Earth processes cause Zr-Hf and Nb-Ta fractionations, but why and how?

Yaoling Niu,*^{*a,b*}

5 Received (in XXX, XXX) Xth XXXXXXXX 20XX, Accepted Xth XXXXXXXX 20XX DOI: 10.1039/b000000x

Zr-Hf and Nb-Ta are two elemental twins, each of which has the ¹⁰ same valence (i.e. 5⁺ for Nb and Ta, and 4⁺ for Zr and Hf) and same ionic size at a given coordination number (e.g. $R_{Nb}/R_{Ta} =$ 1.000 and $R_{Zr}/R_{Hf} =$ 1.006 to ~ 1.026 for coordination numbers of 6, 7, 8 and 12). As a result, it has been the view that Zr does not fractionate from Hf and Nb does not fractionate from Ta in the

- ¹⁵ formation of minerals and in all Earth processes. Recent studies, however, have shown that this traditional view is in error. Up to 2 orders of magnitude fractionations have been observed in rocks from the world ocean floor and also in the seawater on various scales. I discuss some possible processes that may help explain
- 20 such large fractionations, but further work is needed to test the validity of these interpretations. Perspectives from the chemistry community are in compelling need to help address this basic problem of scientific significance in understanding the chemical differentiation of our planet.

25

The theory

Victor Moritz Goldschmidt, Father of Modern Geochemistry (Mason, 1992), recognized that behaviours of chemical elements in the formation of minerals and rocks are largely controlled by

- ³⁰ their ionic size and valence (Goldschmidt, 1937). The type examples are Zr-Hf and Nb-Ta elemental twins. Each of the two twins has the same valence (e.g., Zr^{4+} and Hf^{4+} ; Nb⁵⁺ and Ta⁵⁺) and essentially the same ionic radii for a given coordination number (e.g., with CN = 6, $R_{Zr} = 0.72$, $R_{Hf} = 0.71$; $R_{Nb} =$
- ³⁵ 0.64, $R_{Ta} = 0.64$) in (all or almost all) Earth materials and processes (Shannon, 1976). As a result, there has been wide acceptance that Zr/Hf and Nb/Ta ratios should be the same as, or identical to, those of chondrite meteorites (assumed to represent the composition of the bulk Earth and the solar system) in all
- ⁴⁰ mantle and mantle derived rocks: Zr/Hf \approx 36.30 and Nb/Ta \approx 17.6 (e.g. Jochum *et al.*, 1986; Sun and McDonough, 1989; or Nb/Ta \approx 19 as per Munker *et al.*, 2003) although it has been noted that the Zr/Hf ratio varies in some carbonatites (Andrade *et al.*, 2002) and is superchandric (variably higher than chondritic
- ⁴⁵ ratio) in some alkali basalts (Dupuy *et al.*, 1992) and that the bulk continental crust has a subchondritic (lower than chondritic ratio) Nb/Ta ratio of 11-12 (Rudnick, 1995).

Observations

Modern analytical technologies such as inductively coupled ⁵⁰ plasma mass spectrometry (ICP-MS) has enabled analysis of most elements across the periodic table precisely and accurately in rocks and minerals at all concentration levels, and as low as parts per trillion (10⁻¹²). With this method, Niu and co-authors studied mid-ocean ridge basalts (MORB; Niu and Batiza, 1997; ⁵⁵ Niu and Hekinian, 1997) - rocks that form the world ocean crust and cover ~ 70% of the Earth's surface - solidified from magmas produced by partial melting of mantle rocks along the > 55000 km long ocean ridge system that encircles the globe. They discovered that these rocks show a factor of 2 correlated variation ⁶⁰ between Zr/Hf (~ 25 to 50) and Nb/Ta (~ 9 to 18) (See Fig. 1a).



Fig. 1 Correlated variations of Zr/Hf and Nb/Ta in seafloor basalts (MORB) (*a*) and seafloor mantle peridotites (MORP) (*b*). The grey bands indicate theoretical values of chondrite meteorites that are assumed to represent the values of the solar system and the bulk Earth. These correlations are statistically significant at > 99.9% confidence levels. The blue cluster in (*b*) is the MORB data in (*a*).

Furthermore, Niu (2004) showed that these two ratios in fact show over 2-orders of magnitude correlated variations in midocean ridge peridotites (MORP) (Fig. 1b) that are melting 70 residues complementary to MORB during partial melting of the mantle that produces the ocean crust at ocean ridges. Very recently, Firdaus *et al.* (2011) show very large Zr/Hf and Nb/Ta variations in seawater collected from surface down to varying depths from a number of stations along two longitudinal sections 75 in the Pacific although there is no correlation between the two ratios (Fig. 2). More recently, varying Nb/Ta and Zr/Hf ratios have been observed in basalts erupted on oceanic islands (e.g., Pfander *et al.*, 2007), along island arcs (e.g., Konig and Schuth, 2011), in some mantle xenoliths (e.g., Kalfoun *et al.*, 2002;

Aulbach *et al.*, 2011), granitic rocks (e.g., Dostal and Chatterjee, 2000), and metamorphic rocks including the Archean TTG suites (e.g., Schmidt *et al.*, 2009; Liang *et al.*, 2009; Hoffmann *et al.*, 2011; John *et al.*, 2011). These are important new observations, ⁵ but they have rather limited variation ranges without any Nb/Ta-

Zr/Hf correlation.

I thus focus here on MORB, MORP and seawater because of their huge fractionations of these two elemental twins and because they represent volumetrically significant Earth materials 10 – mantle rocks (MORP) and mantle-derived rocks (MORB) that cover ~ 2/3 of the surface of the solid Earth and the oceans that dominate the Earth's hydrosphere. I hope to inspire testable hypotheses towards revealing the governing principles through a

collaborative effort between the chemical and geological 15 community. All the above observations manifest that the prevailing theory in the geochemistry is invalid, at least in describing processes that

have led to the formation of many rocks, especially MORB and MORP as well as seawater hydro-chemical or hydrodynamic 20 processes. At present, the solid-Earth geochemistry community

- assumes that the bulk Earth must have a chondritic Nb/Ta = 17.6 (or Nb/Ta = \sim 19, see above), and, on the basis of this assumption, interpret the subchondritic Nb/Ta ratio (i.e., < 17.6) of bulk continental crust (Rudnick, 1995) and oceanic crust (Niu
- ²⁵ & Batiza, 1997; Rudnick *et al.*, 2000) as indicating a super chondritic Nb/Ta reservoir present deep in the Earth's mantle that has never been sampled by known magmatism (McDonough, 1991; Niu & O'Hara, 2003). Wade and Wood (2001) proposed that Nb is more siderophile (than Ta) and that the missing Nb in ³⁰ crustal rocks may be in the Earth's core (Kamber *et al.*, 2003).
- All this points to the critical importance in understanding mechanisms that can cause Nb-Ta (and Zr-Hf) fractionation towards an improved understanding on how the Earth works in terms of both chemical and physical processes the planet Earth
- ³⁵ (maybe the entire solar system) may have experienced over its history. The objective of this short article is to draw attention of chemists who may offer new perspectives on what may actually cause Zr-Hf and Nb-Hf fractionations, why and how. In the following, I summarize the current hypotheses on this problem.



Fig. 2 Unexpectedly large Zr-Hf and Nb-Ta fractionations in seawater (data are weight ratios from Firdaus *et al.*, 2011).

Discussion

Zr-Hf and Nb-Ta fractionations in MORB and MORP

⁴⁵ Elements such as Nb, Ta, Zr and Hf are termed "incompatible" elements in geology because they prefer to enter the melt over solid minerals during mantle melting or melt crystallization. This is defined by a parameter called bulk distribution coefficient $D_M^{Solid/Melt} = C_M^{Solid}/C_M^{Melt}$, which is the concentration ratio of ⁵⁰ element *M* in the solid minerals over the melt at equilibrium conditions. If $D_M^{Solid/Melt} < 1$, the element *M* is called incompatible. Obviously, if $D_A^{Solid/Melt} < D_B^{Solid/Melt} < 1$, then elements *A* is more incompatible than element *B*.



55 Fig. 3 MORB data (as in Fig. 1a) show Zr/Hf vs. Zr (top) and Nb/Ta vs. Nb (bottom). As illustrated in the inset in the top panel. The positive trends on these two plots mean that with increasing X, both X and Y increase, but X increases faster, meaning X is more incompatible than Y (Niu & Batiza, 1997).

⁶⁰ Fig. 3 plots MORB data to manifest that Zr is more incompatible than Hf (i.e., $D_{Zr} < D_{Hf}$), and Nb is more incompatible than Ta (i.e., $D_{Nb} < D_{Ta}$) (Niu and Hekinian, 1997; Niu and Batiza, 1997). The same is also true in MORP data (Niu, 2004). This is unexpected because of the same valence and ⁶⁵ similar/identical ionic size. Green *et al.* (2000) showed experimentally that in the basaltic system the distribution between clinopyroxene (an effective mineral) and the melt gives $D_{Zr}/D_{Hf} \approx 0.5$ and $D_{Nb}/D_{Ta} \approx 0.5$, which is qualitatively consistent with the observation and could explain the a factor of 2 Zr/Hf and Nb/Ta variation in the MORB data (Figs. 1a,3), but cannot explain the over 2 orders of magnitude variation in the ⁵ MORP data (Fig. 1b). There are more recent experimental data on

the relative incompatibility of these elements (Schmidt *et al.*, 2004; Fulmer *et al.*, 2010; Xiong *et al.*, 2011), but all these works offer no effective mechanism why Zr-Hf and Nb-Ta elemental twins should fractionate from each other (Foley *et al.*, 2002; ¹⁰ Blundy and Wood, 2003).

Given the identical valences (5⁺ for Nb and Ta, and 4⁺ for Zr and Hf) and essentially the same ionic radii ($R_{\rm Nb}/R_{\rm Ta} = 1.000$ and $R_{\rm Zr}/R_{\rm Hf} = 1.006$ to ~ 1.026 for coordination numbers of 6, 7, 8 and 12) of the two elemental twins, yet a factor of two differences in

- ¹⁵ their respective atomic masses ($M_{Zr}/M_{Hf} = 0.511$; $M_{Nb}/M_{Ta} = 0.513$), it is logical to reason that mass differences may have effects on the observed fractionations (or the apparent relative incompatibility). For example, for two elements of similar or identical chemical properties the lighter element (e.g. $^{90-96}Zr$, $^{20} \, ^{93}Nb$) is more incompatible than the heavier element (e.g. $^{174-}$
- ¹⁸⁰Hf, ¹⁸¹Ta) (Niu & Batiza, 1997). This allowed Niu & Hekinian (1997) to suggest mass-dependent fractionation, either differential diffusion rates or differential mass transfer rates. Such mass-dependent fractionation may also explain why $D_{Rb} < D_{Cs}$
- ²⁵ (mass ratio $M_{\rm Rb}/M_{\rm Cs} = 0.643$) even though Cs (1⁺) is about 10% larger in ionic radius than Rb (1⁺). The 50% mass difference for Zr–Hf and Nb–Ta is significantly greater than commonly considered isotopic fractionations of light stable elements at relatively low temperatures. Indeed, diffusion or mass transfer ³⁰ coefficients are known to be mass-dependent. For example,
- $K_A/K_B = (M_B/M_A)^{1/2}$, where K_A and K_B represent diffusion coefficients of species with mass M_A and M_B , respectively (Lasaga, 1998). For a mass ratio of ~ 2 (e.g., M_{Ta}/M_{Nb} and M_{Hf}/M_{Zr}), the diffusion coefficient ratio would be K_{Nb}/K_{Ta} (or
- ³⁵ K_{Zr}/K_{Hf}) = 1.414, i.e., the lighter element would diffuse, under ideal situations, 41% more efficiently than the heavy element (given all other variables the same: same valence, same ionic radius, same coordination number, etc.). That is, there would be about 41.4% fractionation just from the mass-dependent diffusion 40 coefficients alone.

Such ~ 41% (or 410‰) fractionation contrasts with familiar per-mil level light isotope fractionations. For example, for ¹⁶O and ¹⁸O fractionation, $K^{180}/K^{160} = (16/18)^{1/2} = 0.943$, there would only be a 57‰ (vs. 410‰) fractionation. Is it possible that the

⁴⁵ apparent $D_{Zr}/D_{Hf} < 1$ and $D_{Nb}/D_{Ta} < 1$ may be due to massdependent fractionation under high temperature magmatic (mantle?) conditions? The ultimate test for the hypothesis of mass-dependent fractionation requires well-designed experimental studies and isotopic analysis of MORP samples ⁵⁰ (Fig. 1b) at different mass levels (e.g., ^{46,47,48,49,50}Ti, ^{90,91,92,94,96}Zr,

^{174,176,177,178,179,180}Hf) using multiple collector ICP-MS (Niu, 2004). Alternative possibilities have been reviewed by Huang *et al.* (2011).

While the mass-dependent fractionation hypothesis is stattractive and is worth testing, it is also noted that the two elements, $Y(3^+)$ and Ho (3⁺), which are sufficiently similar in chemical properties ($R_Y/R_{Ho} = 0.9989$ to ~ 1.0039 for coordination numbers of 6, 8 and 9; Shannon, 1976), yet their mass difference of $M_{\rm Y}/M_{\rm Ho} \approx 0.539$ does not seem to cause significant fractionation between these two elements because Y/Ho = 27.33±0.39 for MORB samples and 27.16±4.67 for MORP samples. A recent study on an Indian Ocean basalt suite shows an obvious Y-Ho fractionation (Frey *et al.*, 2011), which may be caused by mass-dependent fractionation.



Fig. 4 Zr/Hf (weight ratio) vs. Zr (ppm) (a) and Nb/Ta (weight ratio) vs. Nb (ppm) (b) of seawater data from Firdaus *et al.* (2011). Solid circles in red are ocean depth interval averages with $\pm 1\sigma$ from the means explained in Fig. 5. The grey bars are chondrite meteorite ratios plotted for 70 reference.

Zr-Hf and Nb-Ta fractionations in seawater

The correlations between total silicate (expressed as Si[OH]₄) and Nb, Ta, Zr and Hf abundances in sweater allow Firdaus *et al.* (2011) to interpret that these elements in seawater must have ⁷⁵ derived from terrigenous sources (i.e., sediments derived from the upper continental crust transported through rivers and wind dust). This interpretation seems reasonable, but cannot explain the unexpectedly large Zr/Hf and Nb/Ta variations (Fig. 2) without additional processes because Zr/Hf \approx 36.4 and Nb/Ta \approx 13.3 in ⁸⁰ upper continental crustal rocks (Rudnick and Gao, 2003).

The positive correlation of Zr/Hf with increasing Zr in Fig. 4a resembles that seen in MORB (Fig. 3a), but the negative correlation of Nb/Ta with Nb in Fig. 4b does not (see Fig. 3b). This observation indicates that there must be different ⁸⁵ mechanisms at work in causing Zr-Hf and Nb-Ta fractionation. Because the authors recognize that the abundances of these elements in seawater vary as a function of both water depth and

geographic locations (latitude and longitude in the Pacific), and because location-dependent variation is more fortuitous, we can examine how these elements and their ratios may vary as a function of water depth so as to gain insights into the mechanism

- ⁵ of Zr-Hf and Nb-Ta fractionation. For convenience and to average out effects including analytical uncertainties and location-dependent variations, we can choose to average the data with respect to water depth intervals as done in Fig. 5 with the averages also plotted in Fig. 4. It is not shown, but significant
- ¹⁰ curvilinear correlations exist between water depth and abundances of Zr, Hf, Nd and Ta, which are best described in terms of a power-law relationship as given in the caption to Fig. 5. However, the silicate nutrient (Si[OH]₄) does not correlate well with the abundances of these elements in having no clear increase
- ¹⁵ with increasing water depth at depths greater than ~ 1200 m below the sea surface. On the other hand, Zr/Hf and Nb/Ta ratios show only change with depth shallower than ~ 1000m, and maintain constant at greater depths. (Fig. 5).

It is obvious in Fig. 5 that the large Zr-Hf and Nb-Ta variation ²⁰ and fractionation in seawater takes place essentially within the upper ~ 1000 m depth interval. If this is indeed the case, then the constant Nb/Ta (18.81±1.01) and Zr/Hf (241±6) ratios at greater

depths may be the net result of shallow level fractionation. It is

possible that the large variation range of Zr/Hf and Nb/Ta (large $^{25}\pm 1\sigma$ values in Fig. 5) at shallow water depth may be associated with varying geographic locations. For example, those near the coastlines such as the Aleutian volcanic arc system to the north and the Tonga-New Zealand system to the south may receive land-derived sediments relative to locations far from the 30 coastlines. There is likely also wind dust (ultimately also from land surface) in shallow water.

However, upper continental crustal rocks have average Nb/Ta \approx 13.3 and Zr/Hf \approx 36.4 (Rudnick & Gao, 2003), which are significantly lower than Nb/Ta = 13 - 85 (with a mean of 32.9)

- as and Zr/Hf = 46.5 312 (with a mean of 137), respectively, in the near surface seawater (Fig. 5). That is, both the extent of Nb/Ta and Zr/Hf fractionation and the large variation range of these two ratios in the near-surface seawater cannot be explained by landderived materials. Shallow water biological (microorganisms) or
- ⁴⁰ biochemical activities are likely the primary cause for such large Nb-Ta and Zr/Hf fractionation, but how they work is unknown and requires chemical and biochemical investigations.

The biochemical effect may be most intense at the surface and declines with depth to ~ 1000m, which would explain the ⁴⁵ systematic Nb/Ta and Zr/Hf variations with depth to that deep level (Fig. 5). If the microorganisms prefer to intake (in non-chondritic ratios) Ta (vs. Nb) and Zr (vs. Hf), this tendency can explain the decreasing Nb/Ta and increasing Zr/Hf in the

- dissolved seawater with depth shallower than ~ 1000 m (Fig. 5). ⁵⁰ If the microorganisms demise at depth > 1000m, the biochemical fractionation would cease. This possibility explains the constant Nb/Ta (18.1 \pm 1.01) and Zr/Hf (241 \pm 6) in seawater at all depths
- deeper than ~ 1000m (Fig. 5). Much of these elements may be absorbed on particles, whether biological remains or surface wind ⁵⁵ dust, and their solubility increase in seawater with depth (pressure
- effect) explains the their increasing abundances with depth (pressure

caption to Fig. 5).

Conclusions

- Contrary to the traditional theory that elemental twins such as ⁶⁰ Nb-Ta and Zr-Hf should not fractionate from each other in all Earth materials and through all Earth processes, in particular the Earth's mantle and mantle-derived rocks, these element twins do fractionate rather significantly as recorded in magmatic rocks such as MORB and mantle rocks like MORP, and as manifested
- 65 in present-day seawater on a global scale. Mass-dependent fractionation has been proposed to explain the Zr-Hf and Nb-Ta fractionation in Earth processes that form MORB and MORP, but this interpretations requires testing through carefully designed experimental investigations as well as isotopic analysis of multi-
- ⁷⁰ isotope elements of varying mass levels like Ti (masses 46, 47, 48, 49, 50), Zr (90, 91, 92, 94, 96) and Hf (174, 176, 177, 178, 179, 180). A correct understanding on the cause or causes of Nb-Ta and Zr-Hf fractionation in Earth's mantle and mantle-derived rocks is of profound importance in understanding the chemical ⁷⁵ differentiation of the Earth over its history.

More recently, significant Nb-Ta and Zr-Hf fractionations have been observed in rocks from other geological environments (though lesser in extent than in MORP, and without Nb/Ta-Zr/Hf correlation), but the origin of such fractionations remains ⁸⁰ unknown. Chemical or biochemical processes at near-surface seawater are suspected to be the primary cause of the observed significant Nb-Ta and Zr-Hf fractionations. While this interpretation is logical, the exact chemical or biochemical mechanism is unknown and requires insightful investigations. ⁸⁵ Firdaus *et al.* (2011) proposed that the observed spatial Nb/Ta and Zr/Hf variations in seawater can be used as an effective proxy to monitor ocean circulation patterns. Although the latter is of fundamental scientific value, it remains in doubt whether these two ratios are useful in this regard without understanding the ⁹⁰ origin of the observed Nb-Ta and Zr-Hf fractionations.

It is well known that chemically to separate Zr from Hf and Nb from Ta is not trivial. For example, one way (without using organic solvents) to separate Zr from Hf is to form Zr-Hf-bearing fluoric complexes (e.g. K₂[Zr,Hf]F₆), and then use the method of ⁹⁵ fractional crystallization because K₂ZrF₆ is less soluble than K₂HfF₆ in water. Similarly, one way (without using organic solvents) to separate Nb from Ta also requires the formation of different fluoric complexes (e.g., K₂[TaF₇] and K₂[NbOF₅]) and uses their different solubility in water. In both cases, formation of fluoric complexes is a prerequisite, but except for the rather rare pegmatitic environments (associated with very late stage granitic magma evolution), it is unlikely to form such fluoric complexes through Earth processes because of the difficulties in concentrating fluorine is all other Earth environments and ¹⁰⁵ processes.

Is there any other method to separate Zr from Hf and Nb from Ta without forming fluoric complexes? Perspectives from the chemistry community are in compelling need to help address this basic problem of scientific significance in understanding the 110 chemical differentiation of our planet.

Notes and references

- ^a School of Earth Sciences, Lanzhou University, Lanzhou 730000, China
- ^b Department of Earth Sciences, Durham University, Durham DH1 3LE, UK. Fax:441913342311; Tel:441913342301; E-mail: <u>niuyl@lzu.edu.cn;yaoling.niu@durham.ac.uk</u>
- ⁵ The research is supported by Durham University and also National Natural Science Foundation of China (91014003, 41130314). Two journal reviewers are thanked for constructive comments.
- 1 B. Mason, The Geochemical Society Special Publication, 1992, 4, 184 pp.
- 2 V. M. Goldschmidt, Journal of Chemical Society, 1937, 140, 655.
- ¹⁰ 3 R. D. Shannon, *Acta Crystallographica*, 1976, A **32**, 751–767.
 - 4 K. P. Jochum, H. M. Seufert, B. Spettel et al., Geochimica et Cosmochimica Acta, 1986, 50, 1173.
 - 5 S. S. Sun, W. F. McDonough, Geological Society Special Publications, 1989, 42, 313.
 - 6. C. Munker, J.A. Pfander, S. Weyer et al., Science, 2003, 301, 84.
 - 7. F. R. D. Andrade, P. Moller, P. Dulski, Revista Brasileira de Geoscienceas, 2002, 32, 361.
- 15 8. R. J. Rudnick, Nature, 1995, 378, 571.
 - 9 Y. L. Niu, R. Batiza, Earth and Planetary Science Letters, 1997, 148, 471.
 - 10 Y. L. Niu, R. Hekinian, Earth and Planetary Science Letters, 1997, 146, 243.
 - 11 Y. L. Niu, Journal of Petrology, 45, 2423.
 - 12 M. L. Firdaus, T. Minami, K. Norosuye, Y. Sohrin, Nature Geoscience, 2011, 4, 227.
- 20 13 J. A. Pfander, C. Munker, A. Stracke, Earth and Planetary Science Letters, 2007, 254, 185.
- 14 S. Kong, S. Schuth, Earth and Planetary Science Letters, 2011, 301, 265.
- 15 F. Klfoun, D. Iono, C. Merlect, Earth and Planetary Science Letters, 2002, 199, 19.
- 16 S. Aulbach, S.Y. O'Reilly, N.J. Pearson, Contributions to Mneralogy and Petrology, 2011, 162, 1047.
- 17 J. Dostal, A.K. Chatterjee, Chemical Geology, 2000, 163, 207.
- 25 18 A. Schmidt, S. Weyer, T. John et al., Geochimica et Cosmochimica Acta, 2009, 73, 455.
- 19 J. L. Liang, X. Ding, X.M. Sun et al., Chemical Geology, 2099, 268, 27.
- 20 J. E. Hoffmann, C. Munker, T. Naerass et al., Geochimica et Cosmochimica Acta, 2011, 75, 4157.
- 21 T. John, R. Klemd, S. Klemme et al., Contributions to Mneralogy and Petrology, 2011, 161, 35.
- 22 R. L. Rudnick, M. Barth, I. Horn et al., Science, 2000, 287, 278.
- 30 23 W. F. McDonough, Philosophical Transaction of the Royal Socety of London, 1991, Series A 335, 407.
- 24 Y. L. Niu, M. J. O'Hara, Journal of Geophysical Research, 2003, 108, 2209, doi:10.1029/2002JB002048, 19 pp.
- 25 J. Wade, B. J. Wood, Nature, 2001, 409, 75.
- 26 B.S. Kamber, A. Greig, R. Schonberg et al., Precambrian Reserch, 2003, 126, 289.
- 27 T. H. Green, J. D. Blundy, J. Adam et al., Lithos, 2000, 53, 165.
- 35 28 M.W. Schmidt, A. Dardon, G. Chazot, R. Vannucci, Earth and Planetary Science Letters, 2004, 226, 415.
- 29 E.C. Fulmet, O. Nebel, W. van Westtrenen, Geochimica et Cosmochimica Acta, 2010, 74, 2714.
- 30 X.L. Xiong, H. Keppler, A. Audetat et al., Geochimica et Cosmochimica Acta, 2011, 74, 1673.
- 31 S. Foley, M. Tiepolo, R. Vannucci, Nature, 2002, 417, 837.
- 32 J. Blundy, B. Wood, Earth and Planetary Science Letters, 2003, 210, 383.
- 40 33 A. C. Lasaga, Kinetic Theory in the Earth Sciences, 1998, Princeton University Press, 811 pp.
- 34 H. Huang, Y. L. Niu, Z. Zhao, H. Hei, D. Zhu, Journal of Earth Sceinces, 22, 52.
- 35 Frey, F.A., M. Pringle, P. Meleney et al., Earth and Planetary Science Letters, 2011, 303, 215.
- 36 R. L. Rudnick, S. Gao, Treatise on Geochemistry, 2003, 3, 1.



Fig. 5 Average total silicate (expressed as Si[OH]₄ umol/l), Nb/Ta and Zr/Hf (weight ratios) in seawater plotted as a function of sampling depths from the surface down to the ocean bed. The averages were derived by averaging all the observations from within each of the 500 meter depth intervals (e.g., N = 5207 observations in the shallowest interval given on the left panel) regardless of actual locations (latitude and longitude) so as to reveal the first-order variations of these chemical parameters as a function of water depth. Not shown here, but Zr, Hf, Nb and Ta (in ppm) concentrations exhibit a nonlinear monotonic increase with increasing water depth (*D* in meter): $Zr = 3*10^{-7}D^{0.4796}$ ($R^2 = 0.987$), Hf = $4*10^{-9}D^{0.3317}$ ($R^2 = 0.996$), Nb = $7*10^{-8}D^{0.1727}$ ($R^2 = 0.986$) and Ta = $2*10^{-9}D^{0.2776}$ ($R^2 = 0.959$). Data from Firdaus *et al.* (2011).