

1       **Palaeoproterozoic Orogenic Gold Style Mineralization at the Southwestern Archaean**  
2       **Tanzanian Cratonic Margin, Lupa Goldfield, SW Tanzania: Implications from U-Pb**  
3                               **Titanite Geochronology**

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6       C.J.M. Lawley<sup>1,2\*</sup>, D. Selby<sup>1</sup>, D. Condon<sup>3</sup> and J. Imber<sup>1</sup>

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<sup>1</sup>Department of Earth Sciences, Durham University, Science Labs, Durham, DH1 3LE, UK

9       <sup>2</sup>Natural Resources Canada, Geological Survey of Canada, Ottawa, Ontario, K1A 0E8, Canada

10       <sup>3</sup>Natural Environment Research Council Isotope Geosciences Laboratory, British Geological  
11       Survey, Keyworth, Nottingham, NG12 5GG, UK

12  
13

\*corresponding author: e-mail: [clawley@NRCan.gc.ca](mailto:clawley@NRCan.gc.ca); Tel: 1 (613) 996-23593

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**Abstract:** The Lupa Goldfield, situated at the southwestern Tanzanian cratonic margin,  
16 comprises a network of auriferous quartz veins and greenschist facies mylonitic shear zones  
17 cutting a suite of Archaean-Palaeoproterozoic granitic-gabbroic intrusions. The existing  
18 geochronological database points to a protracted, but episodic 1.96–1.88 Ga magmatic history  
19 that is broadly coincident with the 2.1–1.8 Ga Ubendian Orogeny. Molybdenite, pyrite and  
20 chalcopyrite samples from mineralized quartz veins and mylonitic shear zones yield Re-Os model  
21 ages that range from 1.95–1.88 Ga, whereas ca. 1.88 Ga pyrite with gold bearing inclusions and  
22 sampled from the host mylonitic shear zone suggest that gold occurred relatively late in this  
23 hydrothermal history. The ca. 1.88 Ga gold event is recorded at all five of the studied prospects,  
24 whereas the relationship between gold and the disparately older 1.95 and 1.94 Ga Re-Os  
25 molybdenite ages is unclear. New U-Pb metamorphic titanite dating of a foliated Archaean  
26 granite sample (ca. 2.76 Ga) suggests that the onset of ductile deformation within the Lupa  
27 Goldfield occurred at ca. 1.92 Ga, and some ca. 40 Myr prior to auriferous and brittle-ductile  
28 mylonitic shear zones at ca. 1.88 Ga. Early ductile deformation is not associated with gold  
29 mineralization, but the ductile deformation fabrics and, in particular the development of  
30 rheologically weak chloritic folia, may have acted as zones of pre-existing weakness that  
31 localized strain and influenced the geometry of later auriferous mylonitic shear zones. The large  
32 age difference between U-Pb zircon and titanite ages for the Archaean granite sample is in  
33 contrast to new U-Pb titanite ages for the Saza Granodiorite ( $1930 \pm 3$  Ma), which are only  
34 slightly outside of analytical uncertainty at the  $2\sigma$  level with a previously reported U-Pb zircon

35 age for the same sample ( $1935 \pm 1$  Ma). These new age results, together with previously reported  
36 U-Pb and Re-Os ages, highlight the protracted magmatic, hydrothermal and structural evolution  
37 of the Lupa Goldfield (1.96–1.88 Ga). They are also consistent with other palaeo-convergent  
38 margins where orogenic gold systems genesis occurs relatively late in the orogen's tectono-  
39 thermal history.

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41 Keywords: Lupa Goldfield, Paleoproterozoic, Orogenic Gold, Tanzania, Ubendian Belt

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## 43 **1.0 Introduction**

### 44 *1.1 Tanzanian gold deposits*

45 Tanzania represents the fourth largest gold producer in Africa and is the subject of  
46 renewed mineral exploration interest since the introduction of reformed mining legislation in the  
47 late 1990s (Roe and Essex, 2009; Yager, 2010; Brown et al., 2013). The vast majority of gold  
48 production and mineral exploration are concentrated in northern Tanzania and the Lake Victoria  
49 Goldfield where world-class deposits (i.e., deposits with  $\geq 3.0$  Moz of contained gold in proven  
50 and probable reserves) such as Geita (12.3 Moz; AngloGold Ashanti Annual Report, 2012) and  
51 Bulyanhulu (8.0 Moz; Barrick Annual report, 2012) are located (Fig. 1). Two other operating  
52 mines are also located in the Lake Victoria Goldfields [North Mara, 2.2 Moz (Barrick Annual  
53 report, 2012); Buzwagi, 2.0 Moz (Barrick Annual Report, 2012)] in addition to past producing  
54 mines (e.g., Golden Pride, Tulawaka) and numerous smaller-scale artisanal mining operations  
55 (Fig. 1). These deposits are hosted by Neoproterozoic greenschist-amphibolite facies granite-  
56 greenstone belts and lesser sedimentary successions comprising the Tanzanian Craton and are  
57 typical of similarly-aged orogenic gold deposits worldwide (e.g., Goldfarb et al., 2001).

58 The Lake Victoria Goldfield represents merely one example of Tanzania's gold endowed  
59 regions (Fig. 1). Metamorphic belts contiguous with the Tanzanian Craton host a variety of lesser  
60 known goldfields (e.g., Lupa, Mpanda, Mbinga and Niassa) that have attracted comparatively less  
61 modern gold exploration. The geologic settings of these goldfields also remain poorly understood  
62 despite their historic importance during Tanzania's colonial period and on-going artisanal mining  
63 activity (van Straaten, 1984; Kuehn et al., 1990). Recent geochronology studies suggest that  
64 several of the metamorphic belts hosting these goldfields have a complex tectono-thermal history  
65 spanning multiple orogenic cycles (e.g., Boniface et al., 2012, Boniface and Schenk, 2012). As a  
66 result, linking deformation, magmatism, and mineralization to a particular orogenic cycle is  
67 equivocal in the absence of precise geochronologic constraints. In this contribution, we report  
68 new U-Pb titanite age constraints on the earliest deformation fabrics ("D1" – see below)

69 recognized within the western Lupa Goldfield, SW Tanzania. Together the available data suggest  
70 that Proterozoic-Palaeozoic metamorphic belts surrounding the Archaean Tanzanian Craton  
71 represent highly prospective regions and possess hitherto unrecognized gold potential.

72

### 73 *1.2 Chronology of gold deposits hosted by metamorphic belts*

74 Determining the absolute ages of magmatism, metamorphism and mineralization at gold  
75 deposits hosted by metamorphic belts is the subject of concentrated study and continuing  
76 controversy (e.g., Groves et al., 2003). Establishing the timing of mineralization is particularly  
77 challenging due to the dearth of suitable minerals for traditional geochronologic methods.  
78 Nevertheless, recent advances in unconventional sulphide (Stein et al., 2000; Morelli et al., 2007),  
79 hydrothermal phosphate (Şener et al., 2005; Vielreicher et al., 2010) and titanite (Lin and Corfu,  
80 2002) geochronometers has dramatically improved our understanding of metallogenic time scales  
81 and processes at paleo-convergent margins. Re-Os geochronology studies (i.e., Re-Os dating of  
82 sulphide minerals presumed to be co-genetic with gold) have proven to be a particularly effective  
83 approach and have provided unequivocal examples of gold mineralization concomitant with  
84 magmatic intrusions at the million-year time scale (Morelli et al., 2007; Ootes et al., 2007; 2011;  
85 Lawley et al., in press-a). These studies add to a rapidly expanding global database of  
86 geochronologic ages that demonstrate broad contemporaneity between the absolute timing of gold  
87 and magmatism at most world-class goldfields (e.g., Witt and Vanderhor, 1998; Kerrich and  
88 Wyman, 1990; Kerrich and Cassidy, 1994; Kerrich and Kyser, 1994; Oberthur et al., 1998; Arne  
89 et al., 2001; Davis and Lin, 2003; Bucci et al., 2004; Bierlein et al., 2009; Dziggel et al., 2010;  
90 McFarlane et al., 2011).

91 The temporal relationship between magmatism and gold mineralization has led to the  
92 intrusion-related deposit model and is cited as evidence for the importance of locally derived  
93 magmatic hydrothermal fluids in the development of epigenetic gold deposits hosted by  
94 metamorphic belts at convergent margins (Sillitoe and Thompson, 1998; Lang and Baker, 2001).  
95 The inference of locally derived magmatic hydrothermal fluids is at odds with orogenic gold  
96 deposits, which are associated with a similar geodynamic setting (i.e., convergent margins), but  
97 are related to distal hydrothermal fluids of probable metamorphic origin (e.g., Groves et al., 1998;  
98 Groves et al., 2003; Phillips and Powell, 2010). Previous attempts at addressing this controversy  
99 have employed a variety of isotopic tracers, but have provided contrasting results due to, in part,  
100 the difficulty in discriminating isotopically distinct fluid sources after the fluid-rock interaction  
101 that characterizes most hydrothermal systems (e.g., Ridley and Diamond, 2000; Salier et al.,  
102 2005; Fu et al., 2012).

103           The recent emphasis placed on the role of metal-rich magmatic fluids emanating from  
104 mid- to lower-crustal magmatic systems also raises the possibility of distally (e.g., lower crustal)  
105 derived magmatic fluids contributing to hydrothermal systems operating at higher crustal levels  
106 (Botcharnikov et al., 2011; Hronsky et al., 2012). Whilst the genetic relationship between  
107 magmatic hydrothermal fluids and gold is satisfactorily demonstrated at several deposits (e.g.,  
108 Hemlo; Davis and Lin, 2003), the exact role of magmatism and magmatically derived  
109 hydrothermal fluids in the development of epigenetic gold mineralization remains uncertain at  
110 most metamorphic belts (e.g., Charter Towers; Kreuzer, 2005). However, precise geochronologic  
111 constraints on magmatism and mineralization in conjunction with isotopic and fluid inclusions  
112 studies remain important tools in further constraining ore deposit models (Rasmussen et al.,  
113 2006).

114           Gold deposits and prospects in the Lupa Goldfield occur within a Palaeoproterozoic  
115 magmatic arc at the Tanzanian cratonic margin and thus share a close spatial and temporal  
116 relationship with granitic-gabbroic intrusions (Manya, 2011; Lawley et al., 2013). This spatial  
117 and temporal association between magmatism and mineralization has led to the suggestion that  
118 the gold prospects of the Lupa Goldfield are typical of the intrusion-related deposit type (Manya,  
119 2012). In contrast, the protracted hydrothermal history (Lawley et al., in press-a) coupled with the  
120 structural setting of the deposits (Lawley et al., in press-b) suggests that these deposits share more  
121 similarities with orogenic style gold mineralization. In this contribution we report new U-Pb  
122 titanite ages, which provide new constraints on the onset of deformation at the Lupa Goldfield.  
123 This, coupled with the available Re-Os and U-Pb ages allows a critical evaluation of ore deposit  
124 models for gold prospects within the Lupa Goldfield.

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## 126 **2.0 Regional Geology**

### 127 *2.1 Tanzanian Craton and contiguous metamorphic belts*

128           The Tanzanian Craton extends from central to northern Tanzania and into southwestern  
129 Kenya and southeastern Uganda (Fig. 1). Previous workers divided the Archaean Tanzanian  
130 craton into three ‘Systems’ (e.g., Cahen et al., 1984) or ‘Supergroups’ (e.g., Borg and Krogh,  
131 1999): 1) the Dodoman of central Tanzania, which comprises high-grade gneiss, migmatitic  
132 rocks, and granitic-granodioritic intrusions; 2) the Nyanzian of northern Tanzania, which  
133 comprises greenschist-amphibolite facies meta-volcanic rocks, lesser meta-sedimentary  
134 successions and granitic intrusions; and 3) the Kavirondian, which unconformably overlies the  
135 Nyanzian and comprises greenschist-amphibolite facies meta-sedimentary successions (for  
136 additional details regarding the geology, age and nomenclature of the Tanzanian Craton readers

137 are referred to reviews by Clifford, 1970; Bell and Dodson, 1981; Cahen et al., 1984; Borg and  
138 Krogh, 1989; Borg and Shackleton, 1997). However, more recent geologic and geochronological  
139 data sets suggest that the Nyanzian System comprises multiple temporally and geochemically  
140 distinct volcano-stratigraphic sequences, which should not be regarded as a single stratigraphic  
141 package (Manya et al., 2006). Moreover, the Tanzanian Cratons's tectonic framework is likely  
142 more complex than previously suggested and can be further subdivided into thirteen broadly  
143 WNW-ESE trending superterrane [Fig. 1; East Lake Victoria Superterrane (ELVST); Mwanza-  
144 Lake Eyasi Superterrane (MLEST); Lake Nyanza Superterrane (LNST); Moyowosi-Manyoni  
145 Superterrane, (MMST); Dodoma Basement Superterrane (DBST); Dodoma Schist Superterrane  
146 (DSST); Eastern Ubendian-Mtera Superterrane (EUMST); Kalenge-Burigi (KBST; not shown);  
147 Tectonic Front (TF; not shown); Mbulu-Masai Superterrane (MAST); Kilindi-Handeni  
148 Superterrane (KHST); Usagara-Ukaguru Superterrane (UKST); Uluguru-Pare Superterrane  
149 (UPST); Kabete et al., 2012a, b]. Although portions of this proposed tectonic framework require  
150 validation, the proposed model provides a potential explanation for the heterogeneous gold  
151 endowment within the Tanzanian Craton (for more details see Kabete et al., 2012a, b).

152         The newly defined tectonic subdivisions also provide a paradigm with which to explain  
153 the extension of gold-enriched superterrane into the metamorphic belts surrounding the  
154 Tanzanian Craton (Kabete et al., 2012a). To the east, the Tanzanian Craton is mantled by the  
155 Neoproterozoic-Palaeozoic East African Orogen (850–550 Ma, including the 650–620 Ma and ca.  
156 550 Ma Mozambique Belt; e.g., Sommer et al., 2005a; Fritz et al., 2005; Thomas et al., 2013).  
157 The southern and western Tanzanian cratonic margins are bordered by the Palaeoproterozoic  
158 Usagaran and Ubendian Belts, respectively (2.1–1.8 Ga; e.g., Lenoir et al., 1994; Möller et al.,  
159 1998; Reddy et al., 2003; Boniface et al., 2012; Boniface and Schenk 2012). Finally, the  
160 northwestern margin of the Tanzanian Craton is bordered by the Mesoproterozoic Karagwe-  
161 Ankole Belt (ca. 1375 Ma; formally known as the Kibaran Belt; see De Waele et al., 2003, 2006,  
162 2009; Tack et al., 2010 and Fernandez-Alonso et al., 2012 for recent discussions). In reality, each  
163 metamorphic belt is likely poly-orogenic and comprises, in part, reworked Archaean-Proterozoic  
164 crust. The details of this complex tectonic evolution is only beginning to emerge and only the  
165 Ubendian Belt, which hosts the Lupa Goldfield, is discussed in detail below.

166

## 167 *2.2 Ubendian Belt*

168         The Palaeoproterozoic Ubendian Belt is over 600 km long and ca. 150 km wide and  
169 comprises granulite-greenschist facies igneous and sedimentary rocks enveloping the western  
170 margin of the Tanzanian Craton. The NW-SE trending Ubendian Belt extends into northern

171 Malawi and Zambia and separates the Tanzanian and Congo Cratons. The belt is divided into  
172 eight litho-tectonic terranes (Katuma, Ikulu, Ubende, Wakole, Ufipa, Nyika, Upangwa, and Lupa;  
173 Daly, 1988) that are each separated by prominent NW-SE trending and steeply dipping shear  
174 zones (Fig. 2). The Palaeoproterozoic tectonic evolution of the belt is divided into two temporally  
175 distinct tectonic phases (Lenoir et al., 1994). The earliest of these tectonic phases, the 2.1–2.0 Ga  
176 Palaeoproterozoic Ubendian tectonic phase, is characterized by rare granulite facies tectonites and  
177 is considered to record the initial collision between the Tanzanian and Congo Cratons (Lenoir et  
178 al., 1994). This early Ubendian tectonic phase is overprinted by the second and younger 1.9–1.8  
179 Ga Ubendian tectonic phase that resulted in the characteristic amphibolite-greenschist facies  
180 metamorphism and the NW-SE trending terrane-bounding shear zones (Lenoir et al., 1994).  
181 However, more recent work has identified disparately younger Mesoproterozoic (ca. 1.09 Ga;  
182 Boniface et al., 2012) and Neoproterozoic-Palaeozoic (593–524 Ma; Boniface and Schenk, 2012)  
183 metamorphic events potentially related to the Irumide (1.05–1.00 Ga; De Waele et al., 2009) and  
184 Pan-African Orogens (950–450 Ma; Kröner, 1984; Stern, 1994). Neoproterozoic-Palaeozoic  
185 eclogite facies rocks dated at ca. 593, 548, 523 Ma provide particularly compelling evidence for  
186 disparately younger and previously unrecognized paleo-sutures within the Palaeoproterozoic  
187 Ubendian Belt (Boniface and Schenk, 2012). Together the available ages suggest that the  
188 Ubendian Belt is the product of at least three temporally discrete orogenic episodes. Moreover,  
189 later rifting associated with the western branch of the East African Rift, also likely contributed to  
190 the current geometry of the Ubendian Belt (Theunissen et al., 1996). The metallogenic  
191 implications of these overprinting metamorphic events are discussed further below.

192

## 193 **2.0 Local Geology**

### 194 *2.1 Geology of Lupa Goldfield*

195 The field area for the current study is located in the western portion of the Lupa  
196 Goldfield, SW Tanzania (Figs. 1–3). Previously reported U-Pb and Re-Os ages are summarized  
197 in Fig. 4 and Table 1. Foliated Archaean granites (2.76–2.72 Ga) are cut by non-foliated  
198 Palaeoproterozoic granites and dioritic-gabbroic intrusions (1.96–1.88 Ga; Lawley et al., 2013).  
199 Archaean granites are generally coarse grained, equigranular to porphyritic (K feldspar  
200 phenocrysts) and represent a range of modal mineralogy (syeno- to monzogranite) and  
201 compositions (Lawley et al., 2013). Primary Fe-Mg minerals have been variably replaced to a  
202 chlorite ± epidote ± muscovite ± calcite ± titanite ± magnetite mineral assemblage.

203 The oldest tectonic fabric (D1), which is the focus of this study, is largely restricted to  
204 these Archaean granites and is characterized by alternating quartzofeldspathic and chloritic folia

205 that locally gives the granites a “gneiss-like” appearance (see legend of Fig. 3; Kimambo, 1984).  
206 The compositional banding is generally steeply dipping and varies in strike from E-W to NW-SE,  
207 but the fabric orientation and intensity can vary remarkably at the scale of individual outcrops.  
208 Cross cutting and overprinting relationships observed in drill core and outcrops imply that the D1  
209 fabric, at least locally, preceded the emplacement of Palaeoproterozoic and non-foliated  
210 intrusions (i.e., D1 is locally  $\geq 1.96$  Ga). However, exposures of the Ilunga Syenogranite (ca. 1.96  
211 Ga) at the Dubwana exploration target (Fig. 4) locally exhibit compositional banding (i.e.,  
212 alternating quartzofeldspathic and chloritic folia) akin to the D1 fabric observed in Archaean  
213 granites and may suggest that D1 occurred diachronously during Palaeoproterozoic magmatism  
214 (discussed further below).

215 We expect that the D1 foliation developed at greenschist facies metamorphic conditions  
216 based on the chlorite  $\pm$  muscovite  $\pm$  epidote  $\pm$  calcite metamorphic mineral assemblage and the  
217 dominant deformation microstructures (e.g., undulose extinction of quartz; crystal plastic  
218 deformation of quartz; Lawley et al., in press-b), which are typical of mid-crustal levels (300–  
219 450°C; 1–3 kbar; Scholz, 1988). However, these inferred P-T conditions are notably lower than  
220 the amphibolite facies metamorphism that characterizes the other Ubendian Terranes and contrast  
221 with the “Acid Gneiss and Schist” lithologies that are reported to constitute large swaths of the  
222 Lupa Goldfield (Fig. 3; Kimambo, 1984). The field area for the current study (Fig. 4) is also  
223 devoid of ‘acid volcanic rocks’ and we suggest that the Lupa Goldfield geology map should be  
224 treated with caution (Fig. 3).

225 Palaeoproterozoic granitic-gabbroic intrusions in the Lupa Goldfield represent part of a  
226 voluminous magmatic arc at the Tanzanian cratonic margin (Figs. 2–4; see Many, 2011 for  
227 further details). Two of these granites, the Saza Granodiorite and the Ilunga Syenogranite, are  
228 regionally extensive and are exposed across large tracts of the Lupa Goldfield and the field area  
229 for the current study (Figs. 3, 4). The Ilunga syenogranite ( $1960 \pm 1$  Ma; Lawley et al., 2013) is  
230 exposed in the northern portion of the field area and consists of a coarse grained and equigranular  
231 mineral assemblage of K feldspar  $\pm$  quartz  $\pm$  plagioclase with lesser amount of chloritized biotite.  
232 In contrast, the Saza Granodiorite ( $1935 \pm 1$  Ma; Lawley et al., 2013) is exposed in the southern  
233 portion of the field area and is characterized by a coarse grained and equigranular mineral  
234 assemblage of plagioclase  $\pm$  quartz  $\pm$  K feldspar with lesser amounts of chloritized hornblende  $\pm$   
235 biotite. Both granitoids possess dioritic-gabbroic enclaves and are also cut by dioritic-gabbroic  
236 dikes, which together suggest that intermediate-mafic magmatism pre- and post-dated the granitic  
237 intrusions. All magmatic phases in the field area have experienced greenschist facies  
238 metamorphism and are cut by D2 brittle-ductile and mylonitic shear zones (discussed further

239 below). Later brittle faults cutting mineralized structures (D3) and faults filled with  
240 unconsolidated gouge likely record younger, but undated deformation event(s) that may be  
241 related to periodic reactivation of Palaeoproterozoic structures to the present day (Theunissen et  
242 al., 1996).

243

## 244 *2.2 Gold Deposit Geology*

245 Most of the exposed gold deposits and prospects are hosted by steeply south-dipping D2  
246 quartz veins and greenschist facies mylonitic shear zones (Figs. 4, 6a–c). The majority of these  
247 are associated with the ENE-WSW trending Saza shear zone and two other dominant structural  
248 trends (NW-SE and E-W; Fig. 4). The Kenge and Mbenge deposits, which are broadly  
249 representative of the majority of prospects studied, contain a measured and indicated resource of  
250 8.7 Mt at 1.33 g/t gold with a 0.5 g/t cut-off resulting in an estimated 0.37 Moz of contained gold  
251 (Simpson, 2012). The deposits are hosted, in part, by mylonitic shear zones cutting foliated  
252 Archaean granite (e.g., Fig. 6a). Hydrothermal alteration, strain and mineralization at these  
253 deposits are largely restricted to the host shear zone and associated laminated-quartz ( $\pm$   
254 carbonate) veins (Fig. 6c), whereas quartz veins outside of the main shear zone represent a  
255 relatively minor contribution to the overall mineral resource.

256 Recent drilling by Helio Resource Corp. has identified a second, geologically distinct,  
257 style of mineralization (e.g., Porcupine) that comprises a moderately to steeply dipping panel of  
258 narrow and discontinuous mylonitic shear zones separated by non-foliated, but hydrothermally  
259 altered (sericitized and silicified) and veined granite (Fig. 6d, e). The Porcupine deposit contains  
260 a measured and indicated resource of 15.4 Mt at 1.31 g/t gold with a 0.5 g/t cut-off for an  
261 estimated 0.65 Moz of contained gold (Simpson, 2012). The highest gold grades (e.g., up to 40 g/t  
262 Au) at Porcupine are associated with shallow dipping quartz veins and intervals of sericitized,  
263 silicified and non-foliated Ilunga syenogranite (Fig. 6d, e). These auriferous and shallow dipping  
264 quartz veins primarily occur within the moderately to steeply dipping panel of hydrothermally  
265 altered granite and significantly widen the mineralized zone at Porcupine, but are notably absent  
266 at Kenge and Mbenge. Regardless of the deposit style, gold is associated with a relatively simple  
267 sulphide mineral assemblage of pyrite  $\pm$  chalcopyrite  $\pm$  molybdenite  $\pm$  galena  $\pm$  sphalerite.

268

## 269 *2.3 Timing of Gold Mineralization*

270 Five gold deposits and prospects were sampled for Re-Os geochronology (Fig. 4; Lawley  
271 et al., in press-a). Individual Re-Os molybdenite, pyrite and chalcopyrite model ages range from  
272 1.95–1.88 Ga and are thus broadly contemporaneous with the entire 1.96–1.88 Ga magmatic



273 history and the 1.9–1.8 Ga Ubendian tectonic phase (Fig. 5; Table 1). The absolute timing of gold  
274 within this temporal framework remains equivocal since gold is preferentially concentrated along  
275 pyrite crystal boundaries and locally occurs as pyrite fracture fills. These paragenetic  
276 relationships may provide evidence for a relatively late, and undated, gold event (Lawley et al., in  
277 press-a). Nevertheless, gold is observed as inclusions within ca. 1.88 Ga pyrite hosted by the  
278 main greenschist facies mylonitic shear zone at Kenge. This proposed ca. 1.88 Ga event is also  
279 recorded at all four of the other dated prospects (Mbenge, Porcupine, Konokono, and Dubwana).  
280 The paragenetic relationship between gold and these ca. 1.88 Ga pyrite and/or chalcopyrite  
281 samples at the other dated prospects is less clear, but the majority of samples were chosen from  
282 mineralized intervals and we suggest that the broad overlap of Re-Os sulphide ages between  
283 deposits and across the study area argues for a regional metallogenic event at ca. 1.88 Ga.

284 Disparately older ca. 1.95 and 1.94 Ga Re-Os molybdenite and pyrite model ages at  
285 Kenge were sampled from laminated and auriferous quartz veins hosted within the main  
286 mylonitic shear zone. The paragenetic relationship between gold and these sulphide samples is  
287 unclear and it remains equivocal whether: 1) the ca. 1.88 Ga gold event merely represents the  
288 youngest metallogenic event within a protracted (1.95–1.88 Ga) hydrothermal and progressive  
289 deformation history; or 2) the anomalously older Re-Os ages provide evidence for a telescoped  
290 deposit whereby the gold event at ca. 1.88 Ga merely overprints earlier, but unrelated style(s) of  
291 mineralization. Unfortunately the available ages and the ambiguous gold paragenesis do not allow  
292 us to rule out either of these possibilities. Nonetheless, the range of Re-Os ages suggest that  
293 laminated quartz veins possess protracted hydrothermal histories characterized by overprinting  
294 sulphidation events and that hydrothermal activity, at least locally, preceded the development of  
295 D2 mylonitic shear zones (Lawley et al., in press-a).

296

### 297 **3.0 U-Pb titanite ID-TIMS Geochronology**

#### 298 *3.1 Sample selection*

299 Foliated Archaean granites, such as the one selected for this study (sample CL109; U-Pb  
300 zircon La-ICP-MS weighted average  $^{207}\text{Pb}/^{206}\text{Pb}$  age  $2757 \pm 10$  Ma; Table 1; Lawley et al., 2013),  
301 are characterized by alternating quartzofeldspathic and chlorite-rich folia that developed during  
302 D1. Titanite crystals are concentrated within the D1 chlorite-rich folia and the greenschist facies  
303 metamorphic assemblage of chlorite  $\pm$  epidote  $\pm$  calcite  $\pm$  titanite overprinting the protolith's Fe-  
304 Mg minerals (Figs. 7a, c). The petrographic association of titanite with the tectonic fabric and  
305 metamorphic mineral assemblage is consistent with titanite neo-crystallization during greenschist

306 facies metamorphism (Essex and Gromet, 2000; Frost et al., 2001; Parrish, 2001). This mineral  
307 association is particularly apparent for the least-deformed examples of Archaean granites where  
308 mineral dissolution and volume loss is unlikely to explain the apparent concentration of  
309 metamorphic minerals. Moreover, euhedral titanite crystals isolated from the metamorphic  
310 mineral assemblage were not observed and suggests that any igneous titanite, which would be  
311 expected given the granitic composition, may have been recrystallized during subsequent  
312 metamorphism. The nature of the metamorphic phase transition is unclear, but could be related to  
313 the breakdown of primary clinopyroxene and/or amphibole by the hydration reaction  
314 clinopyroxene + ilmenite + quartz + H<sub>2</sub>O = amphibole + titanite and/or the oxidation reaction  
315 amphibole + ilmenite + O<sub>2</sub> = titanite + magnetite + quartz + H<sub>2</sub>O (Harlov et al., 2006).  
316 Metamorphic reactions such as these are postulated to play an important role during metamorphic  
317 titanite crystallization at greenschist facies P-T conditions (Frost et al., 2001).

318 Primary Fe-Mg minerals of the non-foliated Palaeoproterozoic Saza Granodiorite (sample  
319 CL1035; U-Pb zircon ID-TIMS weighted average <sup>207</sup>Pb/<sup>206</sup>Pb age at 1935 ± 1 Ma; Lawley et al.,  
320 2013) are also replaced to a greenschist facies metamorphic mineral assemblage of chlorite ±  
321 epidote ± calcite ± titanite. The association of titanite with this metamorphic mineral assemblage  
322 is also consistent with titanite neo-crystallization during greenschist facies metamorphism,  
323 however euhedral titanite crystals isolated from the metamorphic mineral assemblage may  
324 represent relict magmatic titanite (Fig. 7d). Multiple titanite populations are characteristic of  
325 metamorphosed lithologies and are typically distinguished through a combination of  
326 optical/chemical characteristics and/or comparison of titanite ages with independent estimates for  
327 the crystallization age of the sample (Jung and Hellebrand, 2007). In this contribution, we follow  
328 a similar approach and compare new U-Pb titanite ages with previously reported U-Pb zircon  
329 ages for the same samples.

330 The titanite crystals present within the bulk mineral separate from sample CL109 and  
331 CL1035 are translucent, range in colour from brown to clear, and are present as broken fragments  
332 and fine-grained wedge-shaped crystals. For sample CL109, clear and relatively fine grained  
333 titanite crystals devoid of inclusions were chosen (Fig. 7e); whereas relatively much larger clear-  
334 brown titanite crystals devoid of inclusions from sample CL1035 were selected for U-Pb ID-  
335 TIMS analysis (Figs. 7f). The significance of these visually distinct titanite fractions is discussed  
336 further below.

337

### 338 *3.2 Sample Preparation and ID-TIMS Methodology*

339 All of the analysed titanite crystals were ultrasonically cleaned for an hour before being  
340 placed on a hotplate for 30 minutes, photographed in transmitted light and rinsed in ultrapure  
341 acetone. After rinsing, titanite fractions were transferred to 300  $\mu$ l Teflon FEP microcapsules and  
342 spiked with a mixed  $^{233}\text{U}$ – $^{235}\text{U}$ – $^{205}\text{Pb}$  tracer. Titanite crystals were then dissolved in  $\sim$ 120  $\mu$ l of 29  
343 M HF with a trace amount of 30%  $\text{HNO}_3$  within microcapsules, placed in Parr vessels at  $\sim$ 220°C  
344 for 48 hours, dried to fluorides and then converted to chlorides at  $\sim$ 180°C overnight. U and Pb for  
345 all titanite fractions were separated using standard HBr and  $\text{HNO}_3$ -based anion-exchange  
346 chromatographic procedures. Isotope ratios were measured at the NERC Isotope Geosciences  
347 Laboratory (NIGL), UK, using a Thermo-Electron Triton Thermal Ionisation Mass-Spectrometer  
348 (TIMS). Pb and U were loaded separately on a single Re filaments in a silica-gel/phosphoric acid  
349 mixture. Pb was measured by peak hopping on a single SEM detector. U isotopic measurements  
350 were made in static Faraday mode. Age calculations and uncertainty estimation (including U/Th  
351 disequilibrium) was based upon the algorithms of Schmitz and Schoene (2007) and Crowley et al.  
352 (2007), using the updated consensus value of  $^{238}\text{U}/^{235}\text{U} = 137.818 \pm 0.045$  (Hiess et al., 2012) and  
353 the decay constant and its uncertainty of Jaffey et al. (1971).

354

## 355 **4.0 U-Pb titanite ID-TIMS results and interpretation**

### 356 *4.1 Common lead correction*

357 Titanite crystals can incorporate significant concentrations of common lead during  
358 crystallization, which must be corrected for in order to determine accurate U-Pb ages (e.g., Frost  
359 et al., 2001). In general, the available approaches to common lead correction can be divided into  
360 three broad categories: 1) measure the lead isotopic composition of co-genetic and low-U phases,  
361 such as feldspar, and assume this isotopic composition is equivalent to common Pb at the time of  
362 titanite crystallization; 2) apply lead isotopic evolution models, such as the two-stage lead  
363 evolution model of Stacey and Kramers (1975; S-K model), and assume that this composition is  
364 equivalent to the isotopic composition of common Pb at the time of titanite crystallization; and 3)  
365 utilize regression techniques, such as the ‘Total Pb/U isochron’, which requires no *a priori*  
366 assumption of the common lead’s isotopic composition (Ludwig, 2008).

367 The analysis of low-U mineral phases (e.g., feldspar) was deemed inappropriate for our  
368 data set since titanite crystals occurs with the metamorphic mineral assemblage and is therefore  
369 unlikely to be in isotopic equilibrium with igneous feldspar. Furthermore, previous Pb isotopic  
370 studies have demonstrated complex Pb isotopic systematics of feldspar (and galena) for samples

371 taken from the Ubendian and Usagaran Belts, which likely reflects metamorphic overprinting and  
372 open system behavior during multiple orogenic cycles at the Tanzanian cratonic margin (e.g.,  
373 Möller et al., 1998). The ‘Total Pb/U Isochron’ (Ludwig, 1998; Ludwig, 2008) makes no *a priori*  
374 assumption as to the isotopic composition of common Pb and, providing the data set represents a  
375 suite of co-genetic and undisturbed titanite fractions that share a common Pb isotopic  
376 composition, represents a robust approach to common lead correction (e.g., Corfu and Stone,  
377 1998; Storey et al., 2006). Unfortunately, titanite fractions from CL109 and CL0975 show  
378 evidence of disturbance to the U-Pb systematics (discussed further below) and so for the  
379 remaining discussion common lead for both samples was corrected by employing the S-K model  
380 approach. The sensitivity of the analyzed titanite fractions to the S-K model common lead  
381 correction is discussed further below.

## 382 4.2 U-Pb titanite ID-TIMS results and interpretation

### 383 4.2.1 Archaean Granite (CL109)

384 U-Pb titanite data are reported in Table 2 and Fig. 8. For the foliated Archaean granite  
385 (sample CL109), nine titanite fractions yield  $^{238}\text{U}/^{204}\text{Pb}$  and  $^{206}\text{Pb}/^{204}\text{Pb}$  ratios ranging from 110–  
386 545 and 40–190, respectively. Common lead is significant for this dataset as suggested by the low  
387  $\text{Pb}^*/\text{Pb}_c$  ratios (1–9; where  $\text{Pb}^*$  and  $\text{Pb}_c$  correspond to radiogenic and common lead, respectively).  
388 As a result, the choice of the isotopic composition of common lead can have a substantial impact  
389 on the calculated U-Pb ages. We tested the effect of the assumed common lead composition on  
390 titanite ages by varying the S-K model age composition from 2.5–1.5 Ga and found that some  
391 individual  $^{207}\text{Pb}/^{206}\text{Pb}$  titanite ages varied considerably, whereas the weighted average  $^{207}\text{Pb}/^{206}\text{Pb}$   
392 titanite age of all the analyzed titanite fractions did not (weighted average  $^{207}\text{Pb}/^{206}\text{Pb}$  ages range  
393 from 1943–1911 Ma; Fig. 8a inset). Titanite fractions S3 and S4 are considerably more sensitive  
394 to the assumed common lead composition and also possess anomalously younger  $^{207}\text{Pb}/^{206}\text{Pb}$  ages  
395 for S-K model Pb compositions at 1.5 and 1.9 Ga, whereas the other seven titanite fractions were  
396 relatively insensitive to the S-K model lead isotopic composition and possess  $^{207}\text{Pb}/^{206}\text{Pb}$  ages that  
397 overlap within analytical uncertainty at the  $2\sigma$  level (Fig. 8a inset). As a result, the common lead  
398 isotopic composition was assumed by selecting the S-K model lead isotopic composition at 1.9  
399 Ga, which we use as a reasonable first order approximation of its true value.

400 A conventional 2D Terra-Wasserburg regression of the seven common Pb corrected and  
401 insensitive titanite fractions yield a Model 1 York regression solution upper intercept age of  $1921$   
402  $\pm 7$  Ma (MSWD = 0.7; probability of fit = 0.6;  $n = 7$ ; Fig. 8a). It is important to note, that these

403 Palaeoproterozoic ages for titanite are all considerably younger than the LA-ICP-MS weighted  
404 average  $^{207}\text{Pb}/^{206}\text{Pb}$  zircon age for the same sample ( $2757 \pm 10$  Ma; Fig. 8a; Lawley et al., 2013).  
405 The geological significance of these Palaeoproterozoic titanite ages is discussed further below.

406

#### 407 *4.2.2 Saza Granodiorite (CL0975)*

408 The five titanite fractions analyzed from the Saza Granodiorite (sample CL0975) possess  
409 a range of  $^{238}\text{U}/^{204}\text{Pb}$  and  $^{206}\text{Pb}/^{204}\text{Pb}$  from 294–859 and 118–315, respectively. These ratios, and  
410 in particular the relatively higher  $\text{Pb}^*/\text{Pbc}$  ratios (5–13), suggest that titanite crystals from  
411 CL0975 are more U rich and possess less common Pb than titanite analyses from CL109. We  
412 tested the effect of the assumed common Pb composition on titanite ages, as above, by varying  
413 the S-K model age composition from 2.5–1.5 Ga and found that individual  $^{207}\text{Pb}/^{206}\text{Pb}$  titanite  
414 ages and weighted average  $^{207}\text{Pb}/^{206}\text{Pb}$  titanite ages were both relatively insensitive to the inferred  
415 isotopic composition of common lead (weighted average  $^{207}\text{Pb}/^{206}\text{Pb}$  ages range from 1942–1924  
416 Ma; Fig. 8b inset). As above, the common lead isotopic composition was assumed by selecting  
417 the S-K model lead isotopic composition at 1.9 Ga, which we expect is the best first order  
418 approximation of its true value.

419 A conventional 2D Terra-Wasserburg regression of the five common Pb corrected titanite  
420 fractions yield a Model 1 York regression solution upper intercept age of  $1931 \pm 3$  Ma (MSWD =  
421 1.1; probability of fit = 0.4;  $n = 5$ ; Fig. 8b). This upper intercept regression age is comparable to  
422 a weighted average  $^{207}\text{Pb}/^{206}\text{Pb}$  age of the three concordant titanite fractions, which represent the  
423 least disturbed titanite fractions, at  $1930 \pm 3$  Ma (MSWD = 0.6; probability of fit = 0.5;  $n = 3$ ).  
424 Both the weighted average (based on the three concordant and thus least disturbed titanite  
425 fractions) age and the 2D regression age (based on all of the analyzed titanite fractions) ages are  
426 in excellent agreement with each other and are only slightly outside of analytical uncertainty at  
427 the  $2\sigma$  level with a high-precision ID-TIMS weighted average  $^{207}\text{Pb}/^{206}\text{Pb}$  zircon age of  $1935 \pm 1$   
428 Ma (Lawley et al., 2013) for the same sample (Fig. 8b).

429

## 430 **5.0 Discussion**

### 431 *5.1 Geological significance of U-Pb titanite ages*

#### 432 *5.1.1 Archaean Granite (CL109)*

433 Titanite crystals from the foliated Archaean granite sample (CL109) are spatially  
434 associated with the metamorphic fabric/mineral assemblage and provide evidence for titanite neo-  
435 crystallization related to the breakdown of primary Fe-Mg minerals during greenschist facies  
436 metamorphism and the development of the D1 deformation fabric (Fig. 7). The generally low

437 Pb\*/Pb<sub>c</sub> ratios (1–9; Table 2) of the analyzed titanite fractions support this interpretation and are  
438 consistent with the compositional characteristics (i.e., relatively high common lead and low U) of  
439 metamorphic titanite reported in previous studies (e.g., Frost et al., 2001; Jung and Hellebrand,  
440 2007). Whilst the precise age of titanite fractions from CL109 remains open to interpretation due  
441 to, in part, the sensitivity of U-Pb ages to the assumed isotopic composition of common lead,  
442 these new U-Pb data clearly show that titanite crystals are in fact Palaeoproterozoic and therefore  
443 considerably younger than Archaean U-Pb zircon ages for the same sample (Fig. 8a; Table 1).

444 Sample CL109 is cut by a non-foliated gabbroic dyke that is dated by LA-ICP-MS U-Pb  
445 zircon geochronology at ca. 1.88 Ga (Lawley et al., 2013), which is consistent with the ca 1.92 Ga  
446 U-Pb titanite age interpreted to date the timing of D1. However, elsewhere foliated Archaean  
447 granites are cut by ca. 1.96 Ga granodioritic-granitic dykes and intrusions (most notably the 1960  
448 ± 1 Ma Ilunga Syenogranite; Lawley et al., 2013), which is significantly older (ca. 40 Myr older  
449 than the upper Concordia intercept age at 1921 ± 7 Ma ; Fig. 8a) than the inferred D1 timing  
450 reported here. This suggests either: 1) D1 is more complex than previously thought and  
451 developed diachronously during Palaeoproterozoic magmatism (Fig. 9); and/or 2) that U-Pb  
452 titanite dates from sample CL109 are younger than their true age as a result of Pb-loss and/or  
453 inappropriate common Pb correction (Fig. 8b). The latter is supported by the absence of a ductile  
454 deformation fabric within the Saza Granodiorite (ca. 1.93 Ga), which would be expected if D1  
455 had occurred at ca. 1.92 Ga (Fig. 9e). Alternatively, the D1 fabric may have developed  
456 episodically during the emplacement of Palaeoproterozoic granites, which is consistent with  
457 compositional banding akin to the D1 fabric observed in Archaean granites, but locally observed  
458 in the Ilunga Syenogranite at Dubwana.

459 The current data set does not allow us to rule out either of these possibilities and thus the  
460 precise timing of D1 deformation remains unclear. Nevertheless, these new U-Pb titanite ages  
461 provide new evidence to suggest that the earliest identifiable deformation event in the Lupa  
462 Goldfield occurred during the Palaeoproterozoic and some 40 Myr prior to D2 mylonitic shear  
463 zones and gold at ca. 1.88 Ga. This extended Palaeoproterozoic structural history places  
464 important new constraints on the onset of deformation related to Ubendian orogenesis and raises  
465 the possibility that fabric development during D1 may have played an important role in localizing  
466 strain during later deformation (see below for further discussion).

467

#### 468 *5.1.2 Saza Granodiorite (CL0975)*

469 Titanite crystals from the non-foliated Saza granodiorite (sample CL0975) are locally  
470 associated with the greenschist facies mineral assemblage overprinting the Saza Granodiorite and

471 thus exhibit textural similarities in thin section to titanite crystals extracted from the foliated  
472 Archaean granite (sample CL109). However, titanite crystals isolated from the Saza Granodiorite  
473 are visually distinct (i.e., larger and are brown-translucent), which suggests that these titanite  
474 crystals may represent a titanite population dissimilar to metamorphic titanite observed in sample  
475 CL109 (Figs. 7e, f). The higher  $^{206}\text{Pb}/^{238}\text{U}$  ratios and generally higher  $\text{Pb}^*/\text{Pb}_c$  ratios (5–13; Table  
476 2) within the analyzed titanite fractions from the Saza Granodiorite are also unlike titanite  
477 fractions from the foliated Archaean granite and are more consistent with the more radiogenic  
478 composition of magmatic titanite (e.g., Jung and Hellebrand, 2007).

479         The weighted average  $^{207}\text{Pb}/^{206}\text{Pb}$  titanite age of the three concordant titanite analyses  
480 ( $1930 \pm 3$  Ma) are slightly outside of analytical uncertainty of the weighted average  $^{207}\text{Pb}/^{206}\text{Pb}$   
481 zircon age ( $1935 \pm 1$  Ma) of concordant zircons for the same sample and support a magmatic  
482 origin for the analyzed titanite fractions. The broad agreement between both geochronometers  
483 suggests that the Stacey and Kramers (1975) common lead correction approach is broadly  
484 appropriate despite two of the common lead corrected titanite fractions exhibiting severe reverse  
485 discordance (Fig. 8b). The cause of reverse discordance is unclear and could be related to U loss  
486 and/or represent an analytical effect. The slight discrepancy between the interpreted  
487 crystallization age and U-Pb titanite ages could be related to minor Pb-loss in the analyzed titanite  
488 fractions, later closure during cooling, sub-solidus recrystallization, and/or inadequate accounting  
489 for the common lead correction (Frost et al., 2001). If the nominally younger U-Pb titanite ages  
490 represent cooling ages, our results would suggest that the Saza Granodiorite cooled relatively  
491 quickly since U-Pb titanite and zircon ages are only  $<1$  Myr outside of analytical uncertainty at  
492 the  $2\sigma$  level.

493

## 494 *5.2 Implications for ore deposit models*

495         Here we integrate new U-Pb titanite ages with previously reported U-Pb and Re-Os ages  
496 to address some of the outstanding uncertainties (e.g., the temporal relationship between D1 and  
497 D2) at gold prospects within the Lupa Goldfield. At least three temporally distinct hydrothermal  
498 events have been identified in the Lupa Goldfield (1.95, 1.94 and 1.88 Ga; Lawley et al., in press-  
499 a; Fig. 5). Gold is expected to have been introduced at ca. 1.88 Ga, whereas the relationship  
500 between gold and disparately older Re-Os model ages at ca. 1.95 Ga and 1.94 Ga is less clear.  
501 Each of these broadly defined hydrothermal events are represented in detail by complex vein  
502 histories that suggest each event occurred at a time scale that is less than the resolution of the Re-  
503 Os method. This episodic, but protracted hydrothermal history (1.95–1.88 Ga) overlaps with the  
504 Palaeoproterozoic magmatic history of the Goldfield as determined by U-Pb zircon dating of

505 felsic-mafic intrusions/dykes (1.96–1.88 Ga; Fig. 5). Furthermore, high-precision U-Pb zircon  
506 ID-TIMS ages for the Saza Granodiorite ( $1935 \pm 1$  Ma) overlap with Re-Os molybdenite ages  
507 (ca. 1.94 Ga) and provide unequivocal evidence for sulphidation that is concomitant with  
508 magmatism at the million-year time scale.

509 This close temporal relation has led to the suggestion that gold deposits within the Lupa  
510 Goldfield belong to the intrusion-related ore deposit type (Manya, 2012). However, Re-Os model  
511 ages pre- and post-date individual magmatic phases and hydrothermal activity appears to have  
512 occurred at a time scale that is far greater than the expected duration of a single magmatically  
513 derived hydrothermal fluid circulation system (e.g., <1 Myr; von Quadt et al., 2011). Moreover,  
514 the proposed gold event at ca. 1.88 Ga is disparately younger than any of the dated granites and  
515 we suggest that a simple intrusion-related deposit model (e.g., Manya, 2012) whereby  
516 metalliferous and hydrothermal fluids exsolving from the Saza Granodiorite are solely  
517 responsible for gold is unsupported.

518 New U-Pb titanite ages reported here further constrain ore deposit models by  
519 demonstrating that the onset of deformation occurred during the Palaeoproterozoic. It further  
520 suggests that deformation was progressive and broadly overlaps with Palaeoproterozoic  
521 magmatism (Figs. 5, 9). The broad temporal overlap between hydrothermal (1.95–1.88 Ga),  
522 magmatic (1.96–1.88 Ga) and deformation ( $\geq 1.92$ –1.88 Ga) events in the Lupa Goldfield are also  
523 well correlated to the tectono-thermal history recorded by the other Ubendian Terranes (i.e., the  
524 2.1–1.8 Ga Ubendian Orogeny and specifically the 1.9–1.8 Ga Ubendian tectonic phase; Lenoir et  
525 al., 1994). In particular, the Palaeoproterozoic MORB-like chemistry eclogites (ca. 1.89 and 1.86  
526 Ga; Boniface et al., 2012) sampled from the Ubende Terrane demonstrates that high-grade  
527 metamorphism related to subduction of oceanic crust broadly overlaps with the proposed gold  
528 event at ca. 1.88 Ga within the Lupa Goldfield. Palaeoproterozoic eclogitic rocks in the Ubende  
529 Terrane are amongst the oldest eclogites on Earth (Boniface et al., 2012) and therefore represent a  
530 key link between metallogenesis related to modern and ancient subduction zone processes. A  
531 holistic understanding of Ubendian tectonics requires additional constraints on the significance  
532 and distribution of Meso- and Neoproterozoic metamorphic overprints. Nevertheless, the link  
533 between Palaeoproterozoic convergent tectonics, subduction zone processes and gold  
534 mineralization in the Lupa Goldfield is implied.

535 The consistent sulphide mineralogy between gold prospects, in conjunction with a  
536 comparable alteration mineral assemblage and overlapping sulphide ages, suggests that all of the  
537 studied gold prospects are consanguineous and can be considered as part of the same broad  
538 mineralization history related to an interconnected shear zone network that focused fluids at the



539 Tanzanian cratonic margin during Palaeoproterozoic Ubendian orogenesis. Together the timing,  
540 structural setting and geological characteristics suggest that gold prospects exposed in artisanal  
541 workings within the western Lupa Goldfield are typical of the orogenic gold deposit type [i.e.,  
542 shear and quartz-carbonate vein hosted, sulphide poor and Au dominated deposits that result from  
543 structural focusing of low salinity H<sub>2</sub>O-CO<sub>2</sub> ( $\pm$  CH<sub>4</sub>) fluids at convergent margins; *sensu* Groves  
544 et al., 1998]. Our ages are also consistent with other palaeo-convergent margins where systematic  
545 geochronology has demonstrated that orogenic style gold mineralization can occur throughout  
546 orogenesis, but generally occurs relatively late during the orogen's tectono-thermal history (e.g.,  
547 Sarma et al., 2011; Figs. 5, 9f).

548

### 549 *5.3 Archaean versus Proterozoic Tanzanian Gold Deposits*

550 In the preceding section we argue that gold deposits in the Lupa Goldfield share  
551 similarities to the orogenic gold deposit type. In this section we compare and contrast the geology  
552 of goldfields hosted within and at the margins of the Tanzanian Craton. The geologic  
553 characteristics of deposits within each goldfield are taken from the available literature and are  
554 summarized in Table 3. Overall, gold deposits within the Lake Victoria Goldfield share a number  
555 of broad similarities with Palaeoproterozoic (e.g., ca. 1.88 Ga, Lupa), Mesoproterozoic (e.g., ca.  
556 1.2 Ga, Mpanda; Kazimoto and Schenk, 2013), and Palaeozoic (e.g., ca. 483 Ma, Niassa;  
557 Bjerkgard et al., 2009) goldfields situated within metamorphic belts marginal to the Tanzanian  
558 Craton (Fig. 1). These gross geologic similarities include comparable hydrothermal alteration  
559 mineral assemblages (sericite, chlorite, silica flooding, carbonate), sulphide mineral assemblages  
560 (pyrite  $\pm$  base-metal sulphides), metamorphic grade (amphibolite to greenschist) and similar  
561 apparent controls on gold mineralization (predominately shear- and quartz vein-controlled and  
562 lithologic contacts). However, in detail each deposit also possesses distinct geologic  
563 characteristics that make direct comparisons between deposits — even within a single goldfield  
564 — challenging. For example, the BIF-hosted Geita and intrusion-hosted Buzwagi deposits differ  
565 from the other meta-sedimentary and –volcanic rocks hosted deposits in the Lake Victoria  
566 Goldfield (Table 3).

567 Nevertheless goldfield comparisons such as those presented in Table 3 represent an  
568 important exercise since several of the metamorphic belts enveloping that Tanzanian Craton are  
569 now recognized to comprise, in part, re-worked Archaean crust (Ubendian Belt, Kazimoto and  
570 Schenk, 2013, Lawley et al., 2013; Usagaran, Sommer et al., 2005b; Mozambique Belt, Kröner et  
571 al., 2003; Sommer et al., 2005a; Thomas et al., 2013). As a result, these metamorphic belts may  
572 host Archaean gold deposits that have been subsequently re-worked during Palaeoproterozoic-

573 Palaeozoic orogenic episodes. Unequivocal examples of re-worked Archean deposits have not  
574 been documented in Tanzania, but represent hitherto unrecognized gold potential within  
575 Proterozoic-Palaeozoic metamorphic belts surrounding the Tanzanian Craton (Kabete et. al.,  
576 2012a). Moreover, orogenic gold deposits hosted by re-worked Archean crust are scarce globally  
577 although the Mesozoic goldfields within the North China Craton may represent possible  
578 exceptions (Li et al., 2012).

579 Gold deposits within the Lupa goldfield provide other natural examples of this unusual  
580 deposit setting since the inferred ca. 1.88 Ga gold event is hosted, in part, by re-worked Archean  
581 granitoids. Individual Mesoproterozoic Re-Os ages at ca. 1371 Ma and 1057–922 Ma (Lawley et  
582 al., in press-a) and inferred lead-loss events at 514–469 Ma (Lawley et al., 2013), although not  
583 directly related to gold in the Lupa goldfield, are potentially related to later overprinting during  
584 multiple and temporally-discrete orogenic cycles (ca. 1375 Ma, Karagwe-Ankole Belt; 1.05–1.00  
585 Ga Irumide Orogeny; 950–450 Ma Pan African Orogeny). Preliminary U-Pb monazite ages at ca.  
586 1.2 Ga provide evidence for a potential Mesoproterozoic gold event at the Mpanda goldfield,  
587 which is also hosted, in part, by Archean rocks re-worked during Palaeo- and Mesoproterozoic  
588 tectono-thermal events (Kazimoto and Schenk, 2013). However, the precise age of the Mpanda  
589 goldfield remains unclear because previously reported Pb-Pb model ages provided evidence for a  
590 disparately younger gold event at ca. 720 Ma (Stendal et al., 2004). Neoproterozoic-Palaeozoic gold  
591 is also reported in the ca. 743 Ma Niassa Goldfield, which are interpreted to reflect Pan-African  
592 orogenic style gold mineralization in NW Mozambique (Bjerkgard et al., 2009). Similar Pan-  
593 African gold events are also reported in other countries bordering the Tanzanian Craton although  
594 the geologic setting of these deposits remains poorly understood (Burundi, Rwanda, Uganda;  
595 Brinckmann et al., 1994; Fernandez-Alonso et al., 2012).

596 The examples provided above suggest that the metamorphic belts bordering the  
597 Tanzanian Craton are prospective for orogenic style gold mineralization spanning at least three  
598 orogenic episodes. The available geochronologic database highlights the importance of robust  
599 geochronometers that remain closed during overprinting geologic events. It also stresses the need  
600 for robust geochronometers that record and, can be paragenetically linked to, different stages of  
601 the tectono-thermal history. The latter is particularly important in poly-orogenic settings where  
602 the enrichment or possible remobilization of gold through time is expected to be a key process for  
603 gold deposits hosted by metamorphic belts (Groves et al., 2003).

604

605 *5.4 Structural evolution of the Lupa Goldfield*

606 New U-Pb titanite ages from an Archaean granite sample (CL109) demonstrates the onset  
607 of deformation in the Lupa Goldfield occurred during the Palaeoproterozoic and pre-dated D2  
608 auriferous mylonites by  $\geq 40$  Myr. Here we suggest that foliated Archaean granites, developed  
609 during D1, acted as zones of pre-existing structural weakness and/or heterogeneity and may have  
610 played a key role in strain localization during later deformation. This is particularly apparent for  
611 gold prospects hosted by D2 mylonitic shear zones developed within foliated Archaean granites  
612 such as at Kenge and Mbenge (Fig. 6a). At these mineral systems, the orientation of mylonitic  
613 shear zones closely follows the geometry of the D1 fabric (Lawley et al., in press-b). The sharp  
614 contacts between the mylonitic shear zones and foliated wall rock suggests that the D1 fabric was  
615 not passively re-oriented during subsequent deformation, and instead suggests D1 fabric acted as  
616 a structural weakness and/or heterogeneity that localized strain during D2 (e.g., Fig. 6b).  
617 However for several shear zones (e.g., Saza shear zone), and for those gold prospects hosted by  
618 non-foliated Palaeoproterozoic intrusive phases, the importance of D1 deformation on the  
619 development of auriferous shear zones is unclear.

620 Re-Os dating has also identified hydrothermal events (ca. 1.95 and 1.94 Ga; Lawley et  
621 al., in press-a) that pre-date the timing of D1 reported as part of this study (ca. 1.92 Ga). These  
622 anomalously older sulphides were sampled from laminated quartz veins filling and deformed by  
623 D2 mylonitic shear zones at Kenge. The latter suggests that quartz veining may have in fact, at  
624 least locally, pre-dated the D1 fabric (Fig. 9d). Unfortunately, the uncertainty in the common-lead  
625 correction precludes a precise U-Pb titanite age for the onset of ductile deformation and does not  
626 allow us to evaluate this possibility further. Nevertheless, structural preparation has been shown  
627 to play an important role at goldfields where orogenic style gold mineralization is kinematically  
628 late (Groves et al., 2000). Pre-existing structures, such as brittle faults and/or lithologic contacts,  
629 are widely recognized as ‘stress risers’ that facilitate reactivation over the formation of new  
630 structures at the deposit scale (e.g., Dubé et al., 1989; Lin and Corfu, 2002).

631 The importance of structural reactivation is also suggested by the link between orogenic  
632 gold style mineralization, compressional to transpressional settings and “mis-oriented” high-angle  
633 reverse shear zones (e.g., Sibson, 1988). In these geologic settings and in a compressional stress  
634 regime, high-angle reverse faults are more likely to represent reactivation of pre-existing  
635 structures due to high fluid pressure and/or low frictional fault strength rather than newly  
636 developed structures (Lawley et al., in press-b). We suggest that D1 and, particularly the  
637 development of planes comprising rheologically weak chloritic folia, created zones of structural  
638 weakness that may have, at least locally, acted as pre-existing anisotropies that were potentially  
639 reactivated during later D2 brittle-ductile deformation. The abundance of chlorite slickensides on

640 D1 chloritic folia, which are unlikely to have formed during the dominantly ductile D1 event(s),  
641 supports this interpretation and also suggests that early ductile deformation may have been  
642 reactivated during late brittle faulting (D3).

643 Together the available field relationships and absolute ages of deformation fabrics record  
644 a broad progression from dominantly D1 ductile (ca. 1.92), to D2 brittle-ductile auriferous  
645 mylonitic shear zones and quartz veins (ca. 1.88 Ga) and ultimately to D3 cataclasites and  
646 discrete faults during later, but undated, and dominantly brittle deformation. A similar  
647 progression in deformation characteristics, i.e. from dominantly ductile to brittle deformation, has  
648 also been reported in goldfields associated with more modern orogenic settings and has been  
649 attributed to changing P-T conditions during orogenic uplift, denudation and cooling (e.g., Alps;  
650 Pettke et al., 1999).

651

## 652 **6.0 Conclusions**

653 Titanite U-Pb geochronology for a foliated Archaean granite, which represents the  
654 earliest identifiable deformation event in the Lupa Goldfield, suggests that the onset of  
655 deformation occurred during the Palaeoproterozoic at ca. 1.92 Ga. This age raises new questions  
656 regarding the timing of the D1 fabric and suggests either: 1) titanite dates are slightly younger  
657 than their true age due to lead-loss or an inaccurate common lead correction; and/or 2) that D1 is  
658 more complex than previously recognized and occurred diachronously during the emplacement of  
659 Palaeoproterozoic intrusions. New U-Pb titanite ages for the non-foliated Saza Granodiorite at  
660  $1930 \pm 3$  Ma, are only nominally younger than the U-Pb zircon age ( $1935 \pm 1$  Ma) for the same  
661 sample.

662 These U-Pb titanite ages, combined with previously reported U-Pb zircon and Re-Os  
663 sulphide geochronology, constrain the Palaeoproterozoic timing of magmatism (1.96–1.88 Ga),  
664 hydrothermal activity (1.95–1.88 Ga) and deformation/metamorphism ( $\geq 1.92$ –1.88 Ga) in the  
665 Lupa Goldfield. Together the available geochronologic data demonstrate a progression from  
666 dominantly ductile (D1) to brittle-ductile deformation (D2) over  $\geq 40$  Myr, which occurred  
667 diachronously and intermittent with felsic-mafic plutonism during a Palaeoproterozoic orogenic  
668 cycle at the Tanzanian cratonic margin. However, the inferred ca. 1.88 Ga gold event in the Lupa  
669 Goldfield simply represents the earliest episode of orogenic gold deposit formation in western  
670 Tanzania and are superseded by younger orogenic gold style mineralization events related to  
671 temporally discrete orogenic episodes at cratonic margins bordering the Ubendian Belt. The  
672 Tanzanian cratonic margin is therefore highly prospective for orogenic style gold mineralization,  
673 which may have developed during multiple orogenic events.

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682

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1097  
1098 **Figure Captions**

1099  
1100 Figure 1  
1101 Regional geologic map of Tanzania (modified after Pinna et al., 2004; Kabete et al., 2012a).  
1102 Superterrane abbreviations include (Kabete et al., 2012a, b): East Lake Victoria Superterrane,  
1103 ELVST; Mwanza-Lake Eyasi Superterrane, MLEST; Lake Nyanza Superterrane, LNST;  
1104 Moyowosi-Manyoni Superterrane, MMST; Dodoma Basement Superterrane, DBST; Dodoma  
1105 Schist Superterrane; DSST; Eastern Ubendian-Mtera Superterrane, EUMST; Kalenge-Burigi,  
1106 KBST, Mbulu-Masai Superterrane, MAST; Kilindi-Handeni Superterrane, KHST; Usagara-  
1107 Ukaguru Superterrane, UKST; Uluguru-Pare Superterrane; UPST.

1108  
1109 Figure 2

1110 Regional geologic map showing the lithotectonic terranes comprising the Palaeoproterozoic  
1111 Ubendian Belt (modified after Daly, 1988).  
1112  
1113 Figure 3  
1114 Schematic geology map of the Lupa Goldfield (modified after Kimambo, 1984).  
1115  
1116 Figure 4  
1117 Local geology map showing the location of geochronology samples (CL109 and CL0975),  
1118 mineral systems and artisanal mines. Previously reported U-Pb zircon ages are from Lawley et al.  
1119 (2013). Eastings and northings are reported as UTM coordinates (WGS84, Zone 36S).  
1120  
1121 Figure 5  
1122 Diagram summarizing previously reported Palaeoproterozoic U-Pb zircon and Re-Os sulphide  
1123 ages and age ranges, which together with new U-Pb titanite ages, constrain the timing of  
1124 deformation, magmatism and hydrothermal activity in the Lupa Goldfield. Note the broadly  
1125 overlapping magmatic and hydrothermal history in the Lupa Goldfield (U-Pb and Re-Os age  
1126 ranges plotted as horizontal bars; Lawley et al., 2013; Lawley et al., in press-a), whereas  
1127 individual U-Pb zircon ages (U-Pb zircon ages, without analytical uncertainty, are plotted as  
1128 vertical bars) show good agreement with three of the temporally distinct hydrothermal events  
1129 (i.e., ca. 1.95, 1.94, and 1.88 Ga) identified by Re-Os geochronology (weighted average Re-Os  
1130 ages, without analytical uncertainty, for interpreted hydrothermal events plotted as vertical bars).  
1131 The ca. 1.88 Ga hydrothermal event is particularly important as it was recorded at all five of the  
1132 studied gold prospects and also corresponds to the development of D2 auriferous mylonitic shear  
1133 zones. Note Archaean granites and anomalously younger Mesoproterozoic Re-Os ages are not  
1134 shown.  
1135  
1136 Figure 6  
1137 (a) Photo of artisanal working along the Kenge shear zone in section view and looking  
1138 approximately northwest. Fault-fill veins and mylonitic shear zones cut foliated Archaean granite  
1139 at Kenge; (b) closer photo of fault fill vein and mylonitic shear zone contact at Kenge looking  
1140 northwest. Note the sharp contact between the mylonite and Archaean granite wall rock (ruler is  
1141 15 cm in length); (c) core photo showing complex vein textures that are typical of the fault-fill  
1142 vein type. Note the laminated vein appearance due to slivers of mylonitized wall rock intercalated

1143 with the fault fill vein; (d–e) core photos of the mineralized zone at Porcupine. Note auriferous  
1144 quartz veins cutting hydrothermally altered, but non-foliated, Ilunga Syenogranite.

1145

1146 Figure 7

1147 (a) Field photo of foliated Archaean granite sample (CL109); (b) field photo of non-foliated Saza  
1148 Granodiorite. The Saza Granodiorite is cut by auriferous mylonitic shear zones and quartz veins,  
1149 but in turns cuts early foliated Archaean granite (not shown). Note overprinting chloritic  
1150 alteration and en echelon tension gashes; (c) plane polarized light photomicrograph of titanite  
1151 associated with greenschist facies metamorphic mineral assemblage and tectonic fabric from  
1152 CL109; (d) plane polarized light photomicrograph of euhedral, and potentially magmatic, titanite  
1153 crystal from CL0975; (e) stereoscopic photomicrograph showing examples of translucent and fine  
1154 grained titanite fractions analyzed from CL109; (f) stereoscopic photomicrograph showing  
1155 examples of fine grained and brown-translucent titanite fractions analyzed from CL0975.

1156

1157 Figure 8

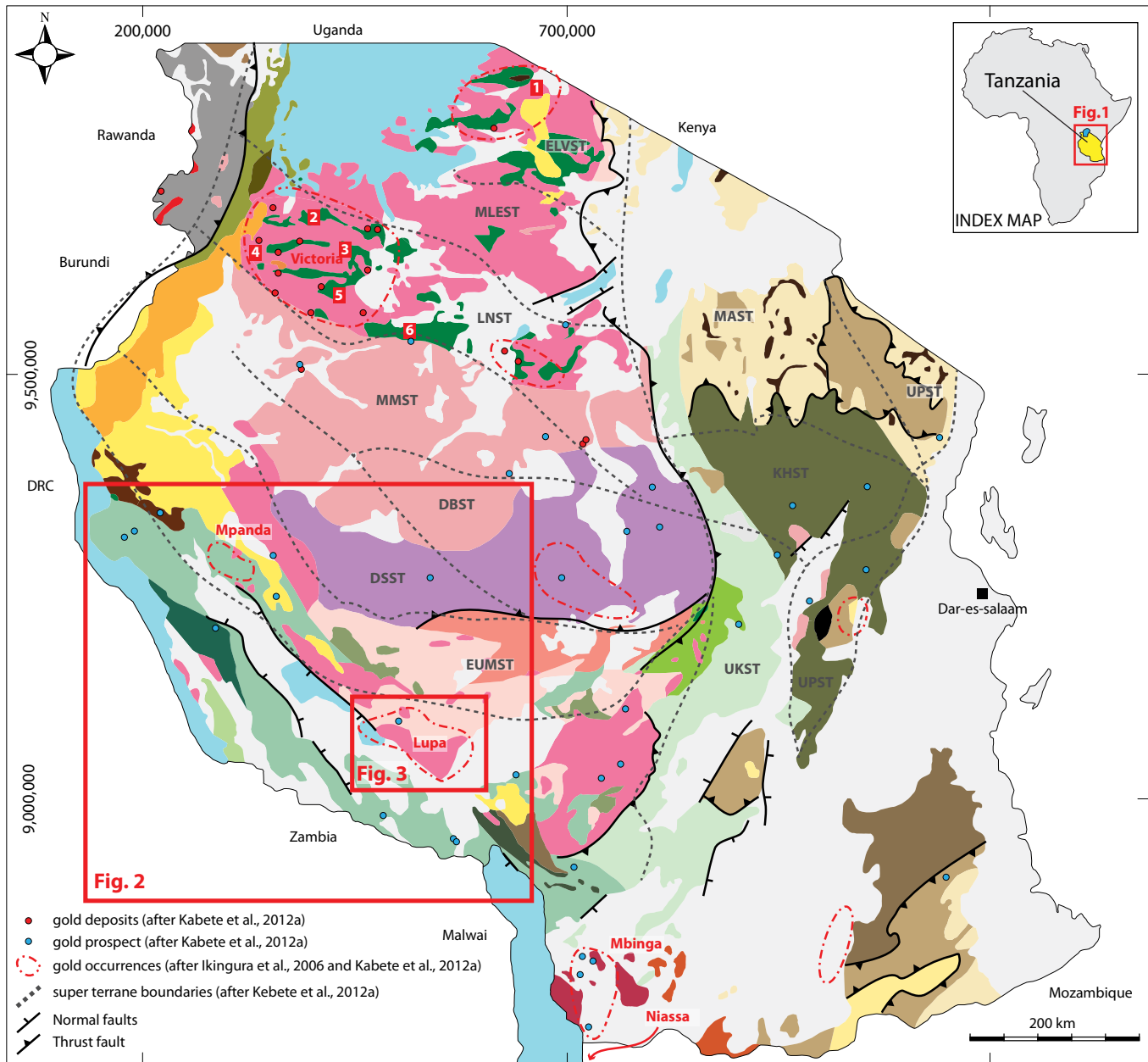
1158 (a–b) Uncorrected (grey ellipses) and corrected [corrected using the Stacey and Kramers (1975)  
1159 Pb evolution model at 1.9 Ga; S-K; purple ellipses] U-Pb titanite data (purple ellipses) from  
1160 CL109 and CL0975 on a 2D Terra-Wasserburg plot. Previously reported LA-ICP-MS and ID-  
1161 TIMS U-Pb zircon ages for CL109 and CL0975, respectively, are also shown for comparison  
1162 (blue ellipses). The inset figures demonstrate the sensitivity of  $^{207}\text{Pb}/^{206}\text{Pb}$  ages to the assumed  
1163 isotopic composition of common Pb by varying the Stacey and Kramers (1975) Pb evolution  
1164 model from 2.5–1.5 Ga. A York Model-1 regression solution of S-K corrected data for CL109  
1165 yields an upper intercept U-Pb titanite age of  $1921 \pm 7$  Ma (MSWD = 0.7; n = 7; regression  
1166 excludes S3 and S4). A York Model-1 regression solution of S-K corrected analyses for CL0975  
1167 yields an upper intercept age of  $1931 \pm 3$  Ma (MSWD = 1.1; n = 5). The zoomed window in Fig.  
1168 6b shows how three of the S-K corrected titanite analyses (purple ellipses) from CL0975 overlap  
1169 with Concordia and yield a weighted a weighted average  $^{207}\text{Pb}/^{206}\text{Pb}$  age of  $1930 \pm 3$  Ma (MSWD  
1170 = 0.6; n = 3).

1171

1172 Figure 9

1173 (a–f) Schematic block diagrams showing the geological evolution of the Lupa Goldfield from the  
1174 Archaean to the Palaeoproterozoic (based on field relationships and U-Pb and Re-Os ages). New  
1175 U-Pb titanite ages from an Archaean granite suggest that the D1 fabric developed during the  
1176 Palaeoproterozoic (potentially at ca. 1.92 Ga; Fig. 7e); however the precise timing of D1 remains

1177 unclear since Palaeoproterozoic granites (e.g., Ilunga Syenogranite and Saza Granodiorite) are  
1178 largely non-foliated, which suggests that the D1 fabric locally pre-dated 1.96 Ga (Fig. 7b) and  
1179 that younger titanite dates may record Pb-loss. Alternatively, D1 may possess a more complex  
1180 history than previously recognized and may have developed diachronously during the  
1181 emplacement of Palaeoproterozoic granites (Fig. 7e). Nevertheless, the New U-Pb titanite ages  
1182 suggest that ductile deformation predated auriferous mylonites by  $\geq 40$  Myr. Note Lupa  
1183 Goldfield's younger geological history is not shown, but anomalous Mesoproterozoic Re-Os ages  
1184 and lower intercept U-Pb ages provide evidence for a tectono-thermal history spanning multiple  
1185 orogenic cycles (Lawley et al., 2013; in press-a). Mineralized structures are also locally offset by  
1186 cataclasites (not shown) and point to a relatively late and dominantly brittle deformation event  
1187 (D3) of unknown age.  
1188  
1189



- Mesoproterozoic**
- Tin granites (~985-950 Ma)
  - Mafic rocks
  - Meta-sedimentary rocks (ca. 1.37 Ga; previously Bukoban Group)
  - Detrital meta-sedimentary rocks
  - Shale, sandstone, quartzite (ca. 1.37 Ga)
  - Orthogneiss (1.19-0.95 Ga)
  - Migmatite and orthogneiss
- Palaeoproterozoic (Ubendian-USagaran Belt)**
- Orthogneissic granulite, meta-gabbro, amphibolite, marble, eclogitic lenses (Isimani Group)
  - Late orogenic (>1.95 Ga granite, granodiorite)
  - Meta-gabbro, anorthosite, ultramafic rocks, Neoproterozoic syenite and granite
  - Felsic igneous suites
  - Eclogite
  - Gneiss, granulite, migmatite, amphibolite, quartzite
  - Igneous-sedimentary rocks (ca. 2.61 Ga granulites)
- Archaean**
- Orthogneiss (TTG), migmatite, granite (Isangan Group, ~3.0-2.85 Ga)
  - Orthogneiss (TTG), migmatite, granite (Dodoman Group ~2.90-2.50 Ga)
  - Granitoid, migmatite, mafic/ultramafic rock
  - Meta-basalt, BIF and phyllite
  - Late- to post-orogenic granite/granodiorite

- Gold deposits discussed in the text**
- 1 North Mara
  - 2 Geita
  - 3 Bulyanhulu
  - 4 Tulawaka
  - 5 Buzwagi
  - 6 Golden Pride
- Post-Precambrian
- Neoproterozoic**
- Sedimentary formations
  - Volcanic-sedimentary formations
  - Marble/quartzite
  - Paragneiss (marble, quartzite, schist), orthogneiss (migmatite, amphibolite), granitoid
  - Anorthosite complexes (ca. 640 Ma)
  - Mafic-felsic granulite, gneiss, migmatite (ca. 640 Ma)
  - Neoproterozoic to Neoproterozoic protoliths re-worked during the tectono-thermal events

Figure 1

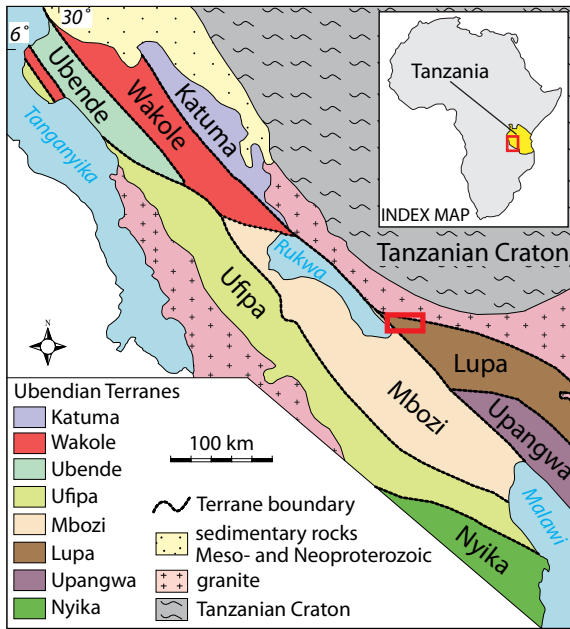


Figure 2

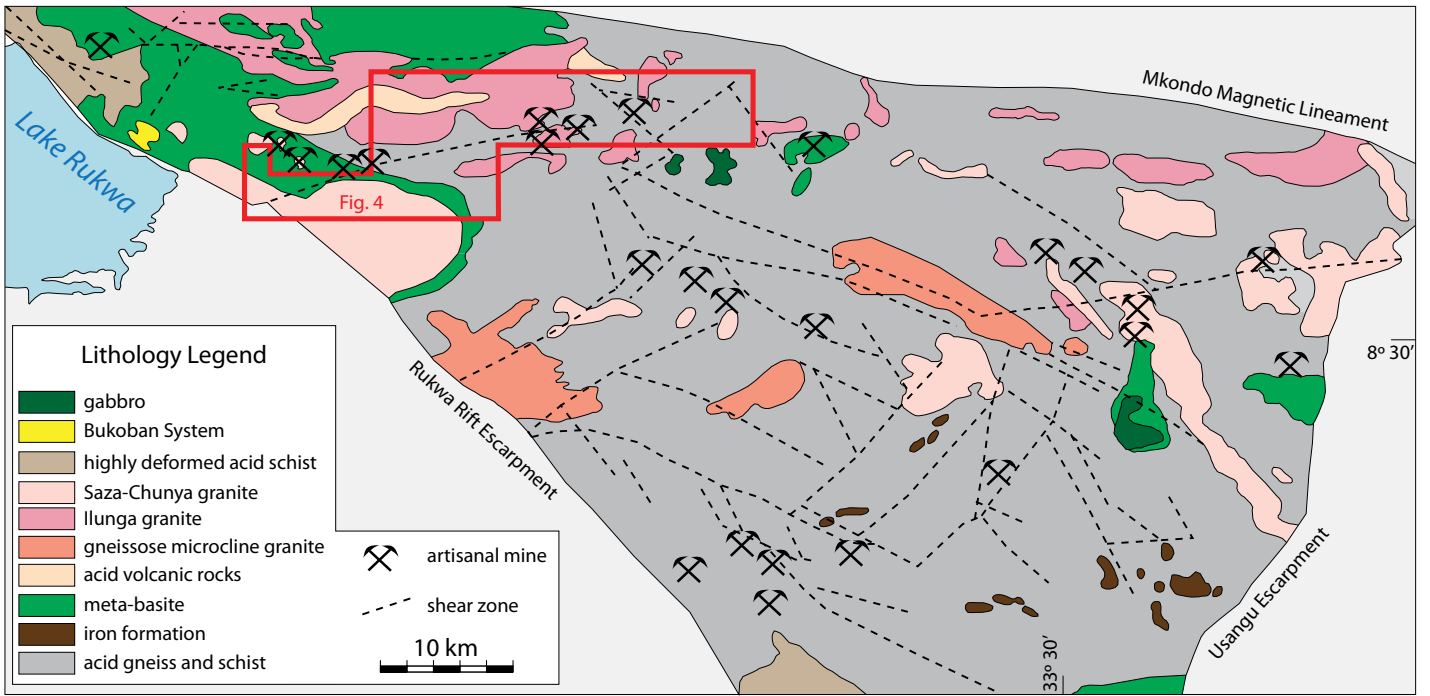


Figure 3

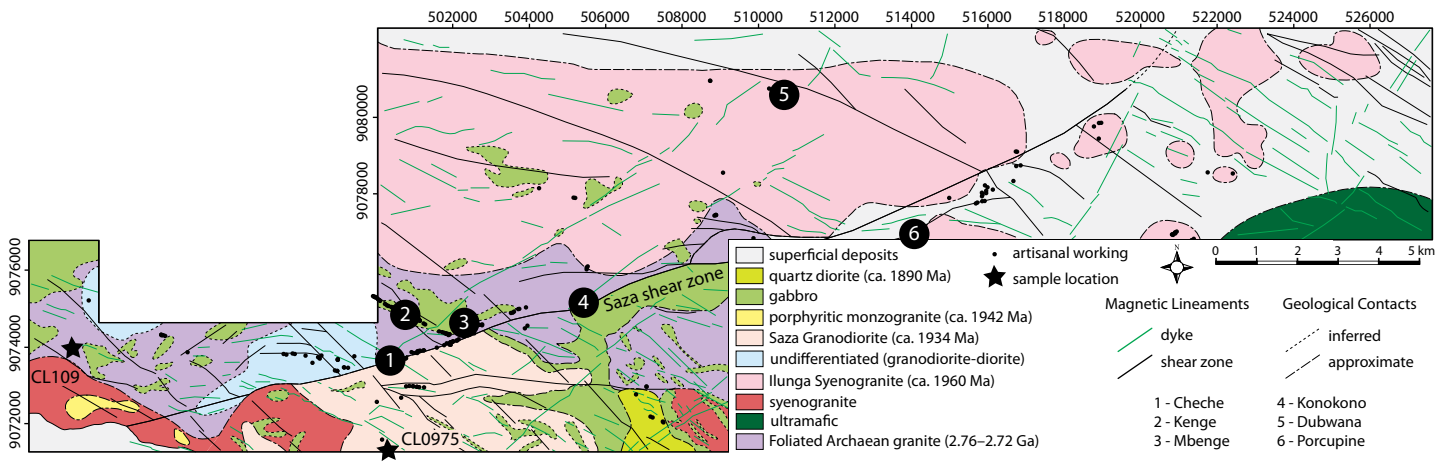


Figure 4



# Palaeoproterozoic Geochronology Summary of the Lupa Goldfield

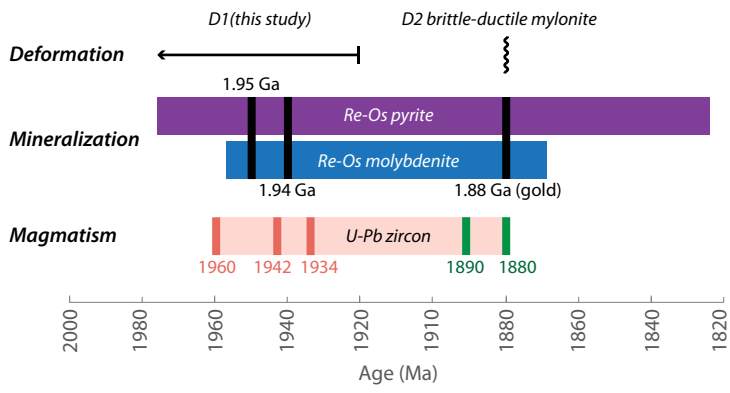


Figure 5

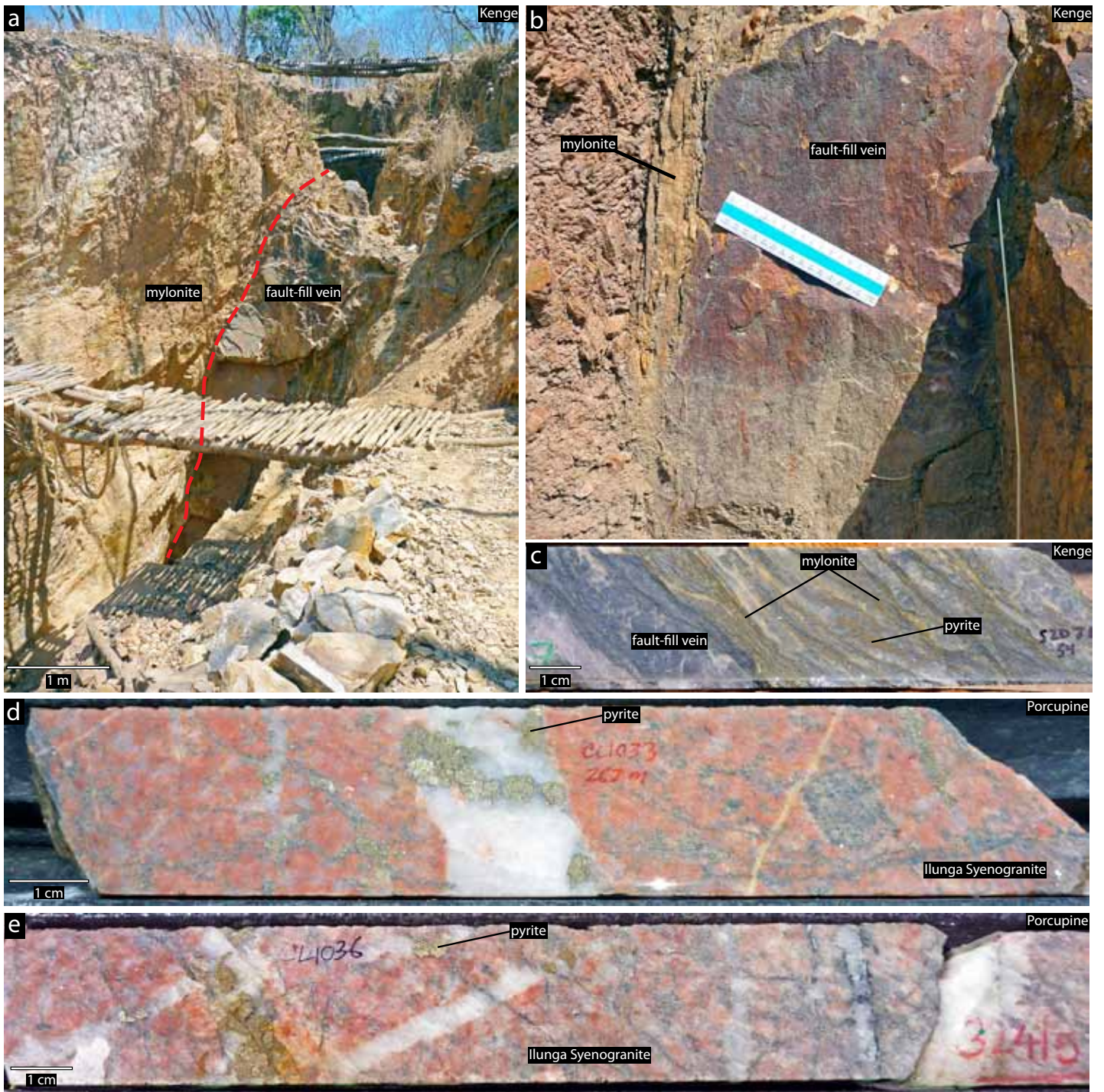


Figure 6



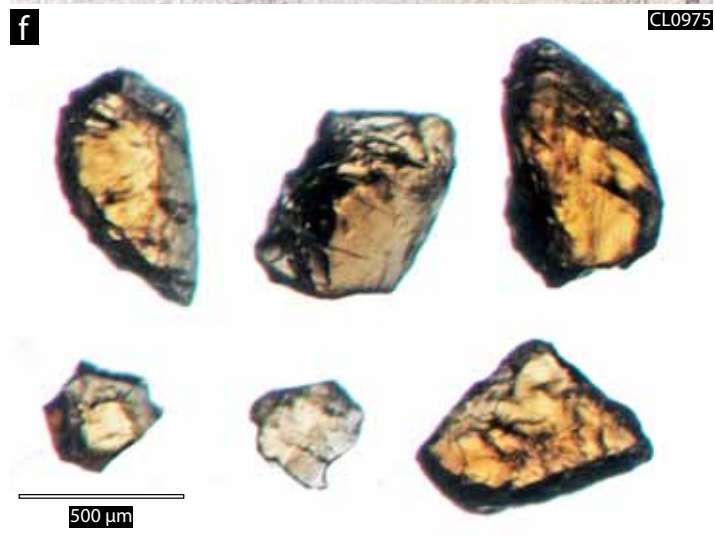
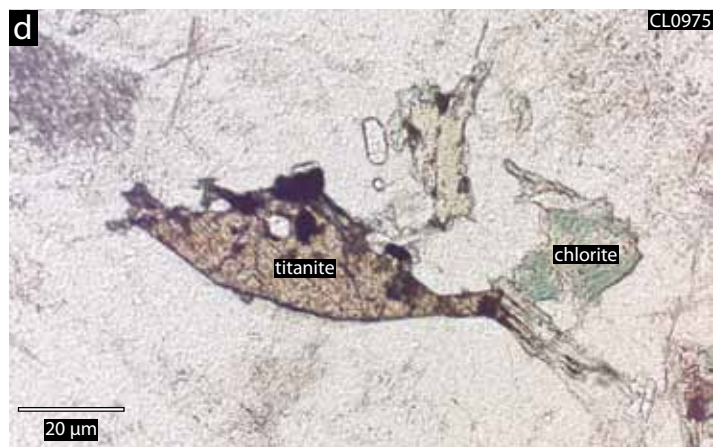
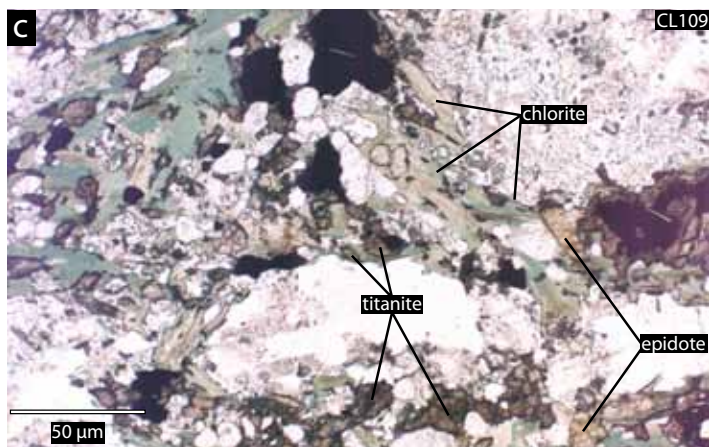
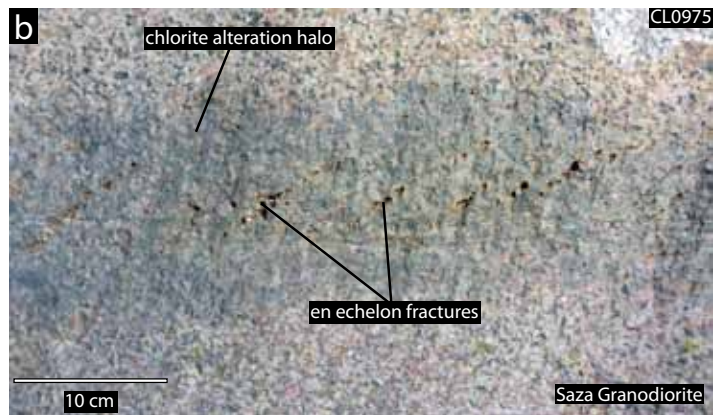
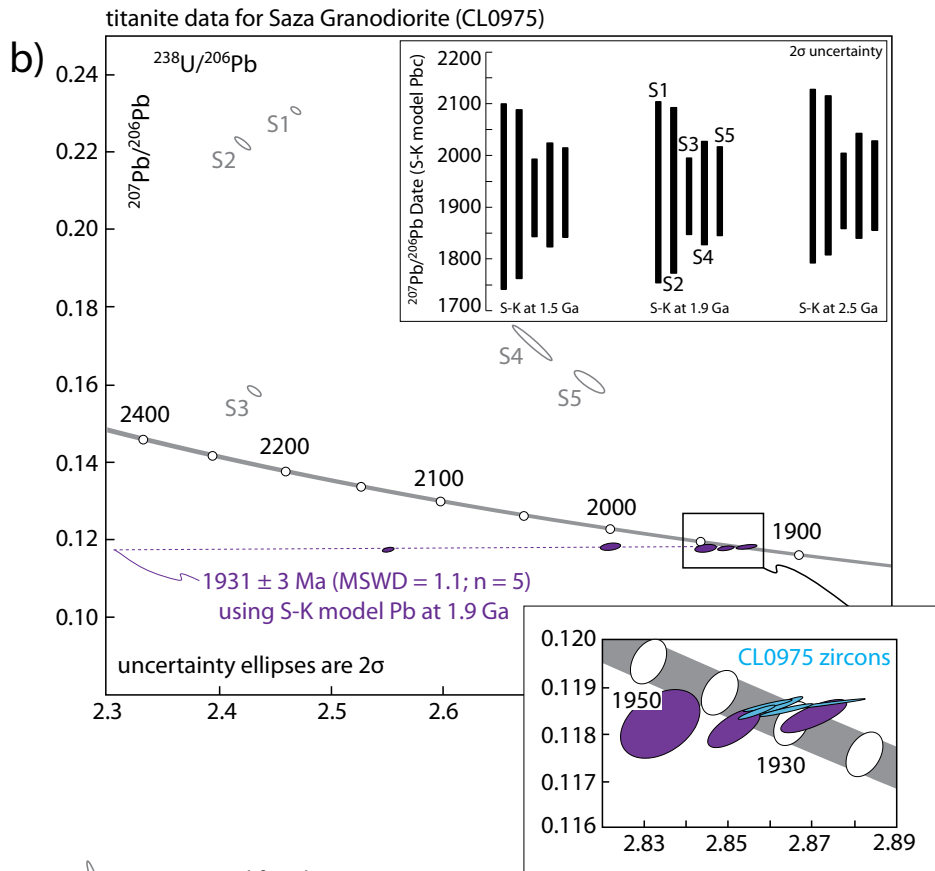
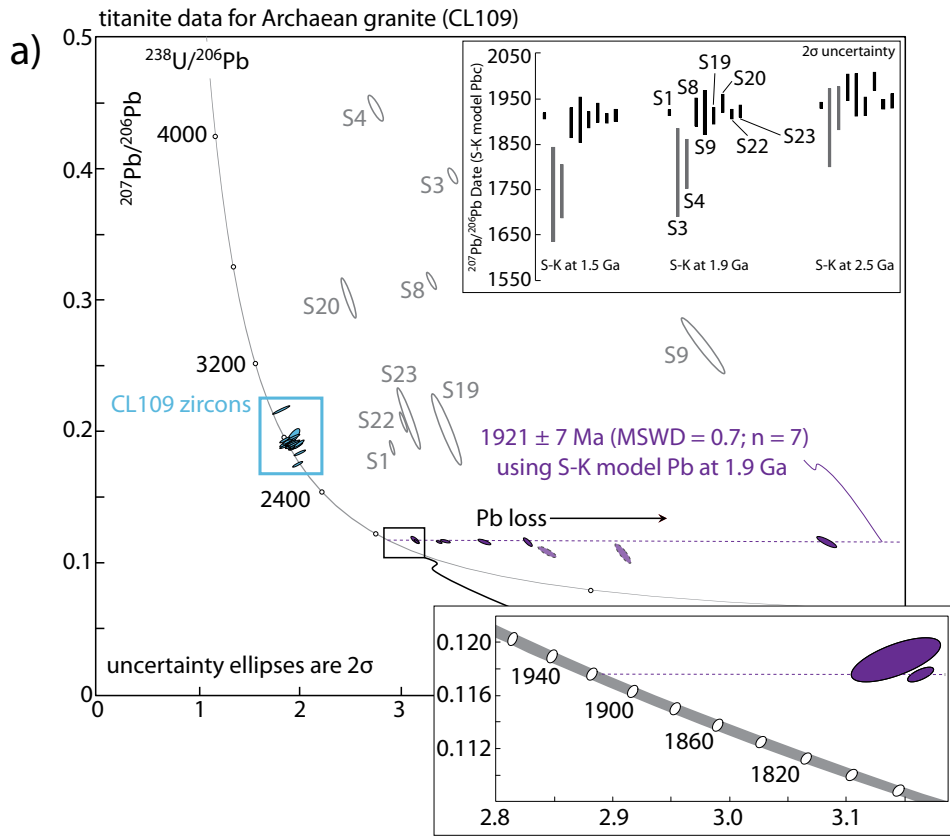
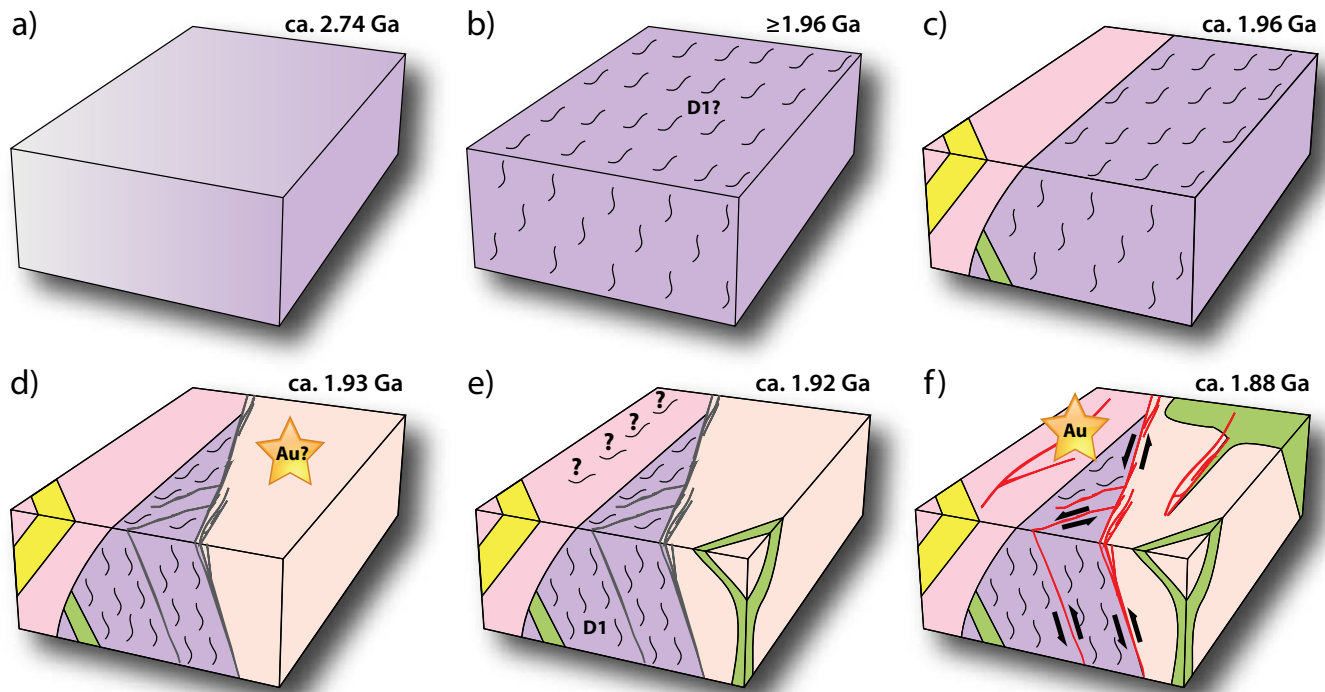


Figure 7



- not corrected for Pb<sub>initial</sub>
- corrected for Pb<sub>initial</sub> using S-K model Pb at 1.9 Ga
- corrected for Pb<sub>initial</sub> using S-K model Pb at 1.9 Ga, but not used in age regression
- zircon analyses (previously reported in Lawley et al., 2013)

Figure 8



- gabbro-diorite
  - Saza Granodiorite
  - granodiorite-tonalite
  - Ilunga Syenogranite
  - Archaean granites (2.76–2.72 Ga)
- 
- D1 - ductile deformation
  - pre-D2 - sulphides
  - D2 - auriferous brittle-ductile mylonites

*Block models are looking approximately northeast and are not to scale*

Figure 9

**Table 1. U-Th-Pb isotopic data**

| Sample        | Compositional Parameters |  |                             |            |             | Radiogenic Isotope Ratios               |   |   |         |  |         |  | Isotopic Ages |                |   | Sample (Radiogenic + Initial Pb) Isotope Ratios |  |       |  |       |   |         |   |          | Sample (Radiogenic + Initial Pb) Isotope Ratios |                        |                        |  |         |   |           |                        |  |          |   |           |                        |           |          |
|---------------|--------------------------|--|-----------------------------|------------|-------------|---|---|---|---------|--|---------|--|---------------|----------------|---|---|--|-------|--|-------|---|---------|---|----------|---|------------------------|------------------------|--|---------|---|-----------|------------------------|--|----------|---|-----------|------------------------|-----------|----------|
|               | Th<br>U                  | <sup>206</sup> Pb*<br>x10 <sup>-13</sup> mol | mol %<br><sup>206</sup> Pb* | Pb*<br>Pbc | Pbc<br>(pg) | <sup>206</sup> Pb/<br><sup>204</sup> Pb | <sup>208</sup> Pb/<br><sup>206</sup> Pb | <sup>207</sup> Pb/<br><sup>206</sup> Pb | % err   | <sup>207</sup> Pb/<br><sup>235</sup> U | % err   | <sup>206</sup> Pb/<br><sup>238</sup> U | % err         | corr.<br>coef. | <sup>207</sup> Pb/<br><sup>206</sup> Pb ± | <sup>207</sup> Pb/<br><sup>235</sup> U ±        | <sup>206</sup> Pb/<br><sup>238</sup> U ± | % err | <sup>238</sup> U/<br><sup>206</sup> Pb | % err | <sup>207</sup> Pb/<br><sup>206</sup> Pb | % err   | <sup>204</sup> Pb/<br><sup>206</sup> Pb | % err    | corr. coef.<br>8/6-7/6                          | corr. coef.<br>8/6-4/6 | corr. coef.<br>7/6-4/6 | <sup>238</sup> U/<br><sup>204</sup> Pb | % err   | <sup>206</sup> Pb/<br><sup>204</sup> Pb | % err     | corr. coef.<br>8/4-6/4 | <sup>235</sup> U/<br><sup>204</sup> Pb | % err    | <sup>207</sup> Pb/<br><sup>204</sup> Pb | % err     | corr. coef.<br>5/4-7/4 |           |          |
| (a)           | (b)                      | (c)  | (c)                         | (c)        | (d)         | (e)                                     | (e)                                     | (f)                                     | (e)     | (f)                                    | (e)     | (f)                                    |               | (g)            | (f)                                       | (g)   | (f)                                      | (g)   | (f)                                    | (h)   | (f)                                     | (h)     | (f)                                     | (h)      | (f)   |                        |                        |  | (h)     | (f)                                     | (h)       | (f)                    |  | (h)      | (f)                                     | (h)       | (f)                    |           |          |
| <b>CL109</b>  |                          |  |                             |            |             |   |   |   |         |  |         |  |               |                |   |   |  |       |  |       |   |         |   |          |   |                        |                        |  |         |   |           |                        |  |          |   |           |                        |           |          |
| s1            | 8.685                    | 1.8765                                       | 91%                         | 9          | 17          | 172                                     | 2.63713                                 | 0.11757                                 | 0.37158 | 5.12124                                | 0.50024 | 0.31607                                | 0.28828       | 0.677          | 1919                                      | 7   | 1840                                     | 4     | 1770                                   | 4     | 2.90842                                 | 0.63551 | 0.18816                                 | 2.11335  | 0.00525   | 5.54664                | -0.94304               | -0.93364                               | 0.99647 | 553.460993                              | 6.144194  | 190.296410             | 5.546638                               | 0.999313 | 4.015883                                | 6.144194  | 35.807065              | 3.445327  | 0.997181 |
| s3            | 5.561                    | 0.9342                                       | 66%                         | 1          | 42          | 46                                      | 2.05308                                 | 0.10859                                 | 5.43178 | 2.87496                                | 5.92031 | 0.19210                                | 1.16198       | 0.501          | 1775                                      | 99  | 1375                                     | 45    | 1133                                   | 12    | 3.51867                                 | 1.01569 | 0.39489                                 | 1.10399  | 0.02109   | 1.74569                | -0.81246               | -0.67796                               | 0.72454 | 166.806967                              | 2.546213  | 47.406231              | 1.745688                               | 0.956041 | 1.210342                                | 2.546213  | 18.720344              | 1.213880  | 0.628353 |
| s4            | 8.675                    | 0.4079                                       | 60%                         | 1          | 25          | 39                                      | 3.00279                                 | 0.10955                                 | 3.04999 | 3.38886                                | 3.34017 | 0.22445                                | 1.52124       | 0.410          | 1791                                      | 56  | 1502                                     | 26    | 1305                                   | 18    | 2.75546                                 | 2.09995 | 0.44626                                 | 1.69109  | 0.02484   | 2.19780                | -0.87946               | -0.86615                               | 0.98221 | 110.950266                              | 4.151525  | 40.265570              | 2.197802                               | 0.967517 | 0.805049                                | 4.151525  | 17.968907              | 0.623712  | 0.793318 |
| s8            | 10.142                   | 1.0647                                       | 76%                         | 3          | 30          | 65                                      | 3.56932                                 | 0.11734                                 | 1.77519 | 3.78833                                | 1.97265 | 0.23427                                | 0.67383       | 0.449          | 1916                                      | 32  | 1590                                     | 16    | 1357                                   | 8     | 3.30306                                 | 0.96507 | 0.31520                                 | 1.52549  | 0.01472   | 2.45117                | -0.91463               | -0.87908                               | 0.95269 | 224.332673                              | 3.331458  | 67.916560              | 2.451167                               | 0.990420 | 1.627746                                | 3.331458  | 21.407120              | 1.100327  | 0.867201 |
| s9            | 6.544                    | 0.3543                                       | 78%                         | 3          | 9           | 75                                      | 2.89817                                 | 0.11743                                 | 2.74106 | 2.23765                                | 3.14245 | 0.13826                                | 1.03516       | 0.528          | 1917                                      | 49  | 1193                                     | 22    | 835                                    | 8     | 5.99911                                 | 2.93770 | 0.26660                                 | 6.28610  | 0.01110   | 11.09868               | -0.97046               | -0.96419                               | 0.99294 | 540.393150                              | 13.952954 | 90.078859              | 11.098684                              | 0.998440 | 3.921064                                | 13.952954 | 24.014967              | 4.913830  | 0.983361 |
| s19           | 6.808                    | 0.3295                                       | 85%                         | 4          | 5           | 107                                     | 2.27226                                 | 0.11711                                 | 1.02463 | 4.21270                                | 1.77968 | 0.26101                                | 1.24656       | 0.827          | 1912                                      | 18  | 1676                                     | 15    | 1495                                   | 17    | 3.45787                                 | 3.49886 | 0.20239                                 | 10.74043 | 0.00634   | 25.09789               | -0.95339               | -0.95246                               | 0.99918 | 545.070753                              | 28.450395 | 157.631917             | 25.097887                              | 0.999298 | 3.955004                                | 28.450395 | 31.902666              | 14.372904 | 0.998682 |
| s20           | 4.450                    | 0.4556                                       | 76%                         | 2          | 13          | 65                                      | 1.35284                                 | 0.11869                                 | 1.16054 | 5.20495                                | 1.56029 | 0.31821                                | 0.99986       | 0.669          | 1936                                      | 21  | 1853                                     | 13    | 1781                                   | 16    | 2.48128                                 | 2.24673 | 0.30248                                 | 3.84972  | 0.01370   | 6.22461                | -0.93974               | -0.93702                               | 0.99762 | 181.144783                              | 8.366722  | 73.004701              | 6.224608                               | 0.995592 | 1.314377                                | 8.366722  | 22.082102              | 2.398797  | 0.987342 |
| s22           | 8.289                    | 1.1981                                       | 88%                         | 6          | 15          | 131                                     | 2.60140                                 | 0.11730                                 | 0.61527 | 4.78081                                | 0.77432 | 0.29572                                | 0.39822       | 0.616          | 1915                                      | 11  | 1782                                     | 7     | 1670                                   | 6     | 3.02946                                 | 0.95581 | 0.20839                                 | 2.82371  | 0.00678   | 6.36678                | -0.95466               | -0.94810                               | 0.99587 | 446.960561                              | 7.279330  | 147.537871             | 6.366777                               | 0.999128 | 3.243122                                | 7.279330  | 30.745556              | 3.563952  | 0.995452 |
| s23           | 11.093                   | 0.3767                                       | 84%                         | 6          | 6           | 105                                     | 3.51968                                 | 0.11774                                 | 0.76732 | 4.71338                                | 1.43760 | 0.29046                                | 1.04548       | 0.855          | 1922                                      | 14  | 1770                                     | 12    | 1644                                   | 15    | 3.07716                                 | 3.01549 | 0.21060                                 | 8.82788  | 0.00691   | 19.70896               | -0.95833               | -0.95740                               | 0.99936 | 445.127104                              | 22.612749 | 144.655189             | 19.708961                              | 0.999258 | 3.229818                                | 22.612749 | 30.464599              | 10.891381 | 0.998668 |
| <b>CL0975</b> |                          |  |                             |            |             |   |   |   |         |  |         |  |               |                |   |   |  |       |  |       |   |         |   |          |   |                        |                        |  |         |   |           |                        |  |          |   |           |                        |           |          |
| s1            | 7.101                    | 12.6005                                      | 87%                         | 5          | 170         | 118                                     | 2.04070                                 | 0.11825                                 | 0.50689 | 5.75101                                | 0.54341 | 0.35288                                | 0.27228       | 0.380          | 1929                                      | 9   | 1939                                     | 5     | 1948                                   | 5     | 2.46803                                 | 0.14137 | 0.23103                                 | 0.24519  | 0.00840   | 0.49173                | -0.73746               | -0.36875                               | 0.75343 | 293.758217                              | 0.559511  | 119.025226             | 0.491732                               | 0.972030 | 2.131494                                | 0.559511  | 27.498842              | 0.346756  | 0.778471 |
| s2            | 7.024                    | 5.5666                                       | 88%                         | 5          | 70          | 126                                     | 1.98811                                 | 0.11854                                 | 0.46212 | 5.94430                                | 0.50466 | 0.36387                                | 0.25145       | 0.411          | 1934                                      | 8   | 1968                                     | 4     | 2000                                   | 4     | 2.42092                                 | 0.21977 | 0.22257                                 | 0.55365  | 0.00775   | 1.16458                | -0.91991               | -0.81574                               | 0.95534 | 312.275906                              | 1.349856  | 128.990334             | 1.164580                               | 0.995556 | 2.265857                                | 1.349856  | 28.709537              | 0.656374  | 0.944823 |
| s3            | 6.545                    | 8.6624                                       | 95%                         | 13         | 40          | 315                                     | 1.77382                                 | 0.11777                                 | 0.21445 | 6.36479                                | 0.29163 | 0.39214                                | 0.14130       | 0.716          | 1922                                      | 4   | 2027                                     | 3     | 2133                                   | 3     | 2.43100                                 | 0.17231 | 0.15861                                 | 0.59326  | 0.00304   | 2.25488                | -0.83437               | -0.74161                               | 0.97755 | 799.518730                              | 2.385468  | 328.884369             | 2.254876                               | 0.998825 | 5.801265                                | 2.385468  | 52.164448              | 1.679597  | 0.993264 |
| s4            | 7.634                    | 2.3294                                       | 93%                         | 10         | 15          | 226                                     | 2.20028                                 | 0.11816                                 | 0.28857 | 5.71131                                | 0.37949 | 0.35072                                | 0.17821       | 0.684          | 1928                                      | 5   | 1933                                     | 3     | 1938                                   | 3     | 2.67886                                 | 0.50225 | 0.17101                                 | 1.96415  | 0.00394   | 6.25903                | -0.98383               | -0.97258                               | 0.99661 | 680.509624                              | 6.748516  | 254.029522             | 6.259025                               | 0.999850 | 4.937741                                | 6.748516  | 43.441625              | 4.304555  | 0.998652 |
| s5            | 6.950                    | 3.5224                                       | 94%                         | 13         | 18          | 285                                     | 2.01251                                 | 0.11841                                 | 0.24778 | 5.68564                                | 0.37252 | 0.34839                                | 0.22256       | 0.765          | 1932                                      | 4   | 1929                                     | 3     | 1927                                   | 4     | 2.73024                                 | 0.38223 | 0.16105                                 | 1.41795  | 0.00318   | 5.28003                | -0.87134               | -0.85159                               | 0.99345 | 859.439228                              | 5.609118  | 314.785585             | 5.280033                               | 0.999362 | 6.236045                                | 5.609118  | 50.696522              | 3.874763  | 0.997856 |

(a) z1, z2 etc. are labels for fractions composed of single zircon grains or fragments; all fractions annealed and chemically abraded after Mattinson (2005).

(b) Model Th/U ratio calculated from radiogenic <sup>208</sup>Pb/<sup>206</sup>Pb ratio and <sup>207</sup>Pb/<sup>235</sup>U age.

(c) Pb\* and Pbc represent radiogenic and common Pb, respectively; mol % <sup>206</sup>Pb\* with respect to radiogenic, blank and initial common Pb.

(d) Measured ratio corrected for spike and fractionation only.

Daly analyses, based on analysis of NBS-981 and NBS-982.

(e) Corrected for fractionation, spike, and common Pb; up to 1 pg of common Pb was assumed to be procedural blank: <sup>206</sup>Pb/<sup>204</sup>Pb = 18.60 ± 0.80%; <sup>207</sup>Pb/<sup>204</sup>Pb = 15.69 ± 0.32%;

<sup>208</sup>Pb/<sup>204</sup>Pb = 38.51 ± 0.74% (all uncertainties 1-sigma). Excess over blank was assigned to initial common Pb.

(f) Errors are 2-sigma, propagated using the algorithms of Schmitz and Schoene (2007) and Crowley et al. (2007).

(g) Calculations are based on the decay constants of Jaffey et al. (1971). <sup>206</sup>Pb/<sup>238</sup>U and <sup>207</sup>Pb/<sup>206</sup>Pb ages corrected for initial disequilibrium in <sup>230</sup>Th/<sup>238</sup>U using Th/U [magma] = 3.

(h) Corrected for fractionation, spike, and blank Pb only.

**Table 3. Summary of geologic characteristics at Tanzanian gold deposits and goldfields**

| Goldfield/R<br>egion                          | Significant<br>Deposits   | Host Rock  | Host Rock<br>Age (Ga) | Ore Controls   | Hydrothermal<br>Alteration Mineral<br>Assemblage | Metamorphic<br>facies      | Metals            | Sulphide<br>Assemblage   | Gold (Ga)               | Method  | Source  |
|---|---------------------------|--|-----------------------|--|--|----------------------------|-------------------|--|-------------------------|---|---|
| <b>Goldfields within the Tanzanian Craton</b> |                           |  |                       |  |  |                            |                   |  |                         |   |   |
| Lake Victoria                                 | North Mara                | sedimentary and volcanic rocks, granodiorite, tonalite, porphyritic andesite/dacite, gabbro                    | 2.76–2.65             | shear and quartz vein hosted; lithologic contact control           | Ser ± Chl ± Carb ± Sil ± Pot ± Sod               | greenschist                | Au ± Cu           | Py ± Cpy ± Po  | 2.69–2.64               | constrained by U-Pb zircon dating of magmatism and/or deformation | Manya et al., 2006; Kabete, 2008; Kazimoto, 2008; Mtoro et al., 2009; Ikingura et al., 2009 |
|   | Geita                     | BIF, felsic-mafic volcanic and sedimentary rocks   | 2.84–2.64             | BIF, fault and quartz vein control                                 | Carb ± Sil ± ?                                   | amphibolite to greenschist | Au                | Py ± Po ± Cpy ± Asp  | ≤2.644                  | U-Pb zircon dating of pre-gold lamprophyre dike                   | Walraven and Borg, 1994; Borg and Krough, 1999; Kabete et al., 2012a                        |
|   | Bulyanhulu                | sedimentary (including graphitic argillite), and felsic-mafic volcanic rocks                                   | 2.84–2.64             | shear (breccia) and quartz vein hosted; lithologic contact control | Ser ± Chl ± Carb ± Sil                           | greenschist                | Au ± Cu ± Pb      | Py ± Po ± Asp ± Cpy ± Gal  | 2.69–2.63               | constrained by U-Pb zircon dating of magmatism and/or deformation | Chamberlain, 2003; Kabete et al., 2012a   |
|   | Tulawaka                  | mafic-felsic volcanic and sedimentary rocks, granite, aplite dikes   | 2.84–2.64             | shear and quartz vein hosted; lithologic contact control           | ?  | amphibolite to greenschist | Au                | ?  | 2.69–2.63               | constrained by U-Pb zircon dating of magmatism and/or deformation | Cloutier et al., 2005; Kabete et al., 2012a   |
|   | Buzwagi                   | granite and lesser mafic volcanic rocks  | 2.84–2.64             | shear and quartz vein hosted                                       | Ser ± Sil ± ?                                    | greenschist                | Au ± Cu           | Py ± Cpy   | 2.69–2.63               | constrained by U-Pb zircon dating of magmatism and/or deformation | Ikingura et al., 2009   |
| Central Tanzania                              | Golden Pride              | sedimentary (predominately sandstone/siltstone) rocks, BIF, dacitic intrusions amphibolite, granitic gneisses; | 2.7 to <2.65          | shear and quartz vein hosted; lithologic control                   | Ser ± Chl ± Carb ± Biot ± Cltd                   | greenschist                | Au                | Po + Asp + Py + Cpy ± Gal ± Sph ± Stb ± Tel ± Co-Ni-Bi sulphides | ca. 2.68                | U-Pb dating of co-genetic magmatic phases                         | Vos et al., 2009; Kwelwa et al., 2012a  |
|   | mainly artisanal workings | mafic volcanic and sedimentary rocks (including BIF)   | 2.82–2.66             | shear and quartz vein hosted                                       | Sil ± Pot ± Chl                                  | amphibolite to greenschist | Au                | Py ± ?   | 2.70–2.66               | constrained by U-Pb zircon dating of magmatism and/or deformation | Kabete et al., 2012a  |
| <b>Goldfields at cratonic margins</b>         |                           |  |                       |  |  |                            |                   |  |                         |   |   |
| Lupa  | Kenge                     |  |                       | shear and quartz vein hosted; lithologic contact control           |  |                            |                   |  |                         |   |   |
|   | Mbenge                    | granite, granodiorite, gabbro  | 2.76–1.88             |  | Ser ± Chl ± Carb ± Sil                           | greenschist                | Au                | Py ± Cpy ± Gal ± Mo ± Sph  | ca. 1.88                | Re-Os pyrite  | Lawley et al., in press   |
| Mpanda  | Porcupine                 |  |                       |  |  |                            |                   |  |                         |   |   |
|   | mainly artisanal workings | orthogneisses, metapelites, metabasites, gabbro  | 2.65–1.93             | shear and quartz vein hosted                                       | Ser ± Chl ± Carb                                 | granulite to greenschist   | Pb ± Cu ± Au ± Ag | Gal + Cpy + Py   | ca. 1.2?<br>ca. 0.72?   | U-Pb monazite Pb-Pb galena  | Kuehn et al., 1990; Stendal et al., 2004; Kazimoto and Schenk, 2013                         |
| Niassa  | mainly artisanal workings | sedimentary rocks, gabbro  | ca. 0.714             | shear and quartz vein hosted                                       | Chl ± Carb ± Grun ± Ser                          | greenschist                | Au                | Py ± Cpy ± Mar ± Po ± Sph  | ca. 0.483               | Re-Os pyrite  | Bjerkgard et al., 2009  |
| Burundi                                       | mainly artisanal workings | volcanic and sedimentary rocks, granite  | 1.78–1.37             | shear, quartz vein, and breccia hosted                             | Ser ± Tour ± Oxides                              | greenschist                | Au ± Sn ± W ± Bi  | Py ± Asp ± Cpy ± Bi-sulphides ± oxides (cassiterite ± hematite)  | ca. 1.0–0.9<br>ca. 0.64 | Rb-Sr whole rock, muscovite, and tourmaline                       | Brinckmann et al., 1994; Fernandez-Alonso et al., 2012                                      |

List of abbreviations used. Hydrothermal alteration mineral assemblage: Ser = sercrite, Chl = chlorite, Carb = carbonate minerals, Sil = silica flooding, Grun = grunerite, Pot = potassic, Sod = sodic, Cltd = chloritoid, Tour = tourmaline, Oxides = oxide minerals. Sulphide Assemblage: Py = pyrite, Cpy = chalcopyrite, Po = pyrrhotite, Asp = arsenopyrite, Gal = galena, Sph = sphalerite, Stb = stibnite, Tel = telurides, Mo = molybdenite, Mar = marcasite.

**Table. 1 - Geochronology Summary**

| Sample ID        | Sample type   | Analysis method                        |              | Interpreted age (Ma) | Analytical Uncertainty at 2σ (Ma) | Age determination method  |
|------------------|---|--|--------------|----------------------|-----------------------------------|---|
| CL0975           | Saza Granodiorite                                     | U-Pb zircon <sup>1</sup>               | ID-TIMS      | 1934.5               | 1.0                               | weighted average <sup>207</sup> Pb/ <sup>206</sup> Pb concordant zircon age   |
| CL0972           | Ilunga Syenogranite                                   | U-Pb zircon <sup>1</sup>               | ID-TIMS      | 1959.6               | 1.1                               | weighted average <sup>207</sup> Pb/ <sup>206</sup> Pb concordant zircon age   |
| CL0911           | granodiorite dike cutting foliated granite (CL098)    | U-Pb zircon <sup>1</sup>               | ID-TIMS      | 1958.5               | 1.3                               | weighted average <sup>207</sup> Pb/ <sup>206</sup> Pb concordant zircon age   |
| CL098            | foliated granite                                      | U-Pb zircon <sup>1</sup>               | LA-MC-ICP-MS | 2723                 | 10                                | weighted average <sup>207</sup> Pb/ <sup>206</sup> Pb concordant zircon age   |
| CL1020           | foliated granite                                      | U-Pb zircon <sup>1</sup>               | LA-MC-ICP-MS | 2739                 | 10                                | weighted average <sup>207</sup> Pb/ <sup>206</sup> Pb concordant zircon age   |
| CL109            | foliated granite                                      | U-Pb zircon <sup>1</sup>               | LA-MC-ICP-MS | 2758                 | 9                                 | weighted average <sup>207</sup> Pb/ <sup>206</sup> Pb concordant zircon age   |
| CL1019           | porphyritic monzogranite                              | U-Pb zircon <sup>1</sup>               | LA-MC-ICP-MS | 1942                 | 14                                | weighted average <sup>207</sup> Pb/ <sup>206</sup> Pb concordant zircon age   |
| CL1021           | quartz diorite  | U-Pb zircon <sup>1</sup>               | LA-MC-ICP-MS | 1891                 | 17                                | upper intercept Concordia age   |
| CL1022           | gabbroic dike cutting foliated granite (CL109)        | U-Pb zircon <sup>1</sup>               | LA-MC-ICP-MS | 1880                 | 17                                | upper intercept Concordia age   |
| multiple samples | ultrafine molybdenite from Kenge                      | Re-Os molybdenite <sup>2</sup>         | N-TIMS       | 1953                 | 6                                 | weighted average Re-Os model age  |
| multiple samples | molybdenite from Kenge                                | Re-Os molybdenite <sup>2</sup>         | N-TIMS       | 1937                 | 4                                 | weighted average Re-Os model age  |
| multiple samples | pyrite from mylonitic shear zone at Kenge and Mbenge  | Re-Os pyrite <sup>2</sup>              | N-TIMS       | 1876                 | 10                                | weighted average Re-Os model age  |
| multiple samples | pyrite and chalcopyrite from quartz veins at Kenge    | Re-Os pyrite chalcopyrite <sup>2</sup> | N-TIMS       | 1953                 | 37                                | weighted average Re-Os model age  |
| multiple samples |   |  |              | 1871                 | 12                                |   |
| multiple samples | pyrite and chalcopyrite from quartz veins at Konokono | Re-Os pyrite chalcopyrite <sup>2</sup> | N-TIMS       | 1885                 | 9                                 | weighted average Re-Os model age  |
| multiple samples |   |  |              | 1371                 | 160                               |   |
| multiple samples | ultrafine molybdenite from Porcupine                  | Re-Os molybdenite <sup>2</sup>         | N-TIMS       | 975                  | 6                                 | weighted average Re-Os model age  |
| multiple samples |   |  |              | 1886                 | 6                                 |   |
| multiple samples | molybdenite from Porcupine                            | Re-Os molybdenite <sup>2</sup>         | N-TIMS       | 1873                 | 5                                 | weighted average Re-Os model age  |
| multiple samples |   |  |              | 1894                 | 45                                |   |
| multiple samples | pyrite from Porcupine                                 | Re-Os pyrite <sup>2</sup>              | N-TIMS       | 1057                 | 56                                | weighted average Re-Os model age  |
| multiple samples |   |  |              | 922                  | 190                               |   |
| multiple samples | pyrite from Dubwana                                   | Re-Os pyrite <sup>2</sup>              | N-TIMS       | 1910                 | 38                                | individual Re-Os model ages   |
| multiple samples |   |  |              | 1900                 | 38                                |   |
| CL109            | foliated granite                                      | U-Pb titanite <sup>3</sup>             | ID-TIMS      | 1921                 | 7                                 | upper intercept Concordia age   |
| CL0975           | Saza Granodiorite                                     | U-Pb titanite <sup>3</sup>             | ID-TIMS      | 1930                 | 3                                 | weighted average <sup>207</sup> Pb/ <sup>206</sup> Pb concordant titanite age |

<sup>1</sup>Data taken from Lawley et al., 2013

<sup>2</sup>Data taken from Lawley et al., in press-a

<sup>3</sup>Data from this study