1	Rapid increase in southern elephant seal genetic diversity after a founder event
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3	Mark de Bruyn <sup>1</sup> , Malin Pinsky <sup>2</sup> , Brenda Hall <sup>3</sup> , Paul Koch <sup>4</sup> , Carlo Baroni <sup>5</sup> , A. Rus Hoelzel <sup>6</sup>
4	1. Molecular Ecology and Fisheries Genetics Laboratory, School of Biological Sciences,
5	Bangor University, Deiniol Road, Bangor, UK
6	2. Department of Ecology, Evolution, and Natural Resources, Rutgers University, New
7	Brunswick, New Jersey, USA
8	3. Climate Change Institute and School of Earth and Climate Sciences, University of Maine,
9	Orono, Maine, USA
10	4. Department of Earth and Planetary Sciences, University of California, Santa Cruz,
11	California, USA
12	5. Dipartimento di Scienze della Terra, Università di Pisa, Pisa, Italy
13	6. School of Biological and Biomedical Sciences, University of Durham, Durham, UK
14	
15	Corresponding authors: Mark de Bruyn, Molecular Ecology and Fisheries Genetics Laboratory,
16	School of Biological Sciences, Bangor University, Deiniol Road, Bangor LL57 2UW, UK Fax:
17	+44 1248 370731, email: markus.debruyn@gmail.com
18	A. Rus Hoelzel, School of Biological and Biomedical Sciences, University of Durham, Durham
19	DH1 3LE, UK. Fax: +44 191 3341201, email: a.r.hoelzel@durham.ac.uk
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#### 21 Summary

22 Genetic diversity provides the raw material for populations to respond to changing 23 environmental conditions. The evolution of diversity within populations is based on the 24 accumulation of mutations and their retention or loss through selection and genetic drift, while 25 migration can also introduce new variation. However, the extent to which population growth and 26 sustained large population size can lead to rapid and significant increases in diversity has not 27 been widely investigated. Here we assess this empirically by applying Approximate Bayesian 28 Computation to a novel ancient DNA dataset that spans the life of a southern elephant seal 29 (*Mirounga leonina*) population, from initial founding ~7000 years ago, to eventual extinction 30 within the past millenium. We find that rapid population growth and sustained large population 31 size can explain substantial increases in population genetic diversity over a period of several 32 hundred generations, subsequently lost when the population went to extinction. Results suggest 33 that the impact of diversity introduced through migration was relatively minor. We thus 34 demonstrate, by examining genetic diversity across the life of a population, that environmental 35 change could generate the raw material for adaptive evolution over a very short evolutionary 36 time scale through rapid establishment of a large, stable population.

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Key words (3-6): ancient DNA, approximate Bayesian computation, extinction, founder event,
genetic diversity, population growth.

#### 40 Introduction

41 Understanding how micro-evolutionary processes lead to changes in evolutionary 42 potential in the form of genetic diversity is a fundamental goal of population genetics. Population 43 genetic diversity can be gained either *in situ* through mutation, through immigration, or some 44 combination of the two (Maruyama 1970, Nei & Feldman 1972). Variation in the form of new 45 mutations is expected to enter a population at a rate of  $2 N_e u$ , where  $N_e$  is the effective 46 population size and *u* the neutral mutation rate per generation, while genetic drift decays 47 variation at the rate of  $1/2 N_e$  (Kimura 1983). New mutations within a population can either be 48 fixed, or more commonly lost (Kimura and Ohta 1969, Crow and Kimura 1970). Fixation time of 49 new unlinked neutral mutations within a randomly mating population is strongly reliant on  $N_e$ , 50 taking some  $4N_e$  generations when  $N_e$  is constant (Kimura and Ohta 1969, 1973). However, 51 change in population size, a fundamental characteristic of many wild populations, can 52 significantly impact on this timeframe (Kimura and Ohta 1969, 1973). Changes in genetic 53 diversity can also result from the movement of individuals (migration) and their genes (gene 54 flow) into or out of a population. Thus, rapid changes in population genetic diversity may result 55 from demographic processes such as immigration, or *in situ* processes such as mutation and 56 genetic drift. These processes will also be influenced by changes in  $N_e$ , which may result from 57 varying physical conditions, resource availability, habitat availability, predator density, human 58 impacts (Otto & Whitlock 1997), or skewed effective breeding ratios. 59 While theoretical models predict changes in levels of genetic diversity within a

population over time, it is uncommon to be able to track such changes in vertebrate populations
 empirically, particularly over timeframes longer than that of pedigree studies. However,
 molecular analyses of ancient DNA allow these processes to be assessed more fully, especially

63	when ancient DNA samples span the life of the population. Here we investigate such a dataset,
64	provided by the founding and subsequent extinction of a colony of southern elephant seals. This
65	population existed along the Victoria Land Coast (VLC), Antarctica for much of the Holocene,
66	resulting from the expansion of suitable habitat due to altered environmental conditions (Hall et
67	al. 2006, de Bruyn et al. 2009). Our study is inspired by the observation that this short-lived
68	population had apparently accumulated extensive novel variation, most of which was lost when
69	the population went to extinction (de Bruyn et al. 2009), but the relative contribution of
70	migration and <i>in situ</i> mutation to this increase in genetic diversity remains unknown.
71	Radiocarbon dates from 223 individual seal remains provided upper- and lower-bound
72	age estimates on the population of $7,200 - 270$ years before present (YBP; Hall et al. 2006, de
73	Bruyn et al. 2009). Subsequent analyses utilizing a southern elephant seal-specific substitution
74	rate estimated from these dated samples supported the notion of an accelerated short-term time-
75	dependent rate (Ho et al. 2005), and agreed with estimates obtained from various other aDNA
76	datasets (e.g. Ho et al. 2007, Burridge et al. 2008), and the mean human mitochondrial control
77	region pedigree rate estimate (Howell et al. 2003). Importantly, population genetic analyses of
78	VLC founder dynamics conducted using this substitution rate fitted well the timeframe recovered
79	from geological dating of Holocene raised beaches along the VLC, and the fate of the Ross Sea
80	Ice Sheet, which indicated that this coastline would have been inhospitable to southern elephant
81	seals prior to around 8,000 YBP (Conway et al. 1999, Hall et al. 2006). Elephant seals require
82	breeding habitat with access to open water, and the VLC coastline was beneath the Ross Sea Ice
83	Sheet prior to this time (Hall et al. 2006). The retreat of the ice shelf was followed by a period of
84	relative warmth when the seal colony apparently thrived and expanded, before more recent

cooling around 1,000 YBP lead to an increase in land-fast sea-ice - a likely cause of the demise
and eventual extinction of this population (Hall et al. 2006, de Bruyn et al. 2009).

87 Molecular analyses of these ancient DNA samples from the VLC, and from 88 contemporary samples from seven putative populations, representing all four major extant 89 southern elephant seal breeding stocks (Leboeuf & Laws 1994), have provided considerable 90 insight into the origin and demographic history of the VLC seals (de Bruyn et al. 2009). Our 91 earlier research demonstrated that the VLC seals were an independent breeding colony - most 92 likely founded by seals from Macquarie Island (MQ) - which grew in size through the mid-93 Holocene, before declining drastically around 1,000 YBP, resulting in eventual extinction and 94 the possible return of some VLC animals to the source population at MQ (de Bruyn et al. 2009). 95 Inferred timings of these demographic events derived from molecular data are consistent with 96 inferences based on fossil radiocarbon dates and geological data. However, it was not clear if this 97 scenario was sufficient to explain the very high levels of diversity of the VLC population, or the 98 relative importance of migration compared to *in situ* mutation.

99 We therefore applied Approximate Bayesian Computation (ABC) methods to the VLC 100 and MQ dataset to test hypotheses about the timing and relative importance of different 101 processes, thereby testing hypotheses about life history and behavior (such as the likelihood of 102 the VLC being a breeding, as opposed to a molting population, and the degree of insularity of the 103 newly founded population). ABC enables the joint assessment of a large number of increasingly 104 complex demographic models, and allowed us to examine in detail processes leading to 105 significant changes in genetic diversity, from initial founding to the eventual extinction of a 106 population.

#### 108 Materials and Methods

109 Data

110 The VLC and MQ dataset analyzed here comprised 223 ancient VLC samples, ranging in 111 age from  $\pm 7,200 - 270$  YBP, and 49 contemporary samples from MQ (Table S1; Genbank 112 accession numbers: FJ168073–FJ168343). Details of laboratory DNA methodology (including 113 extractions, PCR, sequencing of a 325 bp fragment of the HVR1 section of the mitochondrial 114 control region, cloning, 'independent replication' and other strict aDNA protocols), as well as 115 calculation of aDNA calendar dates from radiocarbon dating that incorporated a time-dependent 116 Southern Ocean marine reservoir effect, are presented elsewhere (de Bruyn et al. 2009). 117 118 Coalescent models and summary statistics 119 Based on earlier population genetic results, including Isolation-with-Migration analyses 120 (de Bruyn et al. 2009), a series of increasingly complex models were examined within an ABC 121 framework. These ranged from a simplified one population (Model 1) or model with two 122 populations that diverged 12,000 ybp (Model 2), to those incorporating exponential population 123 growth and bi-directional migration (e.g. Model 7) (Fig. 1). Models of intermediate complexity 124 examined colonization from VLC to MQ (Model 3, the opposite of what we believe occurred), 125 colonization from MQ to VLC (Model 4), ongoing migration from MQ to VLC after 126 colonization (Model 5), and a single migration event from VLC to MQ when the VLC 127 population was extirpated (Model 6). We also added alternative versions of Models 4-7 that 128 included exponential growth (consistent with known demographic potential) in VLC for the first 129 500 years after colonization, followed by a stable population (Models 4b-7b). For models with 130 colonization, we specified that the timing of colonization was 8000 ybp, consistent with the

131	retreat of ice from VLC and the first available breeding habitat. This was a conservative choice
132	given that a more recent colonization would require more rapid population growth.
133	We used serial coalescent theory, as implemented in Bayesian Serial Simcoal (Anderson
134	et al. 2005), to simulate 1 million datasets from each of our eleven models. We set the mutation
135	rate to 9.8x10 <sup>-7</sup> mutations site <sup>-1</sup> yr <sup>-1</sup> , as estimated from dated aDNA samples, with a
136	transition/transversion ratio of 0.8564 and a mutation rate gamma distribution shape parameter of
137	0.2003 (de Bruyn et al. 2009). The sample sizes, sample dates, and length of the locus (325bp)
138	matched our collected data (Table S1). We used uniform priors on population size and migration
139	rates (Tables 1 & S2) so as to specify little prior knowledge about these quantities. Uniform
140	priors for population size and migration rates have been commonly recommended or used in
141	ABC (Bertorelle et al. 2010, Li & Jakobsson 2012, Robinson et al. 2013). The prior on back-
142	migration in Models 6 and 7 was also uniform 0-100% (Table S2). For both on-going migration
143	from MQ to VLC and the back-migration event from VLC to MQ, 100% migration corresponded
144	to full replacement of the sink population with genes from the source population, while 50%
145	migration corresponded to replacement of only half the sink population. In addition, we
146	evaluated the sensitivity of our conclusions to alternative priors by generating 333,333
147	simulations from each of the 11 models with 1) log uniform migration rates from $e^{-9}$ to $e^{0}$ ; or
148	with 2) wider priors on the size of the colonizing population (uniform 0 to 10,000). Colonizing
149	populations up to 10,000 in a single generation are highly unlikely given that MQ only has
150	10,000 females.
151	From the real data and from the simulations, we then calculated a series of summary
152	statistics on each of three statistical groups: MQ ( <i>n</i> =49), VLC 3000-7100 ybp ( <i>n</i> =64), and VLC

153 0-3000 ybp (*n*=159). Statistical groups were chosen to examine temporal dynamics in VLC while

154 maintaining a sufficiently large sample size for the ancient group. However, we also conducted a 155 sensitivity analysis after adding an additional statistical group for the earliest VLC years (6000-156 7100 ybp) and trimming the other ancient VLC statistical group to 3000-6000 ybp. The 157 sensitivity analysis was conducted with 333,333 simulations from each of our 11 models. 158 We used the number of haplotypes (H), number of segregating sites (S), and average 159 number of pairwise differences within each of the statistical groups ( $\pi$ ) as summary statistics 160 (Tables 2 and S3). We also used F<sub>ST</sub> values calculated between MQ and each of the two VLC 161 statistical groups. We chose these summary statistics because they reflect the demographic 162 changes in which we were interested (Bertorelle et al. 2010), namely changes in population size 163 (H, S and  $\pi$  at two time points) (Nei & Takahta 1993, Tajima 1989) and the exchange of 164 migrants through time between MQ and VLC (F<sub>STS</sub>) (Slatkin 1985) (Table S3). Preliminary 165 simulations also suggested that they were sufficient to distinguish among our competing 166 hypotheses.

167

### 168 Approximate Bayesian Computation (ABC)

169 After simulating the models and calculating the summary statistics, we conducted our 170 ABC analysis in three steps: model selection, parameter estimation, and quality control. For 171 model selection, we used a weighted multinomial logistic regression to calculate the relative 172 support for each model (Beaumont 2008) from the 5% of all 11 million simulations for which  $\delta$ 173 (the Euclidean distance between the observed and simulated summary statistics) was smallest. 174 Parameter estimation followed a similar process. From the 0.5% of simulations from each model 175 for which  $\delta$  was smallest, we used weighted local linear regression to estimate the posterior 176 distributions of each parameter (Beaumont et al. 2002). Before the parameter calculations, we

log-transformed MQ and final VLC population sizes, and logit-transformed migration rates and
the size of the VLC colonizing population. Both transformations were applied to facilitate
parameter estimation and projection back onto untransformed axes. We report the median, mode,
and 95% highest probability density (HPD) for each parameter. Calculations were done in R

181 2.15.3 with the abc package (Csillery et al. 2012).

182 For quality control, we followed the recommendations of Bertorelle et al. (2010) and 183 generated pseudo-observed datasets (PODs). We sampled 1000 parameter sets from the posterior 184 distributions of each model and used these to simulate PODs. By sampling from the posteriors, 185 we account for our full knowledge and uncertainty about the true model parameters. Previous 186 authors have used only the best parameter estimates (Bertorelle et al. 2010) or the full prior 187 distributions (Fagundes et al. 2007). We re-ran our model selection calculation on each of 100 188 PODs from each model (1100 total PODs) to evaluate our model selection procedure when the 189 true model was known. This process allowed us to calculate model choice error rates. In addition, 190 these results also allowed us to calculate a *corrected* relative support for each model following 191 the procedure of Fagundes et al. (2007)

192

193 
$$\Pr(Model = X \mid P_{Model A} = a) = \frac{\Pr(P_{Model A} = a \mid Model = X)}{\sum_{i \in Models} \Pr(P_{Model A} = a \mid Model = i)}$$

194

for each potential model X, the originally identified best model A, and the originally observed relative support *a* for Model A. The procedure calculates the probability that X is the correct model given our initial (uncorrected) observation that Model A was chosen as the best model with relative support *a*. The procedure accounts for difficulties differentiating between the models while also accounting for the fact that Model A might have been chosen with strong 200 support (a close to 1). We implemented the procedure by calculating the relative support for 201 Model A from each of the 100 PODs produced by each of our eleven models. For each model, 202 we then calculated a probability density across the range of relative supports (0 to 1). We 203 expected that models that were rarely confused with Model A would have densities centered near 204 zero relative support, while models similar to Model A (including Model A itself) would have 205 densities centered closer to 1. We then evaluated these densities at the observed value (a) of 206 relative support for Model A (e.g.,  $Pr(P_{Model A} = a | Model = X)$  to evaluate the above equation. 207 We also assessed model fit against our observed data using tail-area probability, or p-208 value, tests for each summary statistic (Cornuet et al. 2010). We used false discovery rate (FDR) 209 corrections on the statistics within each model (Benjamini & Hochberg 1995), as well as the 210 method of Ghirotto et al. (2010) to combine *p*-values across statistics. When combining *p*-values, 211 we assigned 0.001 to those *p*-values at the edge of the empirical distribution in order to avoid 212 infinite values in the calculation. Low *p*-values indicated that the observed data were unlikely 213 given a particular model. Finally, we performed a principal components analysis (PCA) on the 214 summary statistics and plotted the observed data along with simulated datasets from the prior and 215 from the posterior (Cornuet et al. 2010). Ideally, the observed data would fall within the posterior 216 for the selected model.

To test the accuracy of our parameter estimation procedure, we estimated model parameter ( $\hat{\theta}_i$ ) for each POD *i* and compared the estimates against the known, true parameters ( $\theta_i$ ) used to simulate each POD. We then calculated the bias and root mean square error (RMSE) for each parameter across all *n* PODs:

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222 
$$bias = \frac{1}{n} \sum_{i=1}^{n} \frac{\widehat{\theta}_i - \theta_i}{\theta_i}$$

224 
$$RMSE = \frac{1}{\theta} \sqrt{\frac{1}{n} \sum_{i=1}^{n} (\hat{\theta}_i - \theta_i)^2}$$

226	We also calculated the Factor 2 statistic, which is the proportion of estimated values that
227	fall between 50% and 200% of the true value (Neuenschwander et al. 2008). We used coverage
228	to assess the accuracy of our confidence intervals, defined as the proportion of simulations that
229	fall within the 50% or 90% HPD around the estimated parameter. Finally, we calculated the
230	coefficient of determination $(r^2)$ for each parameter, which is the proportion of parameter
231	variance that can be explained by the summary statistics (Bertorelle et al. 2010).
232	
233	Results
234	Model Selection
235	The raw posterior probabilities from our ABC analysis suggested the highest support
236	(0.70) for Model 4b (Table 3), the scenario that involved rapid growth of the VLC population
237	after colonization from MQ, but that did not include ongoing migration from MQ or back-
238	migration to MQ. We found lower support for Model 6b (0.24), which did include migration
239	back to MQ near the collapse of VLC. There was very little support for scenarios in which VLC
240	and MQ were panmictic, where they were completely independent, where MQ was colonized by
241	VLC, or where VLC grew slowly.
242	However, when we simulated datasets to validate our model choice results, we found that
243	distinguishing between certain models could be difficult, at least after such models had been fit
244	to the observed data (a more stringent test than used in many previous studies) (Table S3). In
245	particular, Models 1, 3, 5, 6, 7, 5b, and 7b were misidentified as a different model more than

246	50% of the time (Table S3). Model 4b was the most accurately identified scenario, with 90% of
247	simulations correctly chosen as Model 4b. Most of the remaining Model 4b simulations were
248	mistaken for Model 6b, a model that included 4b as a limiting case when back-migration was
249	low. Encouragingly, however, and of particular relevance to our study, few scenarios were
250	misidentified as Model 4b with high support (e.g., $\geq 0.70$ ) (Fig. S1). The exception to this general
251	rule was Model 5b, which, like 6b, included 4b as a limiting case when the migration rate was
252	low. We corrected for the possibility of misidentification by calculating $Pr(Model X   P_{Model 4a} =$
253	0.70) for each model (Table 3, Fig. S1). We found that most of the posterior support, once
254	corrected for errors in model choice, was split between Models 4b and 5b, with lower support for
255	Models 6b and 7b. This result suggested that rapid growth soon after VLC colonization was
256	highly likely, but determining whether or not migration continued after colonization could be
257	more difficult. Back-migration to MQ near the collapse of VLC was less well supported
258	(probability of 0.27 that the true scenario was Models 6b or 7b).
259	Further model checking with tail probabilities indicated that at least one summary
260	statistic failed the test ( $p < 0.05$ ) for each of Models 1-7, but none of the statistics failed for
261	Models 4b-7b (those with rapid growth) (Table S4). The combined <i>p</i> -value statistic indicated
262	particularly poor fits for Models 1 and 2 ( $p < 0.001$ ), but did not reject any of the other models.
263	The PCA analysis on the summary statistics indicated poor fits for Models 1-4, with the observed
264	data clearly falling outside the simulations from the posterior (Fig. S3). The other models
265	provided better fits to the data, according to this test.
266	Under alternative model priors or alternative statistical group definitions, Model 4b
267	continued to be selected as the best model (Table 3). However, under log normal priors for

268 migration rate, the model choice procedure indicated lower support for Model 4b than before,

with the remaining support distributed primarily among Models 5b, 6b, and 7b. This result is to
be expected because very low values of migration rates (close to 0) were commonly selected
under log normal priors, and so the functional differences among Models 4b, 5b, 6b, and 7b were
smaller. Under wider priors for the size of the colonizing population, support for Model 4b
remained high (68%). Support also remained substantially higher for Model 4b (51%) than for
alternative models (≤18%) with four statistical groups.

275

#### 276 Parameter estimation

277 We then estimated the posterior distributions for both of the well-supported models (4b 278 and 5b) and found similar parameter estimates from both (Table 1, Fig. 2). In particular, we 279 found that MQ was about two orders of magnitude smaller than VLC, consistent with the lower 280 genetic diversity in this population (Table 1, Fig. 2). Posterior parameter estimates for the 281 migration rate between MQ and VLC also clarified why Model 5b received similar support to 282 Model 4b: the migration rate posterior for Model 5b was nearly 0 (95% HPD of 0 to 0.003; Table 283 1, Fig. 2). With a very low to zero migration rate, Model 5b was functionally similar to Model 4b, 284 and the two were therefore difficult to distinguish. From the width of the posterior, it was clear 285 that we did not have much power to detect the effective number of females that colonized VLC, 286 though a founding size <200 seals was unlikely (Fig. 2).

In testing the accuracy of our parameter estimates, we found that the relative bias and RMSE error were quite low for the MQ and ancient VLC effective size (<50%), but substantially higher for the final VLC effective size with a positive bias of up to 100% (Table 4). Our estimates of final VLC size, however, were in the right order of magnitude. This accuracy was

also reflected in the factor 2 calculations, which were highest for MQ size and lowest for final

VLC size. The coverage of our 50% and 90% HPDs were quite accurate, with mean values across parameters of 44% and 88%. The high values for our coefficients of determination suggested that our choices of summary statistics were informative. We concluded that our estimates of MQ and ancient VLC population sizes were most accurate, but we could only approximate the final VLC size to within an order of magnitude.

297 When we compared the fit of simulated datasets based on our posterior distributions, we 298 found that they were a substantially better fit to the observed data than were the prior 299 distributions (Fig. 3). However, the best model continued to predict haplotype diversity in the 300 ancient VLC population ( $H_{VLC1}$ ) that was somewhat too low, while the predicted number of 301 segregating sites in the recent VLC population remained somewhat too high ( $S_{VLC2}$ ). On the 302 other hand, the key prediction of final haplotype number in the VLC was a strong match. The fit 303 for Model 4b was also substantially better than for a similar model that did not include rapid 304 initial population growth in VLC (Model 4). The latter model substantially under-predicted 305 haplotype diversity in VLC ( $H_{VLC1}$  and  $H_{VLC2}$ ), suggesting that rapid growth was important for 306 the generation and maintenance of genetic diversity (Fig. S2).

307

#### 308 Discussion

309 Our analysis indicated that the high levels of genetic diversity attained within the VLC 310 southern elephant seal colony (Table 2) most likely arose largely by rapid mitochondrial control 311 region *in situ* mutation as a function of early rapid population growth to a large effective 312 population size, which effectively reduces the probability of extinction for new (single-copy) 313 alleles (Crow and Kimura 1970). This followed the release of a large expanse of suitable 314 breeding habitat along the Victoria Land Coast, after the retreat of the Ross Sea Ice Sheet in the

315 early- to mid-Holocene (Hall et al. 2006). Approximate Bayesian Computation provides strong 316 support for the idea that Macquarie Island was the source of founders of the VLC colony, 317 consistent with previous findings (de Bruyn et al. 2009; though the mtDNA marker provides 318 long-term inference only about the movement of females). Further, our analysis did not support 319 the idea that seals from VLC were purely a non-breeding moulting colony, that VLC was the 320 source of the MQ population, or that seals from the two locations were panmictic. We could also 321 rule out any of the other major extant southern elephant seal breeding colonies as sources of 322 either founders, or observable levels of gene flow (de Bruyn et al. 2009). Thus, the significant 323 levels of genetic diversity observed in the VLC population likely resulted from a combination of 324 rapid substitution rate at the mitochondrial control region, and rapid population growth to a large 325 effective population size, soon after colonization of newly-available habitat resulting from 326 Holocene climate change.

327 The model with the highest posterior probability is one of initial founding of VLC by a 328 cohort of MQ seals (95% HPDs of female effective size: 157 - 953), followed by rapid 329 population growth and a lack of gene flow between the two locations until final extinction of the 330 VLC colony, some 7000 years later (de Bruyn et al. 2009). Interestingly, inclusion of an initial 331 rapid growth phase would appear critical to fitting the models to our data (Table 3, Fig. 3, Fig. 332 S2). The next best-fit model is otherwise the same, also including a rapid growth phase, but 333 distinguished by very low levels of ongoing migration from MQ to VLC (95% HPDs of 334 migration rate 0 - 0.003). Although immigration from an unsampled 'ghost' population cannot 335 be fully excluded, there are no apparent candidate populations not already directly or indirectly 336 sampled, and data from de Bruyn et al. (2009), which sampled all major extant populations, 337 illustrate a clear connection between VLC and MQ in the mtDNA networks. One other possible

338 interpretation is that migrants to VLC were from other now-extinct colonies, not from Macquarie. 339 The most likely candidate, based on geographic proximity, is the population from the north-340 western coast of Tasmania, extinct since prehistoric times (Bryden et al. 1999), though 1500 km 341 further from VLC than MQ. In this context, it seems unlikely that the VLC population would be 342 founded by the Tasmanian (or the even more distant historical colony in New Zealand) rather 343 than by the MQ population. The colonies on Juan Fernandez and St. Helena were likely small and exterminated by hunting in the 19th and 20th centuries. Both were very distant from VLC, 344 345 over 4,000km to Juan Fernandez and over 6,000 km to St. Helena (the northern most record for 346 this species; Fraser 1935). The chance that either of these would form a useful/relevant reference 347 sample for this study seems very remote.

348 Simulated datasets based on posterior distributions fitted the observed data far better than 349 the prior distributions (Fig. S2), including a strong fit between observed recent VLC haplotypic 350 diversity and that predicted (Fig. 3). There were two exceptions, with the model prediction for 351 haplotype diversity being low for the older VLC sample, and high for the number of segregating 352 sites in the more recent VLC sample. The lower than expected observed number of segregating 353 sites may simply reflect higher than expected variance of substitution rate among sites. While the 354 best-supported model was one of no post-founder gene flow, one possible explanation for higher 355 than expected haplotypic diversity early on is a post-founder 'connection event(s)' (see Alcala et 356 al. 2013). Although this may be consistent with our second (Model 5b) and third (Model 6b) 357 best-fitting models (Table 3), the level of migration indicated in those models is very small. 358 Other possible explanations include stochastic sampling effects, and sequence error in the older 359 samples due to post-mortem damage. The latter is unlikely given the controls and replication 360 undertaken to ensure accurate genotyping, including multiple replicate extractions-through-to-

361 sequencing, and cloning (see de Bruyn et al. 2009 for further details).

362 The key result remains the match between the very high haplotype diversity at VLC, 363 including extensive novel diversity compared to the source population MQ (de Bruyn et al. 364 2009), and model expectations derived using ABC and ancient DNA. While the generation of 365 genetic diversity in this case is presumed to be neutral and the mutation rate relatively high 366 (mitochondrial control region), our results suggest that given the right conditions, environmental 367 change could generate the raw material for adaptive evolution (affecting multi-genic phenotypic 368 traits) through rapid population growth over a very short evolutionary time scale. 369 370 Acknowledgements 371 We thank M. Blum for assistance with calculating the coefficient of determination, and Oscar 372 Gaggiotti for helpful comments that improved the manuscript. MLP was supported by National 373 Science Foundation Graduate Research and David H. Smith Conservation Research fellowships. 374 Research funding was provided by National Science Foundation grant PLR-0807292, and 375 support from the Italian Antarctic Programme (PNRA).

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4J/ Figure Capuons	457	Figure	Captions
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- 458 Figure 1. Diagram of hypotheses about migration and colonization between MQ and VLC over
- the period in which VLC was extant (8000 to 1000 ybp). The width of the arrows indicates

460 population size through time, while horizontal arrows indicate migration. Alternate versions of

- 461 Models 4 to 7 (Models 4b to 7b, not shown) specified rapid growth in VLC over the first 500
- 462 years (83 generations) to a plateau in population size.

463

464 Figure 2. Posterior densities for parameters from Model 4b (top row) and Model 5b (bottom

row). Migration rate from MQ to VLC is only present in Model 5b.

466

467 **Figure 3.** Comparison between the posterior distributions (black curve) and the prior

distributions (dashed curve) for Model 4b. The summary statistics are from Macquarie (MQ),

469 ancient Victoria Land Coast (VLC1) and more recent Victoria Land Coast (VLC2) and include

470 the number of haplotypes (H), the number of segregating sites (S), the number of pairwise

471 differences ( $\pi$ ), and pairwise F<sub>ST</sub> values. The observed value of each summary statistic is shown

as a vertical line.

# **Tables and Figures**

Description	Prior	Median	Mode	95% HPD
MQ female effective size	0 to $10^5$	2,050	2,000	1,290 - 3,290
VLC founding female effective size	0 to $10^{3}$	627	719	157 - 953
VLC final female effective size	0 to 10 <sup>6</sup>	158,000	120,000	65,700 - 395,000
MQ female effective size	0 to $10^{5}$	2610	2410	1220 - 4650
VLC founding female effective size	0 to $10^{3}$	527	664	33 - 998
VLC final female effective size	0 to $10^{6}$	474,000	110,000	21,300 – 986,000
MO to VLC migration	0 to 100%	0.0001	0.0004	0 - 0.003
-	Description MQ female effective size VLC founding female effective size VLC final female effective size MQ female effective size VLC founding female effective size VLC final female effective size MO to VLC migration	DescriptionPriorMQ female effective size0 to $10^5$ VLC founding female0 to $10^3$ effective size0 to $10^6$ VLC final female0 to $10^6$ effective size0 to $10^5$ VLC founding female0 to $10^3$ effective size0 to $10^3$ effective size0 to $10^6$ VLC final female0 to $10^6$ effective size0 to $10^6$ effective size0 to $10^6$ effective size0 to $10^6$ effective size0 to $100\%$	DescriptionPriorMedianMQ female effective size0 to $10^5$ 2,050VLC founding female0 to $10^3$ 627effective size0 to $10^6$ 158,000MQ female effective size0 to $10^5$ 2610MQ female effective size0 to $10^3$ 527VLC founding female0 to $10^6$ 527VLC founding female0 to $10^6$ 474,000effective size0 to $100^6$ 474,000	DescriptionPriorMedianModeMQ female effective size $0 \text{ to } 10^5$ $2,050$ $2,000$ VLC founding female $0 \text{ to } 10^3$ $627$ $719$ VLC final female $0 \text{ to } 10^6$ $158,000$ $120,000$ MQ female effective size $0 \text{ to } 10^5$ $2610$ $2410$ VLC founding female $0 \text{ to } 10^3$ $527$ $664$ VLC final female $0 \text{ to } 10^6$ $474,000$ $110,000$ MQ to VLC migration $0 \text{ to } 100\%$ $0 0001$ $0 0004$

**Table 1:** Prior and posterior parameter estimates for the MQ and VLC colonization history.

**Table 2:** Summary statistics for southern elephant seal populations from Macquarie Island (MQ) and from Victoria Land Coast (VLC) during the 3000-7100 ybp period (VLC1) or the 0-3000 ybp period (VLC2) (from de Bruyn et al. 2009). The statistics include the number of haplotypes (*H*), the number of segregating sites (*S*), the number of pairwise differences ( $\pi$ ), and pairwise F<sub>ST</sub> values.

483

summary statistic	value
<i>MQ</i> (n=48)	
$H_{MQ}$	15
$S_{MQ}$	23
$\pi_{MQ}$	6.58
<i>VLC1 3000-7100</i> (n=64)	
H <sub>VLC1</sub>	58
S <sub>VLC1</sub>	49
$\pi_{VLC1}$	7.97
<i>VLC2 0-3000</i> (n=159)	
$H_{VLC2}$	128
$S_{VLC2}$	79
$\pi_{VLC2}$	7.69
F <sub>ST</sub>	
MQ to VLC1	0.20298
MQ to VLC2	0.16809

485

Table 3: Posterior probabilities for each model. Values in bold sum to 90% of posterior model
probability within a column. Corrected probabilities account for errors in scenario identification.
Posterior probabilities under alternative priors and statistical group definitions are also shown,
including log normal migration rate priors (Priors 2), wider priors on the size of the colonizing
population (Priors 3), and four statistical groups (4 Groups).

	Standar	d priors	Priors 2	Priors 3	4 Groups
Model	posterior probability	corrected probability	posterior probability	posterior probability	posterior probability
1. One population	< 0.01	< 0.01	< 0.01	< 0.01	< 0.01
2. Separate populations	< 0.01	< 0.01	< 0.01	0.01	0.04
3. Colonization from	< 0.01	0.06	< 0.01	0.01	< 0.01
VLC					
4. Colonization from	< 0.01	< 0.01	< 0.01	0.07	0.09
MQ					
4b. with rapid	0.70	0.35	0.34	0.68	0.51
growth					
5. Ongoing migration	< 0.01	0.01	< 0.01	< 0.01	0.02
5b. with rapid growth	< 0.01	0.31	0.24	< 0.01	0.02
6. Back-migration	< 0.01	< 0.01	< 0.01	0.03	0.08
6b. with rapid growth	0.24	0.18	0.25	0.18	0.18
7. Ongoing and back- migration	0.04	< 0.01	< 0.01	< 0.01	0.02
7b. with rapid growth	< 0.01	0.09	0.16	< 0.01	0.03

Parameter		True	Estimated	Dias	DMCE	Factor 2	50%	00%	$\mathbf{D}^2$
		Median	Median	Dias	RNDE	Factor 2	30%	90%	K
$N_{\rm MQ}$		2,050	2,180	0.035	0.23	0.983	0.47	0.88	99%
$N_{\rm VLC1}$		627	454	-0.087	0.44	0.763	0.52	0.92	99%
$N_{\rm VLC2}$		158,000	330,000	0.95	1.0	0.578	0.33	0.84	99%

**Table 4:** Accuracy of parameter estimation under Model 4b using 1000 simulated datasets.







**Figure 3:** 

