

1 **Evidence for a 200 km thick diamond-bearing root beneath the**  
2 **Central Mackenzie Valley, Northwest Territories, Canada?**  
3 **Diamond indicator mineral geochemistry from the Horn Plateau**  
4 **and Trout Lake regions**

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18

19 **Abstract** The Central Mackenzie Valley (CMV) area of Northwest Territories is underlain by  
20 Precambrian basement belonging to the North American Craton. The potential of this area to  
21 host kimberlitic diamond deposits is relatively high judging from the seismologically-defined  
22 lithospheric thickness, age of basement rocks (2.2-1.7 Ga) and presence of kimberlite  
23 indicator minerals (KIMs) in Quaternary sediments. This study presents data for a large  
24 collection of KIMs recovered from stream sediments and till samples from two study areas in  
25 the CMV, the Horn Plateau and Trout Lake. In the processed samples, peridotitic garnets

26 dominate (> 25 % at each location) while eclogitic garnet is almost absent in both regions (< 1  
27 % each). KIM chemistry for the Horn Plateau indicates significant diamond potential, with a  
28 strong similarity to KIM systematics from the Central and Western Slave Craton. The most  
29 significant issue to resolve in assessing the local diamond potential is the degree to which  
30 KIM chemistry reflects local and/or distal kimberlite bodies. Radiogenic isotope analysis of  
31 detrital kimberlite-related CMV oxide grains requires at least two broad age groups for eroded  
32 source kimberlites. Statistical analysis of the data suggests that it is probable that some of  
33 these KIMs were derived from primary and/or secondary sources within the CMV area, while  
34 others may have been transported to the area from the east-northeast by Pleistocene glacial  
35 and/or glaciofluvial systems. At this stage, KIM chemistry does not allow the exact location  
36 of the kimberlitic source(s) to be constrained.

37

38 **Key words:**

39 Kimberlite indicator minerals

40 Garnet

41 Ilmenite

42 Hf isotopes

43 Geothermobarometry

44 Diamond exploration

45

46 Character count: 75,799 Characters (with spaces)

47 64,599 Characters (no spaces)

## 48 **Introduction**

49

50 Since the discovery of the first diamondiferous kimberlite in the Lac de Gras area in 1991,  
51 more than 300 kimberlites - many diamondiferous - have been identified in the Northwest  
52 Territories (NT). These discoveries have provided new suites of peridotite and eclogite  
53 xenoliths, as well as diamonds, generating considerable interest in the mantle beneath the  
54 Slave Craton. Despite the economic significance of diamond mines from the Central Slave  
55 Craton, relatively little research has been conducted on the remaining portions of cratonic  
56 lithospheric mantle underlying other parts of northern Canada, such as the area west of the  
57 Slave Craton margin. The possible existence of thick and cold cratonic lithosphere in the  
58 Central Mackenzie Valley (CMV), NT, more than 200 kilometres west of the Slave Craton  
59 margin (Fig. 1), is currently poorly constrained. This area, south of Great Bear Lake, has seen  
60 limited diamond exploration. Olivut Resources Ltd. is currently exploring this area and have  
61 reported the discovery of at least 29 kimberlites (two diamondiferous) since the  
62 commencement of their HOAM project in 1993 (Fig. 1; Pitman 2014). The reported  
63 kimberlite indicator mineral (KIM) chemistry from these kimberlites (Pitman 2014) differs  
64 from data obtained during regional stream sediment and till sampling (Day et al. 2007; Mills  
65 2008; Pronk 2008) suggesting the possible presence of additional, potentially diamondiferous  
66 sources within the CMV or complex mineral transport into the region.

67 Although the CMV region is not a traditional Archean “cratonic” setting in terms of its  
68 crustal geology - a widely accepted pre-requisite for diamondiferous kimberlites (Janse 1994)  
69 - discoveries of primary diamond occurrences in North America have been made in similar  
70 settings (e.g., Buffalo Head Hills). Furthermore, the LITHOPROBE Slave-Northern  
71 Cordillera Lithospheric Evolution (SNORCLE) transect line 1 (e.g., Cook et al. 1999) and  
72 more recent regional-scale surface-wave seismic tomography studies (e.g., Schaeffer and  
73 Lebedev 2014) indicate the likely presence of thick (~ 200 km), cold lithospheric mantle

74 extending into the diamond stability field, underpinning an area of several hundreds of square  
75 kilometres across the CMV (Fig. 1). In this study, we present new geochemical data for a  
76 large collection of KIMs sampled from two regions within the CMV: Horn Plateau (25,233  
77 km<sup>2</sup>) northeast of the Mackenzie River and > 175 km west of Yellowknife and Trout Lake  
78 (10,680 km<sup>2</sup>) south of the Mackenzie River (Figs. 1-2; for river, lake and place locations see  
79 Figures S1-S2). KIM chemistry of grains from these regions will be used to assess their  
80 diamond potential and improving our understanding of potential kimberlite sources.

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82

### 83 **Geological Setting**

84

#### 85 **Precambrian Geology**

86

87 The Precambrian basement of the Wopmay Orogen (ca. 1.95-1.85 Ga; Davis et al. 2015;  
88 Ootes et al. 2015) is known from a few oil and gas exploration wells (Burwash et al. 1993 and  
89 references therein) plus inferences made from geophysical survey data in the CMV (Cook et  
90 al. 1999; Aspler et al. 2003; Cook and Erdmer 2005). From east to west under the  
91 Phanerozoic strata is Precambrian basement from the 1.88-1.85 Ma Great Bear Magmatic  
92 Zone, the 1.95-1.89 Ga Hottah Terrane, the ca. 1.84 Ga Fort Simpson Terrane (magnetic high)  
93 and the cryptic Nahanni magnetic low (Villeneuve et al. 1991; Aspler et al. 2003; Cook and  
94 Erdmer 2005). In the extreme northeast of the CMV, Precambrian crystalline and  
95 metasedimentary rocks outcrop on the exposed Canadian Shield (Fig. 1). In the eastern Horn  
96 Plateau region, Precambrian rocks occur at only 400 m depth (Burwash et al. 1993) and dip  
97 gently towards the west-southwest to depths of greater than two kilometres in the westernmost  
98 areas (Gal and Lariviere 2004).

99

## 100 **Phanerozoic Geology**

101

102 The CMV lies in the northern portion of the Phanerozoic Western Canadian Sedimentary  
103 Basin known as the Northwest Territories Interior Platform. The nearly horizontal, mostly  
104 undeformed Phanerozoic sedimentary rocks are disrupted by highs of underlying Precambrian  
105 basement from episodes of uplift, erosion and subsidence (Dixon 1999). In the extreme  
106 northeast of the CMV, near Lac la Martre, Cambrian to Middle Devonian strata outcrop (Fig.  
107 1). In the northern Horn Plateau, Cambrian rocks overlie the Precambrian basement, except  
108 over the Bulmer Lake gravity high (Fig. 1), where they were either eroded and/or not  
109 deposited. Elsewhere Cambrian sediments appear to have been largely removed or were not  
110 deposited (Meijer Drees 1993). Ordovician and Silurian rocks still exist at depth in the  
111 northern part of the Horn Plateau, north of about 62 °N, while south of 62 °N, Middle-Late  
112 Devonian strata directly overlies the basement (Fig. 1; Williams 1985; Meijer Drees 1993;  
113 Pitman 2014). A prominent series of low escarpments of Devonian limestone located on the  
114 southern edge of the Horn Plateau are an exception to the otherwise monotonous, limited  
115 outcrop landscape (Craig 1965).

116 Cretaceous strata are only preserved in the extreme northwest and in higher areas  
117 where they have not been removed by erosion in the Horn Plateau region (e.g., Horn Plateau  
118 and Ebbutt Hills; Fig. 1) and are covered by glacial till and organics, with few recorded  
119 outcrops in the CMV (Craig 1965; Meijer Drees 1993). On the Horn Plateau, flat-lying Albian  
120 marine shales and minor sandstones unconformably overlie Late Devonian strata (Craig 1965;  
121 Meijer Drees 1993; Dixon 1999) with the Albian strata reaching ~ 60 m thick east of Willow  
122 Lake on the plateau and ~ 100 m thick east of Ebbutt Hills (Dixon 1999). South of ~ 61 °N,  
123 Late Aptian to Campanian strata (Fort St. John Group) are widespread (Fig. 1). East and  
124 southeast of Trout Lake, Cretaceous strata lies unconformably, from north to south, on  
125 Devonian and Carboniferous strata (Stott et al 1993) whereas to the southwest this unit is

126 completely absent (Dixon 1999). Cenomanian-Turonian aged marine shales are only  
127 preserved in the eastern Trout Lake region, with only Early-Middle Cenomanian regressive  
128 sandstone beds (Dunvegan Formation) found capping some hills to the northwest and  
129 southeast of Trout Lake (Dixon 1999).

130

### 131 **Quaternary Geology**

132

133 The CMV is almost entirely covered with a mantle of glacial and post-glacial deposits (Figs.  
134 2, S2). Morainal tills blanket most of the CMV and are locally ridged or hummocky, while  
135 colluvial deposits and local undifferentiated bare rock outcrops are found along the southern  
136 escarpment of the Horn Plateau. To the north and northwest of the Horn Plateau, alluvial  
137 gravel and sand occupy flood plains and terraces along streams (Rutter et al. 1993). South of  
138 the plateau, organic (fen and bog) deposits overlie a massive till plain (Rutter et al. 1993;  
139 Fulton 1995). In the Mackenzie, Liard, Willowlake and other river valleys, fine-grained  
140 organic, eolian, glacio- and/or -fluvial and -lacustrine deposits blanket till, older deposits and  
141 bedrock below 220 m elevation (Fig. 2). North of the Horn Plateau, glacial deposits are up to  
142 81 m thick (Craigie 1991), while in the Fort Simpson area the maximum thickness is 120 m  
143 (Ghaznavi et al. 1986). In the western and eastern parts of the CMV on the Liard and Last  
144 Stop mineral claims of Olivut Resources, below glaciofluvial outwash, till is between 10-60 m  
145 thick (Pitman 2014). In contrast, thinner unconsolidated Quaternary deposits (10-20 m)  
146 typically exist throughout the lowlands and major river valleys in the Horn Plateau, as well as  
147 most of the Trout Lake (Rutter et al. 1993; Huntley et al. 2006, 2008).

148         Highlights of the CMV Quaternary history include intervals of continental glaciation  
149 involving complex frontal retreat and ice stagnation, glacial lake formation and meltwater  
150 drainage (Huntley et al. 2006, 2008). At the last glacial maximum (ca. 18 ka), the entire CMV  
151 was completely covered by the Laurentide Ice Sheet with a thickness > 1000 m, implied by

152 the appearance of granitic erratic boulders at ~ 1500 m elevation to the west of the CMV  
153 (Huntley et al. 2008). The advance of this continental ice-sheet resulted in the erosion,  
154 transport, dispersal and deposition of bedrock and pre-existing sediment within the CMV. The  
155 highlands of Ebutt and Martin Hills and the Horn Plateau are believed to have slightly  
156 deflected the regional advance of the ice-sheet (Rutter et al. 1993; Fulton 1995). To date, the  
157 only ice-flow features reported on top of the Horn Plateau are small drumlinoid features on  
158 the northeast edge of the plateau (Craig 1965), which mark the youngest ice-flow direction  
159 trending west-southwest (~ 255°; Grexton 1995). North and south of the plateau regional ice-  
160 flow was towards the west, while in the west near Martin Hills, it was deflected northwesterly  
161 and southwesterly around it (Fig. 2; Fulton 1995). Throughout the Trout Lake region, a south  
162 to southwest directed ice-flow is consistently preserved by drumlins and fluting (Douglas  
163 1959; Fulton 1995; Huntley et al. 2008).

164         The subsequent retreat of the ice-sheet broadly re-ordered the drainage systems across  
165 the CMV. Between 17-14 ka, the ice-margin on the Horn Plateau retreated towards the east,  
166 forming melt-water channels that eventually drained towards the west into an enlarged glacial  
167 Bulmer Lake (Huntley et al. 2006). After completely retreating from the Horn Plateau, the  
168 ice-margin continued to retreat (12-10 ka) in the Horn Plateau, creating even more meltwater  
169 channels with glaciofluvial deltas parallel to them (Huntley et al. 2006). Around the same  
170 time (12-11 ka), retreat of the ice-margin towards the north-northeast in the Trout Lake area  
171 shifted drainage from the southwest into glacial Lake Liard to the north into the CMV (i.e.,  
172 glacial Lake Mackenzie; Huntley et al. 2008). Rapid retreat of the LIS continued between 11-  
173 10 ka, shifting the location of the glacial lakes and subsequent meltwater channels towards the  
174 east up the valley near Mills Lake (Duk-Rodkin and Lemmen 2000; Huntley et al. 2006).  
175 After ice retreat and final glacial lake drainage, modern fluvial drainage patterns were  
176 established, with post-glacial streams draining runoff and sediment radially from the Horn  
177 Plateau, Ebutt and Martin Hills into the Horn, Liard, Mackenzie and Willowlake river valleys

178 (Duk-Rodkin and Hughes 1994; Huntley et al. 2006). These major river valleys occupy  
179 previously established meltwater channels of the major Late Wisconsinan glacial lakes  
180 (Bulmer, Mackenzie, McConnell; Craig 1965; Huntley et al. 2006).

181 Our understanding of the CMV Quaternary history is further hampered by the nature  
182 of the cap bedrock. The soft, primarily Cretaceous and Devonian sedimentary rocks and pre-  
183 existing glacial deposits were prone to intense erosion, glacial deformation and ice-thrusting  
184 (e.g., Paulen 2009). As a result, most landforms and other evidence of earlier ice-flows were  
185 largely obliterated. Those that remained were extensively eroded, transported and re-  
186 deposited by colluvial, fluvial, lacustrine and aeolian processes during and following ice  
187 retreat. Other potential complexities in understanding the glacial transport and dispersal  
188 distance of KIMs in the CMV include palimpsest landscapes (e.g., Parent et al. 1996; Paulen  
189 2013), ice streams (e.g., Margold et al. 2015) and regional subglacial meltwater storage and  
190 drainage events (e.g., Rampton 2000). Collectively, these many factors result in a complex  
191 and incomplete Quaternary history for the CMV.

192

193

## 194 **Samples**

195

196 In 2003 and 2005, a total of 325 heavy mineral concentrate stream sediment samples (~ 50 m<sup>2</sup>  
197 spacing) were collected in the Horn Plateau region by the Geological Survey of Canada  
198 (GSC; Day et al. 2007). GSC and Northwest Territories Geological Survey (NTGS) personnel  
199 followed this up in 2006 with the collection of 25 till samples weighing only 2-10 kg in the  
200 Horn Plateau region (Mills 2008). Meanwhile, 166 reconnaissance-scale till samples weighing  
201 ~ 10 kg each (~ 100 m<sup>2</sup> spacing) were collected from the Trout Lake region by the NTGS in  
202 2008 (Watson 2010).



203 In total, 3665 (Horn Plateau) and 656 (Trout Lake) potential KIM grains were picked  
204 from the 0.25-2.0 mm size fractions and made into grain mounts. The Horn Plateau KIM  
205 inventory comprises peridotitic garnet (51 %), ilmenite (30 %), chromite (10 %), Cr-diopside  
206 (4 %), olivine (3 %), low-Cr (< 1 wt% Cr<sub>2</sub>O<sub>3</sub>) garnet (< 1 %) and rutile (< 1 %). The Trout  
207 Lake KIM inventory comprises rutile (45 %), peridotitic garnet (23 %), Cr-diopside (16 %),  
208 ilmenite (7 %), chromite (6 %), olivine (2 %) and low-Cr garnet (< 1 %). Sample descriptions,  
209 location, collection, heavy mineral concentrate processing and indicator mineral picking  
210 details are detailed elsewhere for the Horn Plateau (Day et al. 2007; Mills 2008) and the Trout  
211 Lake (Watson 2010). For sample descriptions and locations, the reader is referred to Tables  
212 S1-S3.

213 In addition, a collection of Mg-ilmenites from kimberlites with known emplacement  
214 ages ranging between ~ 700-30 Ma were used to try to constrain possible crystallization ages  
215 or age ranges of Horn Plateau Mg-ilmenites, based on their Hf isotope compositions. The  
216 large (1.0-2.0 mm size fraction) Horn Plateau Mg-ilmenite grains (n = 22) were selected  
217 based on necessary minimum Hf abundances needed for analysis (~ 0.5 ng). Samples used for  
218 Mg-ilmenite Hf isotopic analysis were provided as grain mounts, picked grains, hand samples  
219 or drill core. For a complete list of kimberlite sample localities, emplacement ages and  
220 kimberlite geochronology methods used for reference, the reader is referred to Table S4.

221 A database of KIM chemistry, primarily of peridotitic garnet, was used for comparison  
222 in this study. KIM chemistry for the Southwestern (Drybones Bay), Western (Aquila, Cross,  
223 Orion, Ursa), Southeastern (Snap Lake) and Central (Lac de Gras) Slave Craton kimberlites  
224 were used for comparison (Schulze et al. 1995; Carbno 2000; Kerr et al. 2000; Carbno and  
225 Canil 2002; Creaser et al. 2004; Griffin et al. 2004; Menzies et al. 2004; Aulbach et al. 2007,  
226 2011; Roeder and Schulze 2008; Creighton 2009; Creighton et al. 2010; Bussweiler et al.  
227 2015). The geochemical database has major element data as oxides (SiO<sub>2</sub>, TiO<sub>2</sub>, Al<sub>2</sub>O<sub>3</sub>, Cr<sub>2</sub>O<sub>3</sub>,  
228 FeO, MnO, MgO, CaO) as well as some trace element concentrations (e.g., Ni, Y, Zr, REEs

229 for peridotitic garnet). In addition, a newly filtered (October 2017) NTGS GoData KIMC  
230 database was used to evaluate KIM chemistry from Central and Western Slave Craton  
231 surficial samples.

232

233

## 234 **Methods**

235

### 236 **Sample preparation and elemental analysis**

237

238 Picked indicator minerals from both regions were cleaned, mounted and analyzed for major  
239 and minor elements by electron probe microanalysis (EPMA), either previously at the GSC or  
240 during this study at the University of Alberta. EPMA analytical conditions, including standard  
241 information, are included in Supplementary Methodology. In total, 3153 Horn Plateau and  
242 656 Trout Lake supposed KIM grains were analyzed. For complete KIM EPMA results see  
243 Tables S5-S10. Following EPMA, 947 peridotitic garnet grains and 53 peridotitic olivine  
244 grains from the study regions were analyzed for trace element abundances by in situ sector-  
245 field laser ablation inductively coupled mass spectrometry (LA-ICP-MS) at the University of  
246 Alberta. For LA-ICP-MS instrument setup, analytical conditions and standards used see  
247 Supplementary Methodology and Hardman et al. (this volume). For standard results, see  
248 Tables S11-S12 and Figures S3-S4. All external geochemical data were treated in the same  
249 way as the CMV KIM data, i.e., filtering of oxide totals, limits of detection, calculation of  
250 cations, garnet classification and thermobarometry techniques (e.g., Grütter et al. 2004; Nimis  
251 and Taylor 2000). Well-established methodologies for interpreting KIM chemistry were used  
252 to compare the Horn Plateau geochemical datasets with external data from potential source  
253 regions (i.e., Slave Craton kimberlites).

254

255 **Isotopic analysis**

256

257 Radiogenic isotopes were determined for select Trout Lake rutiles (U-Pb) and Horn Plateau  
258 Mg-ilmenite (Hf) to infer the age of eroded kimberlites that supplied the detrital kimberlitic  
259 material. Trout Lake rutiles grains (n = 53) were analyzed by LA-ICP-MS for U-Pb isotopic  
260 ratios (e.g., Malkovets et al. 2016) on a Thermo Element IIXR. For complete LA-ICP-MS  
261 rutile U-Pb dating instrument setup, analytical conditions and standards used see  
262 Supplementary Methodology and Harris et al. (this volume). For Trout Lake rutile U-Pb  
263 dating standard results, the reader is referred to Table S13 and Figure S5.

264 Nowell et al. (2004) have shown that Mg-ilmenite Hf isotope compositions potentially  
265 are a useful indicator of the broad emplacement age of their kimberlite source. Using a similar  
266 approach on Horn Plateau Mg-ilmenite grains (n = 22), we classify these kimberlitic ilmenite  
267 grains into potential age groups. We recognize that the rutile U-Pb ages offer a more precise  
268 technique (e.g., Cooper et al. 2008; Harris et al. this volume), however Mg-ilmenite Hf  
269 compositions are the only possibility of providing age constraints for the source kimberlites  
270 that supplied the Horn Plateau KIMs. In an attempt to better improve potential age constraints  
271 on Horn Plateau grains, the “kimberlite mantle Hf isotope evolution curve” defined by Mg-  
272 ilmenites was expanded and evaluated with the larger reference dataset of worldwide  
273 kimberlites, primarily from the North American Craton. Details of sample preparation,  
274 analysis using multi-collector (MC-)ICP-MS, analytical conditions and relevant reference  
275 material data are provided in Supplementary Methodology. For standard results, see Table  
276 S14 and Figure S6.

277

278 **Geothermobarometry techniques**

279

280 Thermobarometry of compositionally screened Cr-diopsides from garnet peridotites (criteria  
281 of Grütter 2009) were used to generate paleo-geotherms for each of the two study areas (using  
282 FITPLOT; Mather et al. 2011), based on the assumption that all KIM grains were sampled  
283 penecontemporaneously from their mantle sources. The mean temperature of the Ni-in-garnet  
284 thermometers of Griffin et al. (1989) and Canil (1999) was used to constrain the temperature  
285 of last equilibration of peridotitic garnet xenocrysts. If possible, the Ni-in-garnet temperatures  
286 were projected to their respective pyroxene-based geotherm. The Al-in-olivine thermometer  
287 for garnet peridotites (Bussweiler et al. 2017) was used to evaluate whether any different  
288 mantle sampling information is captured by olivine versus garnet. Similar to garnet  
289 temperatures, the Al-in-olivine temperatures were projected, if possible to their respective  
290 pyroxene-based geotherm.

291

## 292 **Statistical analysis**

293

294 Peridotitic garnet chemistry from Slave Craton kimberlites and the Horn Plateau populations  
295 overlap in many existing bivariate discrimination methods (e.g., Grütter et al. 2004). To  
296 investigate these populations statistically, we recast major element data as natural logarithm-  
297 normalized cation ratios for Ti, Al, Cr, Fe, Mn, Mg and Ca with Si as the denominator (e.g.,  
298  $\ln(\text{Ti}/\text{Si})$ ). The natural logarithm permits elemental values to be expanded past the limited  
299 range created by the unit-sum constraint of geochemical data, shifts distributions closer to  
300 normality (Hardman et al. 2018) and helps alleviate problems with closure in geochemical  
301 data (Aitchison 1994). Peridotitic garnet chemistry populations from Horn Plateau and Slave  
302 Craton kimberlites and surficial samples were tested for normality via the Kolmogorov-  
303 Smirnov (K-S), Anderson-Darling (A-D) and Shapiro-Wilks (S-W) tests. These normality  
304 tests reject the hypothesis of normality when the p-value is  $\leq 0.05$ . Failing the normality test  
305 allows you to state with 95 % confidence the data does not fit a normal distribution while

306 passing the tests only allow you to state no significant departure from normality was found.  
307 Except for  $\ln(\text{Mn}/\text{Si})$  for the Western Slave Craton kimberlite population ( $p = 0.09$ ),  
308 distributions of the Horn Plateau and Slave Craton peridotitic garnet populations are non-  
309 normal for all other elements ( $p \leq 0.05$ ; Table S15).

310 Logistic regression (LR) was selected for comparison of our data with garnet  
311 chemistry from Slave Craton kimberlites, as it makes no underlying assumptions about  
312 distribution or normality. This non-parametric supervised statistical technique linearly  
313 transforms data, reduces dimensionality and maximizes the variance of multivariate datasets  
314 (Pohar et al. 2004). The LR solutions are derived using the freeware statistics package R using  
315 the following log-normalised variables (see Supplementary Methodology for run-stream):  
316  $\ln(\text{Ti}/\text{Si})$ ,  $\ln(\text{Al}/\text{Si})$ ,  $\ln(\text{Cr}/\text{Si})$ ,  $\ln(\text{Fe}/\text{Si})$ ,  $\ln(\text{Mn}/\text{Si})$ ,  $\ln(\text{Mg}/\text{Si})$ ,  $\ln(\text{Ca}/\text{Si})$ . Logistic regression  
317 solutions are derived for pairs of populations (e.g., the Horn Plateau and Central Slave),  
318 which results in a linear equation that assigns a numerical value to all data (Table S15). A  
319 density distribution of garnets in the populations based on these values is assessed for the  
320 degree of separation or overlap between different groups. The advantages and disadvantages  
321 of various statistical techniques are discussed in Supplementary Methodology.

322

323

## 324 **Results: KIM chemistry**

325

### 326 **Cr-diopside**

327

328 Major element variations of Cr-diopside grains from the Horn Plateau ( $n = 138$ ) and Trout  
329 Lake ( $n = 106$ ) indicate multiple petrogenetic sources (Fig. 3). Of these, 124 Horn Plateau and  
330 19 Trout Lake grains have  $\text{mg\#}$  (molar  $\text{Mg}/(\text{Mg}+\text{Fe})$ )  $> 0.88$  and elevated  $\text{Cr}_2\text{O}_3$  ( $> 0.5$  wt%)  
331 typical of mantle peridotites (Fig. 3). The  $\text{mg\#}$  of Horn Plateau peridotitic Cr-diopsides is

332 higher (mean of 0.92) compared to those from Trout Lake (mean of 0.90). Peridotitic Cr-  
333 diopsides from the Horn Plateau can be sub-divided into at least two different groups (criteria  
334 outlined by Ramsey and Tompkins 1994; Nimis 1998; Cookenboo and Grütter 2010). The  
335 first group contains 120 Horn Plateau and 15 Trout Lake Cr-diopside grains with lower Al<sub>2</sub>O<sub>3</sub>  
336 (< 4 wt%) contents and generally higher mg# (mean 0.92), typical of Cr-diopside from garnet-  
337 peridotites (Fig. 3). The second group of peridotitic Cr-diopsides consists of four Horn  
338 Plateau and four Trout Lake Cr-diopside grains with higher Al<sub>2</sub>O<sub>3</sub> (> 4 wt%) contents and  
339 variable mg# (0.89-0.95; Fig. 3). The grains from this group are either from off-craton garnet-  
340 peridotites or from spinel-peridotites (on- or off-craton). Most of the CMV peridotitic Cr-  
341 diopside grains are compositionally indistinguishable from those derived from Central and  
342 Western Slave Craton surficial samples (Fig. 3).

343

#### 344 **Low-Cr garnet**

345

346 Of the 30 Horn Plateau low-Cr garnets analyzed, 20 are undefined (G0) and of potentially  
347 crustal origin, six are low-Cr megacrysts (G1), three are low-Ca eclogitic (G3) and two are  
348 high-Ca eclogitic to pyroxenitic (G4; after Grütter et al. 2004). The crust-mantle method of  
349 Hardman et al. (2018) classifies all four G3 and G4 grains as “mantle-derived” (Fig. S7). Out  
350 of 17 low-Cr garnet grains collected from the Trout Lake, nine classify as G3 with non-  
351 deemed “mantle-derived” using the Hardman et al. (2018) scheme (Fig. S7).

352

#### 353 **Peridotitic garnet**

354

355 Of 1638 peridotitic garnet grains classified using the scheme of Grütter et al. (2004), from the  
356 Horn Plateau, lherzolitic garnets (G9) dominate (51 %), followed by high-TiO<sub>2</sub> peridotitic  
357 garnets (G11, 27 %), harzburgitic garnets (G10, 8 %) and wehrlitic garnets (G12, 5 %). A

358 small portion of the Horn Plateau garnets are Cr<sub>2</sub>O<sub>3</sub>-rich (10-15 wt%) with only 3 % situated  
359 above the graphite-diamond constraint (Fig. 4; Grütter et al. 2006). Conversely, the Trout  
360 Lake peridotitic garnets are all G9 (n = 138), except for one G12 grain and generally plot in  
361 the high-Ca off-craton portion of the lherzolitic field (Grütter et al. 2004), setting them apart  
362 from typically lower Ca lherzolitic garnets from the Horn Plateau (Fig. 4).

363         Based on their chondrite-normalized rare earth element patterns (REE<sub>N</sub>) patterns (Fig.  
364 5), peridotitic garnets from the Horn Plateau can be classified in three principal groups (see  
365 Table S16). (1) HREE-enriched patterns have very low LREE concentrations that increase to  
366 chondritic abundances through the MREEs, becoming ~ 10x chondritic in the HREEs (Fig. 5).  
367 This is the largest group (55 % of the garnets) and is dominated by G9s (n = 309), followed  
368 by 127 G11, four G12 and three G10 grains. Based on their Ni contents, the HREE-enriched  
369 group covers the entire temperature range observed for Horn Plateau garnets and have  
370 variable Zr, Y, Hf abundances (Fig. S8). (2) Garnets from the MREE-depleted group have V-  
371 shaped to extremely sinusoidal REE<sub>N</sub> patterns with variable, typically super-chondritic LREE  
372 abundances, decreasing concentrations from Ce or Nd to Tb or Ho and steep linear positive  
373 slopes in the HREE<sub>N</sub> with near chondritic Lu concentrations (Fig. 5). This group contains 13  
374 G10, seven G11, five G12 and four G9 grains with Ni temperatures < 1100 °C for all garnets  
375 (except one). All garnets from this group have very low Zr, Y and Hf concentrations (< 0.024  
376 ppm), at or below chondrite abundance (Fig. S8). (3) Garnet REE<sub>N</sub> patterns of the sinusoidal  
377 group have positive slopes in the LREE<sub>N</sub> with a peak at Nd, negative slopes in the MREE<sub>N</sub>  
378 with a trough between Dy and Er and positive slopes in the HREE<sub>N</sub> (Fig. 5). This group  
379 consists of 202 G9, 73 G10, 48 G11 and 15 G12 grains with low and high Zr, Y and Hf  
380 concentrations, occurring in approximately equal proportions and again extending across the  
381 entire temperature range observed for Horn Plateau garnets (Fig. S8).

382

383 **Olivine**

384

385 Kimberlite-related (largely mantle-derived xenocrystic) olivine must be differentiated from  
386 that originating from “basaltic” sources recovered in surficial samples. Only 38 of the 109  
387 Horn Plateau olivine grains have chemistry typical of xenocryst cores of olivine from  
388 kimberlites (mg# 0.89-0.94, NiO  $\geq$  0.3 wt%, CaO  $\leq$  0.1 wt%, MnO  $\leq$  0.15 wt%; e.g.,  
389 Bussweiler et al. 2015, 2017). All 15 Trout Lake olivines fall within this mantle range,  
390 although they have a less variable, lower mg# (mean 0.91) compared to those from the Horn  
391 Plateau (mean mg# 0.92). A few olivine grains from the Horn Plateau (n = 13) and Trout  
392 Lake (n = 1) have compositions typical for olivine inclusions in diamond with mg# > 0.91  
393 (Stachel and Harris 2008) and Cr<sub>2</sub>O<sub>3</sub> < 0.03 wt% (Fipke et al. 1995).

394

### 395 **Ilmenite**

396

397 Kimberlite-derived ilmenite can be distinguished from non-kimberlitic magmatic ilmenites  
398 based on the TiO<sub>2</sub> versus MgO discrimination plot of Wyatt et al. (2004). Out of 46 Trout  
399 Lake ilmenite grains, only four are kimberlitic (i.e., plot to the right of the discriminant curve  
400 in Fig. 6) whereas ilmenites from the Horn Plateau are more promising for diamond  
401 exploration, with 793 grains (of 948) being of kimberlitic composition.

402 Mills (2008) noted that Horn Plateau Mg-ilmenites document a complex  
403 crystallization history with both Cr<sub>2</sub>O<sub>3</sub>- and MgO-poor (oxidized) and Cr<sub>2</sub>O<sub>3</sub>- and MgO-rich  
404 (reduced) suites which define two groups: (1) a less prominent group with 5-7 wt% MgO and  
405 very low Cr<sub>2</sub>O<sub>3</sub> and (2) a dominant group with > 10 wt% MgO and Cr<sub>2</sub>O<sub>3</sub> increasing with  
406 MgO-content. All four kimberlitic Mg-ilmenite grains from three Trout Lake samples have  
407 similar Nb<sub>2</sub>O<sub>5</sub>, TiO<sub>2</sub>, Cr<sub>2</sub>O<sub>3</sub>, MnO and MgO contents to the Cr<sub>2</sub>O<sub>3</sub>- and MgO-poor suite of the  
408 Horn Plateau. These four Trout Lake grains and 10-25 % of the Horn Plateau Mg-ilmenites  
409 are characterized by relatively low TiO<sub>2</sub>, MgO and Cr<sub>2</sub>O<sub>3</sub> values that increase at the low end



410 of the MgO-spectrum, similar to those from the Drybones Bay kimberlite (Fig. 6; Schulze et  
411 al. 1995; Kerr et al. 2000). Meanwhile, the majority of Horn Plateau grains (75-90 %) have  
412 similar Nb<sub>2</sub>O<sub>5</sub>, TiO<sub>2</sub>, Cr<sub>2</sub>O<sub>3</sub>, FeO and MgO contents as those from Western and Central Slave  
413 Craton surficial samples (Fig. 6).

414

#### 415 **Rutile**

416

417 Although some studies have attempted to recognize kimberlite-related rutiles based on  
418 chemical composition (e.g., based on Cr<sub>2</sub>O<sub>3</sub> contents; Malkovets et al. 2016), other studies of  
419 rutile from mantle xenoliths show that major elements are less definitive (Harris et al., this  
420 volume). CMV rutiles also allow no clear distinction (Table S10). None of the Trout Lake  
421 rutiles (n = 291) contain  $\geq 1.7$  wt% Cr<sub>2</sub>O<sub>3</sub>, thought to be typical of a cratonic mantle source,  
422 while none of the four Horn Plateau rutiles contain  $> 0.4$  wt% Cr<sub>2</sub>O<sub>3</sub>, thought to be necessary  
423 for distinction from crustal sources (Fig. S9; Malkovets et al. 2016).

424 Only rutile from Trout Lake samples were analyzed for U-Pb isotope systematics and  
425 most grains have extremely low <sup>207</sup>Pb concentrations (in addition to Th and U), as is  
426 especially common for accessory minerals (e.g., rutile, zircon) from Mesozoic or younger  
427 magmas (e.g., Zack et al. 2011). Thus, accurate <sup>207</sup>Pb/<sup>235</sup>U ages were difficult to obtain and  
428 only the <sup>206</sup>Pb/<sup>238</sup>U ratios and apparent ages (Table S17), which have the lowest analytical  
429 uncertainties, were considered for resolving the emplacement age of eroded kimberlites that  
430 supplied these detrital rutile grains.

431 Trout Lake rutile <sup>206</sup>Pb/<sup>238</sup>U ages can be divided into three distinct groups: (1) two  
432 grains (4 %) with Archean <sup>206</sup>Pb/<sup>238</sup>U ages similar to Slave Craton granitoid rocks (2.8-2.5  
433 Ga), (2) 47 grains (89 %) with Proterozoic <sup>206</sup>Pb/<sup>238</sup>U ages similar to the Wopmay Orogen  
434 rocks (2.1-1.6 Ga) and (3) four grains from three samples (8 %) with potential kimberlitic  
435 <sup>206</sup>Pb/<sup>238</sup>U ages between 429-138 Ma (Fig. S10), based on the lack of evidence for

436 metamorphic events of such young ages in the CMV basement. Of these presumed kimberlitic  
437 rutile grains, the two oldest have Silurian and Early Devonian  $^{206}\text{Pb}/^{238}\text{U}$  ages of  $420.6 \pm 8.2$   
438 Ma and  $408.1 \pm 11$  Ma (Silurian and Early Devonian). These two rutile grains are from  
439 separate samples where they occur with other rutiles defining Paleoproterozoic crystallization  
440 ages (Fig. S10). Meanwhile the two youngest kimberlitic rutile grains, again from two  
441 separate samples, have Late Triassic to Early Cretaceous  $^{206}\text{Pb}/^{238}\text{U}$  ages of  $144.6 \pm 6.2$  Ma  
442 and  $227.0 \pm 35$  Ma ( $2\sigma$ ; Fig. S10).

443

## 444 **Spinel**

445

446 At least 70 % of the Horn Plateau (n = 314 total) and only 12 % of the Trout Lake (n = 42  
447 total) spinels have major element compositions ( $\text{TiO}_2$ ,  $\text{Al}_2\text{O}_3$ ,  $\text{Cr}_2\text{O}_3$ ,  $\text{MgO}$ ) typical of spinels  
448 from kimberlites (e.g., Sobolev 1977). On the basis of cr# vs mg# systematics (Roeder and  
449 Schulze 2008), potential Horn Plateau KIM grains with typically low mg# ( $< 0.35$ ) and  
450 moderate cr# (0.4 to 0.8) are indistinguishable from spinels from Central and Western Slave  
451 Craton surficial samples (Fig. S11).

452

453

## 454 **Discussion**

455

### 456 **Geothermobarometry**

457

458 The Horn Plateau Cr-diopsides (n = 54) define a conductive paleo geotherm equivalent to a  
459 lithospheric thickness of  $200 (\pm 10)$  km (Fig. 7), assuming a mantle potential temperature of  
460  $1300$  °C. Although the possibility that this geotherm represents mixed datasets (Cr-diopsides  
461 from various sources) cannot be excluded, the tight correlation (Fig. 7) makes it highly

462 unlikely that sources with significantly variable geotherms could have been sampled. This  
463 geotherm has a wide “diamond window” (Fig. 7) and with increasing depth crosscuts the 35-  
464 38 mW/m<sup>2</sup> model geotherms of Hasterok and Chapman (2011). Meanwhile, only three Trout  
465 Lake Cr-diopside grains passed compositional screening; they cluster in a narrow pressure-  
466 temperature interval within the graphite stability field and do not allow for the derivation of a  
467 reliable paleogeotherm (Fig. 7).

468         Of the 808 peridotitic garnet grains analyzed for trace elements from the Horn  
469 Plateau, the majority have Ni-in-garnet temperatures between 850-1350 °C (Fig. S8; Tables  
470 S16, S18); while the 138 G9 grains from the Trout Lake extend to lower temperatures  
471 spanning a smaller range, between 800-1050 °C (except one grain; Table S19). Projection of  
472 Ni-in-garnet temperatures to the pyroxene-based geotherm for Horn Plateau indicates  
473 significant sampling in the diamond stability field (Figs. 7 and 8). Based on the P<sub>38</sub> Cr-in-  
474 garnet barometer of Grütter et al. (2006), a minimum (as presence of Mg-chromite in the  
475 assemblage cannot be confirmed) maximum pressure of 5.8 GPa can be derived for the most  
476 Cr-rich Horn Plateau garnets, equivalent to a minimum lithospheric thickness of ~ 180 km  
477 (Fig. 4). This result agrees with the intersection of the pyroxene-based geotherm for Horn  
478 Plateau with the mantle adiabat at 200 ± 10 km (Fig. 7). The lack of a reliable geotherm for  
479 the Trout Lake precludes projection of Ni-in-garnet temperatures to depth.

480         Application of the Al-in-olivine thermometer requires screening out of olivines from  
481 spinel- and garnet-spinel facies peridotites, as the thermometer is only confirmed to work in  
482 garnet-facies peridotites. Based on their position on the Al versus V plot for garnet versus  
483 spinel facies discrimination (Fig. S12), only 17 of 38 Horn Plateau olivines classify as garnet-  
484 peridotite-derived. All 15 Trout Lake olivine grains classify as garnet-peridotite-facies (Fig.  
485 S12; Table S20). Al-in-olivine temperatures for 16 of the garnet peridotite-derived olivine  
486 grains from Horn Plateau are between 950-1200 °C (see Table S18). Projection of these Al-  
487 in-olivine temperatures to the Horn Plateau paleogeotherm (Fig. 7) indicates that they most

488 recently equilibrated at or deeper-than the graphite-diamond transition (Day 2012) with a  
489 similar mode in sampling depth as that defined by the garnets (120-180 km; Fig. 8). The Al-  
490 in-olivine temperatures for the 15 Trout Lake garnet-peridotite derived olivine grains are  
491 much higher (1000-1350 °C) than the temperature obtained from garnet-peridotite derived Cr-  
492 diopsides at Trout Lake (650-750 °C).

493

#### 494 **Diamond potential of KIMs**

495

496 Pyroxene-based geothermobarometry for Horn Plateau KIMs indicates derivation from a cold,  
497 deep (190-210 km) lithospheric mantle root, with a large diamond window that was  
498 abundantly sampled (Figs. 7-8). For Trout Lake, the lack of a pyroxene-based geotherm  
499 makes it difficult to interpret the thickness of lithosphere sampled by the kimberlite sources of  
500 the sampled indicators. The three pressure-temperature estimates based on Cr-diopside grains  
501 from this area, however, plot close to the Horn Plateau model-geotherm (Fig. 7) suggesting it  
502 is not significantly different from that for the Horn Plateau Cr-diopside grains.

503  $\text{Cr}_2\text{O}_3$ -CaO contents of peridotitic ( $\geq 1$  wt%  $\text{Cr}_2\text{O}_3$ ) garnets analyzed from both study  
504 areas (Fig. 4) suggest significantly different compositions of lithosphere sampled and  
505 transported to the surface. For Horn Plateau, the peridotitic garnet chemistry suggests  
506 derivation from thick and highly depleted lithosphere that extends into the diamond stability  
507 field. In stark contrast, for Trout Lake, both the absence of G10 garnets and the overall lower  
508 Ni-in-garnet temperatures suggest shallower sampling of a less depleted lithospheric section,  
509 likely with less favourable diamond potential.

510 All four Trout Lake Mg-ilmenites and a smaller portion of Horn Plateau grains (10-25  
511 %) with low  $\text{Cr}_2\text{O}_3$  and MgO contents (Fig. 6) have elevated  $\text{Fe}^{3+}/\text{Fe}^{2+}$  ratios (calculated from  
512 stoichiometry) that indicate less favourable redox conditions for diamond preservation (e.g.,  
513 Schulze et al. 1995). Castillo-Oliver et al. (2017), however, questioned the utility of Fe-Mg

514 systematics in Mg-ilmenite in defining diamond preservation potential. The extremely Cr<sub>2</sub>O<sub>3</sub>-  
515 rich Mg-ilmenite grains from the “reducing” suite (75-90 %) are suggested to indicate a very  
516 depleted and Cr-rich mantle source with favourable diamond potential (Mills 2008 and  
517 references therein). Moreover, 10 Horn Plateau (and no Trout Lake) spinels have > 61 wt%  
518 Cr<sub>2</sub>O<sub>3</sub>, 10-16 wt% MgO, < 0.60 wt% TiO<sub>2</sub>, < 8 wt% Al<sub>2</sub>O<sub>3</sub>, < 6 wt% Fe<sub>2</sub>O<sub>3</sub>, 13-17 wt% FeO  
519 (Fe contents calculated based on stoichiometry), i.e., compositions typical of diamond  
520 inclusion spinels (e.g., Sobolev 1977; Fipke et al. 1995) and hence indicative of favourable  
521 diamond potential.

522         Based on our new data and including previous studies on KIMs from Horn Plateau and  
523 Trout Lake (Day et al. 2007; Mills 2008; Pitman 2014), we conclude that Horn Plateau  
524 appears to have excellent diamond potential (provided that the indicator sources are local)  
525 while Trout Lake does not provide a viable exploration target.

526

### 527 **Potential source ages of Horn Plateau KIMs**

528

529 Although the relationship between ilmenite Hf isotope composition and kimberlite  
530 emplacement age is scattered and hence imprecise, it is very useful in allowing, e.g., to  
531 distinguish between Neoproterozoic-Silurian kimberlites and those of Devonian-Paleogene  
532 ages (Fig. 9, bottom). Based on the current worldwide kimberlitic ilmenite dataset (Fig. 9,  
533 bottom), Horn Plateau Mg-ilmenite Hf isotope compositions can be broadly divided into two  
534 age groups (Table S21).

535         The older age group contains nine Mg-ilmenites (from eight samples) with less  
536 radiogenic and restricted <sup>176</sup>Hf/<sup>177</sup>Hf ratios between 0.2825-0.2826, indicative of  
537 Neoproterozoic to Silurian kimberlite emplacement ages (Fig. 9). All the Mg-ilmenites of the  
538 older age group belong to the Cr<sub>2</sub>O<sub>3</sub>- and MgO-poor oxidizing suite at Horn Plateau (Fig. 6),  
539 as outlined by Mills (2008), with similar Nb<sub>2</sub>O<sub>5</sub>, TiO<sub>2</sub>, Cr<sub>2</sub>O<sub>3</sub>, FeO and MgO contents as those

540 from the Drybones Bay kimberlite (Fig. 6; Schulze et al. 1995; Kerr et al. 2000). The younger  
541 age group contains 12 Mg-ilmenites (from 10 samples) with more radiogenic and less  
542 restricted  $^{176}\text{Hf}/^{177}\text{Hf}$  ratios of 0.2827 to 0.2831, indicative of Devonian to Paleogene  
543 kimberlite emplacement ages (Fig. 9). All younger age Mg-ilmenites lie within the dominate  
544 MgO- and Cr<sub>2</sub>O<sub>3</sub>-rich suite outlined at Horn Plateau by Mills (2008), with similar Nb<sub>2</sub>O<sub>5</sub>,  
545 TiO<sub>2</sub>, Cr<sub>2</sub>O<sub>3</sub>, FeO and MgO contents as those from Central and Western Slave Craton surficial  
546 samples (Fig. 6).

547

### 548 **Horn Plateau KIM populations**

549

550 Results from this study support the conclusions of previous studies (e.g., Mills 2008) which  
551 suggested that the Horn Plateau KIMs are derived from multiple (at least one proximal and  
552 one distal) kimberlite sources. Mg-ilmenites define at least two groups based on distinct  
553 compositions and ages (Neoproterozoic-Silurian versus Devonian-Paleogene). The large  
554 variability in Hf isotope ratios clearly precludes that the Mg-ilmenites came from a single  
555 kimberlite field of restricted age.

556 Peridotitic garnet major element chemical data from the Horn Plateau has a non-  
557 normal distribution for all major and minor elements. This is typical for regional geochemical  
558 data, even after logarithmic transformation and suggests the likelihood of mixing from  
559 multiple kimberlite sources (e.g., Reimann and Filzmoser 2000). Garnet chemistry from the  
560 NTGS GoData data base for Slave Craton surficial samples also reveals widely varying, non-  
561 normal compositional characteristics suggesting they are also from multiple sources.

562 Variations in KIM chemistry with geographic location, as can be observed in the Horn  
563 Plateau (Day et al. 2007; Mills 2008) and HOAM property mineral chemistry data (Pitman  
564 2014), also suggest mixing of KIMs from multiple sources. Mg-ilmenites recovered from till  
565 samples (four grains from three samples) on the Horn Plateau are evenly split into the two

566 compositional/age groups previously outlined. Surficial samples principally containing Mg-  
567 ilmenites of the younger age group were collected from primary or secondary streams with  
568 moderate to good drainage, either on or near the edge of the Horn Plateau, suggesting they are  
569 sourced from the top of the plateau. Whether this source on the Horn Plateau is primary or  
570 secondary remains to be proven. The two stream sediment samples containing Mg-ilmenite  
571 grains with Hf isotope compositions from both age groups were taken from tertiary streams  
572 north of the Horn Plateau, along an ancient meltwater channel, suggesting at least two  
573 kimberlite sources have been mixed within this region.

574

#### 575 **Potential sources of Horn Plateau KIM populations**

576

577 A key question is whether these Horn Plateau KIM populations were derived from proximal  
578 kimberlites within the CMV/Horn Plateau or distal sources, for instance, from diamondiferous  
579 kimberlites on the Slave Craton. Following the discovery of economic diamondiferous  
580 kimberlites in the Lac de Gras area, it was thought that the Horn Plateau KIMs were glacially  
581 transported south-westward from the Central Slave Craton (e.g., Day et al. 2007; Mills 2008  
582 and references therein). Besides distinct similarity in the excellent diamond potential  
583 indicated by Horn Plateau KIM major element chemistry to the Central Slave Craton data set,  
584 one of the main arguments for a distal origin was the large amount of glacial debris, including  
585 granitic boulders of presumed Slave Craton origin, throughout the CMV (Day et al. 2007;  
586 Mills 2008). Most of these granitic glacial erratics have, however, not been dated, leaving the  
587 possibility that they may derive from more proximal source regions, such as the Wopmay  
588 Orogen.

589         The logistic regression (LR) solutions for peridotitic garnet chemistry from the Horn  
590 Plateau and kimberlites and surficial samples from the Central Slave Craton show sufficient  
591 spread in the population density distributions (Fig. 10) to reliably distinguish Horn Plateau

592 and Central Slave Craton derived KIMS where the populations do not strongly overlap. Based  
593 on the major element datasets analyzed using LR, it is likely that at least part of the Horn  
594 Plateau KIMs were not derived from the Central Slave Craton, as their population density  
595 distributions have solutions which do not overlap (Fig. 10). There is, however, significant  
596 overlap between peridotitic garnets from the Horn Plateau and surficial samples and  
597 kimberlites from the Western Slave Craton in their respective LR solution (Fig. 10).  
598 Consequently, the two populations cannot be probabilistically discriminated using LR and  
599 hence we cannot reject the possibility that the Horn Plateau KIMs were derived from the  
600 Western Slave Craton. In context of the inferences we draw from these LR results and other  
601 KIM chemistry results, we explore potential kimberlite sources of the Horn Plateau KIMs.

602

603 Central Slave Craton kimberlites

604

605 The Central Slave Craton kimberlites (> 500 km northeast of the Horn Plateau) were  
606 emplaced largely between Late Cretaceous to Eocene times (75-45 Ma; Heaman et al. 2003;  
607 Creaser et al. 2004; Sarkar et al. 2015). These kimberlites could be the source of the younger  
608 Horn Plateau Mg-ilmenite group. Nine of the 12 analyzed Mg-ilmenites from this age group  
609 have  $^{176}\text{Hf}/^{177}\text{Hf}$  ratios that overlap with ilmenites from Diavik and Ekati kimberlites (Fig. 9;  
610 Table S21). In addition, Cr-diopside chemistry (Fig. 3) and pressure-temperature estimates  
611 (Fig. 8) and peridotitic garnet major, minor and trace element chemistry (Figs. 4-5) are  
612 remarkably similar to those from the Central Slave Craton kimberlites dataset.

613 Paleogeographic reconstructions indicate fluvial systems transported eroded detritus  
614 away from the CMV, east across the Slave Craton towards Hudson Bay, during the Paleogene  
615 and Neogene (Duk-Rodkin and Hughes 1994; Dixon 1999; Duk-Rodkin and Lemmen 2000),  
616 which rules out fluvial transport of kimberlitic material from these kimberlites into the CMV.  
617 Quaternary transport by glacial (-fluvial) activity is the only possible mechanism for



618 transporting material from these kimberlites towards the southwest into the Horn Plateau. Mg-  
619 ilmenites from till samples in the eastern Horn Plateau (Last Stop Claims; n = 56) are all from  
620 the MgO- and Cr<sub>2</sub>O<sub>3</sub>-rich suite, suggesting the source of the “younger” (reduced) Mg-ilmenite  
621 age group may be outside the CMV, to the northeast. However, studies of indicator mineral  
622 glacial dispersal trains/fans indicate typical maximum transport distances of up to 180-200 km  
623 (e.g., McClenaghan et al. 2002; McClenaghan 2005) making transport of large concentrations  
624 of kimberlitic material (even kimberlitic boulders) from these kimberlites over large distances  
625 very unlikely. Ice streams could have influenced dispersal of Horn Plateau KIMs (e.g.,  
626 Margold et al. 2015), although there is no field evidence for these features to the east-  
627 northeast of the Horn Plateau region or the Slave Craton.

628         Distinct spinel major element chemistry variations and lack of eclogitic garnets (G3)  
629 and megacrysts (G1) relative to Central Slave Craton kimberlites, led Mills (2008) to suggest  
630 Ekati Lease Block (Lac de Gras area) kimberlites are not the source of the Horn Plateau  
631 KIMs. In addition, the re-constructed lithospheric stratigraphy of Horn Plateau garnets is  
632 slightly different to that of Central Slave Craton kimberlites; specifically, the Horn Plateau  
633 KIM data lacks the distinct ultra-depleted shallow (120-150 km) “layer” chemistry that  
634 characterises the mantle beneath that region (e.g., Griffin et al. 1999; Menzies et al. 2004).  
635 Further, > 20 % of the Horn Plateau G10 grains equilibrated at depths  $\geq 150$  km ( $> 1000$  °C;  
636 Fig. S8; Table S18), whereas relatively few G10s from the Central Slave lithosphere, in  
637 published datasets (e.g. Grütter et al. 1999), have yielded depths deeper than the ultra-  
638 depleted shallow layer (i.e.,  $\geq 150$  km or  $> 1000$  °C).

639

640 HOAM project Liard properties kimberlites

641

642 Besides the HOAM project kimberlites  $> 20$  km southwest of the Horn Plateau and directly  
643 north-northwest of Trout Lake (Fig. 1), there are no other known kimberlites in the vicinity

644 with possible Devonian to Paleogene emplacement ages. Olivut Resources Ltd. report  
645 intersecting 29 kimberlites (two diamondiferous) from 39 drill holes on their Liard properties  
646 (Pitman 2014), with uppermost diatreme and possibly crater facies (kimberlitic breccias and  
647 tuffs), as well as deeper root zone features (hypabyssal facies and abundant dikes) indicated.  
648 The locations and depths of drill holes that intersected kimberlites in the Liard properties  
649 (Figs. 1; Pitman 2014) when correlated with bedrock stratigraphy suggest Early to Late  
650 Devonian kimberlite emplacement ages.

651         One of the Devonian-Paleogene age Mg-ilmenite grains with the lowest  $^{176}\text{Hf}/^{177}\text{Hf}$   
652 ratio from this group, was recovered just south of the Mackenzie River. This sample collected  
653 alluvium from the same watershed that hosts HOAM project kimberlites (< 10 km south; Fig.  
654 2). The other Devonian-Paleogene age Mg-ilmenite grain (17961-011) with a similar  
655  $^{176}\text{Hf}/^{177}\text{Hf}$  ratio within uncertainty (see Table S21) was collected approximately 120 km  
656 northwest of HOAM project kimberlites (near Ebutt Hills), parallel with regional ice-flow  
657 directions. Correlations between the two samples containing these younger age group Mg-  
658 ilmenites with the glacial transport direction (Fig. 2) and distance from the HOAM project  
659 kimberlites suggest these two grains are likely sourced from these kimberlites.

660

661 Southwestern and Western Slave Craton kimberlites

662

663         The Silurian Southwestern (Drybones Bay) and Ordovician Western Slave Craton  
664 kimberlites (Fig. 1) have emplacement ages consistent with the older age Mg-ilmenites (Fig.  
665 9), suggesting possible derivation from kimberlites from these general areas. KIM chemistry  
666 from surficial samples in the immediate area to the Drybones Bay kimberlite indicate other  
667 undiscovered kimberlites likely exist (Kerr et al. 2000). The older age Horn Plateau Mg-  
668 ilmenites, as well as some of the other KIMs, have similar major and minor chemistry to

669 grains from the Drybones Bay kimberlite and some of the Western Slave Craton surficial  
670 samples (Figs. 3, 4, 6, S8).

671 Present lithosphere thicknesses of 180-200 km, inferred from seismic tomography for  
672 the Southwestern and Western Slave Craton (Fig. 1; Schaefer and Lebedev 2014), are nearly  
673 identical to lithosphere thickness estimates based on Horn Plateau Cr-diopside (190-210 km;  
674 Fig. 7). Viable transport mechanisms (glacial and/or fluvial), distances (< 300 km) and  
675 directions (west-southwest), as well as KIM chemistry similarities suggest some of the Horn  
676 Plateau KIMs could be sourced from kimberlites in the Western and/or Southwestern Slave  
677 Craton areas. Further details of chemical traits between these possible sources, as well as  
678 discussion of viable transport mechanisms of KIMs can be found in the Supplementary  
679 Discussion.

680

681 Undiscovered Horn Plateau region kimberlites

682

683 Nine younger age Horn Plateau Mg-ilmenites are found on the Horn Plateau itself or in  
684 streams directly draining it where till domains with proximal clast lithology are indicated  
685 (Mills 2008 and references therein). In direct proximity to these samples near Willow Lake  
686 (Fig. S1), below the overburden, Aptian-Albian-aged sediments are preserved, which could  
687 host primary source kimberlites for these KIMs. Another viable region for primary source  
688 kimberlites for the younger age Horn Plateau KIMs are within Campanian-Maastrichtian  
689 clastic sediments, which lie below overburden on the southern portion of the Horn Plateau  
690 (Fig. S1). These two regions with preserved Cretaceous sediments have similar present-day  
691 lithospheric thicknesses (180-200 km based on seismic tomography; Fig. 1), in perfect  
692 agreement with the lithosphere thickness (190-210 km; Fig. 7) derived from Horn Plateau Cr-  
693 diopsides.

694           The largest quantities of analyzed peridotitic garnets, with extremely high proportions  
695 of Ni-in-garnet temperatures  $> 950$  °C, are found in samples proximal ( $< 25$  km) to these  
696 Cretaceous clastic sediments on the southern and eastern portions of the Horn Plateau (Fig.  
697 S13). Derived from the same watershed are 14 grains (from 13 samples) giving Al-in-olivine  
698 temperatures between 950-1100 °C and 17 Cr-diopsides grains (from 12 samples) with  
699 temperatures all  $> 950$  °C (Fig. S13). This area coincides with the north-trending Bulmer Lake  
700 gravity high (Fig. 1), which is suspected to be a Precambrian faulted contact that was re-  
701 activated during the Late Cretaceous. Seismic surveys also indicate dominant northeast-  
702 trending and subsidiary northwest-trending faults that were also likely re-activated during this  
703 time (Aitken 1993; Gal and Lariviere 2004 and references therein). Such faults may have  
704 facilitated kimberlite emplacement in the Horn Plateau region.

705           This same area on top of the Horn Plateau could also be a secondary (sedimentary)  
706 source region for the older age KIM population(s). In addition, sub-cropping below  
707 moderately thick glacial deposits (50-80 m; Cragie 1991), Cambrian-Silurian aged  
708 sedimentary rocks to the north-northeast of the Horn Plateau, near Lac la Martre (north of  $\sim$   
709  $62$  °N), could host primary kimberlites and/or be a secondary source of the older age Horn  
710 Plateau Mg-ilmenite group, as they lie directly up-ice of regional ice-flow directions ( $\sim 100$   
711 km northeast of the Horn Plateau).

712

713

## 714 **Conclusions**

715

716 Analysis of KIMs from two Central Mackenzie Valley study areas has resulted in a large, new  
717 geochemical dataset for an underexplored area of the Northwest Territories. These  
718 geochemical data provide new constraints on the diamond potential for future exploration  
719 within the area. The kimberlitic source of the KIMs from the Trout Lake region sampled

720 shallow portions of lithospheric mantle although no constraints on the depth of the lithosphere  
721 can be made from current results. In contrast, KIMs from the Horn Plateau region have  
722 significant diamond potential, with a relatively high proportion of harzburgitic diamond-facies  
723 (G10D) garnets and many of the analyzed KIMs deriving from the diamond stability field. A  
724 ca. 200 km thick lithosphere beneath the kimberlites sourcing the Horn Plateau kimberlites, as  
725 determined from clinopyroxene thermobarometry, is at least consistent with present-day  
726 seismology constraints (Fig. 1).

727         Single grain Hf isotope analyses of Mg-ilmenites from the Horn Plateau show a wide  
728 variety of isotopic compositions indicating at least two major age populations in their source  
729 kimberlites. These results are consistent with past major-element based studies of Horn  
730 Plateau KIMs that suggested the grains were likely sourced from at least two kimberlitic  
731 sources (one proximal and one distal) based on their grain morphology, major element  
732 chemistry and dispersion patterns (Day et al. 2007; Mills 2008). Comparative geochemistry  
733 and statistical tests, coupled with KIM surface textures (Mills 2008), till lithology domains  
734 (Mills 2008), pre-glacial bedrock distributions (Meijer Drees 1993; Duk-Rodkin et al. 1994)  
735 and ice-flow records (Fulton 1995), allow us to exclude some Slave Craton kimberlites  
736 (specifically the Lac de Gras field) as the only possible sources of Horn Plateau KIMs and  
737 instead highlight the southern/western Slave Craton area as a potential source for some of the  
738 KIMs. The KIM populations are, however, clearly mixed and permit more local sources. The  
739 incompletely resolved Quaternary history of the area remains a primary obstacle to locating  
740 the kimberlitic source(s) of the proximal, potentially “younger” Horn Plateau KIM  
741 population(s). Until Horn Plateau chromite, peridotitic garnet and olivine populations can be  
742 distinguished from other Southwestern and/or Western Slave Craton kimberlites, the potential  
743 for undiscovered kimberlites within the CMV remains possible and even likely.

744

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755

756

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1001 **Figure captions:**

1002 **Fig. 1** Bedrock geology map of Northwest Territories showing the Central Mackenzie Valley  
1003 (CMV) sample locations for both study areas. Horn Plateau = Horn Plateau study area,  
1004 Trout Lake = Trout Lake study area. Diamonds = kimberlites, dashed red line ellipse =  
1005 outline Bulmer Lake gravity high. Black line = line of lithosphere seismic tomography  
1006 cross-section (on bottom) from Schaefer and Lebedev (2014). See legend for rest of  
1007 symbology and Fig. S1 for more detailed bedrock geology map. The prominent Horn  
1008 Plateau in the centre of Horn Plateau (it rises 300-450 m above the surrounding plains)  
1009 broadly correlates to the Late Cretaceous sedimentary rock outline (green; see cross-  
1010 section below). Kimberlite locations after Pitman (2014) and NTGS GoData  
1011 Kimberlite Anomaly and Drillhole Data (KANDD). Note pink indicates cold  
1012 lithosphere for cross-section which is ~ 200 km thick lithosphere presently under most  
1013 of CMV and Slave Craton

1014 **Fig. 2** Surficial geology map of Northwest Territories showing the Central Mackenzie Valley  
1015 (CMV) sample locations for both study areas. Horn Plateau = Horn Plateau study area,  
1016 Trout Lake = Trout Lake study area. Diamonds = kimberlites. See legend for the rest  
1017 of symbology and Fig. S2 for more detailed surficial geology map. Surficial deposits  
1018 and glacial flowlines after Fulton (1995)

1019 **Fig. 3** Clinopyroxene  $\text{Cr}_2\text{O}_3\text{-Al}_2\text{O}_3$  discrimination plot (fields after Ramsey and Tompkins  
1020 1994). Grains from both study areas and 95 % contour interval fields for  
1021 clinopyroxene from Central and Western Slave Craton surficial samples (grey dashed  
1022 lines; data from NTGS GoData KIMC database). See legend for symbols. Note  
1023 peridotitic grains have  $\text{mg}\# > 0.88$

1024 **Fig. 4** Garnet  $\text{Cr}_2\text{O}_3\text{-CaO}$  classification plot (after Grütter et al. 2004). Grains from both study  
1025 areas and 95 % contour interval field for peridotitic garnets from CMV mineral claims  
1026 (purple and yellow dashed lines; data from Pitman 2014), Central and Western Slave

1027 Craton kimberlites (black dashed lines; data from variety of sources – see text) and  
1028 surficial samples (grey dashed lines; data from NTGS GoData KIMC database). Solid  
1029 red line = graphite diamond constraint (GDC). Double dashed green line =  $P_{38}$  Cr-in-  
1030 garnet barometer (both after Grütter et al. 2006). Note  $P_{38}$  estimates 5.8 GPa  
1031 (minimum) maximum pressure for Cr-rich garnets from the Horn Plateau

1032 **Fig. 5** Chondrite-normalized rare earth element ( $REE_N$ ) plots for Horn Plateau peridotitic  
1033 garnets. Chondrite normalization values after McDonough and Sun (1995). The  
1034 garnets display three distinct  $REE_N$  patterns: HREE-enriched, an extreme sinusoidal  
1035 MREE-depleted pattern and sinusoidal with variable re-enrichment. Mean values for  $n$   
1036  $< 10$ , median values for  $n > 10$ . Symbol groupings are based on garnet classification  
1037 after Grütter et al. (2004). Green triangles = G9, Blue squares = G10, Pink diamonds =  
1038 G10D, Red circles = G11 and Purple crosses = G12 grains

1039 **Fig. 6** Ilmenite  $TiO_2$ -MgO discrimination plot (after Wyatt et al. 2004). Grains from both  
1040 study areas and 95 % contour interval fields for Mg-ilmenites from Central and  
1041 Western Slave Craton surficial samples (grey dashed lines; data from NTGS GoData  
1042 KIMC database), as well as the Drybones Bay kimberlite (black dashed lines; data  
1043 from Schulze et al. 1995; Kerr et al. 2000). See legend for symbology

1044 **Fig. 7** FITPLOT geotherms (following methodology of Mather et al. 2011) and clinopyroxene  
1045 P–T data for both study areas and Central Slave Craton (data from Hasterok and  
1046 Chapman 2011) using Nimis and Taylor (2000) thermobarometry technique. See  
1047 legend for symbols. LDG = Lac de Gras, Solid red line = graphite diamond transition  
1048 curve (after Day 2012), Solid black line = mantle adiabat with mantle potential  
1049 temperature of 1350 °C. Upper and lower crust thicknesses for Horn Plateau was  
1050 constrained from the SNORCLE line 1 transect (Cook et al. 1999). Depth was  
1051 determined by multiplying pressure (GPa) by 3.15. Note the similarity between  
1052 Central Slave Craton and Horn Plateau Cr-diopside geotherms

1053 **Fig. 8** Mantle sampling profiles of garnet peridotite minerals from the Horn Plateau samples.  
1054 Left = single grain clinopyroxene barometer results (after Nimis and Taylor 2000).  
1055 Centre = projected Ni-in-garnet temperatures (averages of Griffin et al. 1989 and Canil  
1056 1999). Right = projected Al-in-olivine temperatures (after Bussweiler et al. 2017). Red  
1057 line = graphite-diamond transition curve (after Day 2012)

1058 **Fig. 9 Top** –  $^{176}\text{Hf}/^{177}\text{Hf}$  corrected ratios plot for MC-ICP-MS analyzed CMV kimberlitic  
1059 (Mg-)ilmenites. Outlined in respective coloured boxes are the two age groups  
1060 discussed in text: Yellow = Neoproterozoic-Silurian age group, Green = Devonian-  
1061 Paleogene age group. Error bars signify internal standard error ( $2\sigma$ ). **Bottom** –  
1062  $^{176}\text{Hf}/^{177}\text{Hf}$  corrected mean ratios for ilmenites from North American Craton  
1063 kimberlites, as well as those from the Udachnaya East kimberlite and Malaita alnöitic  
1064 breccia. Error bars signify standard deviation at 95 % confidence interval ( $2\sigma$ ).  
1065 Coloured boxes represent the Horn Plateau age groups. See Table S4 for kimberlite  
1066 localities, geochronology techniques used and sources of data

1067 **Fig. 10** Probability density plots for garnets, X-axis values assigned by logistic regression-  
1068 solutions for peridotitic garnet major element data between Horn Plateau grains and  
1069 those from various populations (see text for details). Increasing degrees of overlap  
1070 indicate increasing difficulty in reconciling population, while offset distributions  
1071 indicate population differences (see text for details)

Figure 1

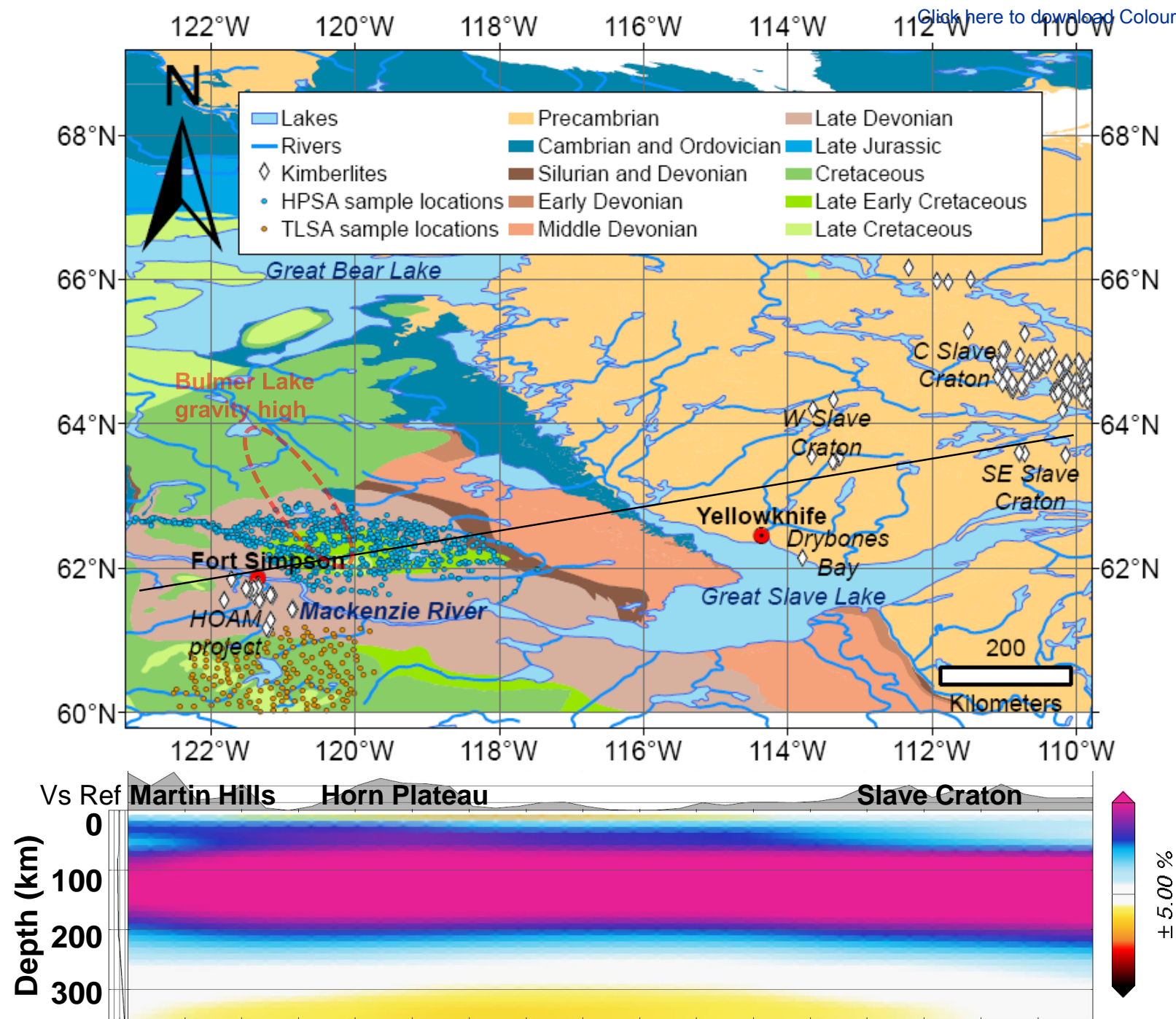
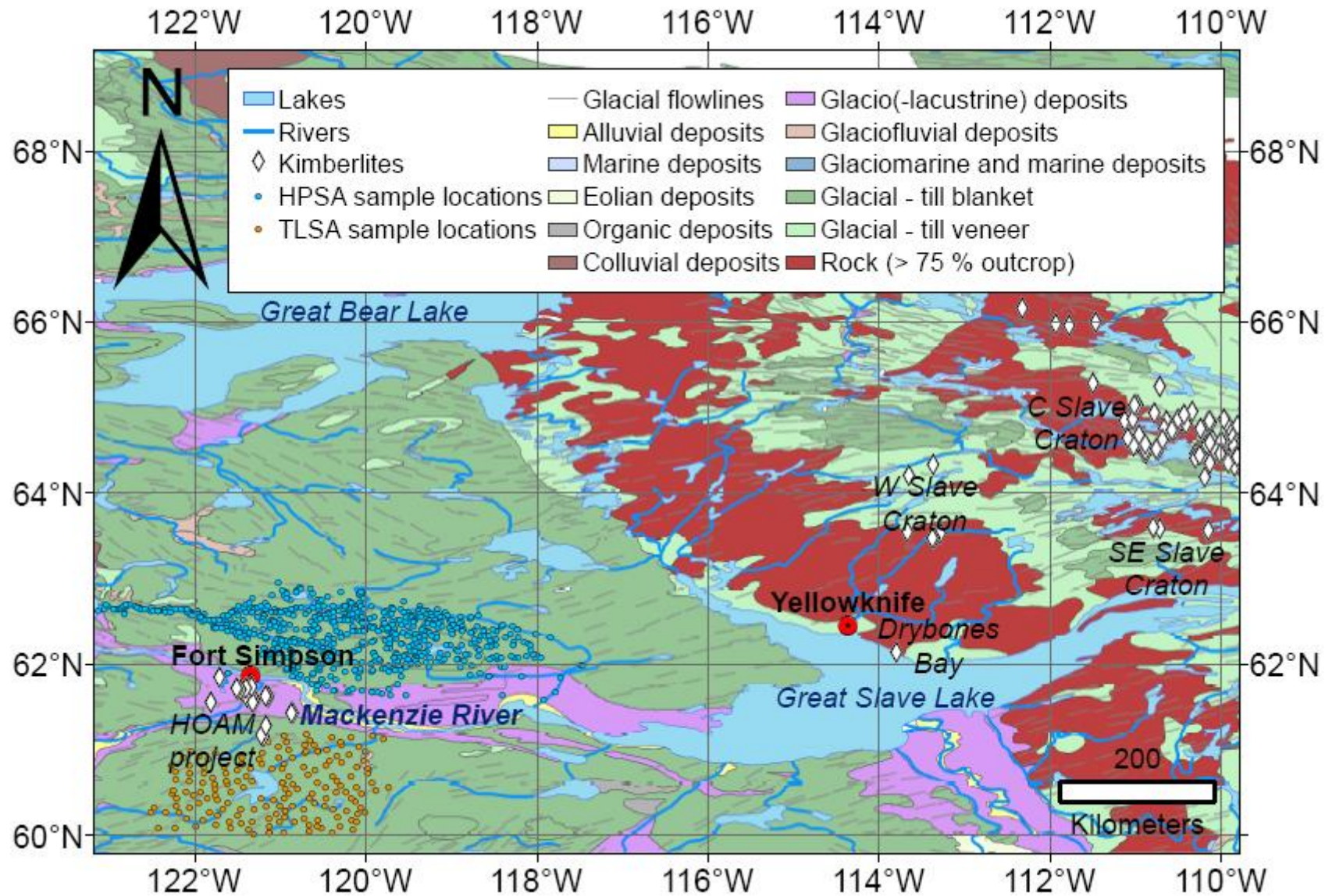


Figure 2



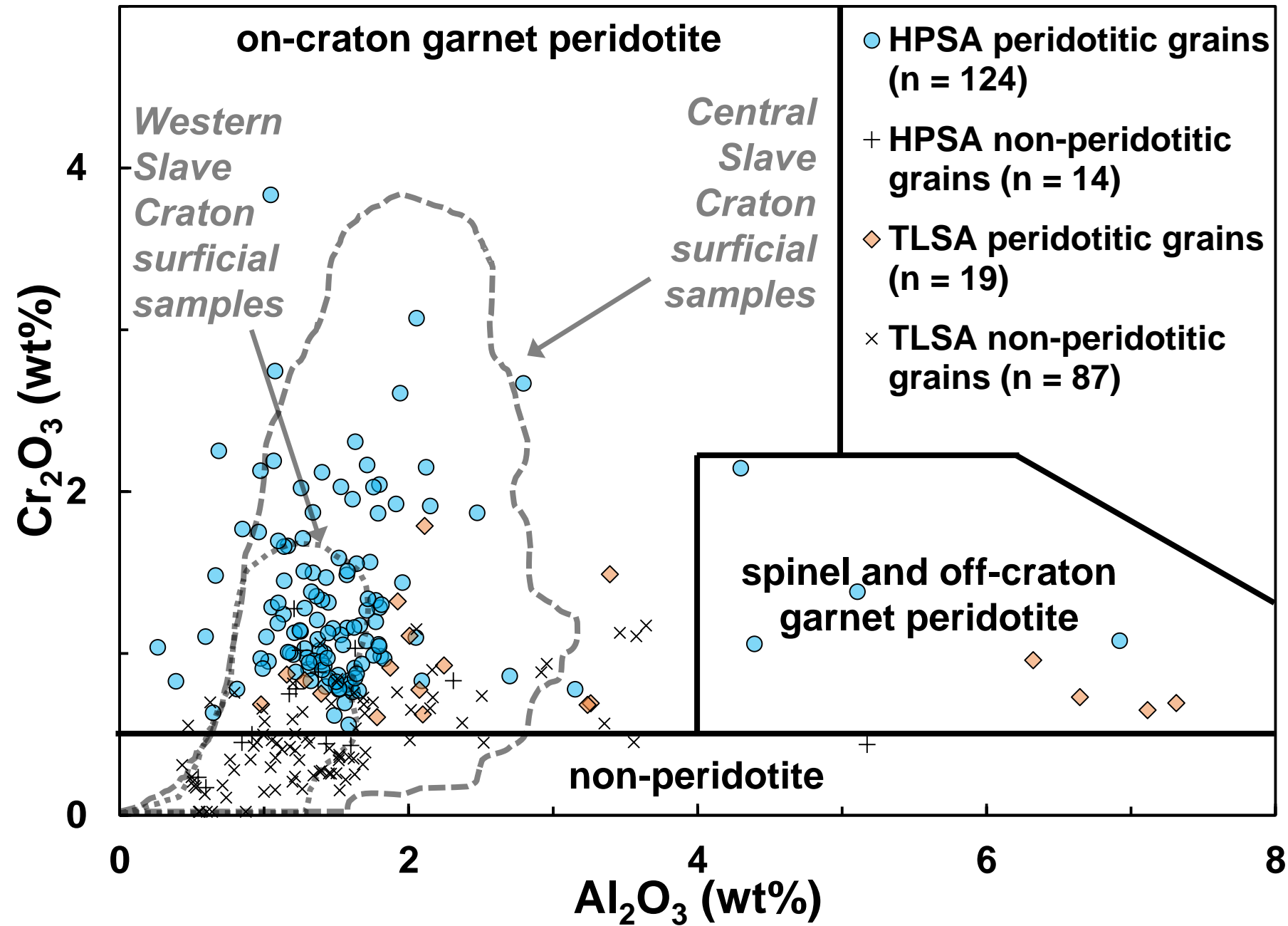


Figure 4

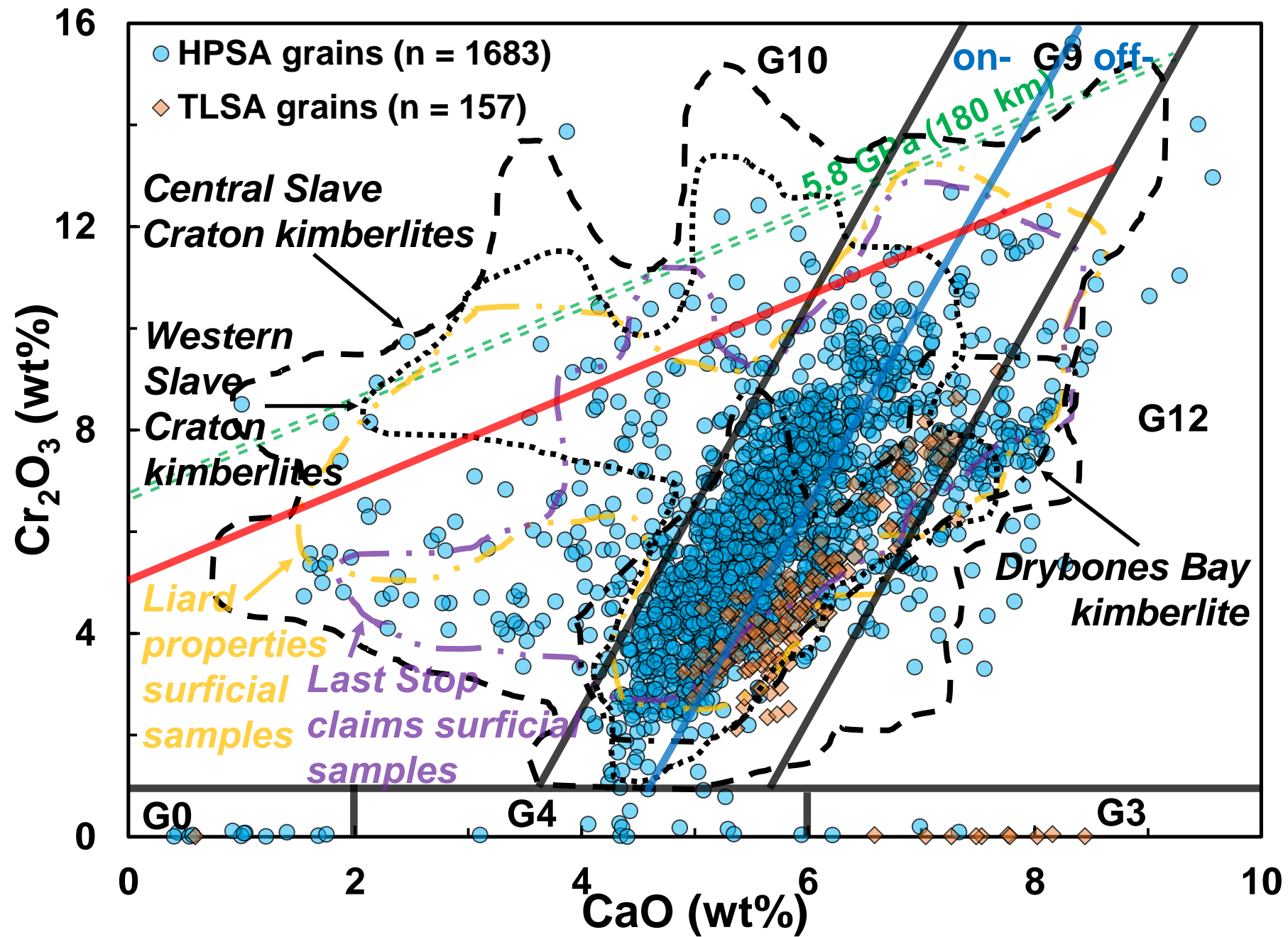


Figure 5

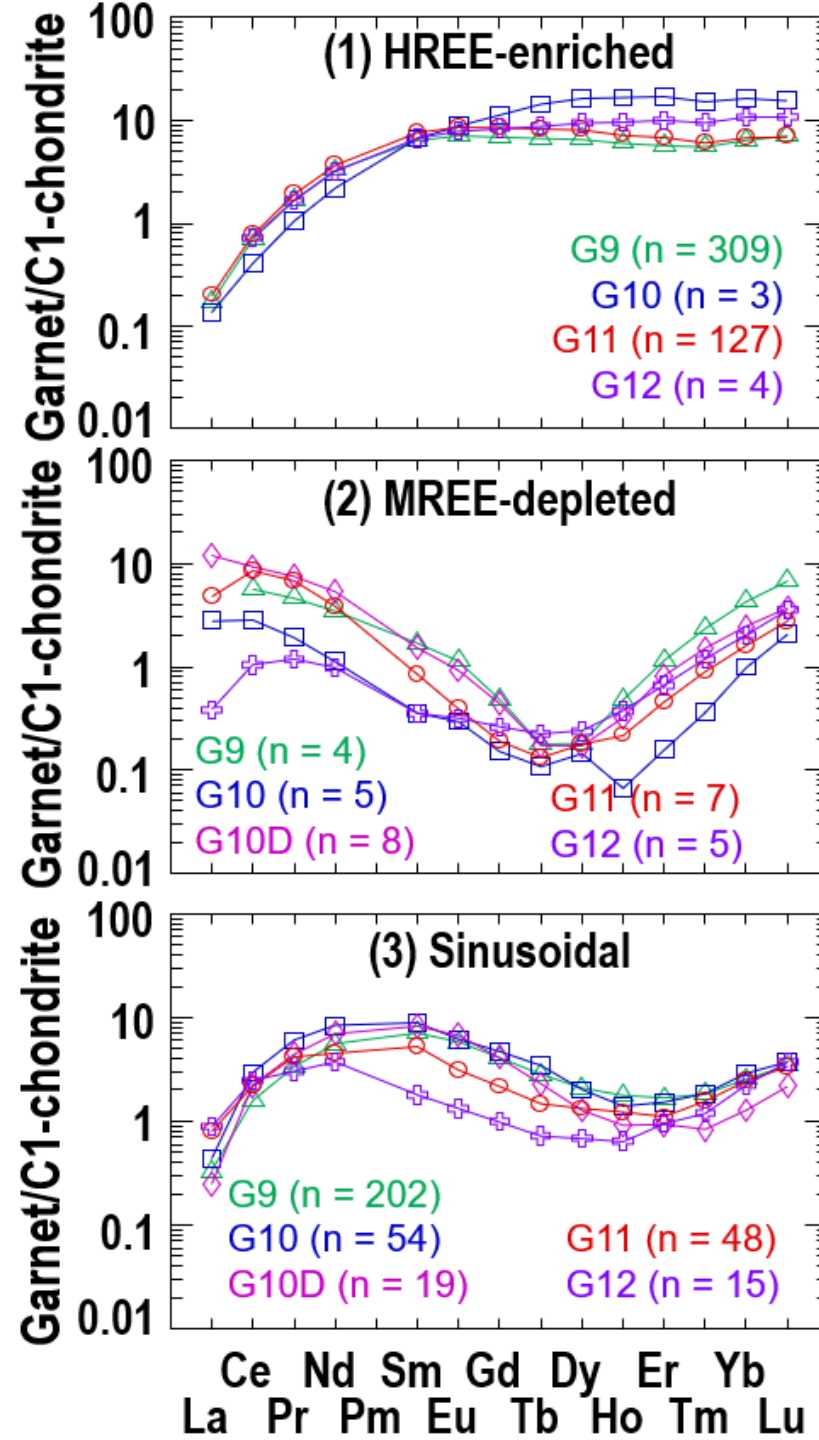
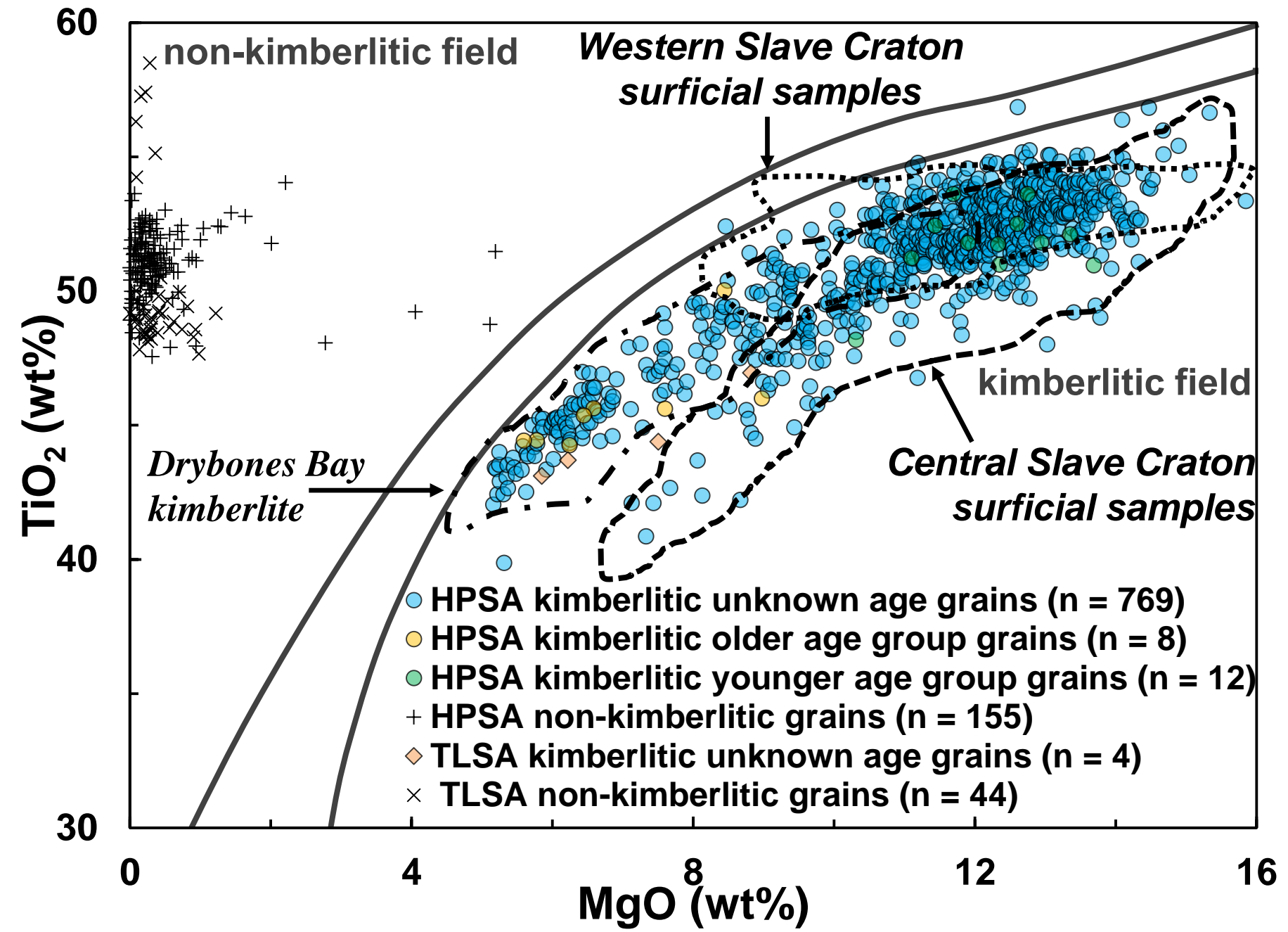




Figure 6



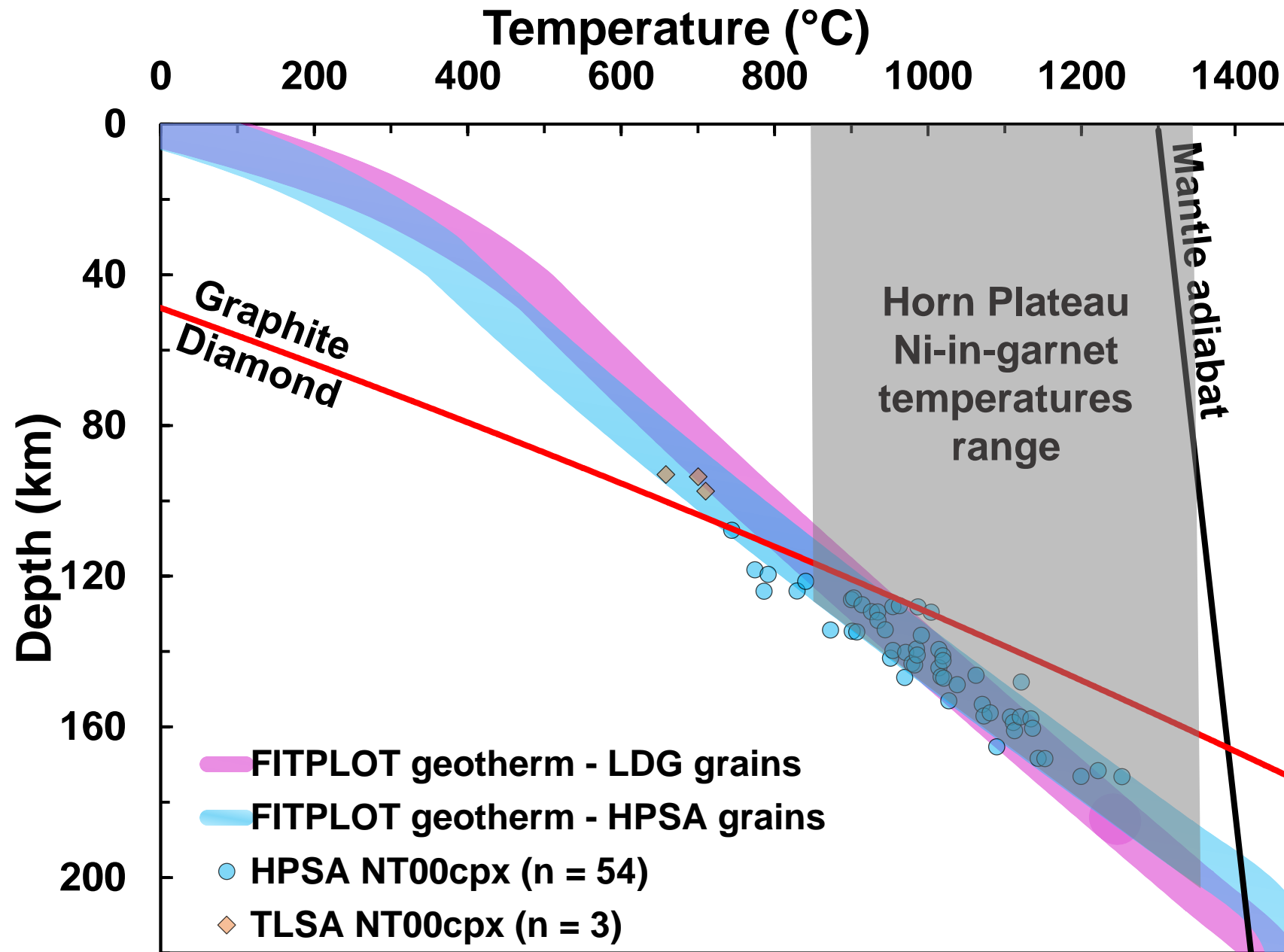


Figure 8

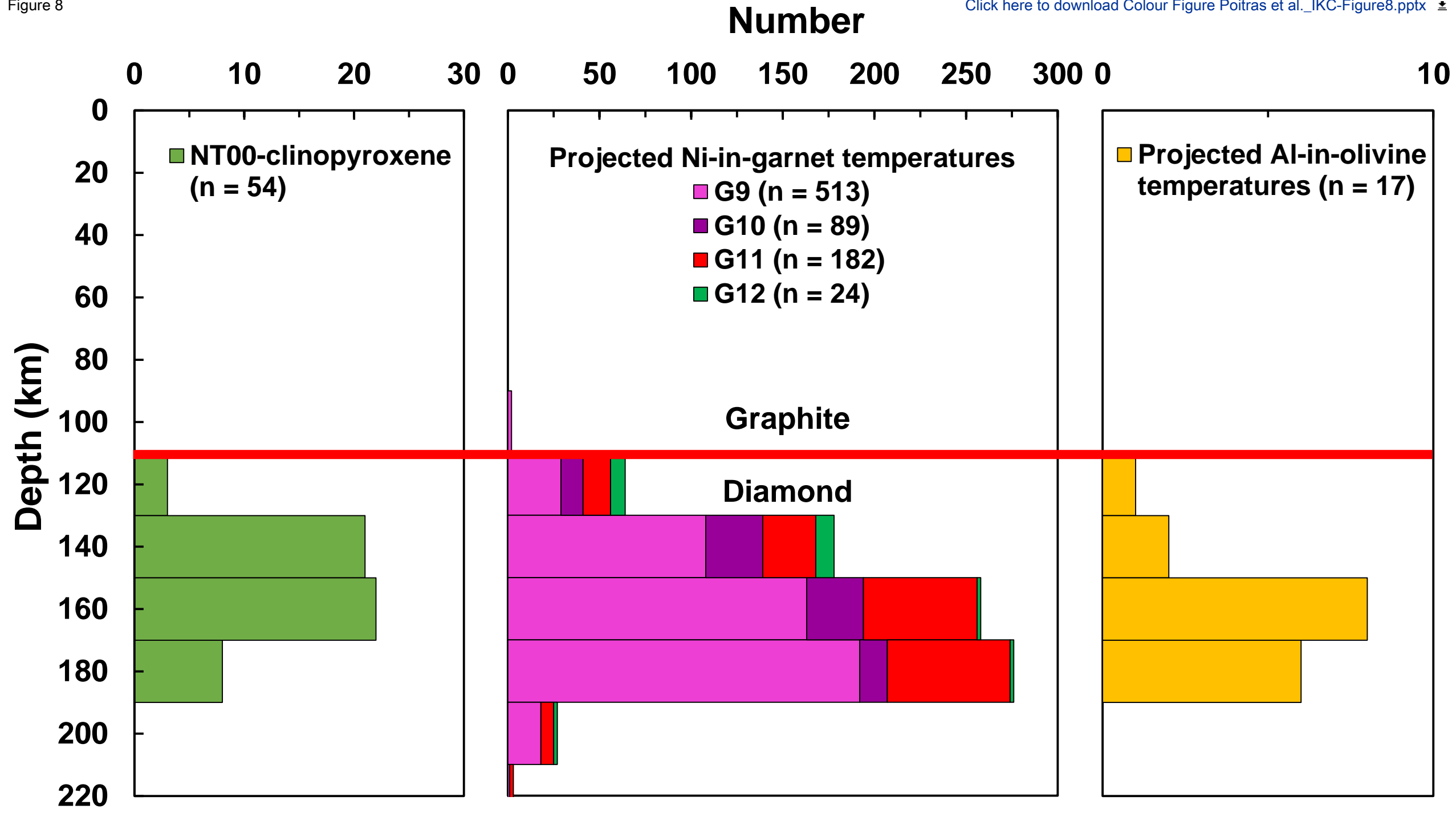


Figure 9

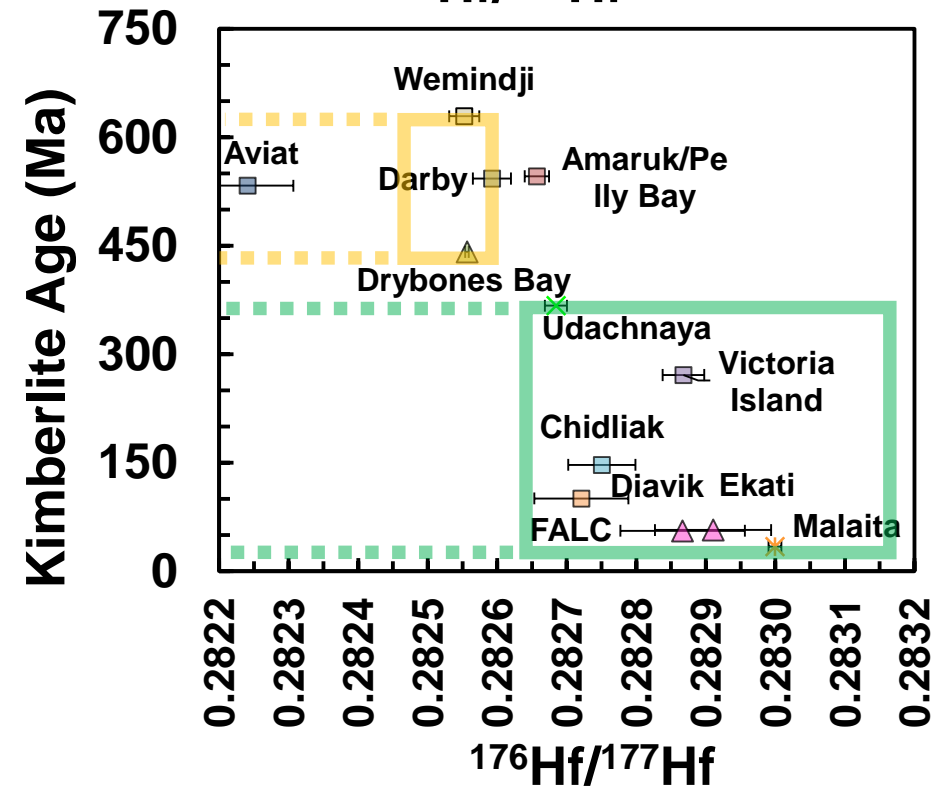
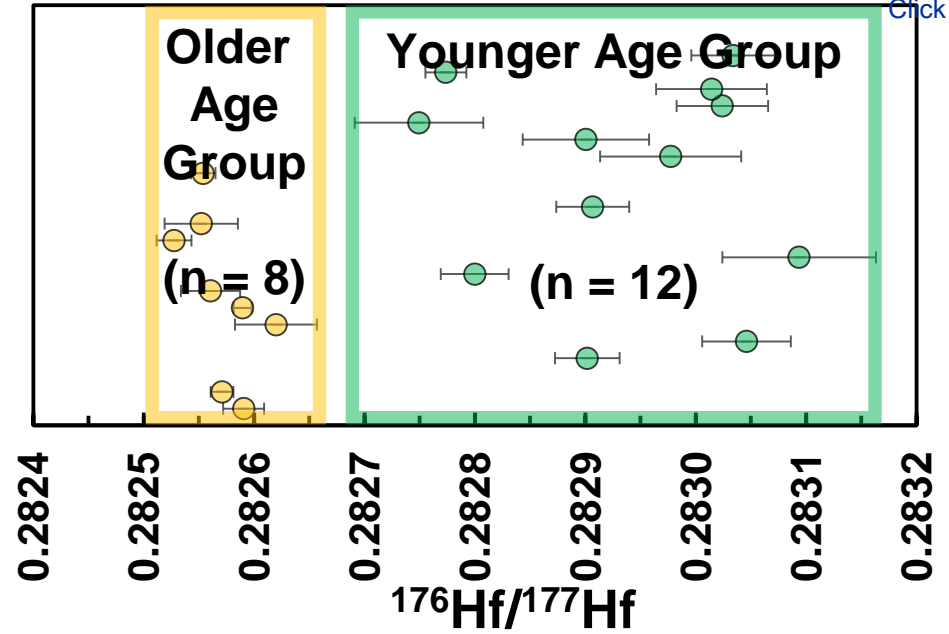


Figure 06

