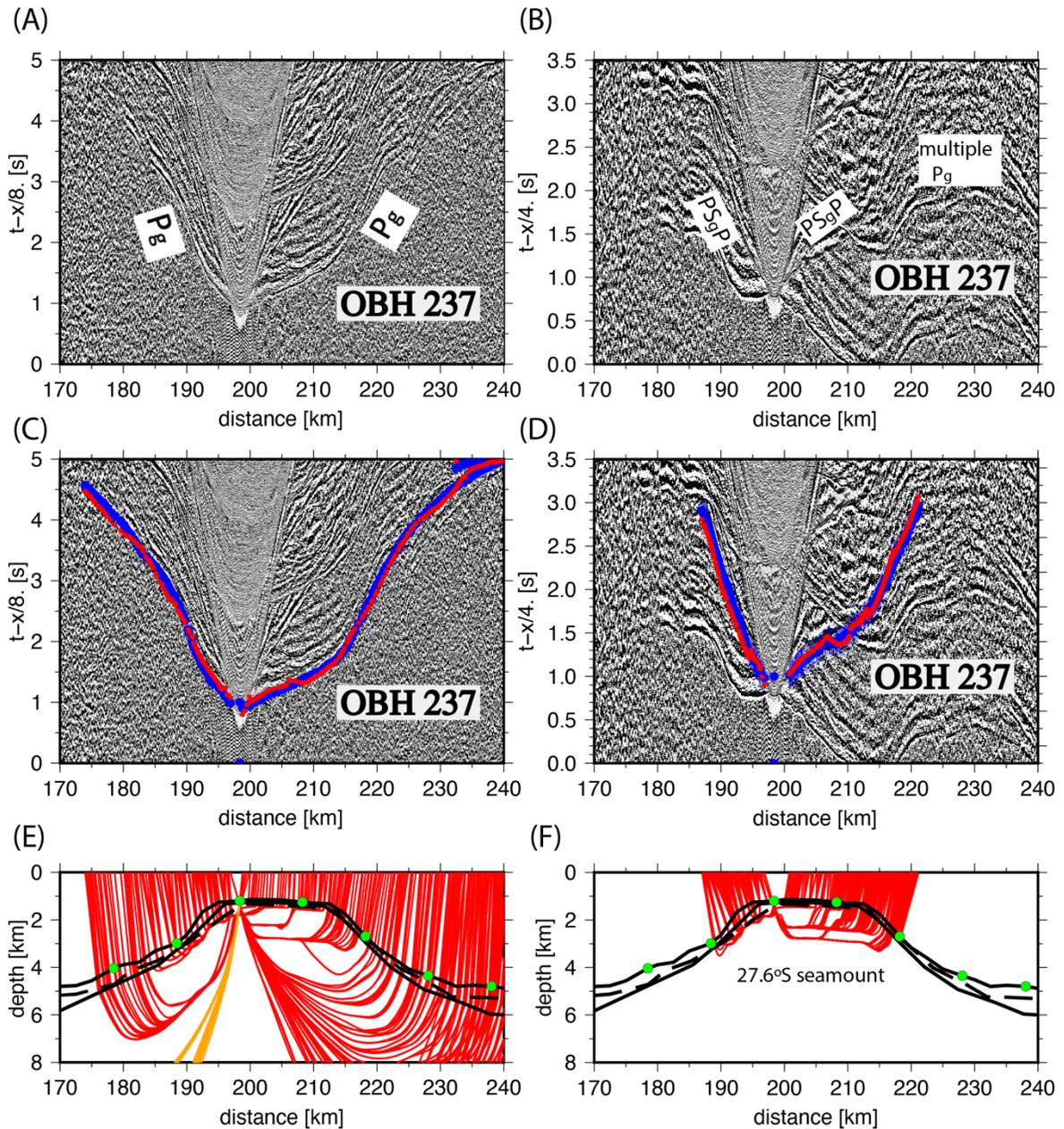
904 Zoom-in as shown by the blue dashed box in (C). (A)-(D) are plotted in two-way-travel (TWT) time. The  
 905 average pelagic sediment thickness is 0.1-0.15 TWT,  $\sim 0.35$ -0.5 km. (E) Main reflectors and bathymetry  
 906 converted from TWT to depth using the 2-D  $V_p$  model of Contreras-Reyes et al. (2010). Pelagic sediments and  
 907 volcanoclastic deposits are differentiated by  $V_p$  ranges of 1.6-1.9 km/s and  $>2.0$ -4.5 km/s, respectively. The  
 908 flexural moat geometry is based on the 2-D  $V_p$  model of Contreras-Reyes et al. (2010). Purple dots depict the  
 909 location of the OBH/S shown in (A), (C) and (E). Red dots mark the bottom of the pelagic sediment layer  
 910 (B) and (D).



911

912 **Fig. 3. (A)** Seismic record section for OBH 253, plotted at a reduction velocity of 8.0 km/s to highlight crustal  
913 ( $P_g$ ) and upper ( $P_n$ ) mantle refractions. The Moho reflections ( $P_mP$ ) were used to constrain the crustal thickness  
914 (Contreras-Reyes et al., 2010). **(B)** Plotted at a reduction velocity of 4.0 km/s to highlight converted P-S-P  
915 crustal ( $PS_gP$ ) and mantle ( $PS_nP$ ) refractions. Mode converted P-S-P Moho reflections ( $PS_mSP$ ) were also  
916 recorded at some stations. In parts (A) and (B) phase identifications are annotated. Predicted travel times,  
917 represented by orange and red curves, are compared with the corresponding interpreted picks (blue curves)  
918 which are scaled to the pick uncertainty for **(C)** P-wave and **(D)** S-wave modelling. Predicted travel times are  
919 calculated using the 2-D  $V_p$  (Fig. 7A) and  $V_s$  (Fig. 7B) inversion models. Ray-tracing of the **(E)** P-wave and **(F)**  
920 converted P-S-P-waves. Green dots in (E) and (F) mark OBH/S locations.

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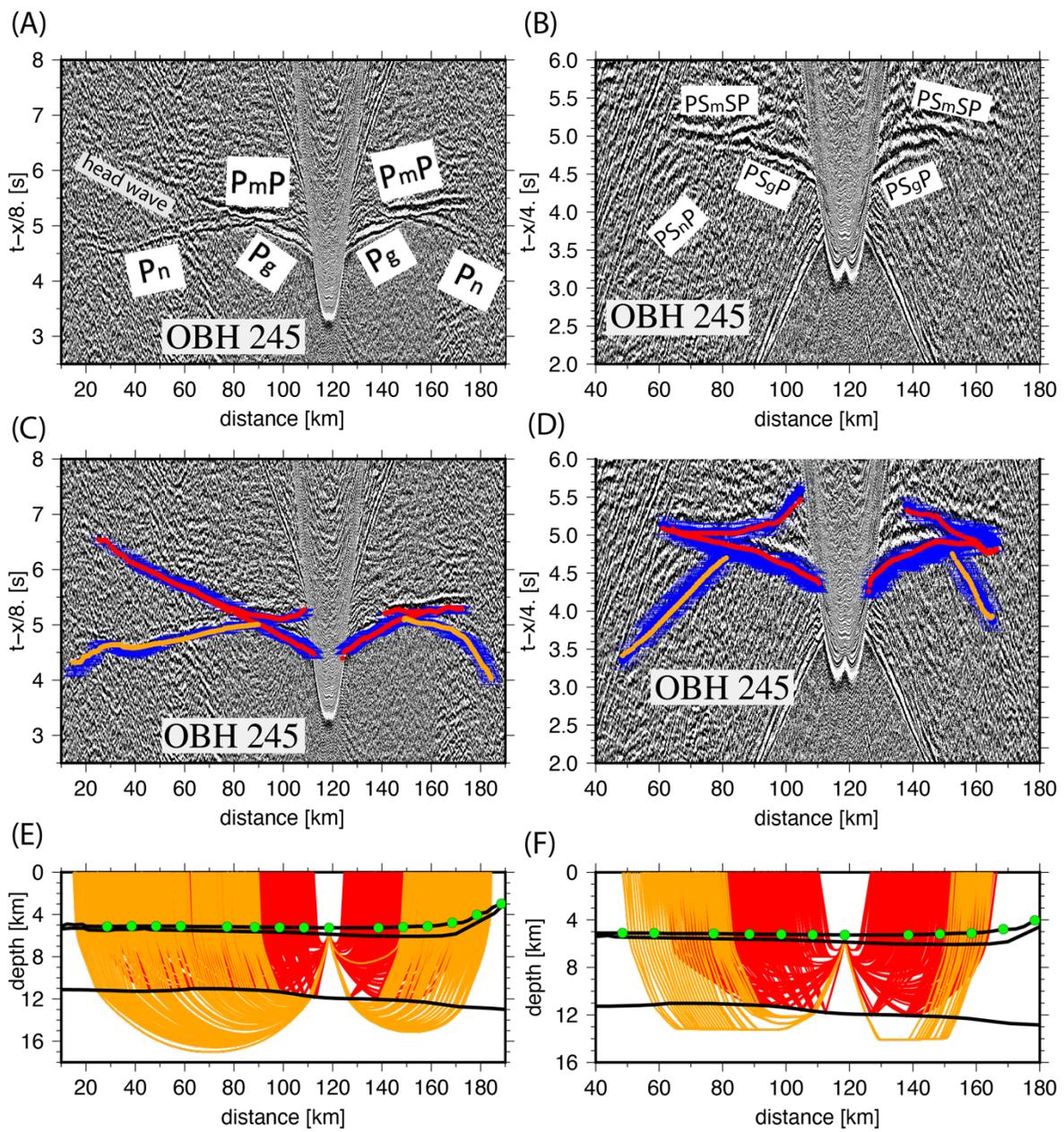
923 **Fig. 4.** (A) Seismic record section for OBH 237, plotted at a reduction velocity of 8.0 km/s to highlight crustal  
 924 ( $P_g$ ) refractions. (B) Plotted at a reduction velocity of 4.0 km/s to highlight converted P-S-P crustal refractions  
 925 ( $PS_gP$ ). In parts (A) and (B) phase identifications are annotated. Predicted travel times, represented by red  
 926 curves, are compared with the corresponding interpreted picks (blue curves) which are scaled to the pick  
 927 uncertainty for (C) P-wave and (D) S-modelling. Predicted travel times are calculated using the 2-D  $V_p$  (Fig.  
 928 7A) and  $V_s$  (Fig. 7B) inversion models. Ray-tracing of the (E) P-wave and (F) converted P-S-P-waves. Green  
 929 dots in (E) and (F) mark OBH/S locations.

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935 **Fig. 5.** (A) Seismic record section for OBH 245, plotted at a reduction velocity of 8.0 km/s to highlight crustal  
936 upper mantle phases. (B) Plotted at a reduction velocity of 4.0 km/s to highlight converted P-S-P phases. In  
937 parts (A) and (B) phase identifications are annotated. (C) Predicted travel times, represented by orange and red

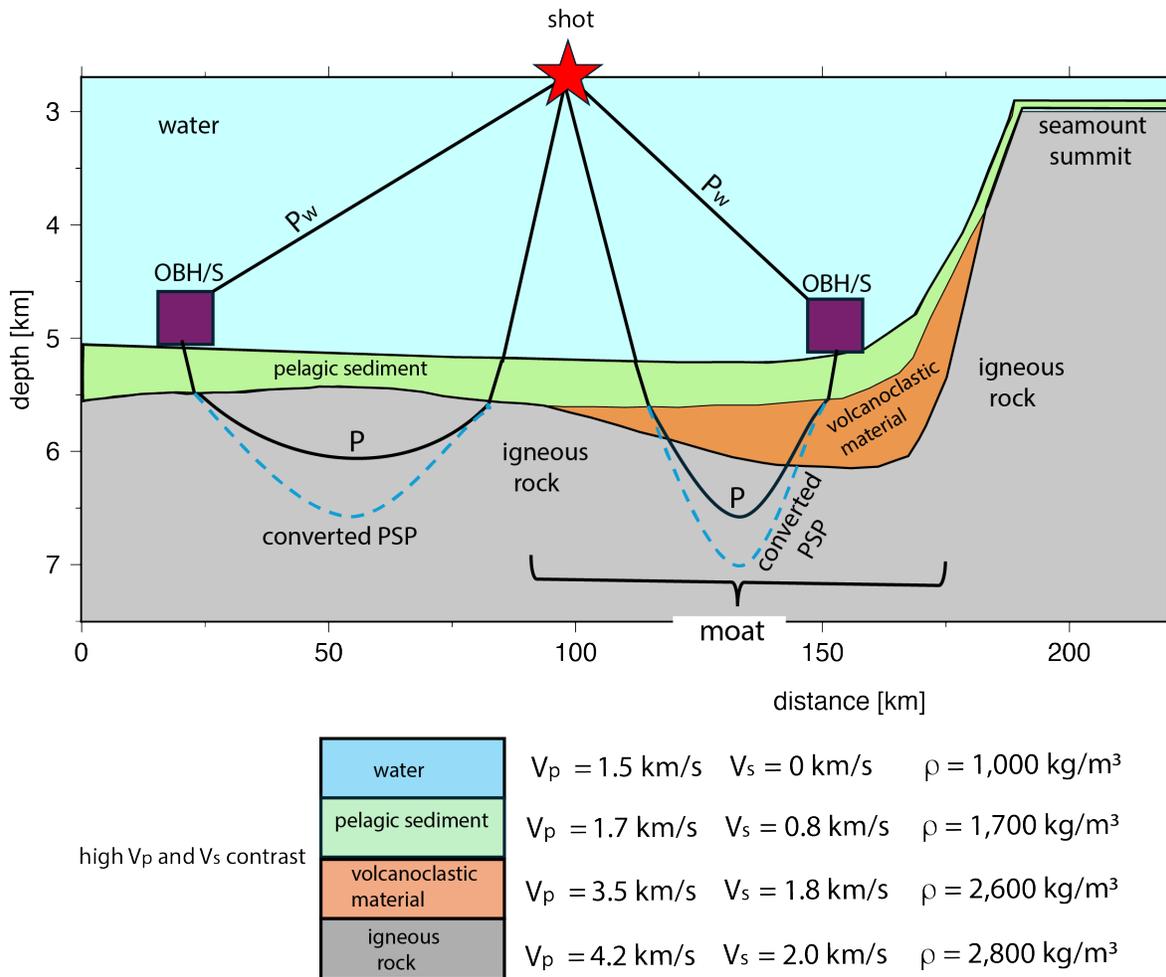
938 curves, are compared with the corresponding interpreted picks (blue curves) which are scaled to the pick  
 939 uncertainty for (C) P-wave and (D) S-modelling. Predicted travel times are calculated using the 2-D  $V_p$  (Fig.  
 940 7A) and  $V_s$  (Fig. 7B) inversion models. Ray-tracing of the (E) P-wave and (F) converted P-S-P-waves. Green  
 941 dots in (E) and (F) mark OBH/S locations.

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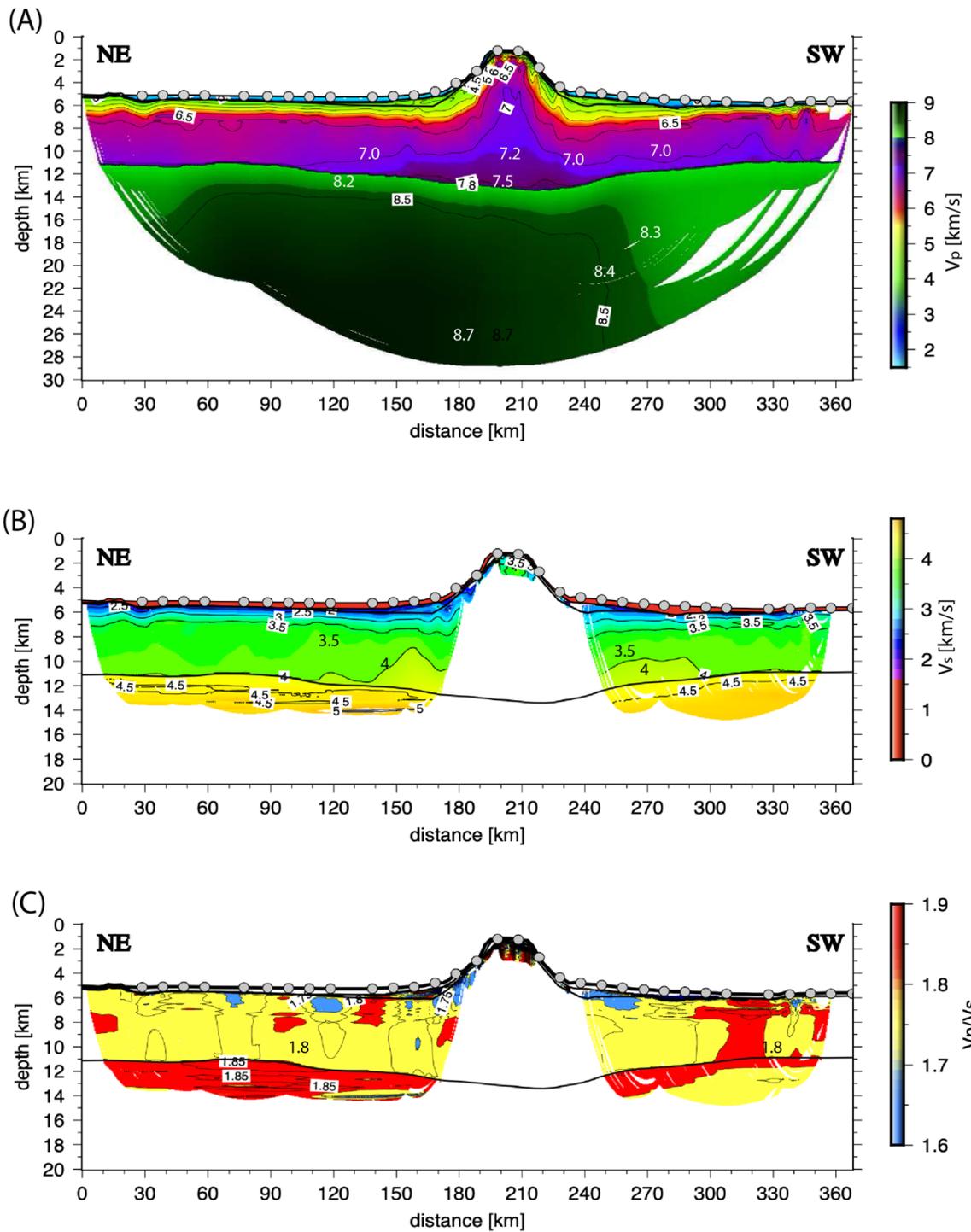
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947 **Fig. 6.** Schematic diagram showing converted P-S-P-waves. Reference values for  $V_p$  and  $V_s$ , and densities ( $\rho$ )  
 948 are shown to highlight the higher velocity and density contrast at the sediment/basement interface compared  
 949 with the water/sediment interface. Typical  $V_p$ ,  $V_s$ , and  $\rho$  values are based on Fulthorpe et al. (1989), Caress et

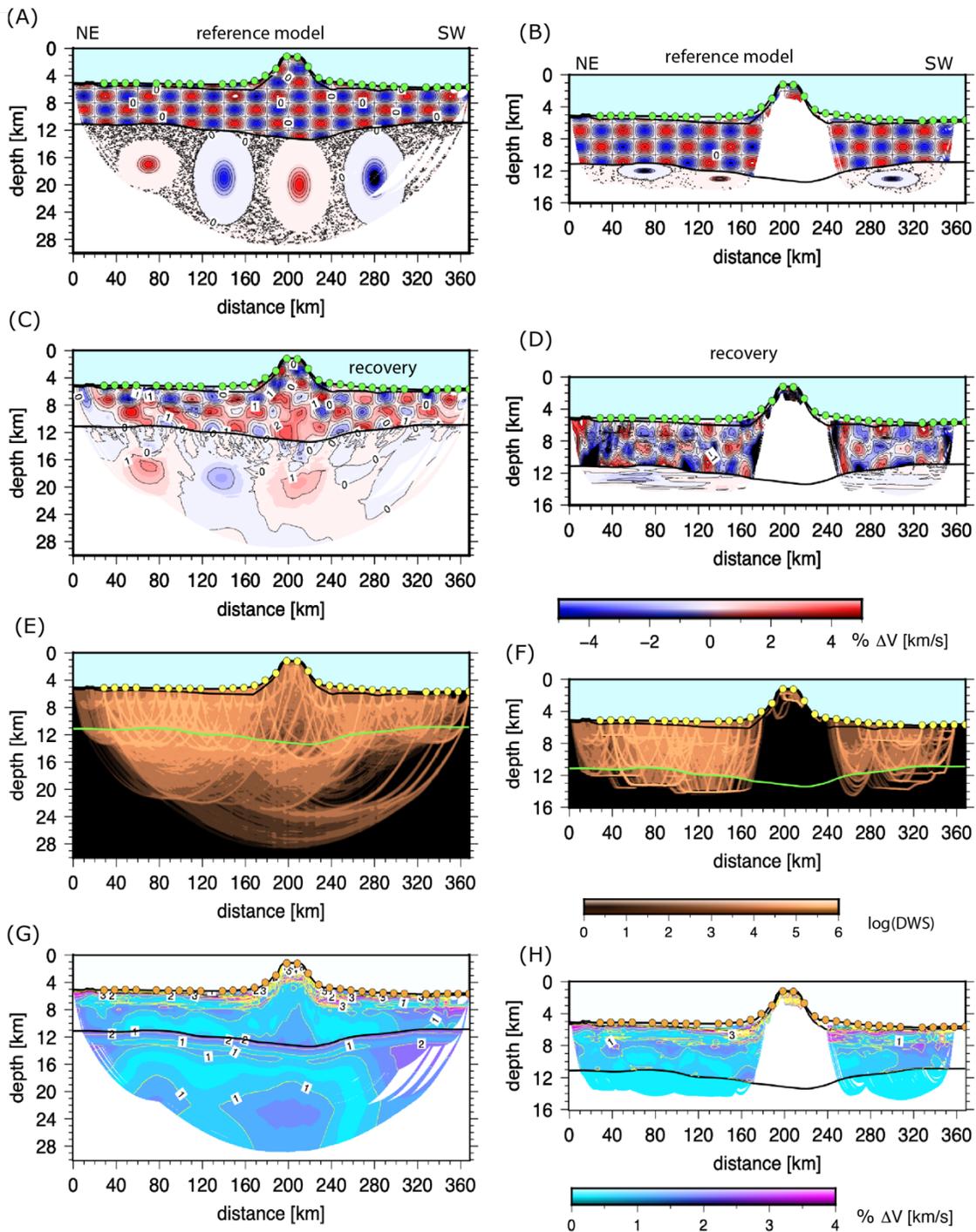
950 al. (1995) and Christeson et al. (2019). Adjacent to the flexural moat, the P-to-S-wave conversion is more likely  
951 to occur at the pelagic sediment/igneous basement interface, while in the flexural moat it is more likely to occur  
952 at the pelagic sediment/volcanoclastic interface that maximizes the acoustic impedance contrast. Seismic energy  
953 propagates as a P-wave through the water and pelagic sediment layer (both upward and downward) and as an  
954 S-wave through the volcanoclastic material, crust, and mantle. P<sub>w</sub>: direct P-wave through the water.



955

956 **Fig. 7.** Final (A) P-wave and (B) S-wave velocity tomographic inversion models masked by the ray coverage  
 957 (see Fig. 8 and Supporting Information). (C) Resultant 2-D  $V_p/V_s$  model derived from (A) and (B).

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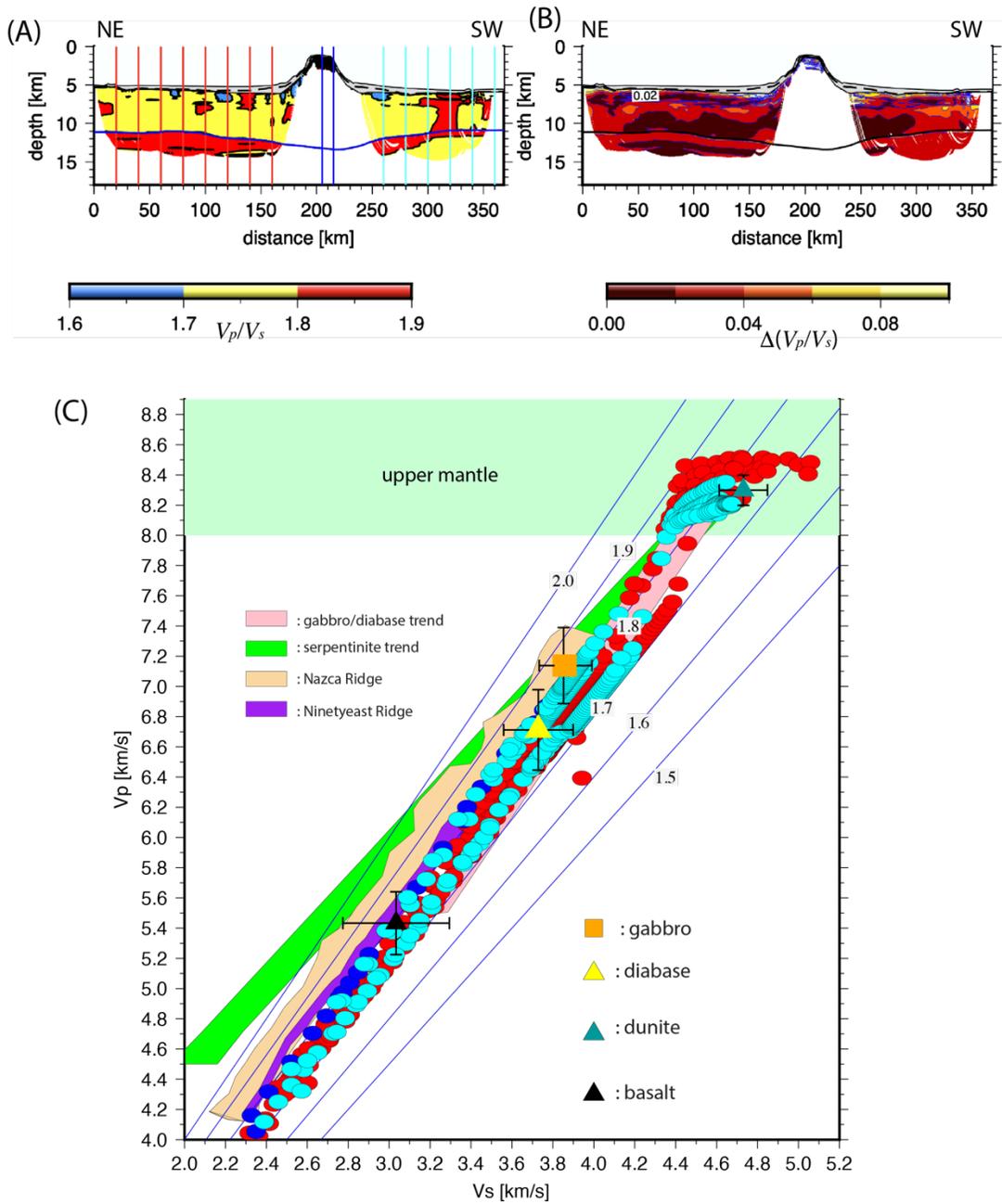


959

960 **Fig. 8.** Results of checkerboard testing. Synthetic reference velocity model for (A)  $V_p$  and (B)  $V_s$ , consisting of  
 961 velocity anomalies of 40 km x 4 km with maximum velocity amplitudes of  $\pm 5\%$  for the oceanic crust. For the  
 962 upper mantle, we use Gaussian anomalies superimposed onto the final velocity model shown in Figs. 7A and  
 963 7B. Recovery for the (C)  $V_p$  and (D)  $V_s$  models. See Supporting Information for details and further resolution

964 tests. Derivative Weight Sum (DWS) in logarithmic scale for rays traveling through the (E)  $V_p$  and (F)  $V_s$   
 965 models. Percentage standard deviation of the (G)  $V_p$  and (H)  $V_s$  models. See enlarged version of these figures  
 966 in the Supporting Information.

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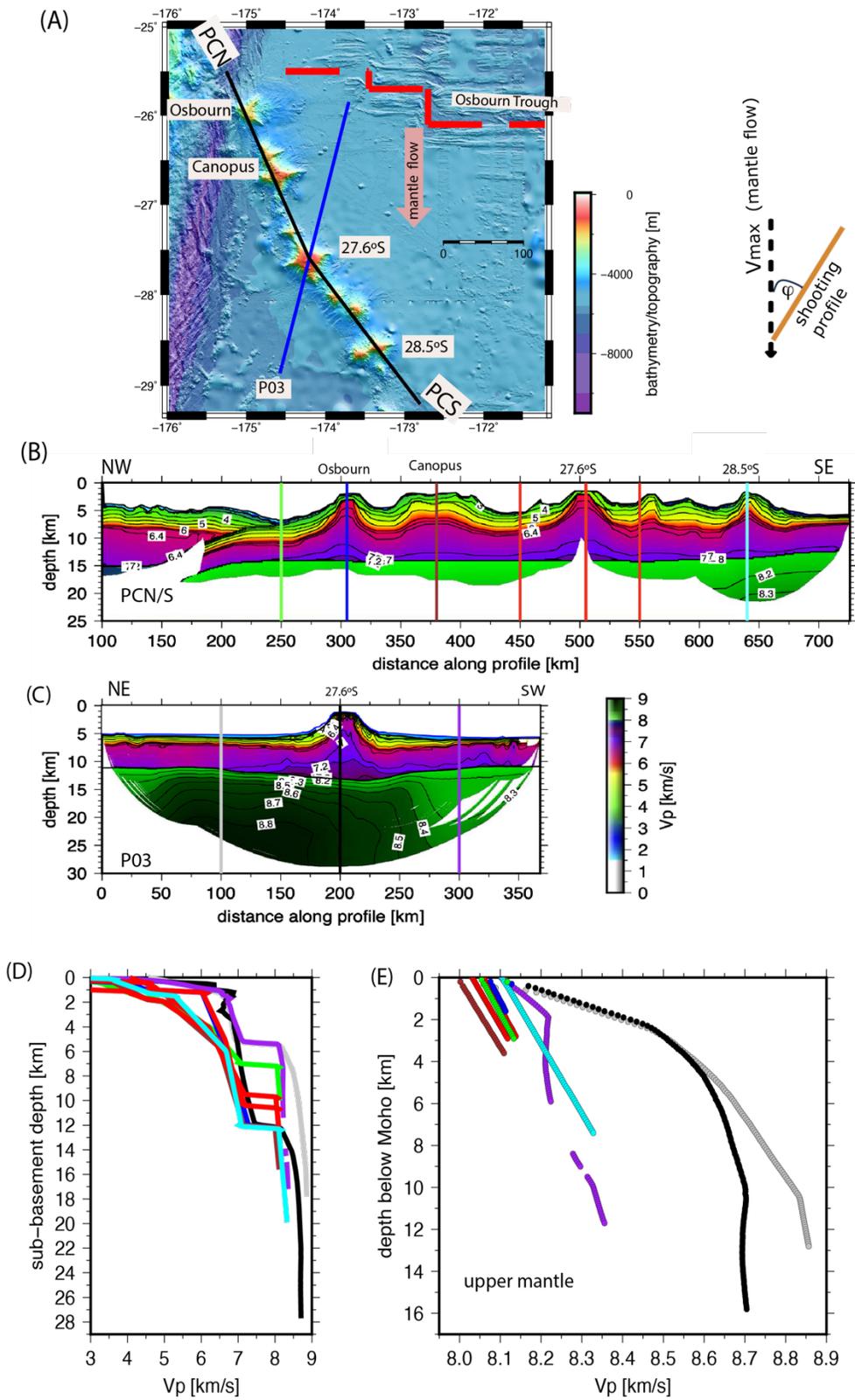


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969 **Fig. 9.** (A) 2-D  $V_p/V_s$  model and its (B)  $\Delta\left(\frac{V_p}{V_s}\right)$  uncertainties (see Supporting Information). (C) Paths of constant  
 970  $V_p/V_s$  ratio (blue lines), which form straight lines in the  $V_p$  -  $V_s$  plane, are plotted at intervals of 0.1. The  $V_p$  and

971  $V_s$  values for gabbro, diabase, dunite, and basalt were determined at 200 MPa by Christensen (1996). The trends  
972 of serpentinite and gabbro/diabase resulted from the optimal alignment of in-situ seismic velocities and  
973 laboratory measurements, as published by Carlson and Miller (2003). The paired ( $V_p$ ,  $V_s$ ) values, depicted as  
974 red, blue, and cyan ellipses, were extracted from the 2-D  $V_p$  and  $V_s$  models at the corresponding horizontal  
975 locations indicated in (A). The semi-minor and semi-major axes of each ellipse is taken as the average  $V_p$  and  
976  $V_s$  uncertainty (see Figs. 8G and 8H).

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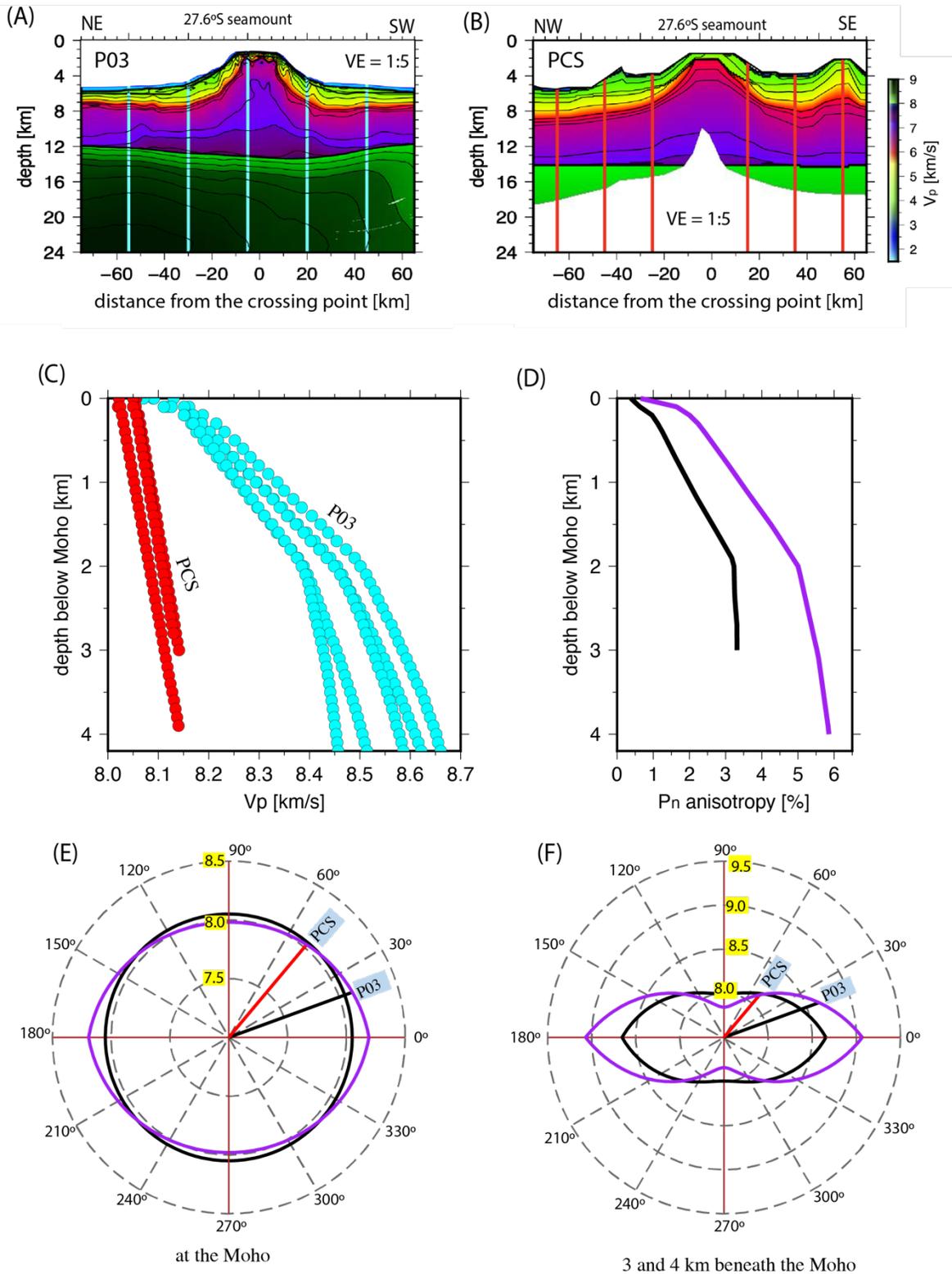
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**Fig. 10.** (A) Location map for profiles PC (annotated in south and north parts – PCS and PCN; Robinson et al., 2018) and P03 (Fig. 1). Seismic profile P03 is oriented at an angle ( $\phi$ )  $\sim 20^\circ$  relative to the expected direction of

981 mantle flow, while profile PCS runs at an angle of  $\sim 50^\circ$ .  $V_p$  models along **(B)** PCS (Robinson et al., 2018) and  
982 **(C)** P03.  $V_p$  profiles plotted as a function depth below **(D)** basement and **(E)** Moho, respectively, for the models  
983 shown in **(B)** and **(C)**.

984



985

986 **Fig. 11.** 2-D  $V_p$  model along (A) P03 and (B) PCS (Fig. 10A; Robinson et al., 2018) beneath the 27.6°S  
 987 seamount. (C)  $V_p$  profiles plotted as a function of depth below the Moho for comparison. (D) Maximum (purple

988 curve) and minimum (black curve)  $P_n$  anisotropic percentage (Eq. 2) as a function of depth. The anisotropic  
989 percentage was calculated using the bounding  $V_p$ -depth profiles shown in (C). **(E)**  $P_n$  velocity measured along  
990 profiles P03 and PCS at Moho depth showing the minimum (black ellipse) and maximum (purple ellipse)  
991 anisotropy. **(F)** The minimum  $P_n$  anisotropic percentage at 3 km below the Moho (black ellipse), and maximum  
992  $P_n$  anisotropic percentage at 4 km below the Moho (purple ellipse) are illustrated. For (E) and (F), profiles P03  
993 and PCS have an azimuth of  $\sim 20^\circ$  ( $\varphi_1$ ) and  $\sim 50^\circ$  ( $\varphi_2$ ) relative to the paleo spreading direction, respectively (see  
994 Fig. 10A). At a specified depth, the  $P_n$  velocity measurements at azimuths  $\varphi_1$  and  $\varphi_2$  enable the determination  
995 of the faster  $V_{max}$ , assumed to align parallel to the paleo spreading direction, and the slower  $V_{min}$ , oriented  
996 perpendicular to it (see Eqs. 3-5).

Supporting Information for

**$V_p/V_s$  structure and  $P_n$  anisotropy across the Louisville Ridge, seaward of the Tonga-Kermadec Trench**

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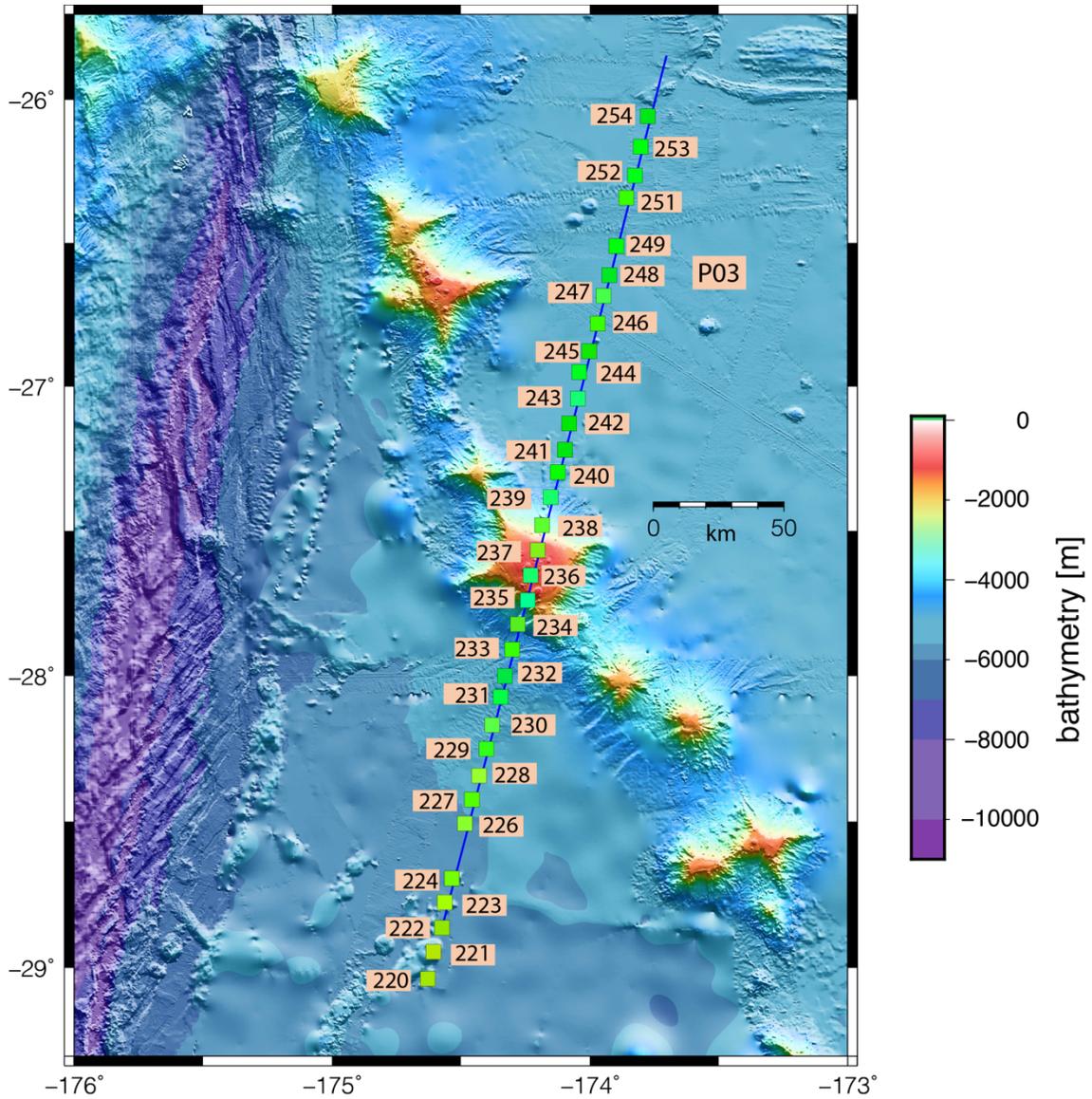
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<sup>2</sup>GEOMAR-Helmholtz Centre for Ocean Research, Kiel, Germany

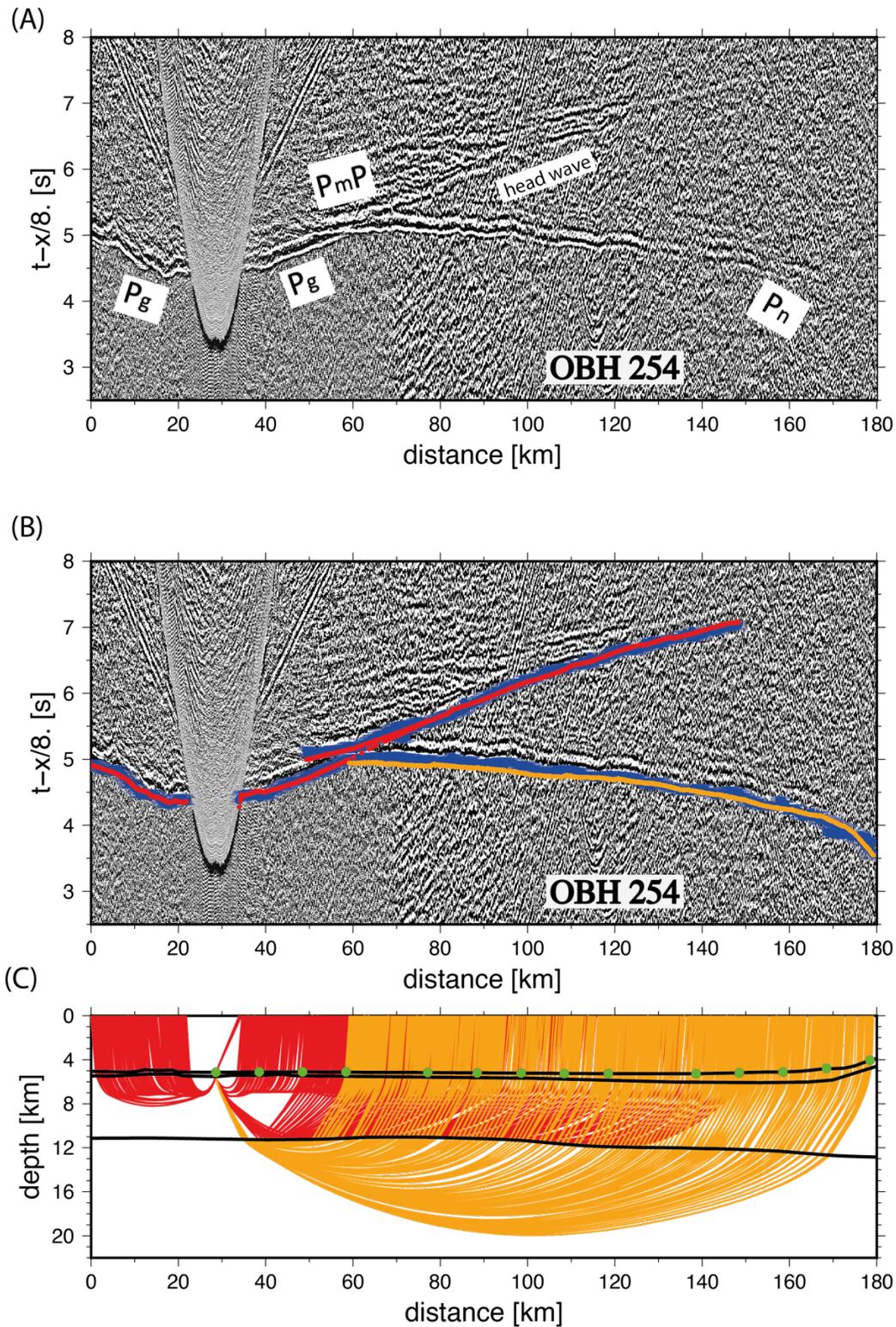
<sup>3</sup>Department of Earth Sciences, Durham University, South Road, Durham, DH1 3LE, UK

## 1. Introduction

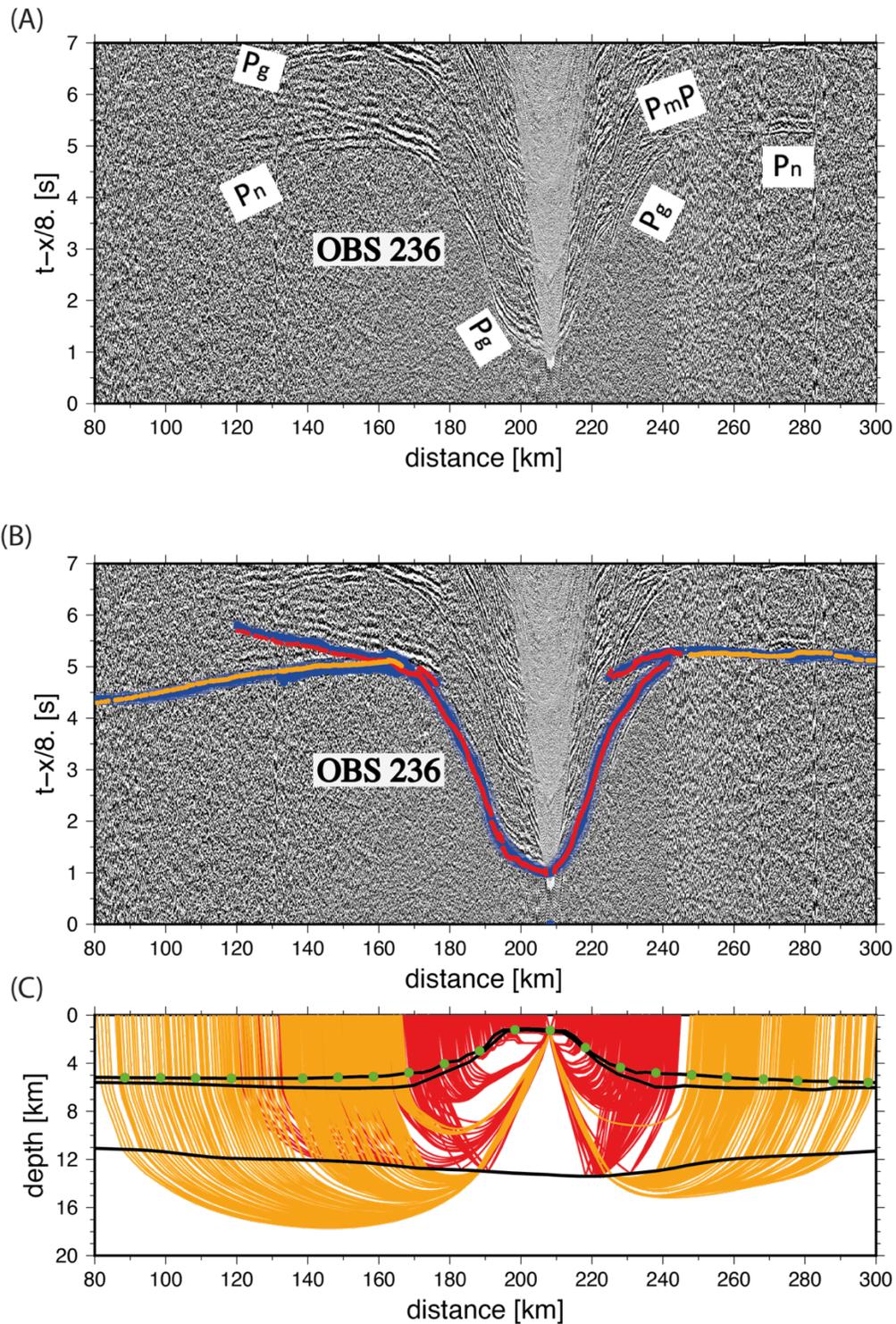
We present example seismic record sections for ocean-bottom hydrophones/seismographs (OBH/S) deployed along profile P03 which obliquely crosses the Louisville Ridge and is orientated sub-parallel to the Tonga-Kermadec Trench axis. For this experiment, we deployed 33 OBH/Ss along the ~368 km-long profile, each with a spacing of ~9 km (Fig. S1). The seismic source comprised a system of G-gun clusters with a total volume of 84 l which was operated at 3000 psi. The shot interval was 60 s (Grevemeyer and Flueh, 2008). Most of the instrument recordings are of good quality, showing clear signals at large offsets. Both P-wave and mode-converted P-to-S-to-P-waves have been recorded. In the following figures we show example seismic record sections with identified phases annotated, uncertainty estimates for both  $V_p$  and  $V_s$  models, resolution tests, and the density of ray coverage.



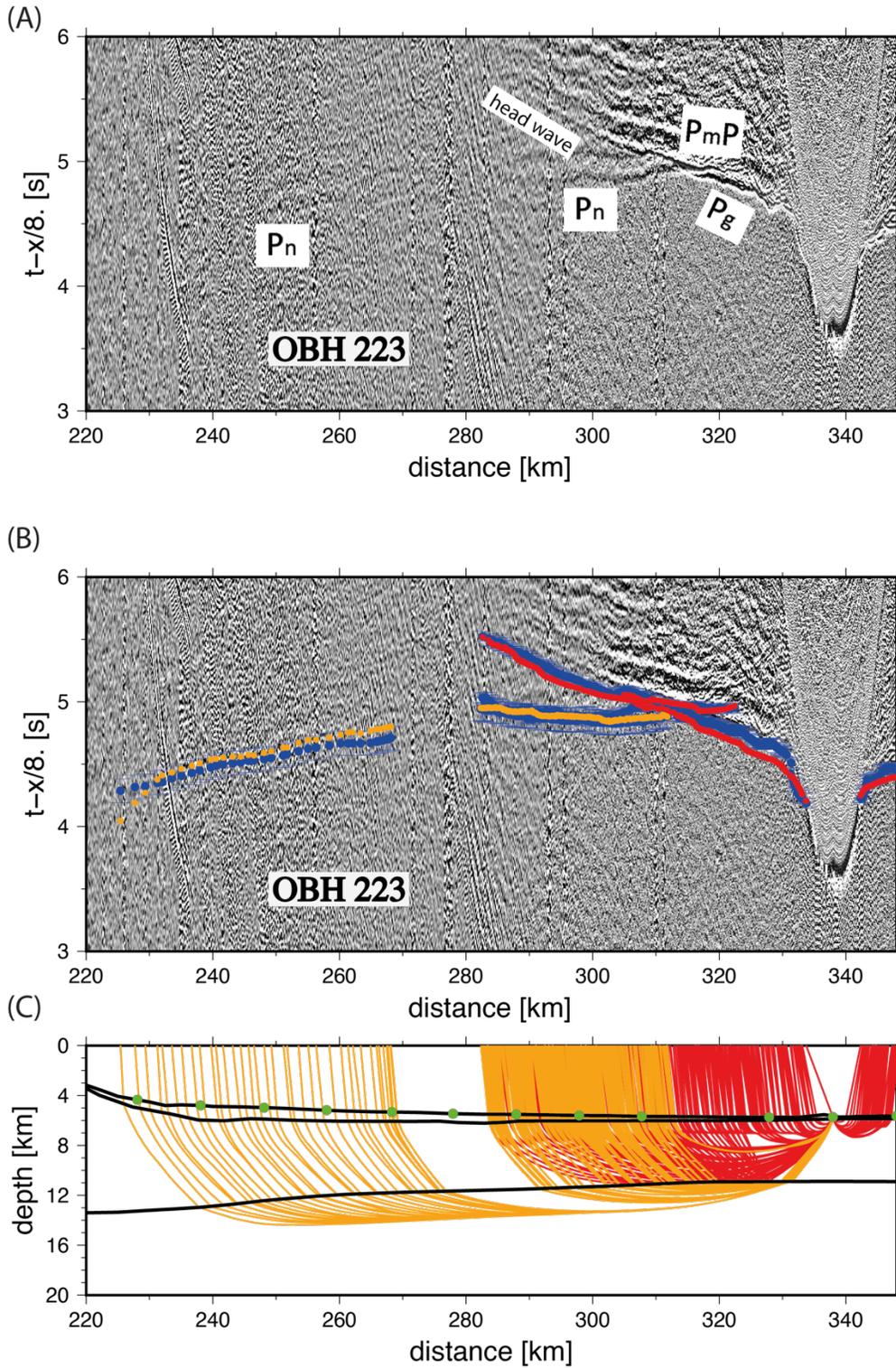
**Fig. S1.** Seismic profile P03, oriented sub-parallel to the Tonga-Kermadec Trench axis and crosses the Louisville Ridge at 27.6°S seamount. Airgun shots are marked by the blue solid line and the locations of deployed OBH/Ss are marked by green dots with the station number annotated.



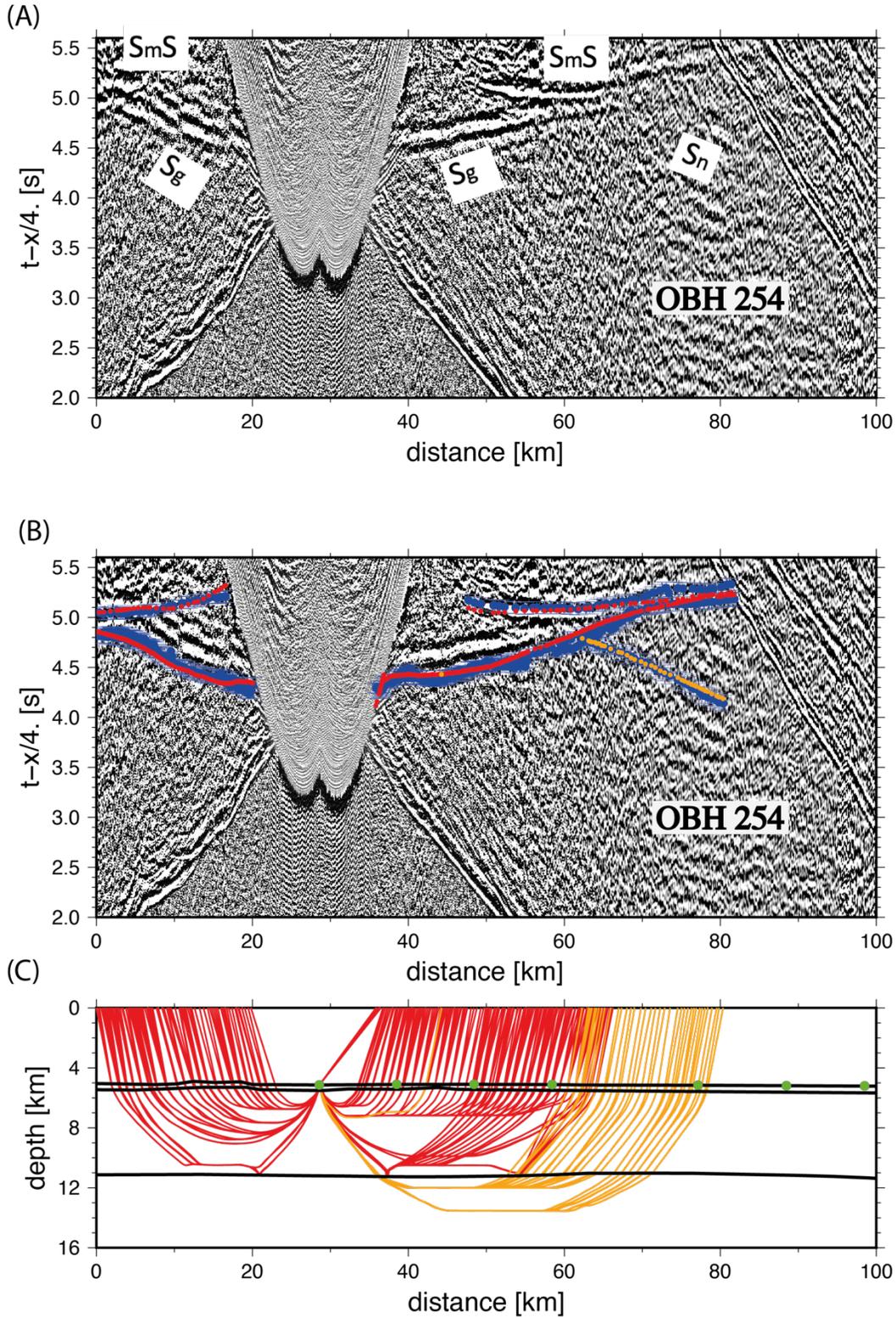
**Fig. S2.** (A) Wide-angle seismic data from OBH 254, plotted at a reduction velocity of 8 km/s and with crustal and mantle P-wave phases identified. (B) Predicted travel times are based on the 2-D  $V_p$  model shown in Fig. 7A. Blue bars corresponds to the travel time picks, scaled to the pick uncertainty. (C) Ray path diagram corresponding to the predicted travel times shown in (B). Green dots mark OBH/S locations.



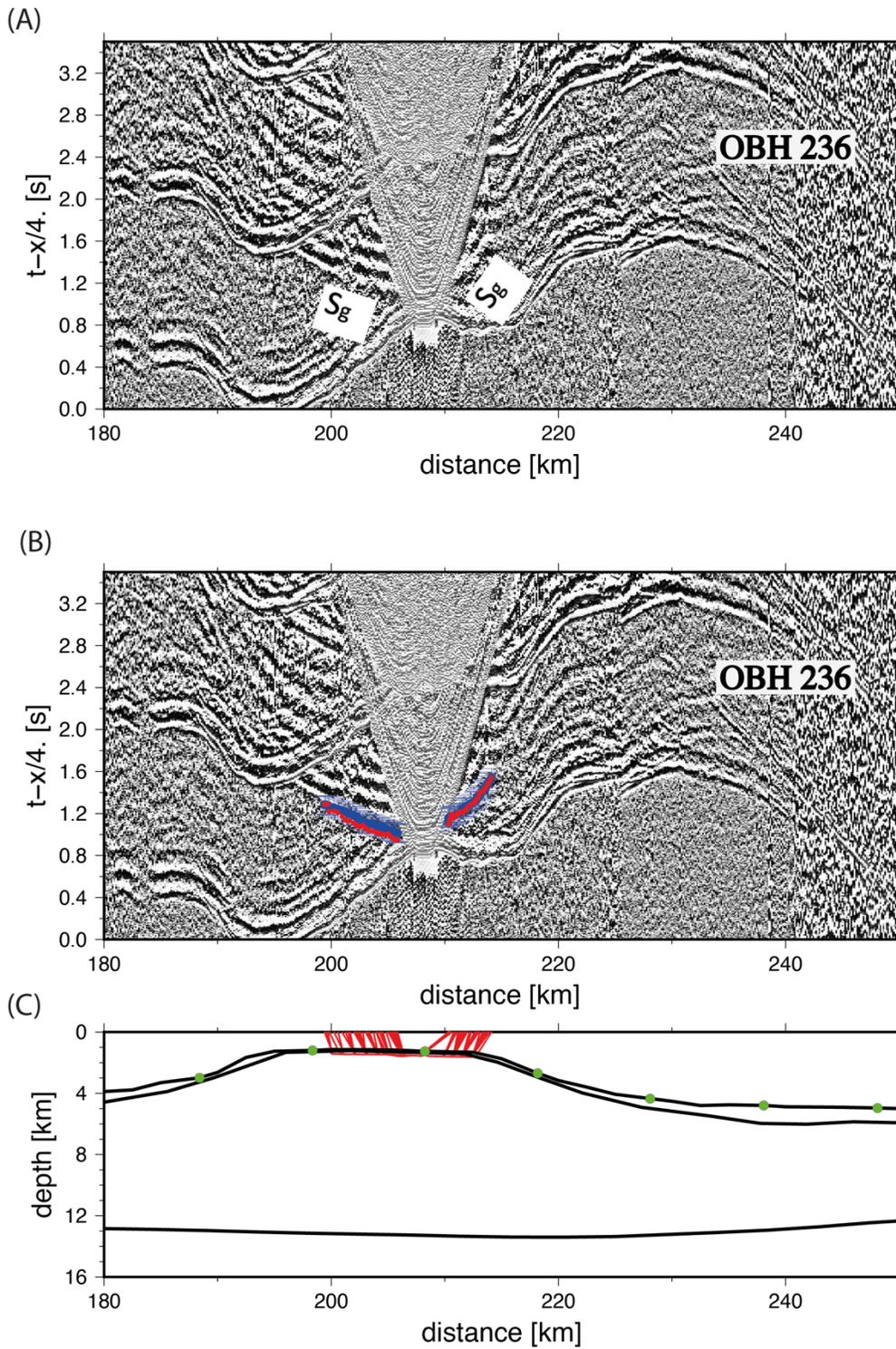
**Fig. S3.** (A) Wide-angle seismic data from OBH 236, plotted at a reduction velocity of 8 km/s and with crustal and mantle P-wave phases identified. (B) Predicted travel time are based on the 2-D  $V_p$  model shown in Fig. 7A. Blue bars corresponds to the travel time picks, scaled to the pick uncertainty. (C) Ray path diagram corresponding to the predicted travel times shown in (B). Green dots mark OBH/S locations.



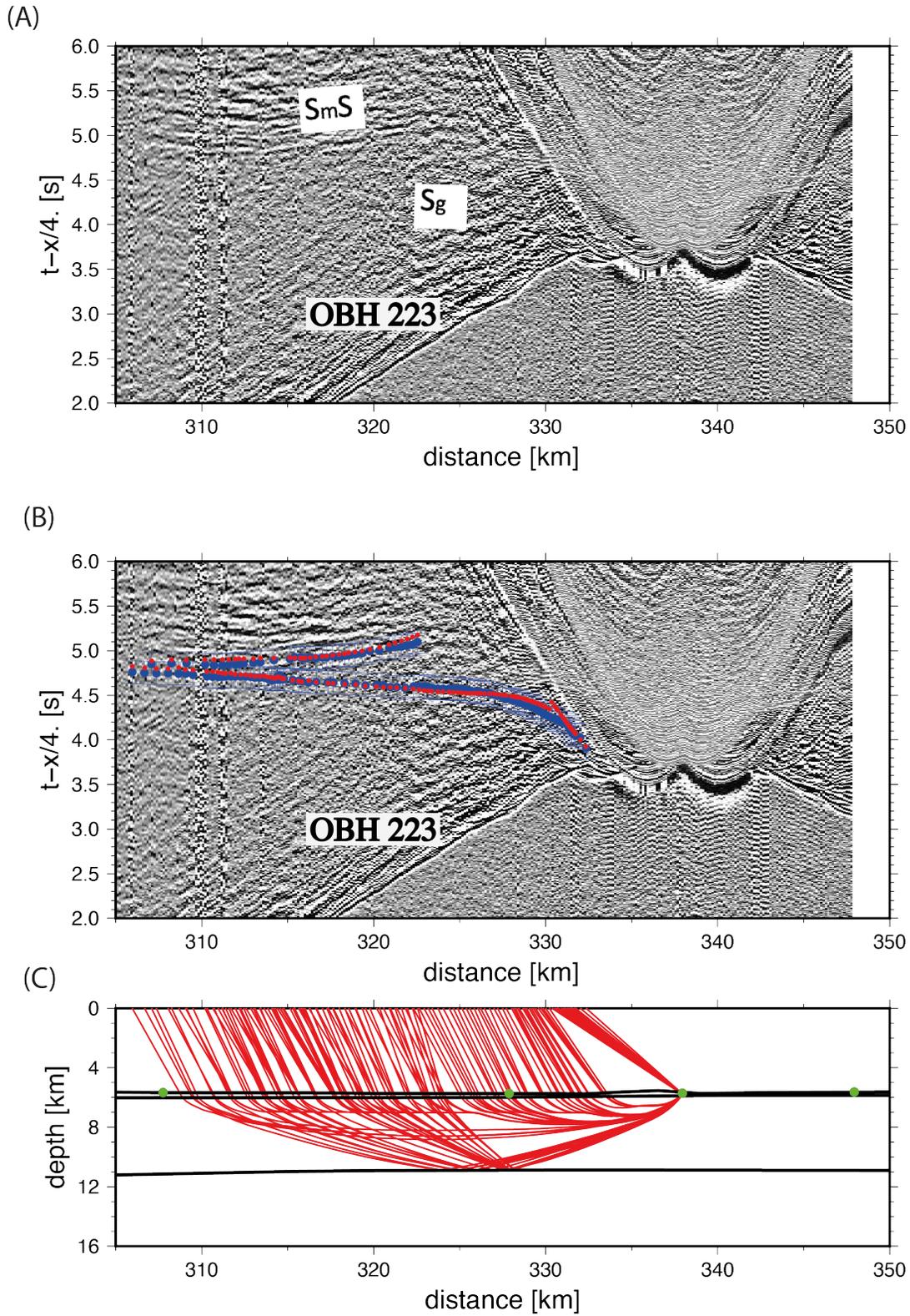
**Fig. S4.** (A) Wide-angle seismic data from OBH 223, plotted at a reduction velocity of 8 km/s and with crustal and mantle P-wave phases identified. (B) Predicted travel time are based on the 2-D  $V_p$  model shown in Fig. 7A. Blue bars corresponds to the travel time picks, scaled to the pick uncertainty. (C) Ray path diagram corresponding to the predicted travel times shown in (B). Green dots mark OBH/S locations.



**Fig. S5.** (A) Wide-angle seismic data from OBH 254, plotted at a reduction velocity of 4 km/s and with converted S-wave phases identified. (B) Predicted travel time are based on the 2-D  $V_s$  model shown in Fig. 7B. Blue bars corresponds to the travel time picks, scaled to the pick uncertainty. (C) Ray path diagram corresponding to the predicted travel times shown in (B). Green dots mark OBH/S locations.



**Fig. S6.** (A) Wide-angle seismic data from OBH 236, plotted at a reduction velocity of 4 km/s and with converted S-wave phases identified. (B) Predicted travel time are based on the 2-D  $V_s$  model shown in Fig. 7B. Blue bars corresponds to the travel time picks, scaled to the pick uncertainty. (C) Ray path diagram corresponding to the predicted travel times shown in (B). Green dots mark OBH/S locations.



**Fig. S7.** (A) Wide-angle seismic data from OBH 223, plotted at a reduction velocity of 4 km/s and with converted S-wave phases identified. (B) Predicted travel time are based on the 2-D  $V_s$  model shown in Fig. 7B. Blue bars corresponds to the travel time picks, scaled to the pick uncertainty. (C) Ray path diagram corresponding to the predicted travel times shown in (B). Green dots mark OBH/S locations.

## 2. Seismic Modeling

The 2-D velocity structure that best fits the observed arrivals was determined using a joint refraction and reflection travel time inversion technique (Korenaga et al., 2000). This method allows the joint inversion of seismic refraction and reflection travel time data for a 2-D velocity field. Travel times and ray paths are calculated using a hybrid ray-tracing scheme based on the graph method and the local ray-bending refinement (van Avendonk et al., 1998). Smoothing constraints using pre-defined correlation lengths and optimized damping constraints for the model parameters are employed to regularize an iterative linearized inversion (Korenaga et al., 2000).

The best-fitting velocity model consists of the following layers: (1) water; (2) sediment/debris on top of the ridge; and (3) igneous crustal basement. To derive the velocity-depth model, the water depth was taken from the swath bathymetry center beam, and remained fixed during the inversion. The pelagic sedimentary layer was picked from the seismic reflection profile shown in Fig. 2 and then converted to depth assuming  $V_p = 1.7$  km/s. The base of the flexural moat was taken from the final  $V_p$  tomographic model of Contreras-Reyes et al. (2010). To obtain the crustal P-wave velocities and the Moho depths, we jointly invert crustal refractions ( $P_g$ ) and reflections ( $P_mP$ ). To obtain the final model, we then inverted the mantle refractions ( $P_n$ ). Moho depth is held fixed in the inversion of S-wave travel time picks.

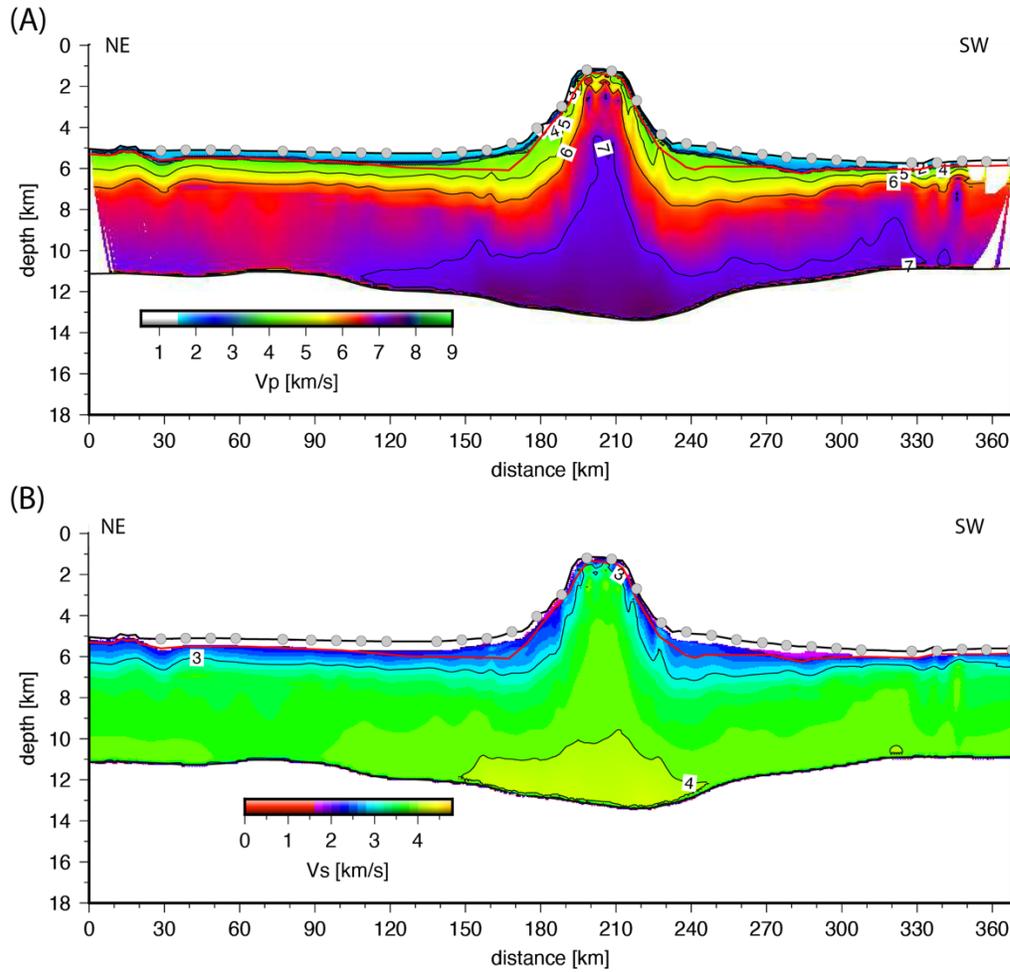
## 3. Initial Model and Inversion Parameters

The 2-D velocity model is ~368 km long and 30 km deep. The initial model for P-wave tomographic inversion is based on the model of Contreras-Reyes et al. (2010). For the S-wave tomographic inversion, we based our initial model on the final P-wave model and converted it to S-wave velocity using a  $V_p/V_s$  ratio of 1.8 (Fig. S8). For both P-wave and S-wave velocity models, the horizontal grid spacing is 1 km, whereas the vertical grid spacing varies from 0.1 km at the top of the model to 0.15 km at the bottom. The horizontal correlation lengths range from 2 km at the top to 5 km at the bottom of the model, and the vertical correlation lengths vary from 0.5 km to 2 km, respectively. Depth nodes defining the Moho reflector are spaced at 4.5 km, with a correlation length of 4 km. The final average P-wave tomographic model is shown in Fig. 7A and the S-wave tomographic model in Fig. 7B. The root mean square (RMS) misfits are shown in Table S1.

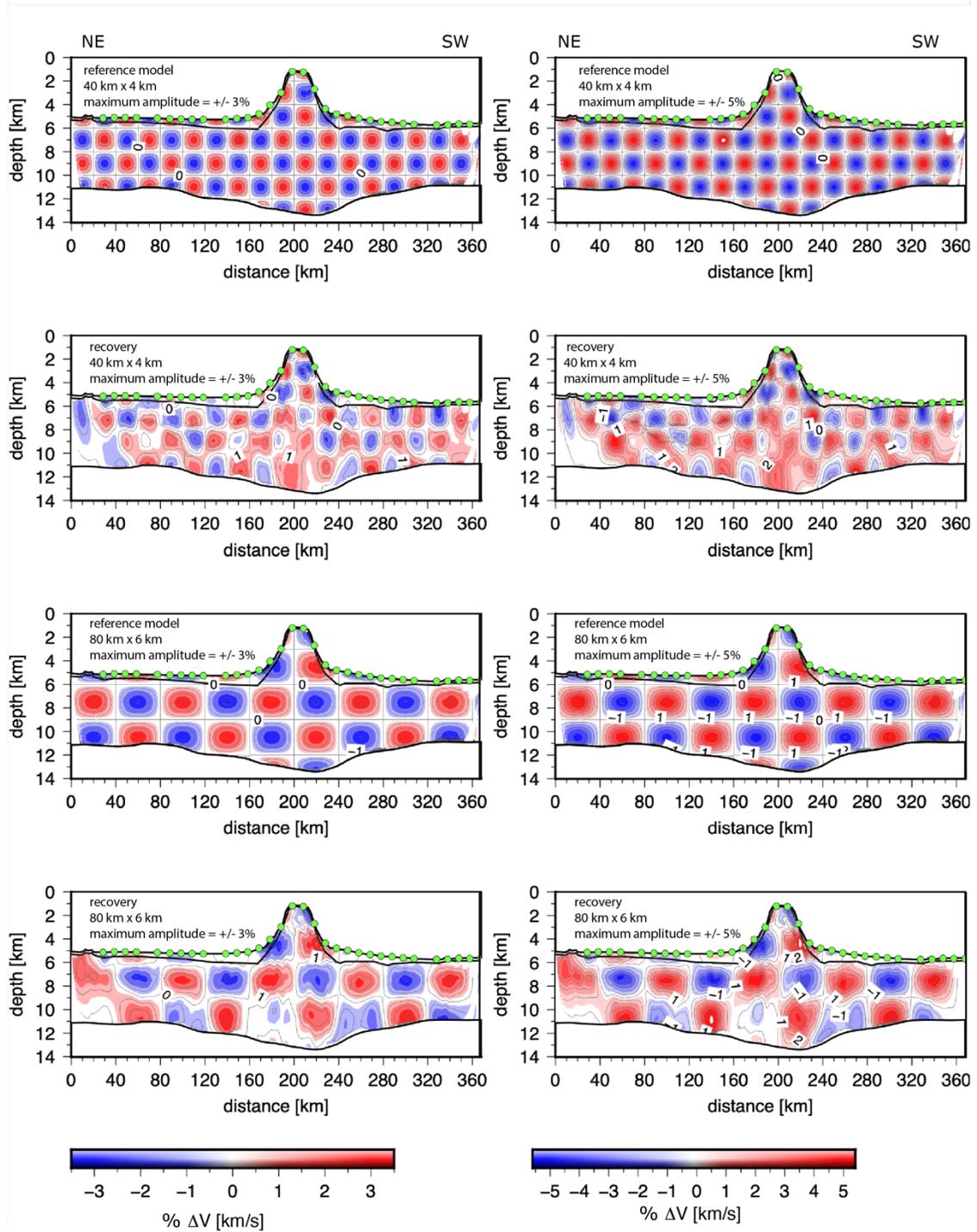
Figs. S9-S14 show different checkerboard tests for the  $V_p$  and  $V_s$  models described in the main text in Section 3.3. Fig. S15 shows the Derivative Weight Sum (a proxy for the ray density), while Figs. S16 and S17 show the velocity uncertainties (as percentages) for the  $V_p$  and  $V_s$  models described in the main text in Section 3.4. Finally, Fig. S18 shows the 2-D  $V_p/V_s$  model and its percentage standard deviation.

Phase	Average travel time uncertainty (ms)	Final model $T_{\text{RMS}}$ (ms)	Final model $\chi^2$
$P_g + P_mP$	50-70	71	1.41
$P_n$	60	96	2.53
$S_g + S_mS$	70 – 90	82	1.89
$S_n$	70	129	4.61

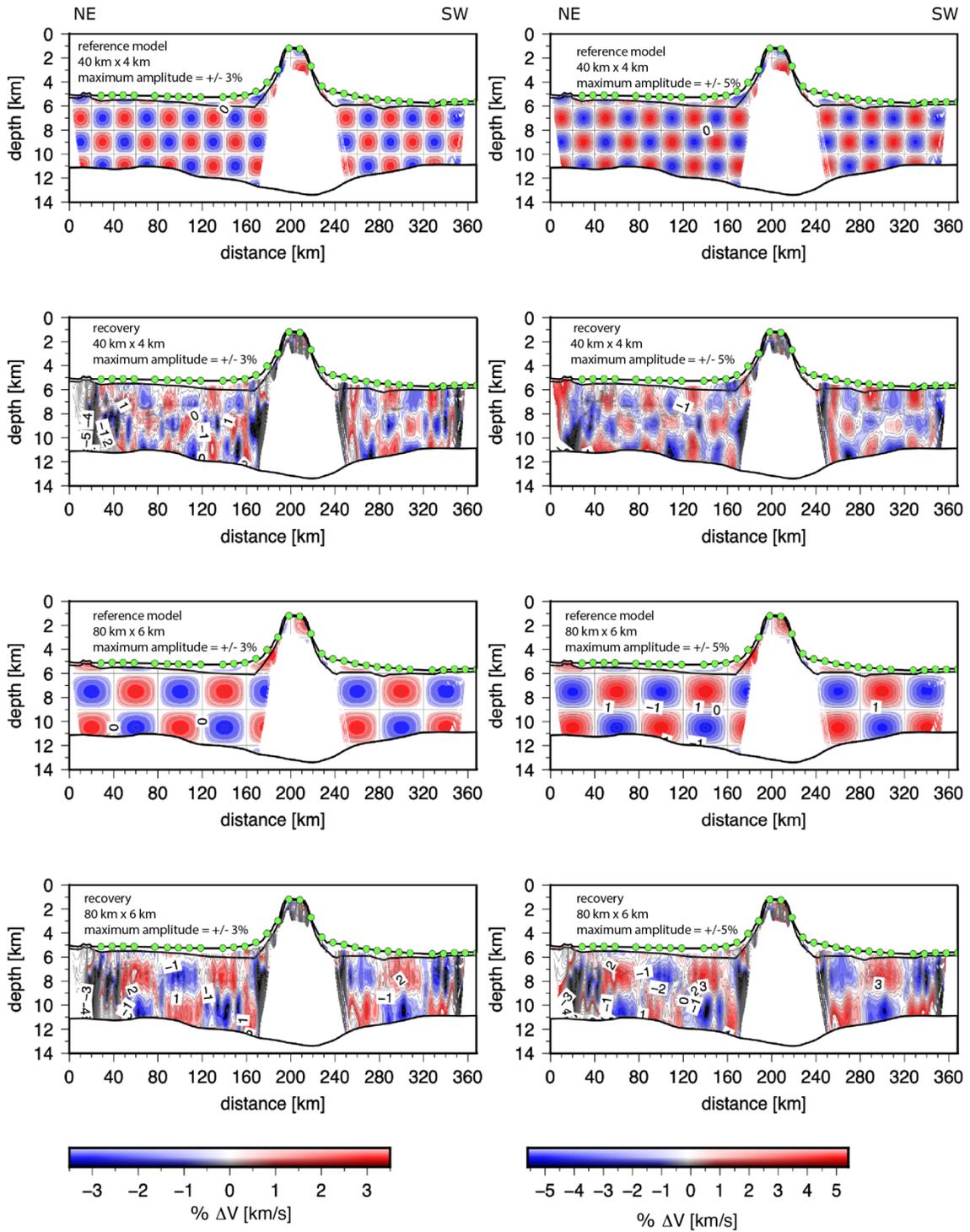
**Table S1.** Travel-time misfit and data uncertainty for seismic profile P03.  $P_g$ : P-wave crustal refraction.  $P_mP$ : P-wave Moho reflection.  $P_n$ : P-wave mantle refraction.  $S_g$ : crustal S-wave refraction.  $S_mS$ : S-wave Moho reflection.  $S_n$ : S-wave mantle refraction.  $T_{\text{RMS}}$ : root-mean-square travel time misfit.  $\chi^2$ : chi-square parameter.



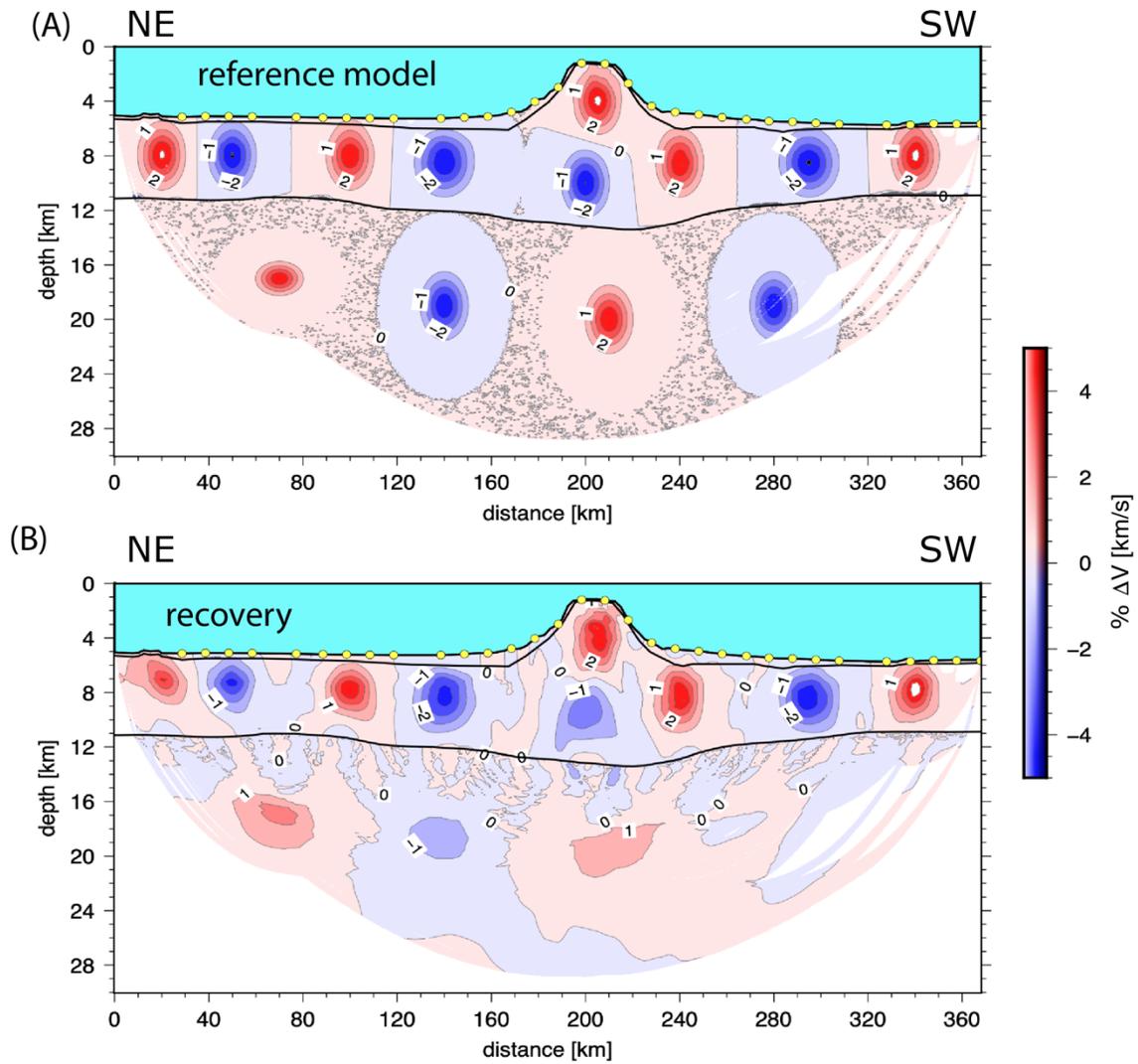
**Fig. S8.** (A) Two-dimensional  $V_p$  tomographic inversion model for the oceanic crust using  $P_g$  and  $P_mP$  phases (see Section 3.2). (B) Initial  $V_s$  tomographic model for the inversion of crustal S-wave travel time picks based on the  $V_p$  model shown in (A) and a  $V_p/V_s$  ratio of 1.8. Gray dots mark OBH/S locations.



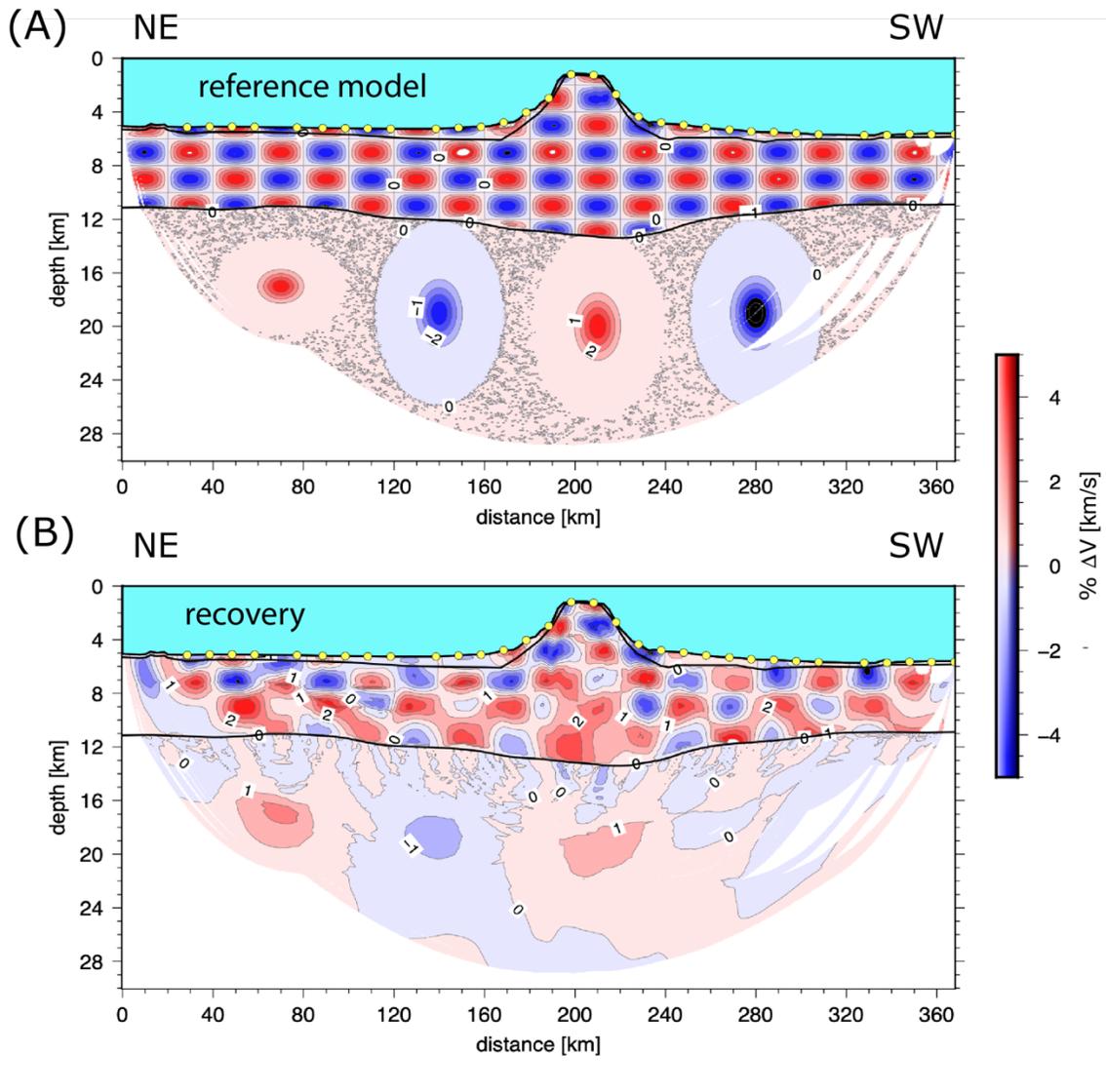
**Fig. S9.** Results of checkerboard tests for the oceanic crust of the  $V_p$  model. The synthetic reference velocity models have  $\pm 3\%$  and  $\pm 5\%$  velocity anomalies of different sizes (as annotated) applied. Green dots mark OBH/S locations.



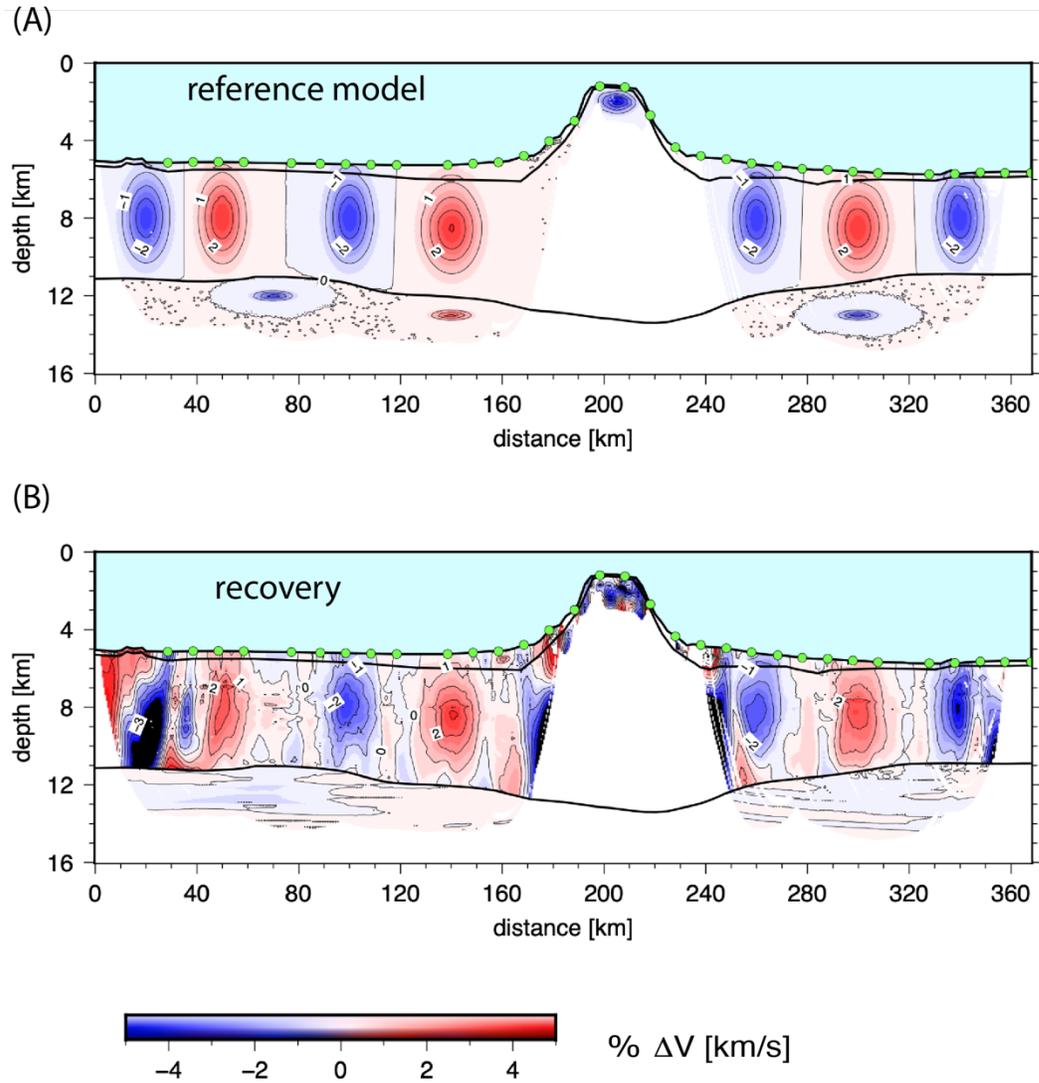
**Fig. S10.** Results of checkerboard tests for the oceanic crust of the  $V_s$  model. The synthetic reference velocity models have  $\pm 3\%$  and  $\pm 5\%$  velocity anomalies of different sizes (as annotated) applied. Green dots mark OBH/S locations.



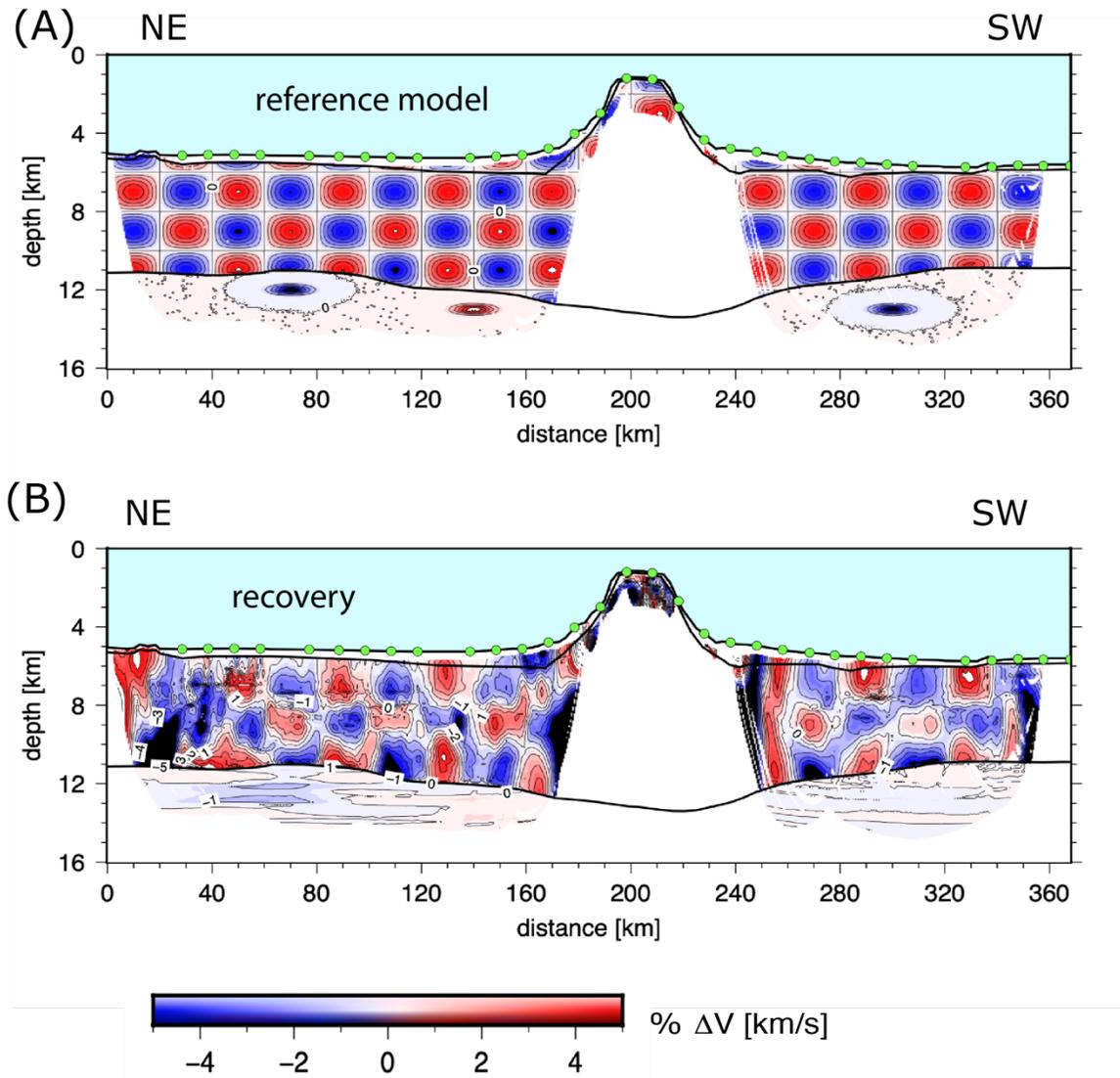
**Fig. S11.** (A) Synthetic  $V_p$  reference velocity-depth model consisting of of  $\pm 5\%$  Gaussian anomalies ( $40 \text{ km} \times 8 \text{ km}$  and  $40 \text{ km} \times 14 \text{ km}$  in the crust and mantle, respectively) superimposed onto the final P-wave velocity model shown in Fig. 7A. (B) Recovery for the  $V_p$  model using the same instrument geometry and ray density (see Section 3.3). Yellow dots mark OBH/S locations.



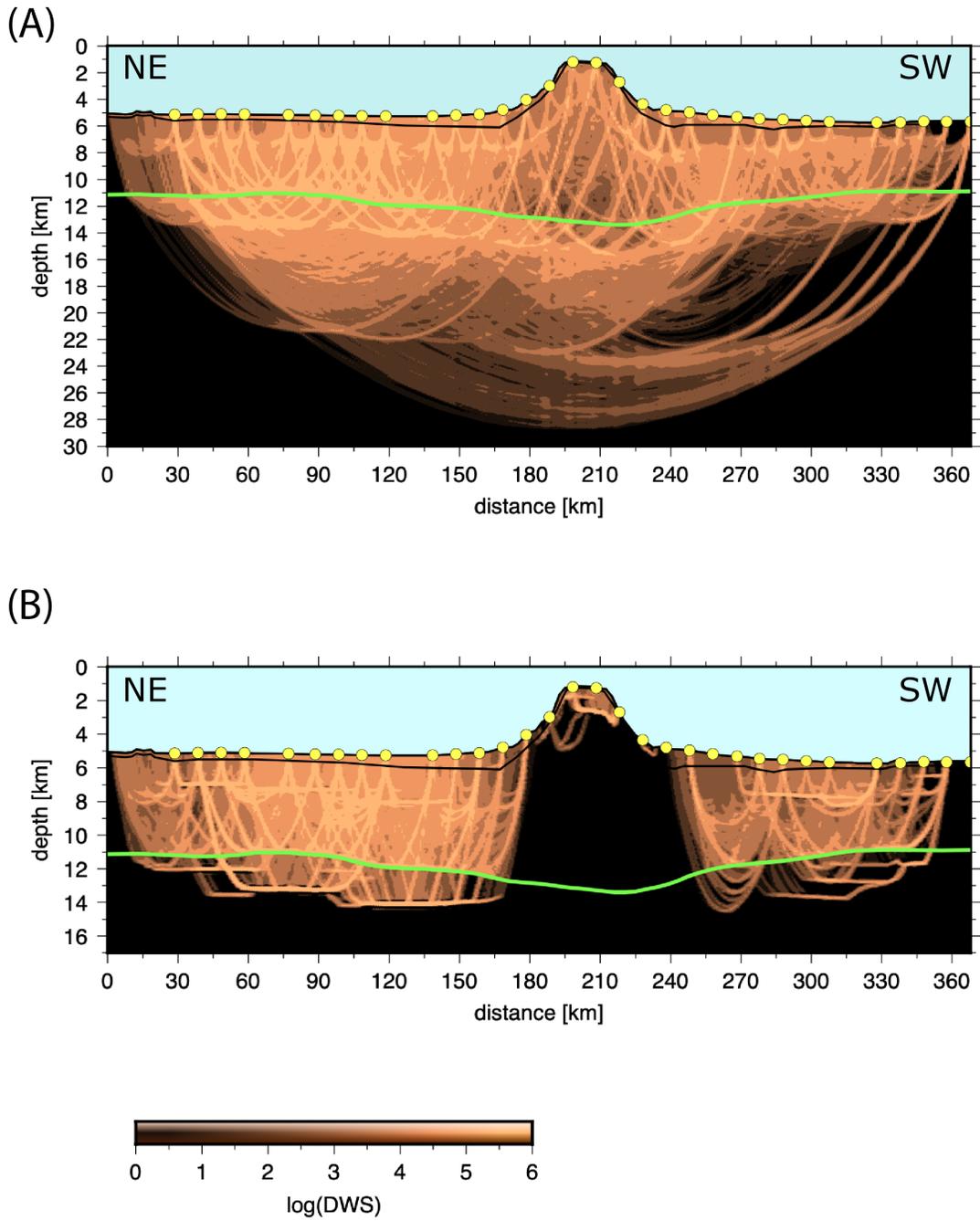
**Fig. S12.** (A) Synthetic  $V_p$  reference velocity-depth model consisting of  $\pm 5\%$  Gaussian anomalies ( $20 \text{ km} \times 2 \text{ km}$  and  $40 \text{ km} \times 14 \text{ km}$  in the crust and mantle, respectively) superimposed onto the final P-wave velocity model shown in Fig. 7A. (B) Recovery for the  $V_p$  model using the same instrument geometry and ray density (see Section 3.3). Yellow dots mark OBH/S locations.



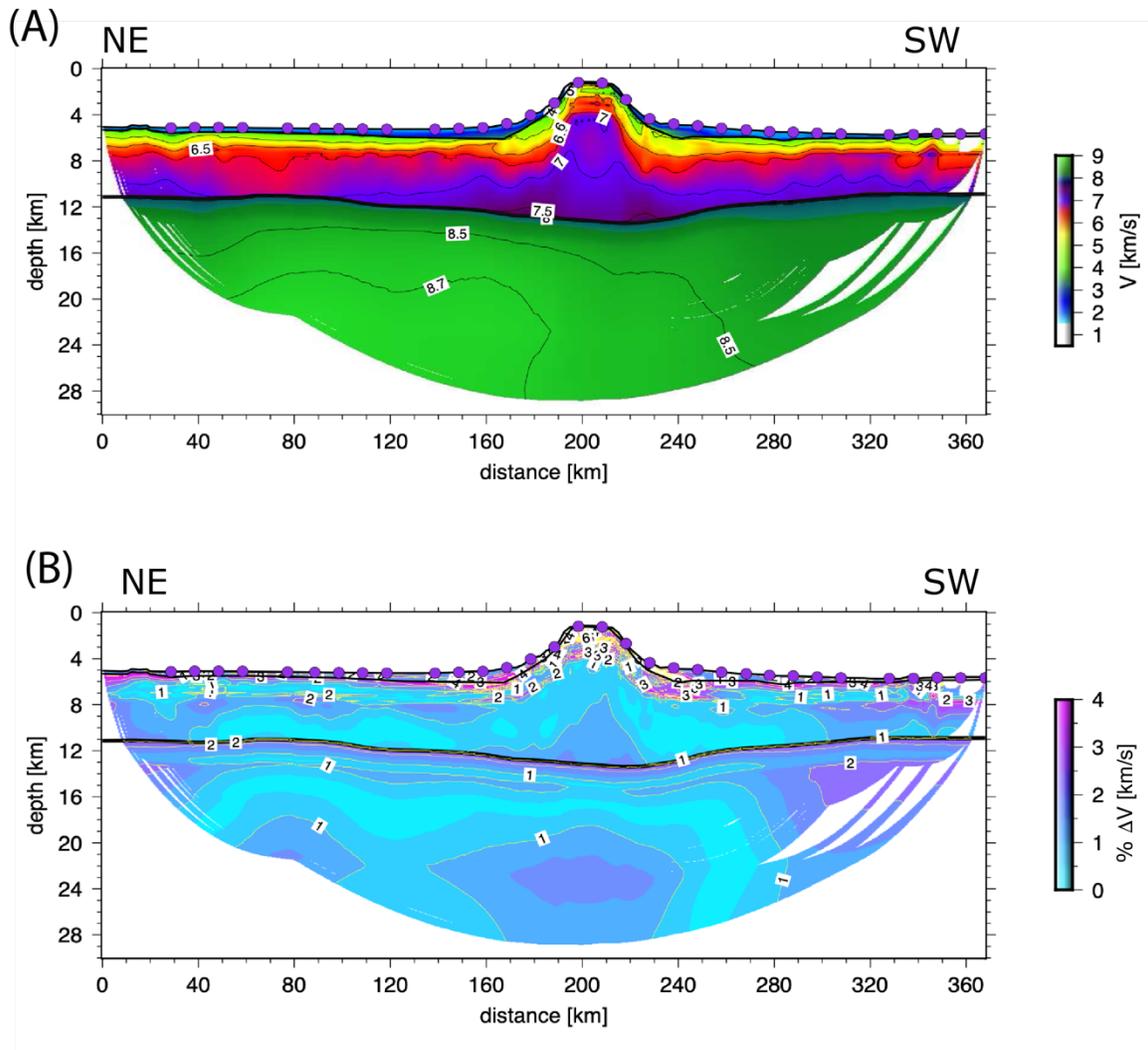
**Fig. S13.** (A) Synthetic  $V_s$  reference velocity-depth model consisting of  $\pm 5\%$  Gaussian anomalies ( $40 \text{ km} \times 8 \text{ km}$  and  $40 \text{ km} \times 4 \text{ km}$  in the crust and mantle, respectively) superimposed onto the final S-wave velocity model shown in Fig. 7B. (B) Recovery for the  $V_s$  model using the same instrument geometry and ray density (see Section 3.3). Green dots mark OBH/S locations.



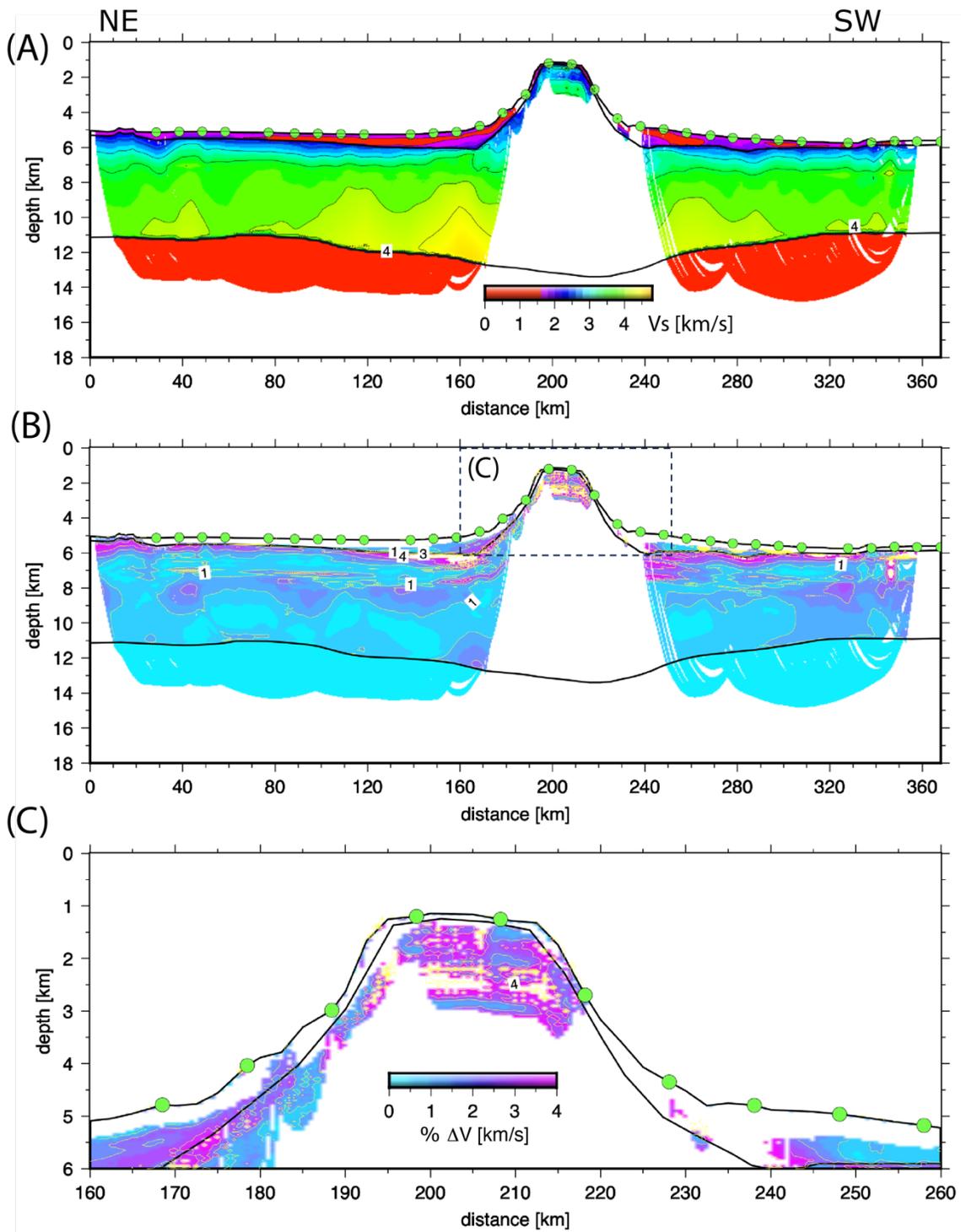
**Fig. S14.** (A) Synthetic  $V_s$  reference velocity-depth model consisting of  $\pm 5\%$  Gaussian anomalies ( $20 \text{ km} \times 2 \text{ km}$  and  $40 \text{ km} \times 4 \text{ km}$  in the crust and mantle, respectively) superimposed onto the final S-wave velocity model shown in Fig. 7B. (B) Recovery for the  $V_s$  model using the same instrument geometry and ray density (see Section 3.3). Green dots mark OBH/S locations.



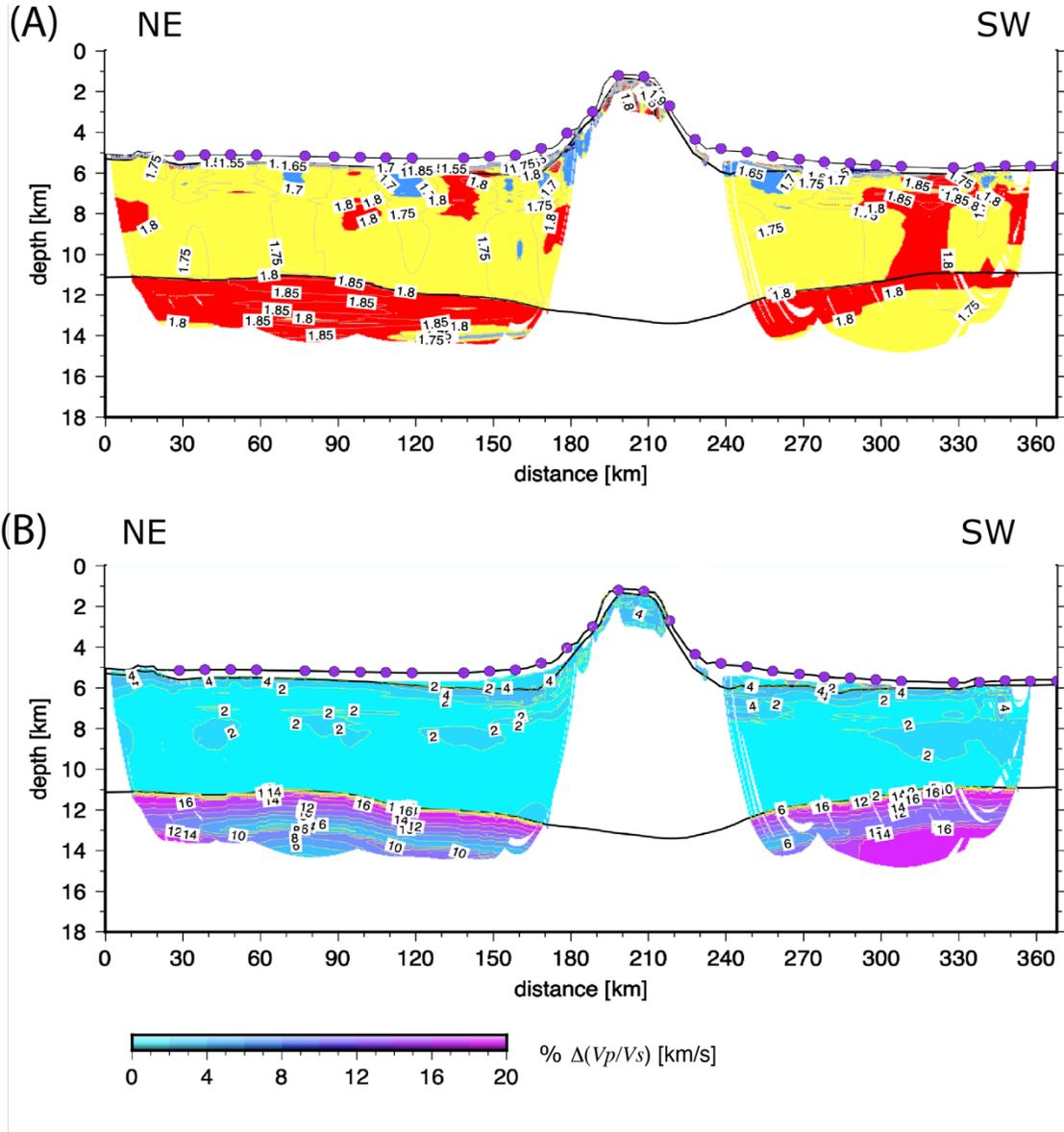
**Fig. S15.** DWS for rays traveling through the (A)  $V_p$  and (B)  $V_s$  models plotted using a logarithmic scale. Yellow dots mark OBH/S locations.



**Fig. S16.** (A) Average P-wave velocity model. (B) Percentage standard deviation model ( $\Delta V$ ) of the  $V_p$  model (see Section 3.4). Green dots mark OBH/S locations.



**Fig. S17.** (A) Average S-wave velocity model. (B) Percentage standard deviation model ( $\Delta V$ ) of the  $V_s$  model (see Section 3.4). (C) Zoom-in of the summit region shown in (B).



**Fig. S18.** (A) Derived 2-D  $V_p/V_s$  model based on the  $V_p$  and  $V_s$  model shown in Figs. S16A and S17A. (B) Percentage standard deviation model  $100 \times \frac{\Delta(V_p/V_s)}{V_p/V_s}$  where  $\Delta(V_p/V_s)$  is calculated using Eq. 1 (see Section 3.4). Purple dots mark OBH/S locations.

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