1	Early to Late Pleistocene history of debris-flow fan evolution in western Death Valley using
2	cosmogenic ¹⁰ Be and ²⁶ Al
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16	Keywords: debris-flow fan; cosmogenic dating; Death Valley; inheritance
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28	Highli	ghts
29	•	Debris-flow fans on the western side of Death Valley preserve a long depositional
30		chronology with early to late Pleistocene surface dates
31	•	¹⁰ Be and ²⁶ Al measurements show continuous exposure history on the fan
32	•	Inheritance complicates interpretation of the true depositional ages
33	•	Surface smoothing with elimination of debris-flow morphology takes a few 10^5 ka
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51 Abstract

Debris-flow fans with depositional records over several 10^5 years may be useful archives for the 52 understanding of fan construction by debris flows and post-depositional surface modification 53 54 over long timescales. Reading these archives, however, requires that we establish the temporal and spatial pattern of debris-flow activity over time. We used a combination of geomorphic 55 mapping of fan surface characteristics, digital topographic analysis, and cosmogenic radionuclide 56 dating using ¹⁰Be and ²⁶Al to study the fan evolution of the Warm Springs fan on the west side of 57 southern Death Valley, California. The ¹⁰Be concentrations yield dates that vary from 989 ± 43 58 to 595 ± 17 ka on the proximal fan and between 369 ± 13 and 125 ± 5 ka on distal fan surfaces. 59 The interpretation of these results as true depositional ages though is complicated by high 60 inheritance with a minimum of 65 ka measured at the catchment outlet and of at least 125 ka at 61 the distal fan. Results from the ²⁶Al measurements suggest that most sample locations on the fan 62 surfaces underwent simple exposure and were not affected by complex histories of burial and re-63 exposure. This implies that Warm Springs fan is a relatively stable landform that underwent 64 several 10^5 years of fan aggradation before fan head incision caused abandonment of the 65 proximal and central fan surfaces and deposition continued on a younger unit at the distal fan. 66 We show that the primary depositional debris-flow morphology is eliminated over a time scale of 67 less than 10^5 years, which prevents the delineation of individual debris flows as well as the 68 precise reconstruction of lateral shifts in deposition as we find it on younger debris-flow fans. 69 70 Secondary post-depositional processes control subsequent evolution of surface morphology with the dissection of planar surfaces while smoothing of convex-up interfluves between incised 71 72 channels continues through time.

74 **1. Introduction**

Debris-flow fans are depositional landforms that record the history of sediment transfer from the 75 source area to the sedimentary sink on the fan. In coupled catchment-fan systems along mountain 76 77 fronts, the production of sediment in the catchment, the occurrence of debris flows, and the subsequent deposition of sediment on the fan are modulated by a variety of different controls. 78 79 These include tectonic, climatic, and internal mechanisms as well as changes in base level. Over time repeated debris-flow activity results in a distinct spatial and temporal pattern of sediment 80 deposition on the fan (e.g. Bull, 1964; Bull, 1977; Gordon and Heller, 1993; Whipple and 81 82 Trayler, 1996). Assessing these sediment records allows us to understand the history of fan construction and how debris-flow fan surfaces evolve through time. 83

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85 One first-order boundary condition for the evolution of a sediment fan and its geometry is set by the pattern and rate of basin subsidence, which controls the amount of sediment accommodation 86 and hence the thickness and spatial extent of the fan (Gordon and Heller, 1993; Whipple and 87 Trayler, 1996). On debris-flow fans, debris-flows migrate across the fan surface by lateral and 88 radial shifts in deposition. Lateral shifts result from avulsion of individual flows (Hooke, 1967; 89 Whipple and Dunne, 1992), which over time can lead to resurfacing of a whole alluvial fan. 90 Radial shifts in the depocenter can either occur as backstepping or basinward migration of the 91 depocenter. The latter is often caused by fan head incision, initiated by base level lowering, 92 93 changes in the sediment transport capacity of the alluvial fan system, or autogenic processes (e.g., Harvey, 2002; Kim and Jerolmack, 2007; Nicholas and Quine, 2007; Dühnforth et al., 94 2008; Clarke et al., 2010), and leads to abandonment and preservation of older fan units. Both, 95 96 lateral and radial changes in the depocenter, result in the construction of a debris-flow fan

characterized by multiple individual surface units, each of which was active at a different time 97 (Denny, 1965; Hooke, 1967; Wells et al., 1987; Blair and McPherson, 1994; Ritter et al., 2000; 98 Dühnforth et al., 2007; Harvey, 2011; Ventra and Nichols, 2014; d'Arcy et al., 2015). Cross-99 100 cutting relationships between individual debris-flow channels and levees as well as differences in 101 the degree of weathering of debris-flow boulders are useful criteria to identify individual fan 102 surface units (Whipple and Dunne, 1992; Dühnforth et al., 2007; Le et al., 2007; Schürch et al., 2016). The history of fan occupation in time and space including the migration of the sediment 103 depocenter can therefore be reconstructed by using the morphological properties on the fan 104 105 surfaces in combination with numeric dating techniques (Bierman et al., 1995; Zehfuss et al., 106 2001; Dühnforth et al., 2007; Le et al., 2007; Schürch et al., 2016). While many studies on debris-flow fan activity have established fan chronologies with ages younger than 100 ka, it is 107 not clear how far back in time we can reconstruct debris-flow fan activity as previous results 108 have shown that primary depositional morphology becomes obscured or removed by secondary 109 processes over time (e.g., Wells et al., 1987; Bull, 1991; Matmon et al., 2006; Frankel and 110 111 Dolan, 2007; Regmi et al., 2014). Models of fan surface relief evolution suggest that the degradation of the original depositional morphology leads to smoothing of alluvial fan surfaces 112 after abandonment. With ongoing time these smooth surfaces become dissected by channels with 113 an increase in the relief between channel floors and interfluves (Wells et al., 1987; Bull, 1991; 114 Matmon et al., 2006; Frankel and Dolan, 2007; Regmi et al., 2014). It would therefore be useful 115 to estimate the age range over which a fan surface preserves a record of debris-flow deposition, 116 which in turn requires analysis of fans with chronologies that extend back for several 10^5 years. 117 This time scale is long enough to examine the long-term evolution of a debris-flow fan without 118 119 the inherent variability in the tectonic and climatic boundary conditions. So far only a limited

number of field studies are available that address the pattern of debris-flow deposition in combination with numeric age constraints to explore the long-term signal of debris flow activity preserved on the alluvial fan.

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Arid-region alluvial fans in the southwestern United States are a well-studied target for 124 sedimentary processes on alluvial fans and have a good record of numeric age control. Results 125 from numeric dating techniques such as cosmogenic ¹⁰Be and ²⁶Al dating, ¹⁴C dating, Optical 126 Stimulated Luminescence (OSL), or tephronochronology show that the majority of alluvial fan 127 128 surfaces in the southwest of North America have ages younger than 100 ka (Owen et al., 2014 and references therein). Only a limited number of alluvial fans in the southwestern United States 129 show dates close to or older than 10⁵ years (e.g. Bierman et al., 1995; Reheis et al., 1996; 130 131 Matmon et al., 2005; Matmon et al., 2006; Frankel et al., 2007a; Frankel et al., 2007b; Machette et al., 2008; Spelz et al., 2008; Owen et al., 2011; Gray et al., 2014; Owen et al., 2014). Debris-132 flow dominated fans on the west side of southern Death Valley show exceptionally old surfaces 133 134 with dates extending back to around 400 ka (Nishiizumi et al., 1993; Liu and Broecker, 2008; Machette et al., 2008). Therefore, these debris-flow fans represent ideal targets to study the long-135 term dynamics of debris-flow fan activity and surface modification over a period that includes 136 multiple 10^5 years. We aim to build upon the existing dataset of numeric alluvial fan 137 chronologies in the Death Valley/Mojave Desert region and use these time constraints as 138 139 reference for a comprehensive study of numeric ages across an individual fan in Death Valley. A relatively dense pattern of multiple sampling sites across an entire fan would allow the 140 construction of a numeric age chronology that records times and durations of fan activity and 141 142 abandonment. The chronology also enables us to assess the time scale over which postdepositional change of the fan surfaces occur. Both, the timing on fan deposition and postdepositional modification of fan surfaces allows us to validate the time scale over which existing
models of fan evolution with surface smoothing and dissection operate.

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The strategy is first to establish a relative age chronology of fan deposition, which in turn guides 147 our sampling for a numeric age chronology using cosmogenic ¹⁰Be and ²⁶Al. Especially on 148 depositional surfaces with ages older than several 10⁵ years, the influence of surface erosion or 149 surface reworking may limit our ability to measure true depositional ages. In addition, the 150 amount of inheritance can make up a significant fraction of the total nuclide concentration in a 151 152 sample, which challenges the interpretation as depositional age (Anderson et al., 1996; Hancock et al., 1999). Therefore, the amount of inheritance will have to be determined, as inheritance on 153 154 alluvial fans in Death Valley can be up to 100 ka (Machette et al., 2008; Owen et al., 2011; Frankel et al., 2015). Based on a combination of field work, analysis of digital airborne swath 155 mapping (ALSM) topographic data, and cosmogenic surface exposure dating using in-situ ¹⁰Be 156 and ²⁶Al, our goal is (a) to establish a relative and numeric age chronology on an individual 157 158 alluvial fan, (b) to investigate the temporal and spatial pattern of numeric age constraints, and c) to evaluate the time scale over which primary debris-flow morphology is removed by post-159 depositional fan surface modification. 160

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162 **2. Setting**

We focus on the Warm Springs fan located on the western side of Death Valley in the southwestern United States (Fig. 1). Death Valley is an extensional half-graben bounded on its eastern margin by the oblique-normal Black Mountain fault zone (BMFZ). To the north and 166 south, the BMFZ merges with the right-lateral Northern Death Valley fault zone (NDVFZ) and 167 Southern Death Valley fault zone, respectively (Machette, 2001). The slip rate of the NDVFZ is 168 approximately 4-5 mm yr⁻¹ averaged over the past 100 ka (Frankel et al., 2007b). For the BMFZ, 169 the minimum normal slip rate calculated over that past 120-186 ka is 0.15-0.2 mm yr⁻¹ (Knott 170 and Wells, 2001), while Holocene vertical slip rates have been estimated at 1-3 mm/yr (Klinger 171 and Piety, 2001) and 0.5-2 mm/yr (Frankel et al., 2015). Minor normal faulting also occurs along 172 the western margin of Death Valley on the West Side Fault Zone (Machette et al., 2008). Normal 173 slip on the BMFZ leads to large accommodation on the east side of the valley, and less 174 accommodation on the west side (Hooke, 1972), so that fan deposits on the western side are generally thinner but larger in surface area (~20-40 km²) than those on the east, with long radial 175 176 lengths from apex to toe (e.g. Hunt and Mabey, 1966; Hooke, 1972; Gordon and Heller, 1993; 177 Whipple and Trayler, 1996). As the rate of resurfacing is high on the eastern alluvial fans in 178 Death Valley, these alluvial fans show younger age chronologies compared to fan surfaces on the 179 west side (Machette et al., 2008; Owen et al., 2011; Frankel et al., 2015).

The Warm Springs fan, in the southwestern corner of the basin, stands out from all other debrisflow fans on the west side of southern Death Valley with its prominent surface relief, expressed as the elevation difference between interfluves and incised channel floors (Fig. 1). Therefore, the Warm Springs fan is well-suited to test models of post-depositional surface evolution (Matmon et al., 2006; Frankel and Dolan, 2007; Regmi et al, 2014). Based on such model we expect to find the oldest surface dates at the proximal fan, where the relief between channel floors and interfluves is highest while undissected distal fan areas would be younger.

The Warm Springs fan is sourced in the Panamint Range and has a surface area of about 30 km²
(Fig. 1). The topographic elevation of the fan varies from about 60 m below sea level at the distal

189 fan margin to about 360 m above sea level at the fan head. Several NNE-SSW-trending antithetic 190 and synthetic normal faults of the West Side Fault Zone run across the fan (Hooke, 1972; Blair, 191 1999a), with fault scarp heights that range between a few meters to several tens of meters. Blair 192 (1999a) pointed out that this local faulting pattern strongly controls sedimentation and surface 193 preservation on the Warm Springs fan. Sediment is supplied to the fan by debris flows 194 originating in the associated Warm Springs catchment, as shown by the fan stratigraphy (Blair 195 1999a; Blair, 1999b). Sediment is stored in small alluvial cones in the catchment area, and these 196 cones are the sediment source for debris flows exiting onto the Warm Springs fan. The debris 197 flows contain cobbles and boulders of Precambrian to Paleozoic quartzite, dolomite, and shale, 198 Mesozoic granite, and Triassic to Jurassic andesite (Johnson, 1957; Jennings et al., 1962; Miller, 199 1985). Debris-flow deposits exposed in incised channel walls on the alluvial fan are commonly 200 preserved as meter-thick, matrix-rich units with abundant cobble-sized clasts (2-30 cm) and low 201 to moderate concentrations of boulders with sizes up to 1 to 2 m.

202 As with other debris-flow fans on the western side of Death Valley, the Warm Springs fan 203 head is incised by an active channel, in this case about 30 m deep and 300 m wide (Fig. 2). The 204 incision resulted in a basinward shift of the active depocenter to the distal part of the fan, and has 205 been attributed to tectonic activity and eastward tilting of the fan surface (Hunt and Mabey, 206 1966; Hooke, 1972; Hooke and Dorn, 1992). Older depositional surfaces on the fan are also 207 dissected by channel incision, resulting in the formation of elongated ridges (Fig. 2). Interfluves 208 are generally planar with very low cross-sectional curvature, except where channel spacing is 209 narrow and diffusional channel margins have coalesced into gently-curved, convex-upwards 210 ridge crests. These planar interfluve surfaces often have well-developed desert pavements 211 composed of closely-interlocking, varnished clasts of sizes similar to those found in the channel walls (Fig. 3). Boulders are rare on both proximal and medial fan surfaces, but are relatively
 frequent in the incised channels and in the more recent deposits at the distal fan.

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215 While the present-day climate in Death Valley is extremely arid, with an annual mean 216 precipitation rate of less than 65 mm yr⁻¹ (Western Region Climate Center – 217 http://www.wrcc.dri.edu), paleoclimatic studies from Death Valley and adjacent basins have 218 documented significant hydrologic variations since 1.2 Ma (Smith, 1984; Winograd et al., 1985; 219 Jannick et al., 1991; Smith and Bischoff, 1997; Lowenstein et al., 1999). Evidence for large 220 paleolakes, indicating generally wetter climate conditions, has been found in sediments from 221 Searles, Panamint, Owens, and Death Valley basins (Smith, 1984; Jannick et al., 1991; Li et al., 222 1996; Ku et al., 1998; Lowenstein et al., 1999; Forester et al., 2005). In Death Valley, the 223 Badwater sediment core combined with outcrop dates provide direct evidence for a large lake 224 between 186 and 120 ka and a smaller lake between 35 and 10 ka (Ku et al., 1998; Lowenstein et 225 al., 1999). According to micropalaeontologic data (Forester et al., 2005), the highest lake level in 226 the past 200 ka occurred at ~155 ka. Machette et al. (2008) determined that the Lake Manly lake 227 level was at 30 m a.s.l. at 130 ka, and noted that higher shoreline deposits on the Hanaupah 228 Canyon fan at >67 m a.s.l. may be related to a lake level highstand during marine isotope stage 229 (MIS) 6 or possibly MIS 8. On the Warm Springs fan, a paleoshoreline is cut into the distal fan 230 deposits at a topographic elevation of ~65 m a.s.l. (Hooke, 1972; Blair, 1999b) (Fig. 2). It is very 231 likely that this shoreline is related to the lake level highstand during MIS 6 (Hooke and Dorn, 232 1992; Blair, 1999a; Knott et al., 2002; Machette et al., 2008; Owen et al., 2011).

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3. Prior fan surface chronologies in Death Valley

236 Debris-flow fans on the western side of Death Valley have been the subject of numerous 237 investigations of depositional processes and surface evolution on arid-region fans (e.g. Denny, 238 1965; Hunt and Mabey, 1966; Hooke, 1972; Hooke and Dorn, 1992; Blair, 1999b). Many of 239 these studies have mapped different depositional units and established a relative fan chronology. 240 Denny (1965) and Hunt & Mabey (1966) were the first to publish a relative age chronology for 241 alluvial fan surfaces in Death Valley based on stratigraphic relationships, soil development, the 242 degree of desert pavement and rock varnish development, and the distribution of boulder sizes, 243 and this approach has since been widely applied to other alluvial surfaces beyond Death Valley 244 (e.g. Bull, 1991).

245 On the Warm Springs fan, Hunt and Mabey (1966) mapped three different alluvial fan units, 246 termed Qg2, Qg3, and Qg4 in order of decreasing age. We note that Hunt and Mabey (1966) also 247 described an even older alluvial fan unit QgT1 (Q1 according to Knott et al., 2005), which was 248 characterized as a degraded form of surface unit Qg2 showing a lack of desert pavement and 249 rock varnish (Menges et al., 2001) or only patches of well-developed desert pavement (Hooke, 250 1972). Even though this unit has not been identified on the Warm Springs fan (Hunt and Mabey, 251 1966), this older unit has been mapped for example on the Hanaupah Canyon fan (Fig. 1) 252 (Hooke, 1972; Menges et al., 2001). According to Hunt and Mabey (1966) deposits of unit Qg2 253 generally form the highest surfaces above the present-day washes. They are characterized by 254 well-developed, smooth desert pavement of highly varnished clasts underlain by a several-255 centimeter thick Av-horizon consisting mainly of eolian silt and clay. Underneath the Av-256 horizon, cementation of clasts by calcium carbonate (caliche) is ubiquitous. Unit Qg3 is located 257 more towards the toe of the Warm Springs fan and overlaps unit Qg2 at its lowest position (Hunt 258 and Mabey, 1966). Surfaces of unit Qg3 show subdued bar and swale morphology and an 259 intermediate degree of desert pavement development. Similarly, clasts on Qg3 are less cemented 260 than those on Og2 and the thickness of the Av-horizon is similar or thinner compared to Og2 261 (Hunt and Mabey, 1966). Deposits of unit Qg4 are present in modern washes and differ clearly 262 from units Qg2 and Qg3 by the lack of desert pavement and desert varnish (Hunt and Mabey, 263 1966). Deposits of unit Qg4 show a well-developed bar and swale morphology originating from 264 primary debris-flow depositional forms. Hooke (1972) introduced additional criteria such as the 265 degree of fan dissection and surface relief, which led to a further subdivision of the units. A 266 number of subsequent studies have sought to assign absolute ages to this relative chronology 267 (e.g., Klinger, 2001; Frankel et al., 2007a; Frankel et al., 2007b; Sohn et al., 2007; Machette et 268 al., 2008; Owen et al., 2011; Frankel et al., 2015). Indirect dating of fan surfaces by dating 269 shoreline tufas developed on bedrock along the Black Mountains south of Badwater using U-270 series techniques (Ku et al., 1998), and correlation with Hunt and Mabey's unit Qg2, led to the 271 suggestion that Qg2 is younger than MIS 6 (Knott et al, 2002). Other correlations of Hunt and 272 Mabey's stratigraphic fan units with numeric dates (e.g., Klinger, 2001; Sohn et al., 2007; 273 Machette et al., 2008; Owen et al., 2011; Sohn et al., 2014 Frankel et al., 2015) suggest that units 274 Qg4 and Qg3 are Holocene in age, and support the hypothesis that Qg2 was deposited in late 275 Pleistocene time.

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The first cosmogenic ¹⁰Be and ²⁶Al measurements on western Death Valley fans were published by Nishiizumi et al. (1993). They analyzed samples from alluvial fan units Q3a, Q2a, and Q1b on the Hanaupah Canyon fan and from units Q1b and Q2a on the Galena Canyon fan (Fig. 1; Table 1), based on an alluvial fan unit map by Dorn (1988). Minimum ages on the Hanaupah fan 281 are 145.8 ± 10.9 ka (Q3a), 325.5 ± 24.7 ka (Q2a), and 383.4 ± 30.8 ka (Q1b). On the Galena Canyon fan the two measured minimum ages are 403.1±36.8 ka (Q1b) and 410.1±29.2 ka 282 283 (Q2a). These dates have been recalculated with CRONUS calculator (http://hess.ess.washington.edu/math/; version 2.2; Balco et al., 2008) applying the production 284 rate calibration by Heyman (2014) and the time-variant scaling model by Lal (1991) and Stone 285 (2000). Machette et al. (2008) sampled depth-profiles for cosmogenic ³⁶Cl to measure numeric 286 ages on the Hanaupah Canyon, Trail Canyon, and Galena Canyon fans. They took six samples 287 from one single fan unit (Qai) and one sample from the slightly older fan unit Qlm on the 288 289 Hanaupah Canyon fan. In addition, they took two samples from unit Qai on the Galena Canyon fan and one sample each from units Qai and Qaio on the Trail Canyon fan (Fig. 1; Table 1). The 290 291 two profiles on the Galena Canyon fan yielded best-fit profile ages of 47 ka and 84 ka for unit 292 Qai. Five profiles on the Hanaupah Canyon fan gave best-fit ages of 39 ka, 66 ka, 72 ka, 96 ka 293 (all unit Qai), and 130 ka (unit Qlm), and one rock sample had an age of 118 ± 11 ka (Qai). Unit 294 Qaio on the Trail Canyon fan vielded the oldest best-fit depth-profile age at 171 ka and the rock 295 sample from unit Qai had an age of 102 ± 7 ka (Machette et al., 2008). Another set of numeric 296 results comes from Liu and Broecker who applied the varnish microlamination (VLM) dating 297 technique on the Galena Canyon, Hanaupah Canyon, Six Spring Canyon and Warm Springs fans. They measured 13 samples from Bull's (1991) fan units O2a, O2b, and O2c (Table 1) and 298 299 obtained VML ages of 276 ka and 295 ka for unit Q2a, three VML ages between 74 and 165 ka 300 for unit Q2b, and VML ages between 30 and 60 ka for unit Q2c. Among the latter two samples 301 were obtained from the Warm Springs fan.

303 On the eastern side of Death Valley, cosmogenic radionuclide and OSL dates on several alluvial fan surfaces near Badwater and Mormon Point (Fig. 1) are generally younger compared to 304 measurements on the west side alluvial fans (e.g. Machette et al., 2008; Owen et al., 2011; Sohn 305 306 et al., 2014; Frankel et al., 2015). Most of the dated alluvial fan surfaces yield exposure ages that are younger than 20 ka, while a few samples have dates between 40 and 80 ka (Sohn et al., 2007; 307 Owen et al., 2011; Frankel et al., 2015). Inheritance on these dated alluvial fan surfaces, as 308 309 measured in samples from the active channels and from cosmogenic depth-profiles, ranges between 6 and 94 ka (Machette et al., 2008; Owen et al., 2011; Frankel et al., 2015), and is thus 310 fairly high compared to the exposure ages of the surfaces themselves. 311

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313	Table 1. Correlation of mapped fan surface units in Death V	Valley
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		Hunt and	Machette et al.	Nishiizumi et	Bull $(1991)^2$
		Mabey (1966)	(2008)	al. $(1993)^{4}$	
					Q4b
ne		Qg4			Q4a
Ce	(0-10 ka)		Qay		
olo	(0 10 114)	Qg3			Q3 (a-c)
Ħ			Oavo		
		_			O2c
	Late				
	(10-130 ka)		Oai	O3a	
o	()		Zui	254	
cen		Οσ2	Olm		O2h
toe			Qaio	O_{2a}	220
eis	Middle		Qaio	Q2a	
Ы	(120, 780 kg)		Qau		Ω_{2}
	(130-700 ka)			011	Q2a
		_		QID	
	F_{1} (7001)	OT 1	07		
	Early (>780 ka)	QIgI	Qla		
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315 ¹The fan surface units are based on Dorn (1988).

² Units were originally based on the Lower Colorado River region, but applied to the Death Valley region.

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4. Methods

³²⁴ 4.1. Mapping of fan surfaces

325 In order to establish a relative chronology to guide our sample collection, we mapped distinct 326 depositional fan surface units using similar criteria to those applied in previous studies by Denny 327 (1965) and Hunt and Mabey (1966): channel spacing, degree of pavement development, varnish 328 development, and short- and long-wavelength relief on the fan surfaces. Fan areas with similar 329 surface characteristics were mapped as a single surface unit. Surface relief and channel spacing 330 were derived from high-resolution Airborne Laser Swath Mapping (ALSM) data obtained from 331 the National Center for Airborne Laser Mapping. The survey area comprised approximately 25 332 km² and contained 25 flight lines each with a scan width of 582.4 m and swath overlap of 260 m. 333 The data were interpolated to a digital elevation model with a spatial resolution of 1 m/pixel and 334 a vertical accuracy of 5-10 cm. Due to the lack of vegetation the data were not filtered for 335 vegetation removal. Short-wavelength relief was defined as elevation variation over a lateral 336 distance of 10 m and reflects roughness on the scale of individual debris-flow channels and 337 deposits. Long-wavelength relief was assessed over a radius of 1000 m and reflects topographic 338 variations due to post-depositional incision of the fan surface units. The degree of pavement and 339 varnish development were qualitatively assessed in the field.

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³⁴¹ **4.2.** Cosmogenic radionuclide exposure dating

We dated debris-flow fan surfaces on the Warm Springs fan using the in-situ produced cosmogenic radionuclides ¹⁰Be and ²⁶Al. The use of both nuclides allows the identification of complex exposure and burial histories which can potentially complicate the interpretation of ages from depositional surfaces (e.g., Repka et al., 1997; Hancock et al., 1999; Wolkowinsky and 346 Granger, 2004; Matmon et al., 2005; Frankel et al., 2007b; Machette et al., 2008). We collected 347 28 samples from different fan surface units on the Warm Springs fan (Fig. 4; Table 2). Surfaces 348 on the Warm Springs fan have few boulders which could be used as sampling targets (Fig. 3B). 349 Therefore, we amalgamated 30 to 40 individual surface clasts from each sample location (e.g., 350 Anderson et al., 1996; Repka et al., 1997; Hetzel et al., 2002; Marchetti and Cerling, 2005) in 351 order to statistically average out potential inheritance outliers (Anderson et al., 1996; Repka et 352 al., 1997). We sampled clasts with diameters between 5 and 10 cm. At all sampling sites we 353 collected the same lithology: a fine-grained quartzite. We chose sampling sites on planar fan 354 surfaces with well-developed pavement, well away from hillslopes and channel margins, except 355 in the active channel where we took clasts directly from the wash.

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The exposure of rock in the source area during erosion and transport, prior to final deposition on the fan can result in a significant concentration of inherited cosmogenic nuclides. Thus, the assessment of nuclide inheritance is absolutely essential for the interpretation of surface dates. Measurements of nuclide concentrations from multiple depths at the same site are one way to account for inheritance (Anderson et al., 1996; Repka et al., 1997). When depth sampling is not an option, a less accurate approach is to estimate the nuclide concentration in a sample from the

TABLE 2. ANALYTICAL DATA AND RESULTS FOR ¹⁰BE AND ²⁶AL SAMPLES ON THE WARM SPRINGS FAN, DEATH VALLEY 363 364

Sample Name ETH	Sample Name	Latitude (°N)	Longitude (°W)	Elevation (m)	Quartz (g) ^a	Be carrier mass (mg)	¹⁰ Be/ ⁹ Be x 10 ^{-12 b,c,d}	Al carrier mass (mg)	²⁶ Al/ ²⁷ Al x 10 ^{-12 b,d}
70221	10202 1	25.067	116.854	208	20.2	0 2011	4.70 ± 0.16	2 5100	× 46 ± 0.25
ZD2021 ZD2012	10505-1	35.907	110.634	508	30.5	0.3911	4.70 ± 0.10	3.3100	8.40 ± 0.23
ZB3812	80303_2	35.991	110.813	-13	27.4	0.3973	1.09 ± 0.05	3.1500	2.49 ± 0.11
ZB5430	80303_3	35.9837	116.813	57	26.5	0.4061	2.33 ± 0.05	3.3950	4.76 ± 0.14
ZB3813	80303_4	35.984	116.819	41	27.4	0.3968	1.80 ± 0.06	3.2700	4.17 ± 0.19
ZB2822	90303-2	35.982	116.829	124	29.8	0.3622	3.36 ± 0.12	3.2900	6.88 ± 0.21
ZB2820	190303-1	35.962	116.857	349	30.0	0.4161	4.79 ± 0.14	3.0450	8.79 ± 0.28
ZB3814	110304_6	35.955	116.846	288	27.0	0.3973	3.20 ± 0.10	2.9500	7.88 ± 0.37
ZB3298	120304_1	35.997	116.846	128	29.0	0.3904	2.87 ± 0.09	3.1900	5.37 ± 0.19
ZB3820	120304_2	35.998	116.859	276	25.8	0.3673	3.75 ± 0.14	3.0350	6.88 ± 0.35
ZB3815	120304_4	35.982	116.857	276	27.6	0.3978	3.54 ± 0.16	3.0950	7.55 ± 0.35
ZB3821	130304_1	35.975	116.816	102	31.4	0.3861	3.57 ± 0.11	3.4200	6.89 ± 0.28
ZB3816	130304 3	35.974	116.809	73	26.6	0.3979	2.66 ± 0.08	2.9200	6.47 ± 0.27
ZB3822	150304 1	35.999	116.803	-56	22.6	0.3951	1.15 ± 0.03	3.4750	2.44 ± 0.12
ZB3817	150304 3	35.994	116.808	-21	26.4	0.3969	1.07 ± 0.03	3.1500	2.54 ± 0.11
ZB3299	170304 1	35.998	116.834	73	28.4	0.3988	0.83 ± 0.03	3.2950	2.01 ± 0.09
ZB3818	1703042	36.014	116.830	-34	27.8	0.3954	1.09 ± 0.03	3.9350	2.05 ± 0.11
ZB3823	180304 1	36.005	116.813	-6	33.5	0.3879	0.93 ± 0.04	4.9050	1.32 ± 0.07
ZB3823	180304^{-2}	36.002	116.812	0	24.2	0.3964	0.71 ± 0.03	2.6350	2.05 ± 0.09
ZB3824	180304_5	35,988	116.842	171	25.4	0.3917	3.51 ± 0.11	2.8700	8.19 ± 0.32
ZB5421	280207 1	35,9611	116.813	415	26.4	0.3993	1.00 ± 0.05	3,7700	2.00 ± 0.09
ZB5422	280207 1A	35,9611	116.813	415	29.2	0.3971	0.43 ± 0.02	3,9400	0.92 ± 0.08
ZB5423	280207_1B	35 9611	116 813	415	27.5	0 3998	0.38 ± 0.02	2.9450	1.17 ± 0.06
ZB5424	280207 2	35 9749	116 813	235	27.1	0.4040	0.60 = 0.02 0.61 ± 0.02	3 4450	1.37 ± 0.06
ZB5425	280207_3	35 9687	116 813	310	25.8	0.4048	0.01 = 0.02 0.41 ± 0.02	3 9800	0.77 ± 0.04
ZB5426	20307_1	36.002	116.813	-13	28.0	0 4070	0.62 ± 0.02	3 4350	1.39 ± 0.05
7B5/127	20307 14	36.002	116.813	-13	25.2	0.3083	1.03 ± 0.02	3 1950	2.74 ± 0.05
ZB5429 7B5428	20307_1R	36.002	116.813	-13	23.6	0.3783	1.03 ± 0.04 0.50 ± 0.02	3 5000	2.74 ± 0.10 1 30 ± 0.06
ZB5420 7B5420	20307_10	36.002	116.813	-15	32.7	0.4035	0.39 ± 0.02 0.73 ± 0.02	1 7950	1.30 ± 0.00 1.10 ± 0.05

365 366

^a We used a density of 2.7 g cm⁻³ for all samples. ^b Sample ratios were normalized to S555 (10 Be; Kubik and Christl, 2010) and ZAL94 (26 Al; Nishiizumi, 2004) standardizations. ^c Blank corrected 10 Be/ 9 Be ratio including a correction of 2.54 x 10⁻¹⁴ ± 2.12 x 10⁻¹⁵ based on the weighed mean of nine process blanks. No correction was applied to the 26 Al/ 27 Al 367 368 ratio.

369 ^d Uncertainty includes 1_o AMS measurement errors, statistical counting error and the uncertainties due to the normalization to the standards and the blank. 370 active channel, and to use that as a proxy for accumulated inheritance during sample deposition 371 (e.g., Brown et al., 1998; Hetzel et al., 2002; Machette et al., 2008; Owen et al., 2011). This 372 approach assumes constant inheritance accumulation through time, or equivalently a constant 373 mean time required for sample erosion and transport from the sediment source to the fan. Due to 374 National Park restrictions we were unable to excavate alluvial fan surfaces to a sufficient depth 375 for profile sampling. Therefore, we assessed inheritance by collecting 10 amalgamated samples 376 from the active channel at different distances from the catchment outlet. The sampling procedure 377 was the same as for the fan surface samples. To check reproducibility, we took several duplicate 378 samples from the active channel at proximal, mid, and distal fan positions: three at the catchment 379 outlet, two each at 3 and 4.5 km downstream from the outlet, and five at the distal part of the fan, 380 \sim 7.5 km from the outlet.

381

All samples were processed at the PSI/ETH Zürich cosmogenic laboratory between 2003 and 382 2007 using standard techniques according to Kohl and Nishiizumi (1992), Ivy-Ochs (1996), and 383 Ochs and Ivy-Ochs (1997). For our analysis we only took clasts with a thickness of ~2 cm, and 384 we used a rock saw to cut those with larger thicknesses. We crushed and sieved each clast to a 385 grain size of 0.25 to 0.5 mm, from which we took the same mass to contribute to the 386 amalgamated sample. The inherent Al concentration was determined from an aliquot of the 387 solution of the dissolved quartz using ICP-OES measurements carried out at the University of 388 Jena. No additional ²⁷Al was added to our samples. The isotopic ratios were determined by 389 accelerator mass spectrometry (AMS) at PSI/ETH Zürich before April 2010. Concentrations 390 were referenced to ¹⁰Be standard 07KNSTD (Kubik and Christl, 2010) and ²⁶Al standard 391 KNSTD (Nishiizumi, 2004). For ¹⁰Be we ran nine process blanks and used their weighted mean 392

ratio of 2.54 x $10^{-14} \pm 2.12$ x 10^{-15} for the blank correction. We did not apply a blank correction to 393 the ²⁶Al/²⁷Al ratio. We calculated ¹⁰Be and ²⁶Al exposure ages using the CRONUS Earth online 394 exposure age calculator (http://hess.ess.washington.edu/math/, Balco et al., 2008; version 2.2 395 396 with updated constants based on version 2.2.1) and the time dependent Lal (1991)/Stone (2000) scaling model. We applied the production rate calibration of Heyman (2014) with production 397 rates of 3.94 ± 0.2 atoms g⁻¹ SiO₂ yr⁻¹ for ¹⁰Be and 26.59 ± 1.35 atoms g⁻¹ SiO₂ yr⁻¹ for ²⁶Al, and 398 assumed a production ratio ²⁶Al/¹⁰Be of 6.75 (Lal, 1991; Stone, 2000; Heyman, 2014). A ¹⁰Be 399 half-life of 1.39 x 10⁶ yr (Chmeleff et al., 2010; Korschinek et al., 2010) and a ²⁶Al half-life of 400 7.08 x 10^5 yr (Nishiizumi, 2004) were used for the age calculation. The effect of topographic 401 shielding is less than 0.5%, and therefore we did not include a correction. Given the present and 402 past climatic setting of Death Valley, we did not correct for either snow cover or vegetation 403 404 cover.

405

406 **5. Results**

407 **5.1. Relative chronology of fan surface units**

Following Hunt and Mabey (1966), we divide the Warm Springs fan into three different fan surface units, named Qg2, Qg3, and Qg4 in order of decreasing age. Qg4 is the most recently active unit and shows well-developed bar-and-swale morphology with clear channels and depositional levees. Qg2 and Qg3, in contrast, have surfaces that are modified by postdepositional processes, and do not reflect the original depositional morphology.

Fan surface unit Qg2 is located at the proximal part of the fan (Figs. 2 and 4), and comprises the largest fraction of the entire fan area. This unit is incised by channels, which have an average spacing of about 200 m (Figs. 2 and 5) and relief between the channel floors and the interfluves of up to 50 m (Figs. 2, 5). The surfaces of Qg2 are smooth, and lack obvious bar-and-swale
morphology or any debris-flow depositional characteristics, such as channel-levee complexes.
Away from the incised channels, the surfaces are planar and have very low short-wavelength
relief of several centimeters to a few decimeters.

420

421 Well-developed pavements with darkly varnished clasts are partly inset into the underlying fine-422 grained Av- horizon. Surface unit Qg2 is cut by several faults of the West Side Fault zone, which 423 appear to dip generally eastwards into the basin and show normal displacement. Offsets of the 424 alluvial fan surface are between ~ 10 and ~ 60 m across individual fault strands, with the largest 425 offset on a fault located approximately 4 km from the fan head (Fig. 2). To the east of this fault, 426 surface unit Qg2 is substantially less incised, with a long-wavelength relief of less than 20 m 427 (Fig. 2). Near the distal limit of unit Qg2, the paleoshoreline is cut into the surface at an 428 elevation of ~65 m a.s.l. and thus lies topographically higher than unit Qg3 (Fig. 2) (Hooke, 429 1972; Blair, 1999a; Machette et al., 2008).

430

431 Alluvial fan surface unit Qg3 is relatively small in area and forms several individual surface 432 fragments at the distal part of the fan that have been isolated from each other by later incision. 433 These fragments stand a few meters above the presently-active channels. The long-wavelength 434 relief on surface unit Qg3 is significantly smaller than that of surface unit Qg2 and ranges 435 between 2 and 5 m (Fig. 5). In contrast, the short-wavelength surface relief is higher compared 436 with Qg2, on the order of a few decimeters. The surface of unit Qg3 has pavements with darkly-437 varnished clasts. As fan unit Qg3 is inset below the Warm Springs fan paleoshoreline, Qg3 must 438 be younger, and is therefore probably younger than MIS 6.

Fan surface unit Qg4 represents the active depositional surface at the distal part of the Warm Springs fan, including the incised trunk channel that funnels sediment from the catchment outlet downstream into the basin. The long-wavelength relief on this modern surface is less than 2 m (Fig. 5); short-wavelength surface relief is on the order of several decimeters. There is no pavement and no Av horizon developed on this lobe and most clasts lack any apparent coating of desert varnish.

445

446 **5.2. Measurements of ¹⁰Be and ²⁶Al from amalgamated clast samples**

447 If we assume no erosion, the ¹⁰Be exposure data from the Warm Springs fan surfaces yield 448 apparent exposure dates between 989 ± 43 ka and 125 ± 5 ka, whereas the results from ²⁶Al range 449 from 957 ± 74 ka to 124 ± 6 ka (Figs. 4 and 6; Table 3; all ages reported to $\pm 1\sigma$ uncertainty). 450 Dates obtained from both nuclides compare relatively well, especially for samples younger than 451 500 ka. Eleven ¹⁰Be dates from the oldest fan unit Qg2 range between 989 ± 43 ka and 628 ± 17 452 ka and are also the oldest measurements on the Warm Springs fan (Fig. 6). Two other dates from 453 Qg2 (272±13 ka and 448±18 ka), however, are younger than all other measurements on this 454 unit. The younger sample was collected from an elevation of -13 m, well below the maximum 455 lake level highstand during MIS 6 (Hooke, 1972; Blair, 1999a; Forester et al., 2005), and 456 therefore could have been affected by reworking through wave action during the lake phase. The 457 older sample originates from the hanging wall of a west-dipping normal fault with a vertical 458 surface offset of about 15 m. It is possible that this sample has been affected by reworking 459 associated with fault motion and fault scarp degradation. ¹⁰Be exposure dates on unit Qg3 range 460 from 369 ± 13 ka to 180 ± 8 ka and four out of five samples have dates between 280 ± 10 ka and 461 180 ± 8 ka (Figs. 4 and 6).

462 The ten samples from the presently active, incised channel of unit Og4 vield ¹⁰Be concentrations 463 ranging from $3.69 \pm 0.2 \times 10^5$ to $10.7 \pm 0.4 \times 10^{5}$ ¹⁰Be atoms g⁻¹ and ²⁶Al concentrations that range 464 between $26.5 \pm 1.4 \times 10^5$ and $75.7 \pm 2.9 \times 10^5$ atoms g⁻¹ (Table 3). Figure 7 shows the pattern of 465 measured ¹⁰Be concentrations along the active channel. While the spread of concentrations at the 466 catchment outlet and at the distal part of Qg4, where we took multiple samples, varies by a factor 467 of 2 to 3, measured concentrations tend to be lower at the fan head compared to those at the fan 468 toe. When we scale ¹⁰Be concentrations to the local production rate at the sampling site, the calculated ¹⁰Be exposure dates would range between 63 ± 3 ka and 174 ± 9 ka at the catchment 469 470 outlet and between 125 ± 5 ka and 275 ± 12 ka at the fan toe. The equivalent ²⁶Al dates range 471 between 70 ± 7 ka and 168 ± 9 ka at the fan head and between 124 ± 6 ka and 310 ± 15 ka at the 472 distal position of the fan.

473

474 Two-nuclide erosion island plots elucidate the erosional history of the dated landform (or 475 boulder) and complex histories through burial and exposure (Fig. 8). The upper curve of a two-476 nuclide erosion island plot delineates the trajectory of a sample with increasing age but 477 undergoing no erosion. As soon as the sample is affected by erosion the trajectories move below 478 the zero-erosion line and show lower ²⁶Al/¹⁰Be ratios, respectively. These trajectories will 479 continue until an isotopic equilibrium is reached, where production and decay of cosmogenic 480 nuclides are balanced. Samples reveal a simple exposure history, when they fall between the 481 non-erosion and equilibrium curve, or they plot exactly on both curves. The history of a sample 482 is more difficult to reconstruct if the sample falls below the 'erosion-island' in the 'complex 483 field'. Samples within the complex field have experienced at least one period of burial, coverage, or significant (on the order of several 10 cm) erosion by spalling. The plot of ²⁶Al/¹⁰Be ratio 484

485	against ¹⁰ Be shows that 22 out of 26 samples fall within $\pm 1\sigma$ of the model curve for constant
486	exposure (Fig. 8). Four samples (20307_2, 120304_1, 120304_2, and 190303_1), with calculated
487	dates of 152.2±5.9 ka, 640.2±24.6 ka, 811.4±41.9 ka, and 989.1±42.9 ka, fall below the
488	equilibrium curve (Fig. 8, Table 3). Two samples (280207_1B and 20307_1A), with apparent
489	exposure dates of 62.7 ± 2.9 ka and 274.8 ± 12.2 ka, plot above the constant exposure curve in the
490	'forbidden field', which may be related to the chemical processing of the samples. When
491	including the model curve for erosion, all samples except 280207_1B, 20307_1A, 120304_1, and
492	190303_1 fall within $\pm 1\sigma$ of the 'erosion island' (Fig. 8, Table 3), and all samples fall within the
493	erosion island when we also consider the errors on both production rates (Stone, 2000). We infer
494	from these results that the majority of our samples record a single-stage, continuous exposure
495	history (Fig. 8). Samples 190303_1 and 120304_1 indicate a limited period of burial, because
496	both clearly plot below the 'erosion island'.
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499	
500	
501	

TABLE 3. ¹⁰BE AND ²⁶AL MEASUREMENTS ON THE WARM SPRINGS FAN, DEATH VALLEY 504 505

Sample Name	Fan unit	¹⁰ Be (10 ⁵ atoms g ⁻¹ qtz) ^a	¹⁰ Be date (ka) ^{a,b}	²⁶ Al (10 ⁵ atoms g ⁻¹ qtz) ^a	²⁶ Al date (ka) ^{a,b}	²⁶ Al/ ¹⁰ Be
10303 1	Og2	40.68 ± 1.4	923.9 ± 45.6 (72.5)	219.05±6.7	871.0 ± 49.2 (81.8)	5.39 ± 0.25
80303 2	Ög2	10.58 ± 0.5	$272.5 \pm 13.3 (19.4)$	63.9±2.9	255.2 ± 14.0 (19.7)	6.04 ± 0.38
80303 3	Õg2	23.82 ± 0.5	628.4 ± 16.8 (40.6)	135.81±4.3	595.4 ± 2.2 (48.4)	5.71 ± 0.22
80303 4	Og2	17.40 ± 0.6	447.8 ± 17.8 (30.5)	110.96 ± 0.5	468.1 ± 29.1 (40.5)	6.38 ± 0.36
90303 2	Og2	27.33 ± 1.0	$689.7 \pm 31.3(50.9)$	169.67±5.2	753.6 ± 38.1 (66.1)	6.21 ± 0.29
190303 1	Õg2	44.42 ± 1.3	989.1 ± 42.9 (75.5)	198.75±6.5	712.8 ± 38.3 (62.2)	4.48 ± 0.20
110304_6	Og2	31.47 ± 1.0	687.7 ± 27.1 (48.5)	192.33±9.2	733.5 ± 58.2 (74.8)	6.12 ± 0.35
120304 1	Qg2	25.79 ± 0.8	640.2 ± 24.6 (44.6)	131.69±4.8	527.2 ± 27.5 (43.2)	5.11 ± 0.24
120304 2	Qg2	35.68 ± 1.4	$811.4 \pm 41.9(63.6)$	180.19±9.3	677.2 ± 56.0 (69.7)	5.06 ± 0.32
120304 4	Qg2	34.11 ± 1.5	$770.5 \pm 46.3 (64.0)$	188.77±9.1	724.1 ±57.5 (73.8)	5.54 ± 0.37
130304 1	Qg2	29.33 ± 0.9	$772.0 \pm 30.6(55.7)$	167.41 ± 7.0	761.1 ± 52.4 (74.3)	5.71 ± 0.29
130304 3	Qg2	26.63 ± 0.8	$706.0 \pm 27.6(50.1)$	158.70 ± 6.8	$729.1 \pm 50.6(70.7)$	5.97 ± 0.31
150304 1	Qg3	13.38 ± 0.4	$368.8 \pm 13.1(24.0)$	83.75±4.1	$368.4 \pm 23.0(31.1)$	6.27 ± 0.36
150304 3	Qg3	10.78 ± 0.3	280.4 ± 9.8 (17.9)	67.56±2.9	274.3 ± 14.6 (21.0)	6.27 ± 0.33
170304 1	Qg3	7.80 ± 0.3	$181.6 \pm 8.6 (12.6)$	51.98±2.5	$185.5 \pm 10.3(14.1)$	6.67 ± 0.42
170304 2	Qg3	10.37 ± 0.3	$272.5 \pm 9.5(17.3)$	64.71±3.7	$265.1 \pm 18.3 (23.1)$	6.25 ± 0.40
180304 1	Qg3	7.20 ± 0.3	$180.4 \pm 8.2 (12.3)$	43.14±2.2	$164.0 \pm 9.7 (12.9)$	6.00 ± 0.40
180304 2	Qg4	7.71 ± 0.3	$192.2 \pm 8.7 (13.1)$	49.65±2.1	189.8 ± 9.4 (13.8)	6.44 ± 0.38
180305 5	Qg2	36.15 ± 1.1	924.5 ± 38.8 (69.4)	206.39±8.3	957.3 ± 74.0 (102.9)	5.72 ± 0.29
280207 1	Qg4	10.07 ± 0.5	$173.8 \pm 9.1 (12.5)$	63.67±2.9	$167.8 \pm 8.9 (12.3)$	6.33 ± 0.41
280207 IA	Qg4	3.85 ± 0.2	65.4 ± 3.4 (4.7)	27.59±2.3	$70.1 \pm 6.5 (7.1)$	7.18 ± 0.70
280207 ¹ B	Qg4	3.69 ± 0.2	62.7 ± 2.9 (4.2)	27.97±1.4	$71.1 \pm 4.0 (5.3)$	7.60 ± 0.50
280207_2	Qg4	6.07 ± 0.2	$120.1 \pm 4.7(7.7)$	38.78±1.7	$115.9 \pm 5.7 (8.2)$	6.39 ± 0.36
280207_3	Qg4	4.24 ± 0.2	$78.8 \pm 4.4 (5.8)$	26.52±1.4	73.7 ± 4.3 (5.6)	6.27 ± 0.46
20307_1	Qg4	5.97 ± 0.2	$149.2 \pm 5.8 (9.6)$	37.76±1.3	$142.7 \pm 5.8 (9.5)$	6.33 ± 0.32
20307_1A	Qg4	10.66 ± 0.4	274.8 ± 12.2 (18.9)	75.68±2.9	310.0 ± 15.2 (23.2)	7.11 ± 0.39
20307_1B	Qg4	5.02 ± 0.2	$124.5 \pm 4.9 \ (8.0)$	33.10±1.4	$124.0 \pm 6.1 \ (8.8)$	6.60 ± 0.37
$2030\overline{7}$ 2	Qg4	6.04 ± 0.2	152.2 ± 5.9 (9.8)	36.00±1.5	$136.6 \pm 6.4 \ (9.6)$	5.97 ± 0.33

^a Uncertainty includes 1σ AMS measurement errors, statistical counting error, uncertainty due to the normalization of the standards and the blank, and for ²⁶Al the ICP-OES 506 507 measurement uncertainty.

508

^b Exposure ages were calculated with the CRONUS-Earth exposure age calculator (<u>http://hess.ess.washington.edu/math/</u>, Balco et al., 2008 and update from v. 2.1. to v. 2.2) and time-dependent Lal (1991)/Stone (2000) scaling model using the production rate calibration set by Heyman (2014). A ¹⁰Be half-life of 1.39 x 10⁶ years (Korschinek et al., 2010; 509

Chmeleff et al., 2010) and a ²⁶Al half-life of 7.08 x 10⁵ years (Nishiizumi, 2004) were used for the age calculation. Uncertainties in parentheses include propagated production rate 510

511 scaling uncertainty.

512 **6.** Discussion

⁵¹³ 6.1. Surface unit chronology on the Warm Springs fan

514 The results from measured ¹⁰Be and ²⁶Al concentrations show that surfaces of the Warm 515 Springs fan record a depositional history that extends back into the early Pleistocene. Based 516 on the observation that most samples fall within $\pm 1\sigma$ of the 'erosion island' (Fig. 8, Table 3), 517 we infer that the fan surfaces mostly underwent a simple exposure history. This implies that 518 the early and middle Pleistocene surfaces have been relatively stable since they were 519 abandoned and that they have been affected only by slow erosion. If we assume that our 520 oldest sample (190303 1) with the highest measured ¹⁰Be concentrations is in secular 521 equilibrium, the maximum erosion rate that we could infer from the ${}^{26}Al/{}^{10}Be$ ratio is 522 0.6 ± 0.03 m Ma⁻¹. We have, however, no constraints to determine whether our samples are in 523 secular equilibrium or not. Therefore, we interpret our apparent exposure dates conservatively 524 as minimum ages as even small amounts of erosion will shift the results towards older ages.

525

Our interpretation of apparent surface exposure dates as minimum ages for surface 526 abandonment must also account for the possible influence of inheritance (e.g., Anderson et 527 al., 1996; Repka et al., 1997; Hancock et al., 1999; Wolkowinsky and Granger, 2004; 528 Matmon et al., 2005). The results from the active channel samples show that fan surface 529 samples on the Warm Springs fan contain a significant inherited radionuclide component, as 530 previously documented for other localities in Death Valley (Machette et al., 2008; Owen et 531 532 al., 2011; Frankel et al., 2015). Inheritance in the Warm Springs fan could be derived from 533 three different potential sources. First, depending on the residence time of sediment in the headwaters, the fan catchment could contribute a significant amount of inheritance. The upper 534 Warm Springs catchment contains a low-relief area that has likely acted as temporary 535 sediment storage (Fig. 1). Second, inheritance could be derived from exposure during 536 537 transport in the active channel. Third, incision and erosion of older fan surface units could contribute high-inheritance clasts to sediments in the active channel. This last mechanism
might be particularly important where the channel cuts through old fan surface deposits with
high radionuclide concentrations at the surface.

541

542 Measurements of ¹⁰Be in the active channel show that, while radionuclide concentrations 543 in replicate samples vary by a factor of 2 to 3, the concentrations generally increase with 544 increasing distance from the catchment outlet (Fig. 7). Substantial concentrations of 545 cosmogenic ¹⁰Be and ²⁶Al measured at the catchment outlet suggest that pre-exposure of the 546 sediment in the catchment is a likely source for inheritance. The downstream increase in 547 concentration along the active channel could be due either to exposure during transport, or to 548 progressive incorporation of old, high-concentration sediment derived from erosion of the fan 549 surface units. Comparison of cosmogenic concentrations in samples from the fan surfaces 550 with samples from the active channel at 3 and 4.5 km distance from the catchment outlet, 551 however, show that channel samples contain only 10-20% of the concentrations measured in 552 the fan surface samples. While the incision of the debris-flow fan indicates that channel wall 553 clasts must have contributed to the channel sediments, the nuclide signal suggests that the 554 contribution of old clasts from the fan surfaces into modern alluvium does not significantly 555 increase the amount of inheritance. One reason could simply be that most channel wall clasts 556 are shielded below the depth of the cosmogenic production zone, and therefore these clasts do 557 not carry a substantial cosmogenic signal when they enter the channel floor. Another reason 558 could be that clasts derived from the channel incision have already largely been removed from 559 the active channel and have been incorporated into the deposits of unit Qg3. This would 560 imply that most clasts in the modern channel are derived directly from the catchment area. 561 Based on our data, we cannot explicitly determine the source of inheritance, but most likely 562 all three components (pre-exposure in the catchment, exposure during transport, and input 563 from the channel walls) contribute to the total inherited signal.

564 We can use the measured inherited concentrations from the channel samples to define 565 boundaries for the timing of fan surface activity. Correcting the dates on fan unit Qg2 with the 566 minimum inheritance value of 63 ± 3 ka (280207 1B) measured at the catchment outlet would 567 lower the measured exposure dates by less than 10% for samples that are 600 ka or older. For 568 the maximum inheritance correction of 174 ± 9 ka (280207 1) also measured at the catchment 569 outlet, the calculated dates would decrease by 20 to 30%. Thus, the measured exposure dates 570 are likely to be close to the true depositional ages on Qg2, as long as the amount of 571 inheritance measured at the catchment outlet has not varied significantly through time.

572

573 On unit Og3, we can approximate depositional ages based on the mean inheritance of our five 574 distal samples 180304 2, 20307 1, 20307 1A, 20307 1B, and 20307 2. This yields a 575 correction of 178 ± 59 ka, which shifts measured fan surface dates from unit Qg3 to 576 significantly younger dates. In comparison, the mean surface date of unit Qg3 yields 256+78 577 ka. After the subtraction of mean inheritance from each individual Qg3 sample, four samples 578 postdate MIS 6 (101 ka, 93 ka, 3 ka, and 1 ka) and one sample has an age of 190 ka. Two 579 samples dated on unit Qg3 using varnish microlaminations as dating technique yield 60 ka 580 (Liu and Broecker, 2008). These results fit the expectation that unit Qg3 postdates the MIS 6 581 paleoshoreline on the Warm Springs fan. The uncertainty in the appropriate inheritance 582 correction, and the large inherited concentrations relative to measured sample concentrations, 583 mean that the exact timing of fan activity cannot be determined, and therefore at present we 584 can only limit the period of fan activity during Qg3 deposition to the late Pleistocene and 585 Holocene. For unit Qg4, these results imply that fan activity occurred throughout the 586 Holocene.

587

⁵⁸⁸ 6.2 Comparison with previous surface chronologies on the Warm Springs fan

⁵⁸⁹ Our surface mapping largely agrees with the relative chronology published by Hunt and

590 Mabey (1966). With a few differences on the distal fan, we placed the boundary of unit Qg2 591 on the proximal and central part of the Warm Springs fan very near the boundary defined by 592 Hunt and Mabey (1966). Our unit Qg3, however, covers significantly less surface area 593 compared to their interpretation. The major difference is the large area at the fan toe, which 594 we map as unit Qg4 while Hunt and Mabey (1966) identified several individual Qg3 "islands" 595 separated by unit Qg4. This difference in the separation of individual fan units results from 596 our analysis of surface roughness using the Lidar dataset, which shows lower long-597 wavelength relief on Qg4 compared to Qg3.

598

599 The suggested age boundaries for Hunt and Mabey's stratigraphic units Qg2, Qg3 and Qg4 600 range from the late Pleistocene (10-130 ka; Qg2) to the early Holocene (Qg3) and the late 601 Holocene (Qg4) (Klinger, 2001; Machette and Crone, 2001; Knott et al., 2002; Machette et 602 al., 2008; Owen et al., 2011; Frankel et al., 2015). Given that on Warm Springs fan units Qg3 603 and Qg4 the uncertainties due to inheritance and measurement errors are too high to allow a 604 reliable correction and comparison with the proposed stratigraphic ages, we can only use our 605 measured surface dates on unit Qg2 to roughly correlate these dates with the stratigraphic 606 ages. For unit Qg2, the majority of dates, when we include a correction for inheritance based 607 on the measurements in the active channel, overlaps with early to middle Pleistocene time. 608 This implies that surfaces that we mapped as unit Qg2 are much older compared to the 609 proposed stratigraphic late Pleistocene age of Qg2 (Knott et al., 2002). In fact, the dates on 610 unit Qg2 agree with the suggested timing of the older stratigraphic unit Q1 (Knott et al., 611 2005). Certainly, given that unit Qg2 must be younger than MIS 6 (Knott et al., 2002), the 612 time constraints measured on our mapped unit Qg2 would fit the stratigraphic time boundaries 613 for unit Q1 much better (Knott et al., 2005; Machete et al., 2008). However, the morphologic 614 description of surface unit Q1 does not entirely fit our field observations. While the lack of 615 bar-and-swale morphology and the high degree of surface dissection on our mapped unit Qg2

616 is also characteristic for the morphologic description of unit Q1, rock varnish and desert 617 pavement are well-developed on the oldest Warm Springs fan surfaces but are described to be 618 only weakly preserved on the Q1 surfaces (Hunt and Mabey, 1966; Knott et al., 2005). When 619 we ignore the morphologic descriptions of the fan units, the existing time constraints on unit 620 Qg2 elsewhere in the region (Knott et al., 2002) and the numeric data we obtained from our 621 mapped Qg2 unit appear to suggest that Warm Springs fan unit Qg2 would correctly be 622 labeled as the stratigraphically older unit Q1. This is turn would imply that all stratigraphic 623 units on the Warm Springs fan would have to be revised.

624

625 6.3 Evolution of fan morphology and depositional pattern on the Warm Springs fan

626 Overall, the numeric dates on our oldest mapped fan surface unit Qg2 are consistent with 627 deposition over a long period of more than 4×10^5 years until fan activity shifted basinward to 628 form a new debris-flow fan unit. Assuming that our measured cosmogenic dates are a 629 reasonable proxy for an approximate time window of fan surface activity, the spatial 630 distribution of these dates across Qg2 shows that fan activity probably shifted laterally over 631 time as one would expect for a debris-flow dominated fan (e.g. Suwa and Okua, 1983; 632 Dühnforth et al., 2007, Ventra and Nichols, 2014; D'Arcy et al., 2015; Schürch et al., 2016). 633 The uncertainty on the cosmogenic measurements and the unknown inheritance at each 634 sampling location, however, do not allow us to interpret our results as true depositional ages. 635 In addition, the lack of distinct morphologic differences on the oldest surface prevents any 636 more detailed identification of individual debris-flow lobes or units as typically found on 637 younger debris-flow fans, where lateral shifts in deposition and cross-cutting relationships 638 allow temporal reconstruction of debris-flow activity (e.g. Dühnforth et al., 2007; Schürch et 639 al., 2016). Clearly, evidence for radial shifts in the locus of fan deposition (from unit Qg2 to 640 Qg3 to Qg4), as also seen at Illgraben (Schürch et al., 2016) and in debris-flow fan 641 experiments by de Haas et al. (2016), are preserved on the Warm Springs fan. The radial shift 642 in deposition results from the incision of the Warm Springs fan head. The incision represents 643 a major change in the geologic history of the fan as the sedimentary system shifted from 644 aggradation to entrenchment and the depocenter shifted out into the basin leading to the 645 preservation of the older depositional surfaces (e.g. Hooke, 1967; Blair, 1999a, 1999b; 646 Harvey, 2005; Dühnforth et al., 2007). The driver for fan head incision in this setting has been 647 the subject of much debate, and unfortunately our age constraints are not sufficiently precise 648 to discriminate between different potential models. According to Denny (1965) and Hooke 649 (1972), the eastward tilting of the valley floor in the hanging wall of the BMFZ led to base 650 level lowering with the subsequent entrenchment of the fan heads on the western side of 651 Death Valley. It may also be possible that drawdown of the MIS 6 lake level highstand drove 652 incision of the Warm Springs fan and of other debris-flow fans along the west side of Death 653 Valley, as has been proposed for other alluvial fans along the shoreline of Lakes Lahontan 654 and Mojave in the western United States (Harvey et al., 1999; Ritter et al., 2000; Harvey, 655 2005). Following incision of the Warm Springs fan head, the depocenter during late 656 Pleistocene und Holocene time has been located near the fan toe leading to surface unit 657 deposition and reworking. It is clear that the reoccupation of the oldest proximal and central 658 surfaces on the Warm Springs fan, and resurfacing across the proximal fan, will only occur if 659 there is substantial aggradation in the incised channel and backstepping of the depocenter 660 (e.g., de Haas et al., 2016; Schürch et al., 2016).

661

Based on field observations and analysis of the Lidar dataset, we find that the longwavelength relief of the fan surfaces increases from the youngest to the oldest unit while the short-wavelength relief decreases. For example, while the youngest unit Qg4 shows clear barand-swale morphology, the next oldest unit Qg3 lacks well-preserved channel-levee morphology, and surfaces instead are relatively planar with incipient development of desert pavement and rock varnish. We interpret these observations as an evidence for progressive ⁶⁶⁸ surface smoothing by aeolian inflation and diffusive processes and surface dissection, which
⁶⁶⁹ have likely varied in effect and intensity through time (Wells et al., 1987; Bull, 1991; Ritter et
⁶⁷⁰ al., 1993; Matmon et al., 2006; Frankel and Dolan, 2007; Regmi et al., 2014).

671 Existing numerical constraints on the time scale of surface smoothing of primary debris-flow 672 morphology on an alluvial fan suggest that it takes more than 280 ka on alluvial fans in the 673 Mojave Desert (Matmon et al., 2006) and about 70 ka on fans in northern Death Valley 674 (Frankel and Dolan, 2007). When we compare surface relief on the Warm Springs fan units 675 Qg4 to Qg2, we find that debris-flow morphology is completely eliminated within a few 10^5 676 years. Due to inheritance we are not able to resolve a more precise time scale it takes for 677 smoothing. Even though our results are comparable to Matmon et al. (2006), we argue that 678 local difference in the climatic situation or other local factors, for example differences in the 679 grain size distribution, could be responsible for the difference in the timescale of smoothing.

680

⁶⁸¹ **7. Conclusions**

682 Dating of debris-flow surfaces on the Warm Springs fan in Death Valley, California, using 683 cosmogenic ¹⁰Be and ²⁶Al shows that this fan preserves one of the longest chronologies in the 684 southwestern United States. The samples yield early to late Pleistocene dates ranging between 685 990-630 ka for the oldest mapped fan surface unit (Qg2) and 370-180 ka for the younger fan 686 surface unit Og3, excluding a correction for inheritance. The amount of inheritance, measured 687 in the active channel, increases from ~65 ka at the catchment outlet to up to 275 ka at the 688 distal fan, which implies that our measured ¹⁰Be and ²⁶Al dates cannot be interpreted as true 689 depositional ages.

While the youngest, presently active fan surface unit Qg4 shows clear bar-and-swale morphology, the older surfaces Qg2 and Qg3 have undergone post-depositional, shortwavelength smoothing of original, primary debris-flow depositional forms. At the same time, older fan surfaces are progressively more dissected by a tributary channel network that leaves 694 the oldest surfaces increasingly isolated and protected from younger deposition and 695 resurfacing. We propose that debris-flow fan surfaces on the west side of Death Valley are 696 undergoing continuous surface smoothing, albeit at ever-decreasing rates as the surface 697 becomes smoother through time, while dissection of fan surfaces occurs. As such the Warm 698 Springs fan represents a good example for existing models of surface modification, in which 699 surfaces initially become smoother through time while dissection of the surfaces becomes 700 deeper with increasing age. This system of surface modification through time can only be 701 reset by aggradation in the active channel and backstepping of the active depocenter onto the 702 formerly abandoned part of the fan, such that renewed debris-flow deposition can occur. Our 703 results highlight the complications and limitations inherent in extracting information on long-704 term fan evolution and palaeo-flow behavior from old debris-flow fans, especially in cases 705 where the primary depositional morphology is eliminated and estimation of the timing of fan 706 deposition is complicated by high uncertainties and high inheritance.

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708 **8. Acknowledgement**

This project was funded by the Swiss National Science Foundation (grants 2100-067624 and 200020-105225/1). We would like to thank B. Bookhagen, and C. Matter for field assistance. We greatly appreciate the support of Death Valley National Park service for permitting our research activities on the Warm Springs fan. We also thank Jeffrey Knott, Martin Stokes, two anonymous reviewers, and the editor Scott Lecce for constructive reviews that helped to improve and focus the manuscript.

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720 9. References

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1030 **10. Figure Captions**

1031 Figure 1:

Shaded-relief map of the study location showing the Panamint Range and Death Valley,
California. Topographic data are based on USGS 10 m National Elevation Data. The white
box indicates the location of Fig. 2 and 4. Note the difference in surface texture on the Warm
Springs fan compared to all other fans on the west side of Death Valley. BMFZ, Black
Mountain Fault Zone.

1037

1038 Figure 2:

1039 Shaded-relief map of the Warm Springs fan showing the textural fan surface characteristics. Topography is derived from Airborne Laser Swath Mapping (ALSM) topographic data with 1040 resolution of 1 m/pixel. Solid lines labeled with letters indicate the location of the topographic 1041 1042 profiles shown in Fig. 5. The original boundaries of fan units mapped by Hunt and Maybe 1043 (1966) are indicated by labels Qg2, Qg3, and Qg4. Dashed lines delineate fault offsets on the 1044 fan (for clarity only selected offsets are shown); dotted line indicates the 65 m a.s.l. contour 1045 line, where a paleoshoreline cut into fan deposits on the Warm Springs fan. X-Y coordinates are in UTM projection (zone 11). 1046

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1048 Figure 3:

Field photos of fan surfaces on the Warm Springs fan. A, overview over the fan surfaces of unit Qg2 showing the planar surfaces with darkly varnished desert pavement. View is to the north/northeast. B, fan surface Qg2 at sampling location 10303_1 (~925 ka). View is to the south/southeast. C, example for well-developed desert pavement with darkly varnished clasts at sampling location 10303_1. D, Surface characteristics in the active channel (Qg4) with well-preserved channel-levee morphology. View is to the west/northwest.

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1057 Figure 4:

Relative and absolute chronology of fan units on the Warm Springs fan. Colored fan surfaces represent the relative chronology based on geomorphological criteria such as surface dissection and pavement development. Note that the boundary of our Qg2 is similar to Hunt and Maybe's original map (1966) while unit Qg3 differs significantly from the original map (see Fig. 2 for comparison). Red dots show sample locations for cosmogenic ¹⁰Be surface exposure dating. See Tables 2 and 3 for sample details. X-Y coordinates are in UTM projection (zone 11).

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1066 Figure 5:

Topographic profiles across the Warm Springs fan surfaces showing the long-wavelength
relief of the fan surfaces. Note the differences in incision depths on fan units Qg4 to Qg2.
Topographic data were extracted from 1 m-resolution Airborne Laser Swath Mapping data.
For the profile locations see Fig. 2.

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1072 Figure 6:

Box-and-whisker plots of ¹⁰Be ages plotted for surface units Qg2, Qg3, and Qg4. The box encloses the area between the first and third quartiles and the horizontal line represents the median. Whiskers show one standard deviation. Sample point outside box Qg2 lies one standard deviation outside from the mean.

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1078 Figure 7:

1079 Plot of ¹⁰Be concentrations from incised channel samples vs distance from the catchment 1080 outlet. The minimum inherited component at the catchment outlet makes about 10-20% of the 1081 measured ages on surface unit Qg2, which is within the error of our ages. The significantly higher inherited concentrations at the distal fan overlap or even exceed measured concentrations on the fan unit Qg3. This implies that we cannot provide reliable estimates on the timing fan surface abandonment. Note the plotted concentration from surface samples for comparison.

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1087 Figure 8:

¹⁰⁸⁸ Plot of 26 Al/ 10 Be vs 10 Be (see explanation of plot in text). All measured sample concentrations ¹⁰⁸⁹ are normalized to sea level after Stone (2000).

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1	0	9	1	
1	0	9	2	
1	0	9	3	
1	0	9	4	
1	0	9	5	
1	0	9	6	
1	0	9	7	
1	0	9	8	
1	0	9	9	
1	1	0	0	
1	1	0	1	
1	1	0	2	
1	1	0	3	
1	1	0	4	
1	1	0	5	
1	1	0	6	
1	1	0	7	
1	1	0	8	
1	1	0	9	
1	1	1	0	
1	1	1	1	
1	1	1	2	
1	1	1	3	
1	1	1	4	
1	1	1	5	
1	1	1	6	
1	1	1	7	
1	1	1	8	
1	1	1	9	
1	1	2	0	
1	1	2	1	
1	1	2	2	
1	1	2	3	
1	1	2	4	
1	1	2	5	





Figure 2





Figure 3



Figure 4



Figure 5



Figure 6









Figure 8

