

1 Radiogenic Isotopes Document the Start of Subduction in the Western Pacific

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ABSTRACT

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21 Subduction initiation is one of the least understood aspects of plate tectonics. In an effort to  
22 obtain the first *in situ* magmatic record of subduction initiation, the International Ocean  
23 Discovery Program Expedition 352 drilled at four sites in the inner trench wall of the Bonin  
24 Trench to recover 1.22km of oceanic upper crust accreted within a few m.y. of subduction  
25 initiation. The two sites nearer to the trench (U1440 and U1441) yielded axial and off-axis  
26 fore-arc basalts (FAB), while those c. 15km further from the trench (U1439 and U1442)  
27 yielded axial low-silica boninites and high-Mg andesites overlain by off-axis high-silica  
28 boninites. This study uses Hf-Nd-Sr-Pb isotope analysis from c. 50 stratigraphically  
29 representative core samples to trace the evolution of the mantle source during the brief period  
30 of FAB-through-boninite magmatism immediately following subduction initiation. Results  
31 show that: 1) the FAB have high  $\epsilon_{\text{Hf}}$  relative to  $\epsilon_{\text{Nd}}$  and were derived from variably depleted  
32 mantle of 'Indian' provenance with no detectable subduction input; 2) the axial boninites  
33 follow mixing trends between a residual FAB mantle source and a subduction component  
34 derived from shallow (amphibolite facies) melting of oceanic crust of 'Pacific' provenance;  
35 and 3) the off-axis boninites define mixing trends between a hybrid mantle wedge (residual  
36 mantle + slab melt) and an additional subduction component with lower  $\epsilon_{\text{Nd}}$  and higher  
37  $^{207}\text{Pb}/^{204}\text{Pb}$  that requires a significant contribution from pelagic sediment. This incoming of  
38 pelagic sediments may signify a change from an accretionary to non-accretionary margin as

39 subduction evolves. The results thus indicate a rapidly evolving system in terms of  
40 geodynamics, magma genesis and crustal accretion immediately following subduction  
41 initiation.

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43 Keywords: Radiogenic isotopes, Subduction initiation, Forearc basalts, Boninites,  
44 International Ocean Discovery Program (IODP) Expedition 352.

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## 47 **1. Introduction**

48

49 Unlike most types of plate tectonic process, regional-scale subduction initiation is not taking  
50 place at the present-day. In consequence, subduction initiation is one of the least understood  
51 aspects of plate tectonics. Much of our limited geological knowledge to date derives from the  
52 Izu-Bonin-Mariana (IBM) outer forearc, which carries a full record of magmatic and tectonic  
53 activity from the start of subduction in the Eocene to the start of normal arc volcanism some  
54 10 m.y. later (e.g. Ishizuka et al., 2011). It also hosts the type localities for the two rock types  
55 characteristic of subduction initiation and fore-arc terranes: forearc basalts (FAB), which are  
56 MORB-like tholeiitic basalts (Reagan et al., 2010), and boninites, which are distinctive high-  
57 Si, high-Mg, low-Ti volcanic rocks (e.g. Crawford, 1989).

58

59 The presence, and close association in space and time, of FAB and boninites during the birth  
60 of the IBM system have provided key supporting evidence for geodynamic models of

61 subduction initiation. Notably, the presence of FAB supports the original Stern and Bloomer  
62 (1992) hypothesis that subduction initiation was immediately succeeded by subsidence and  
63 roll-back of the embryonic subducted slab resulting in an episode of near-trench extension or  
64 sea-floor spreading. In addition, the presence of boninites supports the need, in such a model,  
65 for a period of anomalous magmatism between initial subsidence and rollback and the  
66 establishment of the stable down-dip subduction required for normal island arc volcanism.  
67 The type of model proposed by Stern and Bloomer has since been supported by both  
68 ophiolite studies (e.g. Shervais, 2001; Dilek and Flower, 2003) and numerical experiments  
69 (e.g. Leng et al., 2012).

70

71 Despite much support for this general model, there is, however, much debate over the detail.  
72 For the starting conditions, Stern and Bloomer (1992) base their model on the reconstruction  
73 of Hussong and Uyeda (1981) and others in which subduction initiation began at a pure  
74 strike-slip, transform plate boundary, a setting that continues to be supported by tectonic  
75 reconstructions (e.g. Wu et al., 2016). However, Casey and Dewey (1994) argue for a  
76 transtensional boundary (a 'leaky' transform fault), which, if correct, may have led to  
77 episodes of ridge subduction following subduction initiation. Loci not involving a transform  
78 fault have also been proposed. These include a thermal anomaly in the mantle (Macpherson  
79 and Hall, 2001), the edge of an oceanic plateau (Niu et al., 2003) and a Mesozoic continental  
80 margin (Ishizuka et al. 2018).

81

82 Moreover, although the concept of slab roll-back as the cause of spreading, and hence the

83 mechanism for FAB genesis, is supported by numerical models, it still requires ground-  
84 truthing. To do this, and to constrain better the pre-subduction tectonic setting, one of the  
85 optimal approaches is to establish the extent, stratigraphy and genesis of the volcanic rocks  
86 erupted immediately following subduction initiation. Because much of this evidence is  
87 sediment-covered, drilling is needed. Thus, in 2014, IODP Expedition 352 to the Bonin  
88 forearc had the objective of documenting the changing nature and composition of the crust  
89 formed immediately after subduction initiation (Reagan et al., 2015). Here, we present the  
90 first radiogenic isotope (Hf-Nd-Sr-Pb) data from the Expedition, and use these data to 1)  
91 establish the evolution of mantle sources and slab fluxes following subduction initiation and  
92 2) use the resulting information to test, and develop new, hypotheses for the plate  
93 configurations and driving forces involved.

94

## 95 **2. Materials**

96

97 IODP Expedition 352 drilled at four sediment-covered sites between the ophiolitic sequence  
98 exposed in the inner trench wall in the east and the Bonin Ridge embryonic volcanic arc in  
99 the west (Fig. 1 a-b). These subdivide into two deeper, more trench-proximal sites (U1440  
100 and U1441: the ‘FAB sites’) and two shallower, trench-distal sites (U1439 and U1442: the  
101 ‘boninite sites’). Representative basement samples taken on-board the drilling ship from each  
102 of the petrologic units defined and described at these four sites (Reagan et al., 2015) form the  
103 basis for this isotope study.

104

105 Fig. 1c summarizes the stratigraphies for the holes with the deepest penetration, namely  
106 U1440B, U1441A, U1439C, and U1442A. All four recovered lavas, but one of the FAB  
107 holes (U1440B) and one of the boninite holes (U1439C) also rooted in sheeted intrusions  
108 (Reagan et al., 2015). The probable explanation is that upper oceanic crust was penetrated at  
109 both FAB and boninite sites, the former representing crust in the more trench-proximal  
110 location. Lava sequences at FAB Site U1440 and boninite Sites U1439 and U1442 have been  
111 divided into lower and upper units, each with a number of sub-units. In both FAB and  
112 boninite sites, the lower units compositionally match the underlying dikes and henceforth will  
113 be termed ‘axial’, while the upper units are compositionally distinct and will be termed ‘off-  
114 axis’.

115  
116 According to bulk rock Ar-Ar, and CA-TIMS zircon U-Pb, dating (Reagan et al., 2019), the  
117 FAB and boninites in this paper formed from an intermediate-rate, sea-floor spreading event  
118 that took place between approximately 52 and 50Ma immediately following subduction  
119 initiation (Fig. 1c). The Bonin Ridge boninites and their differentiates subsequently erupted  
120 onto this oceanic crust between about 50 and 44 Ma to form an embryonic arc, followed by  
121 normal tholeiitic and calc-alkaline arc magmatism (e.g. Cosca et al., 1998, Ishizuka et al.,  
122 2011).

123  
124 In terms of rock type, the FAB are tholeiites and resemble MORB in all but setting and the  
125 greater degree of depletion of their mantle source. Shervais et al. (2019) subdivide axial- and  
126 off-axis FAB into a main, ‘normal’ group entitled N-FAB, less common depleted and

127 enriched groups respectively entitled D-FAB and E-FAB, as well as primitive variants of  
128 these not distinguished here. The axial boninites are predominantly made up of low-Si  
129 boninites accompanied by their fractionation products, high-Mg andesites (HMA), while off-  
130 axis boninites are primarily made up of high-Si boninites (Reagan et al., 2015, 2017).

131

132 The preliminary model of Reagan et al. (2017) produces the FAB by decompression melting  
133 with little to no slab flux. Melting of an extremely depleted (harzburgitic) mantle flushed  
134 with fluids and/or melts from subducted sediments and oceanic crust then generated  
135 boninites. Shervais et al. (2019) refine the petrogenesis of the FAB, providing evidence that  
136 their mantle source had higher potential temperatures than normal MORB but was more  
137 depleted due to an episode of prior, garnet-facies depletion.

138

### 139 **3. Methods**

140

141 Table 1 provides a subset of the data used in the geochemical plots, and the text below  
142 provides a brief summary of analytical methods and quality controls. Appendix A contains  
143 sample locations and the full element and isotope data, set and Appendix B gives full  
144 analytical detail.

145

#### 146 *3.1 Preparation procedure*

147

148 Following crushing and agate milling, we analysed these samples for major elements by XRF

149 at Utah State University and trace elements by ICP-MS at Guizhou Tongwei Analytical  
150 Technology Co. Ltd. (Appendix B, Section 2). We heavily leached altered samples for  
151 analysis of isotopes of the alteration-mobile elements Sr and Pb (see Appendix B, Sections  
152 3.3 and 5 for leaching procedures), but some analyses retained evidence of metasomatism,  
153 with Pb isotopes in FAB particularly affected. In such cases, if there was fresh glass or rock  
154 nearby, we reanalysed Pb isotopes using the fresh material.

155

### 156 *3.2 Analytical methods.*

157

158 We analysed Hf isotope ratios at the Guangzhou Institute of Geochemistry, Chinese  
159 Academy of Sciences (GIGCAS) on a Neptune MC-ICPMS (Appendix B, Section 3). We  
160 took particular care to check for complete sample dissolution and hence avoid the potential  
161 problem of small, residual zircons (Appendix B: Section 7 and Table B.5). We report the Hf  
162 isotopic ratios relative to a  $^{176}\text{Hf}/^{177}\text{Hf}$  ratio of 0.282189 for JMC14374 (corresponding to a  
163 value for JMC475 of 0.282158).

164

165 We also carried out neodymium isotope ratio analyses at GIGCAS on the Neptune MC-  
166 ICPMS (Appendix B, Section 3). We report the Nd isotopic ratios relative to  $^{143}\text{Nd}/^{144}\text{Nd}$  of  
167 JNdi-1=0.512115. We carried out further Nd isotope analyses in the MAGIC laboratories at  
168 Imperial College London with a Nu Instruments HR MC-ICPMS, also using neodymium  
169 standard, JNdi. (Appendix B, Section 4). Analyses of USGS reference materials, BCR-2,  
170 BIR1a and BHVO-2, run as unknowns demonstrate excellent agreement both between the



171 two laboratories and with accepted values (Appendix B, Table B.3). We determined Sr  
172 isotope ratios at both GIGCAS and the MAGIC using a Thermo Finnigan Triton thermal  
173 ionization mass spectrometer (TIMS) (Appendix B, Sections 3 and 4). Both laboratories  
174 report the Sr isotopic ratios relative to  $^{87}\text{Sr}/^{86}\text{Sr}$  of NBS-987 = 0.710248. Appendix B, Table  
175 B.2 gives international standard data obtained by the two laboratories.

176

177 We measured Pb isotope ratios on a Thermo Neptune MC-ICP-MS at the University of  
178 Southampton UK, using a double spike to correct for instrumental mass fractionation  
179 (Appendix B, Section 5). Procedural blanks range between 30-95 pg Pb. NBS SRM 981  
180 values achieved during the measurement period were  $^{206}\text{Pb}/^{204}\text{Pb}=16.9404\pm 32$ ,  $^{207}\text{Pb}/^{204}\text{Pb} =$   
181  $15.4969\pm 32$ ,  $^{208}\text{Pb}/^{204}\text{Pb} = 36.7149\pm 90$  (2s.d.; n=44). Propagated uncertainties for combined  
182 natural sample and spiked sample analyses were always less than the reproducibility  $2\sigma$  of  
183 NBS SRM 981.

184

185 Several diagrams in this paper combine our data with published data from other laboratories.  
186 Only Nd isotopes exhibited significant inter-laboratory variation. All data used in this paper  
187 have thus been normalized to  $^{143}\text{Nd}/^{144}\text{Nd}$  of 0.511858 for the La Jolla standard, which is  
188 equivalent to the JNdi value of 0.512115. We provide checks for internal consistency of these  
189 various data sets in Appendix B, Section 6. In the text that follows, we describe and explain  
190 the principal isotopic features under 'Results' and the more detailed isotopic component  
191 modeling under 'Interpretations'.

192

193 **4. Results**

194

195 *4.1 Isotope stratigraphy*

196

197 Fig. 2 gives the isotope stratigraphies for the four principal Exp. 352 holes. Off-axis FAB  
198 from U1440B have distinctively high  $\epsilon\text{Hf}_i$ , low  $\epsilon\text{Nd}_i$  and high  $^{206}\text{Pb}/^{204}\text{Pb}$ . The axial-FAB  
199 (both lavas and dikes) are predominantly homogenous N-FAB compositions. Exceptions are  
200 the single analysed sample of E-FAB, which has distinctly lower  $\epsilon\text{Nd}_i$  and slightly lower  
201  $\epsilon\text{Hf}_i$ , and the single analysed sample of D-FAB, which has distinctly higher  $\epsilon\text{Nd}_i$  and slightly  
202 higher  $\epsilon\text{Hf}_i$ . There is no significant difference between the axial N-FAB compositions from  
203 the two FAB holes.

204

205 Off-axis boninites have lower  $\epsilon\text{Nd}_i$  and lower or similar  $\epsilon\text{Hf}_i$  compared to axial-boninites, but  
206 higher or similar Sr and Pb isotope ratios. The axial-boninites are more variable than the  
207 axial-FAB, likely reflecting their more complex genesis, which includes magma mingling  
208 (Reagan et al., 2015). There is no significant compositional difference between the off-axis  
209 boninites from the two holes, but there are small differences between axial-boninites (e.g.  
210 lower average  $^{206}\text{Pb}/^{204}\text{Pb}$  and higher average  $\epsilon\text{Nd}_i$  in Hole U1442A). Compared to the FAB,  
211 all boninites have lower  $\epsilon\text{Hf}_i$  and higher  $^{206}\text{Pb}/^{204}\text{Pb}$ .

212

213 *4.2 Principal Hf, Nd, Sr and Pb element characteristics*

214

215 Figure 3 depicts chondrite-normalized REE patterns extended to include Pb, Sr and Hf in  
216 positions that maintain the order of incompatibility during melting of the upper mantle.  
217 Appendix A provides the full data set used for these patterns and for the means and standard  
218 deviations reported below. It also provides values for other petrogenetically significant trace  
219 elements,

220

221 For the FAB sites, mantle sources were variably depleted, as reflected in the variable slopes  
222 of the REE patterns in Fig. 3a. Of the axial-FAB, the E-FAB sample has the least LREE-  
223 depletion ( $Ce/Yb=2.46$ ) and D-FAB ( $Ce/Yb=0.96$ ) the greatest. The axial N-FAB have  
224 intermediate levels of depletion ( $Ce/Yb=1.50\pm 0.13[1\sigma]$ ) as do off-axis N-FAB  
225 ( $Ce/Yb=1.63\pm 0.05[1\sigma]$ ). However, all the FAB are significantly more depleted than average  
226 N-MORB ( $Ce/Yb=3.8\pm 0.2[1\sigma]$ ) and D-MORB ( $Ce/Yb=3.0\pm 0.2[1\sigma]$ ) (data of Gale et al.,  
227 2013). Strontium and Pb show no selective element enrichments that might indicate a  
228 subduction component, provided the patterns are restricted to glass and the least-altered  
229 rocks.

230

231 For the boninite sites (Fig. 3b-c), all samples are significantly enriched in Sr and Pb relative  
232 to the surrounding REE, a characteristic of all arc-related lavas. Unlike most arc lavas,  
233 however, they exhibit pronounced positive Hf anomalies, features commonly associated with  
234 subduction initiation (e.g. Hickey-Vargas, 1989; Pearce et al., 1999). Their low  
235 concentrations of the subduction-immobile HREE indicate very depleted mantle sources,  
236 significantly more depleted than FAB sources. Their LREE enrichments

237 Ce/Yb=3.05±0.44[1σ]) are, however, greater than most FAB, so requiring modification of  
238 this depleted mantle by a LREE-enriched subduction component. Thus Nd must have been  
239 subduction-mobile during boninite genesis. Relative to Sr, Pb and Hf, however, chondrite-  
240 normalized Nd concentrations are low, and hence the mobility of Nd during subduction must  
241 also have been comparatively low.

242

243 The difference between axial-boninites (Fig. 3b) and off-axis boninites (Fig. 3c) is less  
244 obvious than that between FAB and boninites. Primarily, the off-axis boninites have the  
245 larger Pb, Sr and Hf peaks, lower and variable concentrations of the HREE for a given MgO  
246 content, and slight LREE enrichment (Ce/Yb=3.82±1.23[1σ]) rather than depletion. These  
247 features indicate that the off-axis boninites likely experienced greater degree of mantle  
248 depletion and a greater contribution from a subduction component containing all daughter  
249 elements (Pb, Sr, Hf and Nd).

250

#### 251 *4.3 Principal Sr, Nd, Hf, Pb isotope ratio covariations*

252

253 Fig. 4 summarizes the radiogenic isotope evolution from 1) axial-FAB to off-axis FAB which  
254 defines the ambient mantle before subduction and then 2) from FAB to axial-boninite which  
255 defines the first stage of subduction and finally 3) from axial-boninite to off-axis boninite  
256 which defines the beginning of embryonic arc.

257

258 For 1), both the  $\epsilon\text{Hf}-\epsilon\text{Nd}$  plot (Fig. 4a) and the  $^{208}\text{Pb}/^{204}\text{Pb}-^{206}\text{Pb}/^{204}\text{Pb}$  plot (Fig. 4d) are  
259 effective discriminants between mantle from ‘Pacific’ and ‘Indian’ mantle domains. On these  
260 plots, all FAB clearly plot well within the ‘Indian’ domain, a well-known characteristic of all  
261 basalts erupted within the Philippine Sea Plate (e.g. Hickey-Vargas, 1998).

262

263 For 2), the key characteristic is the decrease in  $\epsilon\text{Hf}$  and  $^{208}\text{Pb}/^{204}\text{Pb}$  from all FAB to axial-  
264 boninite. These features have already been identified in rocks from the nearby Bonin Ridge  
265 and attributed to the interaction of melt of ‘Pacific’ provenance from the newly-subducting  
266 Pacific Plate with overlying ambient mantle of ‘Indian’ provenance (Pearce et al., 1992; Li et  
267 al., 2013).

268

269 For 3), the key characteristic is the subsequent decrease in  $\epsilon\text{Nd}$  accompanied by an increase  
270 in Sr and Pb isotope ratios from axial to off-axis boninites. This indicates the incoming of a  
271 further subduction component containing a significant contribution from Pacific Pelagic  
272 Sediment (PPS).

273

274 Fig. 4 thus highlights a remarkable variation over such a small range in space (c.15km: Fig.  
275 1) and time (< 2m.y.: Reagan et al., 2019), and perhaps holds the key to understanding the  
276 complex and rapid evolution of ophiolite complexes believed to have a subduction initiation  
277 origin. In the Sections that follow, we interpret the trends in Fig. 4 in more detail in an  
278 attempt to establish how they may take place in the context of Western Pacific subduction  
279 initiation.

280

## 281 **5. Interpretations**

282

### 283 *5.1 Characterizing the mantle source using $\epsilon_{\text{Hf}}$ - $\epsilon_{\text{Nd}}$ covariations*

284

285 Applying the  $\epsilon_{\text{Hf}}$ - $\epsilon_{\text{Nd}}$  projection (Fig. 4a) to FAB is a useful way to understand the nature  
286 and provenance of the ambient mantle at the start of subduction. Fig. 5a depicts MORB and  
287 OIB from the Western Pacific region. Our reference line is the Indian-Pacific mantle domain  
288 boundary (Pearce et al., 1999). This boundary has the equation  $\epsilon_{\text{Hf}} = \epsilon_{\text{Nd}} * 1.6$  and hence is  
289 almost parallel to, but displaced by c. 1.25 epsilon units from, the principal axis of dispersion  
290 of the global terrestrial array ( $\epsilon_{\text{Hf}} = \epsilon_{\text{Nd}} * 1.59 + 1.26$ ) of Chauvel et al. (2009). The boundary  
291 also bisects the type Pacific-Indian mantle domain boundary from the Australian-Antarctic  
292 Discordance (AAD), a long distance further to the south-east (Kempton et al., 1992).

293

294 As many authors have noted, individual lava suites commonly form arrays (termed ‘ambient  
295 mantle arrays’ by Woodhead et al., 2012), which typically run sub-parallel to the Pacific-  
296 Indian boundary and global terrestrial array. Behind the trench, the most relevant ambient  
297 array is that of the Philippine Sea (Fig. 5a), which began spreading at about the time of  
298 subduction initiation and continued through the evolution of the Bonin Ridge embryonic arc.  
299 Its ambient array lies well within the ‘Indian’ domain.

300

301 In Fig. 5b, the axial-FAB plot in the upper 50-90 percentile of the Philippine Sea MORB-OIB

302 array. They form an ambient FAB mantle array, which extends from E-FAB, through N-FAB  
303 to D-FAB and runs sub-parallel to other ambient mantle arrays within the Philippine Sea  
304 Plate. In contrast, the off-axis FAB form a near-vertical trend, which extends to  $\epsilon\text{Hf}$ - $\epsilon\text{Nd}$   
305 values above the Philippine Sea array.

306

307 ‘Indian’ characteristics, such as those exhibited by Exp. 352 FAB, have been attributed to  
308 mantle depletion during a partial melting event in garnet facies (as garnet fractionates Lu  
309 from Hf, but not Sm from Nd) perhaps coupled with enrichment by a Nd-rich, Hf-poor  
310 subduction component (e.g. Kempton et al, 2002; Janney et al., 2005, Salters et al., 2011).  
311 This interpretation supports the hypothesis of Gurnis et al. (1998) that the ‘Indian’ domain  
312 province contains relics of subduction-modified, depleted lithosphere from a long history of  
313 sub-Gondwana subduction.

314

315 Yogodzinski et al. (2018) explore further the evolution of the mantle beneath the Philippine  
316 Sea Plate at about the time of subduction initiation by focusing on the 50-49Ma basaltic  
317 basement at Site U1438. They use plots of  $^{176}\text{Hf}/^{177}\text{Hf}$  v Lu/Hf and  $^{143}\text{Nd}/^{144}\text{Nd}$  v Sm/Nd to  
318 demonstrate that the mantle source experienced a long history of depletion, which includes a  
319 significant melt extraction event in the 400-500 Ma time window.

320

321 On their base diagram (Fig. 5c-d), we note that our main N-FAB magmas extend along a  
322 52Ma errorchron in Fig. 5c, indicating fractionation of Lu/Hf during the melting event from  
323 which they formed. The D-FAB and E-FAB form two-point 195Ma and 235Ma errorchrons

324 for the Lu-Hf and Sm-Nd systems respectively, a potentially significant observation but one  
325 requiring further analyses. The mean N-FAB composition and the off-axis FAB plot close to  
326 the Palaeozoic errorchrons plotted by Yogodzinski et al. (2018) for IODP Site U1438: the  
327 combined basalt dataset from Sites U1438 and U1439-1442 gives similar 450Ma errorchrons  
328 on both diagrams (Fig. 5c-d), though with considerable scatter. Overall, therefore, our isotope  
329 data support the concept of Yogodzinski et al (2018) and Shervais et al. (2019) that  
330 subduction initiated within a mantle domain that had already experienced one or more ancient  
331 depletion events, though more work is needed to ascertain precise ages and details. In any  
332 event, this mantle domain provides the mantle wedge end-member needed to explain the  
333 compositions of the subsequent subduction events.

334

### 335 *5.2 Characterizing the subduction component using isotope-element ratio covariations*

336 Plots of the form  $\epsilon_{\text{Hf}_i-x/\text{Hf}}$  are particularly effective at interpreting Hf anomalies on the  
337 extended REE diagram (Fig. 3b-c) in terms of subduction components, as any mixing lines  
338 are then linear and so easier to interpret (e.g. Barry et al., 2006).

339

340 In Fig. 6a,  $x=\text{Sm}$ , chosen because Sm is closest to Hf in its bulk partition coefficient during  
341 mantle melting and fractional crystallization. The axial-FAB data plot at the upper (MORB)  
342 end of the ambient Philippine Sea array, while the off-axis FAB lie on an extension of this  
343 array. As can be seen in all four diagrams (Fig. 6a-d), an average axial-FAB composition is  
344 the optimum choice for the isotopic mantle end-member for the subsequent generation of  
345 boninites. Being a boninite source, this end-member is depleted in incompatible elements



346 compared to the FAB mantle source (Fig. 3) but, importantly, it is isotopically similar. We  
347 term this end-member 'Residual FAB Mantle' (M in Figs 6-9).

348

349 During subduction in most active present-day arcs, Sm is more mobile than Hf. In  
350 consequence, arc lavas typically lie within, or adjacent to, the high Sm/Hf (or Sm/Zr) side of  
351 any MORB array (e.g. Hickey-Vargas, 1989; Barry et al., 2006). In contrast, samples from  
352 the boninite sites plot towards lower Sm/Hf, reflecting their positive Hf anomalies. They have  
353 been interpreted as reflecting an amphibolite-facies slab-melt component in which Sm is  
354 retained by amphibole, while the slab-top temperature is high enough to dissolve sufficient  
355 zircon in the slab-derived melt to release significant concentrations of Hf (Pearce et al., 1992;  
356 Tollstrup et al., 2010; Li et al., 2013).

357

358 The  $\epsilon_{\text{Hf}}$  value of the first subduction component ( $S_1$ ) must lie on an extension of the axial-  
359 boninite trend. The minimum  $\epsilon_{\text{Hf}}$  value of the subduction component (c. 8) is the intersection  
360 with the Sm/Hf axis - the value at which Sm in  $S_1$  is negligible. The maximum value (c. 12) is  
361 the lowest  $\epsilon_{\text{Hf}}$  sample of the boninite trend - the value at which mantle contribution is  
362 negligible. This places the  $\epsilon_{\text{Hf}}$  value of  $S_1$  between c.8 and c.12. This range may further be  
363 constrained using the less subduction-mobile (though more compatible) element, Ti, in place  
364 of Sm (Fig. 6b). This increases the minimum  $\epsilon_{\text{Hf}}$  value to c.10.7, the value at which Ti/Hf in  
365  $S_1$  is negligible.

366

367 The location and origin of  $S_1$  can be further constrained using experimental data and slab

368 fusion modeling. Zr/Sm ratios have been studied in particular detail for amphibolite melting  
369 (Pearce et al., 1992; Foley, 2002, 2008) based on published, experimentally-derived phase  
370 proportions and partition coefficients, and taking into account minor phase solubility.  
371 Because Zr/Hf is not significantly fractionated during the experiments, these values convert  
372 simply to Sm/Hf ( $\text{Sm/Hf} = c.35 / (\text{Zr/Sm})$ ). Our best estimate for the average Sm/Hf in a  
373 shallow slab melt is 0.5, which equates to a  $\epsilon_{\text{Hf}}$  value for  $S_1$  of 11.5 (Fig. 6a). Pearce et al.  
374 (1992: Fig. A2) obtain this ratio at c.900°C for water-deficient melting and c.950°C for  
375 water-saturated melting.

376  
377 Foley's (2002, 2008) models report these ratios as a function of degree of melting rather than  
378 temperature, reaching  $\text{Sm/Hf} = 0.5$  after c.5% batch melting. In both cases, reducing  
379 temperature (decreasing zircon solubility) increases Sm/Hf (decreases Zr/Sm), and hence  
380 increases our  $\epsilon_{\text{Hf}}$  estimate for the slab melt. Similarly, increasing temperature decreases  
381 Sm/Hf and the estimated  $\epsilon_{\text{Hf}}$ . An error bar of  $\pm 0.5$  epsilon units covers the range of 800-  
382 1000°C and 1-10% batch melting. Foley's mean values for eclogite melting equate to Sm/Hf  
383 of 0.6 (rutile-free) to 1 (rutile-bearing), while Rapp et al. (1999, 2003) obtain Sm/Hf of 0.64  
384 in one experimental charge (AB-1) for 3.8GPa and 1100°C. These values lie above the  
385 empirical upper bound for  $\epsilon_{\text{Hf}}$ , therefore making deep (eclogite facies) slab melts less likely  
386 than shallow (amphibolite facies) slab melts as the source of  $S_1$ .

387  
388 The same approach gives a best estimate for amphibolite melts of c.500 for Ti/Hf (Fig. 6b),  
389 thus supporting our  $\epsilon_{\text{Hf}}$  estimate for the  $S_1$  end-member. In contrast, Rapp's (1999) eclogite

390 facies melting experiment gives a Ti/Hf ratio 1100, too high for eclogite facies slab melts to  
391 be a likely source for  $S_1$ .

392

393 In Fig. 6c,  $x=Nd$ , allowing us to determine the Hf/Nd component ratios needed for the  
394 modeling of the  $\epsilon_{Hf}$ - $\epsilon_{Nd}$  diagram. Most of the axial-FAB samples, including D-FAB, plot in  
395 the upper part of the ambient mantle array, at  $Nd/Hf=c. 3.6$  and  $\epsilon_{Hf}=19$ . The off-axis FAB  
396 are distinct in having slightly higher  $Nd/Hf$  of 4.5 and higher  $\epsilon_{Hf}$  of c. 21. The value of 11.5  
397 for  $\epsilon_{Hf}$  in the slab component from Fig. 6a-b gives  $Nd/Hf$  of 1.8.

398

399 The inverse ratio,  $Hf/Nd$ , can then be used to delimit the  $\epsilon_{Nd}$  value for the subduction  
400 component (Fig. 6d). The axial-boninites form a trend from residual FAB mantle (M) to  
401 higher  $Hf/Nd$ . Taking the crustal end member as 0.55 (the reciprocal of  $Nd/Hf=1.8$  from Fig.  
402 6c) then gives  $\epsilon_{Nd}= 7.5$  for the slab melt component,  $S_1$ .

403

404 Importantly, Fig. 6d also defines better the composition of the components forming the off-  
405 axis boninites. The high  $Hf/Nd$  end-member of the off-axis boninite trend lies within and at  
406 the lower end of the axial-boninite M- $S_1$  trend. We interpret this end-member as a mixture of  
407 slab melt and depleted mantle, which we term hybrid mantle wedge (H). The off-axis  
408 boninites extend from H towards an  $S_2$  component at lower  $Hf/Nd$ . A projection of this trend  
409 intersects the  $\epsilon_{Nd}$  axis at a value of 4 for  $Hf/Nd=0$ , making this the minimum value of  $\epsilon_{Nd}$   
410 for this component. The maximum value, the lowest analysed  $\epsilon_{Nd}$  on the trend, is 6.5.

411

412 Fig. 6d also highlights the fact that the compositions of pelagic sediments ( $\epsilon\text{Nd}=-2$  to  $-9$ ) lie  
413 well below the H-S<sub>2</sub> trend. This requires that the second subduction component (S<sub>2</sub>) also  
414 includes a high  $\epsilon\text{Nd}$  component and this is most likely an altered ocean crust (AOC)  
415 component. However, its  $\epsilon\text{Nd}$  value is poorly defined on this projection. If it is as low as 4,  
416 the AOC contribution must have a very low Hf/Nd ratio, which more typical of aqueous  
417 fluid. In contrast, if it is as high as 6.5, it must have high Hf/Nd, which is more characteristic  
418 of a higher temperature slab melt or supercritical fluid.

419

### 420 *5.3 Characterizing the subduction component using $\epsilon\text{Hf}$ - $\epsilon\text{Nd}$ covariations*

421

422 Two drill sites outboard of the IBM system (ODP Sites 801 and 1149; Fig. 1a) provide the  
423 most complete isotope and element data set available for the (Jurassic) crust and sediment of  
424 subducting Pacific plate and hence for interpreting the boninite site subduction components  
425 (Fig. 7a).

426

427 At Site 1149, the section comprises a Pacific-MORB basement overlain by pelagic sediments,  
428 the latter divided into a series of sedimentary sub-types (e.g. Plank et al., 2007; Chauvel et  
429 al., 2009; Vervoort et al., 2011). Site 801 is more complex. Its oceanic crust basement  
430 comprises MORB overlain by a relatively thin unit of OIB (termed ‘Top Alkali Basalts’, or  
431 TAB) (Hauff et al, 2003). The pure basalt compositions form an ambient mantle trend well  
432 within the Pacific Field in Fig. 7a. However, the composites, made up of interstitial materials  
433 in addition to lavas, are displaced to lower  $\epsilon\text{Nd}$  and plot close to the Indian-Pacific boundary.

434 In addition to pelagic sediments, the Site 801 sedimentary section contains volcanogenic  
435 sediments resulting from Jurassic intraplate magmatism. Plotted on Fig. 7a are a series of  
436 component mixing hyperbolae for Sites 1149 and 801 based on these published compositions.  
437 Bulk subducted compositions for Sites 1149 and 801 should plot on the bulk crust - bulk  
438 sediment curves A and D respectively.

439

440 In Fig. 7b, we plot the Exp. 352 data together with the subduction components  $S_1$  and  $S_2$   
441 inferred from Fig. 6. The obvious interpretation is that component  $S_1$  lies between Site 801  
442 MORB and TAB composites (i.e. along trend B on Fig. 7a). Component  $S_2$  lies on the trend  
443 from the inferred hybrid mantle wedge (H) and subducted Site 801 sediment (pelagic +  
444 volcanogenic). A model based on the Site 801 section can therefore best explain the  
445 components produced at the start of subduction.

446

447 Fig. 7c focuses on mixing models for M- $S_1$  and H- $S_2$ . The former has been constructed using  
448 a standard mixing equation with Nd/Hf ratios obtained from the  $\epsilon_{\text{Hf-Nd/Hf}}$  and  $\epsilon_{\text{Nd-Hf/Nd}}$   
449 plots (Pearce et al., 1999). We use axial-FAB to define the isotopic composition of residual  
450 FAB mantle (M), as in Fig. 6.

451

452 Mantle: E-FAB source (for trend 3):  $\epsilon_{\text{Nd}_M}=7.8$ ;  $\epsilon_{\text{Hf}_M}=18$ ;  $(\text{Nd/Hf})_M=4.3$

453 Mantle N-FAB source (for trend 4):  $\epsilon_{\text{Nd}_M}=8.8$ ;  $\epsilon_{\text{Hf}_M}=19.5$ ;  $(\text{Nd/Hf})_M=3.6$

454 Mantle: D-FAB source: (for trend 5):  $\epsilon_{\text{Nd}_M}=9.7$ ;  $\epsilon_{\text{Hf}_M}=21$ ;  $(\text{Nd/Hf})_M=3.8$

455 Subduction component ( $S_1$ =‘Pacific’ slab melt):  $\epsilon_{\text{Nd}_{S_1}}=7.5$ ;  $\epsilon_{\text{Hf}_{S_1}}=11.5$ ;  $(\text{Nd/Hf})_{S_1}=1.8$

456

457 It is apparent from resulting mixing hyperbolae (Fig. 7c, trends 3-5) that the trends from  
458 axial-FAB to  $S_1$  encompass almost all of the axial-boninite data, so supporting the concept  
459 that the axial-boninites are the product of mixing of ‘Pacific’ crustal slab melt and the  
460 ambient ‘Indian’ mantle source. To annotate the hyperbolae in terms of mass fractions of  
461 subduction zone components requires a knowledge not just of ratios but also of the absolute  
462 values of  $(Nd)_M/(Nd)_{S1}$  or  $(Hf)_M/(Hf)_{S1}$ . Because these are not well constrained, we plot mixing  
463 lines as hyperbolae for the particular optimum value of  $r$ , where  $r = (Nd/Hf)_M / (Nd/Hf)_{S1}$ . For  
464 example, the value of  $r$  for Trend 4 is then  $3.6/1.8=2.0$ .

465

466 Note that it is very rare to be able to define mantle-slab melt mixture so clearly; it is only  
467 possible here because of the distinction between ‘Indian’ provenance mantle and ‘Pacific’  
468 provenance crust and the absence of a large sediment component. Most Western Pacific arcs  
469 are cooler and mixing lines run between ‘Indian’ mantle and pelagic sediment. Notable  
470 exceptions are in the rear-arc volcanoes (e.g. Tollstrup et al., 2010), which lie above deeper,  
471 and hence hotter, slabs.

472

473 To model the off-axis boninite variations, we continue to infer that they involve the mixing of  
474 mantle already containing the slab melt component (the hybrid mantle wedge, H) and an  
475 added subduction component ( $S_2$ ) with lower  $\epsilon Nd$  made up of a mixture of a sediment-  
476 derived component and an AOC-derived component. Here, we base our model on the lower  
477 temperature  $S_2$  end-member from the H- $S_2$  trend in Fig. 6d, made up of an AOC component

478 with Hf/Nd=c.0 and a pelagic sediment component with Hf/Nd=c. 0.5. Values used are:

479

480 Hybrid (slab melt-modified) residual mantle (H):  $\epsilon\text{Nd}_M=8.0$ ;  $\epsilon\text{Hf}_M=12$ ;  $(\text{Nd}/\text{Hf})_M=1.8$

481 Additional subduction component ( $S_2$ ):  $\epsilon\text{Nd}_{S_2}=4.0$ ;  $\epsilon\text{Hf}_{S_2}=12$ ;  $(\text{Nd}/\text{Hf})_{S_2}=20$

482

483 In fact, Fig. 6d showed that this  $S_2$  could have  $\epsilon\text{Nd}$  as high as 6.5, although this higher value  
484 would have little effect on the general interpretation. We can, however, constrain this  
485 component further by incorporating isotopes of Pb and Sr, which partition more readily than  
486 Nd and Hf into crust-derived fluids and sediment-derived fluids and melts (Fig 8).

487

488 *5.4 Identifying and tracing subduction sources during subduction initiation using plots of*  
489 *isotope ratios v  $\Delta 8/4$*

490

491 In Fig. 8, we plot  $\epsilon\text{Hf}$ ,  $\epsilon\text{Nd}$ ,  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $^{207}\text{Pb}/^{204}\text{Pb}$  against  $\Delta 8/4$ , where  $\Delta 8/4$  is the  
492 orthogonal deviation from the Northern Hemisphere Reference Line (NHRL) in  $^{208}\text{Pb}/^{204}\text{Pb}$ -  
493  $^{206}\text{Pb}/^{204}\text{Pb}$  space, as defined by Hart (1984) (Fig. 4d). In all the plots, we can recognize the  
494 two subduction components that were apparent in the Hf-Nd projections. For the purpose of  
495 this paper (the tracing of mantle and subduction components following subduction initiation),  
496 we continue to interpret these trends in terms of simple mixing of two end members, but  
497 realise that this is an approximation given the complex nature of the mantle wedge.

498

499 The first trend, as modeled in Fig. 7, marks the addition of ‘Pacific’ slab melt ( $S_1$ ) to ‘Indian’

500 Residual FAB Mantle (M). Given that the  $\epsilon_{Nd}$  value of this component was established at c.  
501 7.5, we can use Fig. 8b to fix the  $S_1$  component at  $\Delta 8/4 \approx -30$ , i.e. in the expected field of  
502 Pacific altered oceanic crust (AOC).

503

504 The second trend runs from the hybrid mantle wedge (H) to  $S_2$ .  $S_2$  cannot be precisely  
505 defined, but constraints are sufficient to demonstrate that AOC and pelagic sediments are its  
506 principal contributors. First, an extrapolation of the linear regression in Figure 8d places  $S_2$   
507 on a line between H and the pelagic sediment field. As we inferred from Fig. 6d that the  $\epsilon_{Nd}$   
508 of  $S_2$  lies between 4 and 6.5, this in turn ties  $\Delta 8/4$  to a value between c.5 and 20, some way  
509 from being pure pelagic sediment. Thus,  $S_2$  requires a combination of AOC and Pacific  
510 pelagic sediment (PPS), possibly with some volcanogenic sediment. The former could have  
511 the same isotopic composition as  $S_1$  if it was derived from the same subducted basalt as that  
512 supplying the slab melt for the axial-boninites. Changing the  $\epsilon_{Nd}$  value of  $S_2$  changes the  
513 ratio of these components but not the nature and composition of the end-members.

514

515 The elemental concentrations of Nd, Hf and Sr relative to Pb in  $S_1$  and  $S_2$  may be gleaned  
516 from the curvatures of the axial and off-axis boninite trends respectively. These shapes of the  
517 mixing hyperbolae are a function of the ratios  $r_1 [(x/Pb)_M/(x/Pb)_{S_1}]$  and  $r_2 [(x/Pb)_H/(x/Pb)_{S_2}]$   
518 where  $x = \text{Hf, Nd or Sr}$ . These shapes match the patterns in Fig. 3, by indicating that  $S_1$   
519 contains high Pb and Sr, significant Hf and relatively little Nd and that  $S_2$  carries further Pb  
520 and Sr but undetectable Hf and significant Nd. As already discussed,  $S_1$  can be explained by  
521 residual amphibole during slab melting,  $S_2$  by the incoming of pelagic sediment.



522

523 **6. Discussion: isotopic constraints on magma genesis following subduction**  
524 **initiation**

525

526 Fig. 9 presents a conceptual model that we believe best satisfies the isotopic constraints  
527 derived from this study. Further details and discussion of this model are given below.

528

529 *6.1 Constraints from the absence of a subduction component in the axial fore-arc basalts*  
530 *(Sites U1440 and U1441)*

531

532 The FAB drilled at IODP Exp. 352 Sites U1440 and U1441 show no clear isotopic evidence  
533 for a subduction component. There are therefore two principal modes of origin: spreading  
534 immediately after subduction; or spreading immediately before subduction. Both models are  
535 consistent with the ‘Indian’ character of the ambient mantle given that the position of the  
536 ‘Pacific-Indian’ mantle domain boundary lay outboard of the transform plate boundary  
537 (Miyazaki, 2015) and would have been sampled in either case. Based on regional-scale  
538 sampling, however, Reagan et al. (2010) find that subduction-free FAB similar to that drilled  
539 during Exp. 352 comprises one end-member of a spectrum that extends to FAB with clear  
540 subduction enrichment. From this, they conclude that spreading following subduction was the  
541 more likely, i.e. that extension was due to slab roll-back as proposed by Stern and Bloomer  
542 (1992). In that case, rapid mantle upwelling relative to the rate of heating of the slab is  
543 required to minimize the subduction input of the subduction-free end-member compositions.

544

545 *6.2 Constraints from the shallow slab-melting component in the axial-boninites (Holes*  
546 *U1439C and U1442A)*

547

548 The isotopic evidence presented here supports a model in which the axial-boninite sources  
549 are the products of variable interaction between ambient, ‘Indian’ domain, residual FAB  
550 mantle and shallow, amphibolite facies, slab melts ( $S_1$ ). These shallow slab melts are likely  
551 tonalitic in composition, lacking the depletion in heavy REE relative to middle REE of  
552 deeper (adakitic) slab melts. This model is similar to that developed for boninites in the  
553 Bonin Ridge and its northward extension (Pearce et al., 1992; Li et al, 2013). However, our  
554 boninites from Exp. 352 are older than the Bonin Ridge boninites (c. 51Ma versus c. 46Ma),  
555 so providing the first evidence that shallow slab melting started very soon after subduction  
556 initiation before continuing for at least 5 m.y. Early slab melting is consistent with thermal  
557 models, where the first crust to subduct melts at shallow depths because it encounters mantle  
558 uncooled by subsequent subduction (Pearce et al., 1992).

559

560 An alternative model that can be ruled out is the basalt melt contributing to  $S_1$  was not the  
561 product of slab melting, but instead was derived from fusion of basalt (FAB) veins within the  
562 mantle wedge (Pearce et al., 1999). This model was put forward to explain the genesis of  
563 ODP Leg 125 boninites, where both end-members of the equivalent M- $S_1$  trend lie within the  
564 ‘Indian domain’. In this study,  $S_1$  has a clear ‘Pacific’ provenance, so strongly supporting the  
565 slab-melt model.

566

567 It is significant that our new data effectively rule out two previous models for the origin of  
568 the slab melts. Casey and Dewey's (1984) model of a leaky transform fault attributes slab  
569 melting to subduction of ridges within the transform zone. If true, the spreading ridges would  
570 need to lie within the 'Pacific' mantle domain. However, isotopic studies of accreted basalts  
571 show that the Izanagi-Pacific Ridge had crossed a static 'Indian-Pacific' mantle domain  
572 boundary by 80m.y. (Miyazaki et al., 2015) and thus that the present boundary lies beneath  
573 the Pacific plate. In fact, Straub et al. (2015) found that oceanic crust of 'Indian' provenance  
574 did become the source of subduction components within the Izu-Bonin system, but not until  
575 at least 10 m.y. after subduction initiation. The same arguments negate the likelihood of  
576 larger-scale ridge subduction events being the cause of boninite magmatism (Seton et al.,  
577 2015).

578

579 *6.3 Start of sediment input and the origin of the off-axis boninites (Holes U1439C and*  
580 *U1442A Upper lava Units)*

581

582 A key question arising from Section 5 is why sediment input begins between the eruption of  
583 axial-boninites and the eruption of off-axis boninites. There are at least three possibilities: 1)  
584 a sediment component is also present in S<sub>1</sub> but is swamped by the slab melt signal; 2) the first  
585 subducted crust originated in a transform zone where deeper sea-floor, less high-temperature  
586 hydrothermal activity and more detachment faulting resulted in minimal sedimentary cover;  
587 and 3) the sediment is initially accreted rather than subducted. The argument against first

588 option is that Pb isotopes proved highly sensitive to the sediment content of  $S_2$ , and so would  
589 have detected a sediment component in  $S_1$  were it present in significant proportion. The  
590 second option cannot be ruled out, but transform faults are rarely sediment-free and often  
591 contain debris flows made up of igneous clasts in sediment matrices, which would contribute  
592 a sedimentary signal to the subduction component. Thus, we consider the accretionary option  
593 to be most likely.

594

595 In the accretionary model, sediments would have to be accreted at the start of subduction so  
596 that the slab that initially rolls back is sediment free. For this to happen, subduction dynamics  
597 must change between the input of  $S_1$  and  $S_2$  in a way that first inhibits sediment subduction  
598 and then permits it. Critical taper theory (e.g. Dahlen, 1990) supported by subduction zone  
599 comparisons (Clift and Vannucchi, 2004) reveal that shallow and slow subduction favor  
600 development of accretionary margins and hence off-scraping of sediment, while steep and  
601 fast subduction favor non-accretionary margins and hence sediment subduction.

602

603 In Fig 9 (a) subduction starts slowly and at a shallow angle up to the point at which oceanic  
604 crust converts to eclogite, after which (b) rapid roll-back takes place of the sediment-free  
605 subducting slab culminating in addition of  $S_1$ , followed by (c) normal, rapid and steep  
606 subduction (likely with a roll-back component) in which the accreted sediment is able to  
607 subduct. Introduction of fluid and/or melt from sediment and AOC ( $S_2$ ), either separately or  
608 together as a *mélange* with altered basaltic debris (Nielsen and Marschall, 2017), then allows  
609 the sub-solidus hybrid mantle wedge to undergo the additional hydration to produce the off-

610 axis boninites trench-side of the original ridge.

611

## 612 **7. Conclusions**

613

614 1. The first stage of magmatism following subduction initiation is associated with sea-floor  
615 spreading following subduction initiation and results in the eruption of forearc basalts (FAB)  
616 at IODP Sites U1440 and U1441. Isotope data confirm that these FAB lack a subduction  
617 component and so can be used to characterize the ambient mantle reservoir prior to the input  
618 of the subduction component. Notably, this reservoir had ‘Indian’ provenance and had  
619 experienced an ancient (possibly Palaeozoic), garnet facies depletion event, similar to that  
620 proposed by Yogodzinski et al. (2018) based on the slightly younger, and more trench-distal,  
621 basalts drilled at IODP Site U1438.

622

623 2. The first evidence of subduction is recorded by the axial boninites at IODP Sites U1439  
624 and U1442. Their isotopic compositions follow mixing trends between residual mantle of  
625 ‘Indian’ provenance and oceanic crust of ‘Pacific’ provenance. The latter had undergone  
626 shallow (amphibolite facies) fusion to give positive Hf anomalies on extended REE plots, so  
627 providing further evidence for slab fusion at the start of subduction. The absence of sediment  
628 in the isotope signal may be explained by initial sediment accretion, although other options  
629 are possible.

630

631 3. The final stage is the eruption of the off-axis boninites. Their isotopic compositions follow

632 mixing trends between a hybrid mantle wedge similar to the axial boninite source (slab melt  
633 plus residual mantle) and a subduction component derived from altered oceanic crust (AOC)  
634 and Pacific pelagic sediment (PPS). Our favored explanation for the sudden appearance of the  
635 sediment component is subduction of previously accreted material resulting from increased  
636 subduction rate and slab dip. As this event takes place <2m.y after the genesis of the forearc  
637 basalts (Reagan et al., 2019), the results provide evidence that subduction initiation was  
638 followed by a rapidly evolving system in terms of geodynamics, magma genesis and crustal  
639 accretion.

640

#### 641 **Acknowledgements**

642

643 We thank the International Ocean Discovery Program (IODP) for samples and infrastructure.

644 We appreciate the contributions of the Staff Scientist (K. Petronotis) and remaining members

645 of the IODP Exp. 352 Scientific Party as well as the officers and crew, and scientific support

646 staff, on the JOIDES Resolution. We thank Gene Yogodzinski and an anonymous referee for

647 their helpful recommendations and Tamsin Mather for her editorial support. Hong-Yan Li

648 additionally thanks the Strategic Priority Research Program of the Chinese Academy of

649 Sciences (Grant No. XDB18000000) and the National Programme on Global Change and

650 Air-Sea Interaction (GASI-GEOGE-02). Shervais, Ryan and Reagan thank the USSSP for

651 initial Expedition funding and the NSF for grants OCE1558689, OCR1558855 and

652 OCE1558647 respectively. Taylor & Pearce and Prytulak thank the NERC (UK) for grants

653 NE/M012034/1 and NE/M010643 respectively. Godard thanks CNRS-INSU Terre Solide

654 'Soutien Post-campagne à la Mer'. Hong-Yan Li thanks Xiang Li, Le Zhang and Jinlong Ma  
655 for their contributions to the analyses in GIGCAS laboratories. Prytulak thanks K. Kressig,  
656 B.J. Coles, M. Mangler and F. Wei for their contributions to the analyses and running of the  
657 MAGIC laboratories. This is contribution No. IS-2640 of GIG-CAS

658

#### 659 **Appendix A and B. Supplementary material**

660

661 Supplementary material related to this article can be found on-line at  
662 <https://doi.org/10.1016/.....>

663

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812 *Acta.* doi: <https://doi.org/10.1016/j.gca.2018.02.047>.

813

814 Table Caption

815 Table 1. Representative isotope analyses from fore-arc basalt IODP Exp. 352 Holes  
816 U1440B and U1441A, and boninite Holes U1439C and U1442A. For locations, see  
817 Fig. 1. For the full data table, including major and trace element analyses and errors,  
818 see Appendix A. For full analytical details, see Appendix B. For Pb and Sr isotope  
819 analyses: the four samples marked by asterisks in the Table were separates of fresh  
820 glasses collected close to Hf-Nd sample locations; non-glass samples that gave non-  
821 reproduceable Pb isotope ratios despite heavy leaching are marked as 'altered'. Rock  
822 type abbreviations: FAB= Forearc Basalt (D= depleted; N= Normal, E= enriched);  
823 LSB= Low-Si Boninite; HSB= High-Si Boninite; HMA= High-Mg Andesite. All  
824 samples are lavas except U1440B Unit 15 and U1439C Unit 10, both of which are  
825 sheeted intrusions (superscript D).

826

827

828 Figure Captions

829

830 Figure 1. Location and geological setting of the IODP Exp. 352 drill sites in the Bonin  
831 forearc, Western Pacific (modified from Reagan et al., 2015). (a) Location relative to  
832 key geographic features and drill sites. (b) Geological setting of the four drill sites.  
833 Sites U1440 and U1441 are the ‘FAB sites’, and Sites U1439 and U1442 are the  
834 slightly younger and more trench-distal ‘boninite sites’. Base map from submersible  
835 sampling and dredging coupled with age dating (Reagan et al., 2010; Ishizuka et al.,  
836 2011). (c) Simplified logs for the four principal basement-penetrating holes from the  
837 FAB and boninite sites. Modified from Reagan et al. (2017) with ages from Reagan et  
838 al. (2019). FAB=forearc basalt; LSB=low-Si boninite; HSB=high-Si boninite;  
839 HMA=high-Mg andesite.

840

841 Figure 2. Isotope stratigraphies for the FAB sites (Holes 1440B and U1441A) and  
842 boninite sites (Holes U1439C and U1442A). Both of the FAB and boninite  
843 stratigraphies may be subdivided into axial (lower) and off-axis (upper) groups. The  
844 axial-FAB subdivide on trace element criteria into D(depleted)-FAB, the dominant  
845 N(normal)-FAB and E(enriched)-FAB: these differ in  $\epsilon\text{Nd}_i$  values, which increase  
846 from D- through N-FAB to E-FAB. Off-axis FAB have distinctly lower  $\epsilon\text{Nd}_i$  than  
847 axial-FAB, and off-axis boninites mostly have lower  $\epsilon\text{Nd}_i$  and  $\epsilon\text{Hf}_i$  than axial-  
848 boninites as well as higher  $^{206}\text{Pb}/^{204}\text{Pb}$  isotope ratios. All FAB have higher  $\epsilon\text{Hf}_i$  and  
849 lower  $^{206}\text{Pb}/^{204}\text{Pb}$  than all boninites. Note that the Pb analysis in parentheses is



850 anomalous (off-axis composition in an axial sequence), which may indicate fluid-flux  
851 from above, and so is omitted from later diagrams. mbsf=meters below sea-floor.

852

853 Figure 3. Extended chondrite-normalized REE plots highlighting the changes in the  
854 radiogenic daughter isotope elements, Pb, Sr, Nd and Hf during the evolution from (a)  
855 axial- and off-axis FAB through (b) axial-boninites to (c) off-axis boninites. Note the  
856 absence of significant subduction-related anomalies in the FAB (the alteration-related  
857 Sr anomaly in the D-FAB sample has been omitted), in contrast to the positive Pb, Sr  
858 and Hf anomalies in the boninites. Note also that the LREE (and hence Nd)  
859 progressively increase from FAB through axial boninite to off-axis boninite, as seen  
860 in the trend from LREE-depleted to slight LREE enrichment. Chondrite-normalizing  
861 factors are from Sun and McDonough (1989), with the Pb value amended to remove  
862 the metallic component of the chondrite. For full elemental analyses, see Appendix A.

863

864 Figure 4. Standard Nd, Hf, Sr and Pb isotope projections, showing the isotopic  
865 evolution of magmatism from axial-FAB through off-axis FAB through axial-boninite  
866 to off-axis boninite. Principal features, investigated in more detail using Figures 5-8,  
867 are: 1) an increase in  $\epsilon_{\text{Hf}_i}$  from axial-FAB to off-axis-FAB; 2) a decrease in  $\epsilon_{\text{Hf}_i}$ , and  
868 Pb isotope ratios from FAB to axial-boninites; and 3) a decrease in  $\epsilon_{\text{Nd}_i}$  coupled to an  
869 increase in Sr and Pb isotope ratios from axial to off-axis boninites. The general  
870 explanations for these observations are respectively: for 1), the arrival of a long-term  
871 depleted mantle source for the off-axis FAB; for 2), incoming of a subduction  
872 component derived from fusion of 'Pacific' domain, altered oceanic crust (AOC) for  
873 the axial boninites and for 3), incoming of a further subduction component containing

874 a significant contribution of Pacific Pelagic Sediment (PPS) for the off-axis boninites.  
875 This is a remarkable variation for such small differences in space and time.

876

877 Figure 5. Isotopic interpretations of FAB used to establish mantle compositions  
878 immediately following subduction initiation. (a) Contrast in ‘Pacific’ and ‘Indian’  
879 mantle sources in the Western Pacific and Australian-Antarctic Discordance (AAD).  
880 #801, #1149, #1438 refer to drill core from ODP Leg 129 Site 801, ODP Leg 185 Site  
881 1149 and IODP Exp. 351 Site U1438 respectively (see Fig. 1 for locations). Data are  
882 from Chauvel et al. (2009), Heydolph et al. (2014), Hickey-Vargas (1998), Miyazaki  
883 et al. (2015), Pearce et al., (1999), Savov et al. (2006) and Yogodzinski et al. (2018).  
884 (b) Plot of Exp. 352 FAB data, demonstrating that the axial-FAB plot within the  
885 Philippine Sea ‘Indian’ mantle array in a similar position to Site U1438 Unit 1e, while  
886 the off-axis FAB lie above this array. Fig. 5c-d explore the nature of Exp. 352 FAB  
887 mantle sources, taking as base diagrams the plots of Yogodzinski et al. (2018) for Site  
888 U1438 Units a-e, the closest lavas from the Philippine Sea in space (at the time of  
889 eruption) and time. The plots indicate that Lu/Hf fractionation during mantle melting  
890 likely explains the variations within N-FAB, and that the isotopic variations from E to  
891 N to D-FAB within axial-FAB, and those within off-axis FAB, likely require older  
892 (possibly both Mesozoic and Palaeozoic) enrichment and depletion events. Note that  
893 boninites have not been plotted because Hf and Nd are both subduction-mobile (Fig.  
894 3). The AAD field marks the compositions of lavas transitional between Indian and  
895 Pacific provenance along the Australian-Antarctic Ridge.

896

897 Figure 6. Use of  $\epsilon\text{Hf-x/Hf}$  and  $\epsilon\text{Nd-y/Nd}$  plots (where x is Sm, Ti or Nd and y is Hf)  
898 to constrain the  $\epsilon\text{Hf}$  and  $\epsilon\text{Nd}$  values of mixing end-members. The diagrams help  
899 define two trends: the first from M (residual FAB mantle) and a shallow (amphibolite-  
900 facies) slab melt (at  $S_1$ ); and the second from a point on that trend representing a  
901 hybrid mantle wedge (H) to a component containing pelagic sediment and an altered  
902 oceanic crust (AOC) component (at  $S_2$ ). Data sources for the ambient (Philippine Sea)  
903 mantle array are given in the caption to Fig. 5. The gradient of the ambient mantle  
904 array reflects the difference in incompatibility between x and Hf or Nd (positive if x is  
905 less compatible).

906

907 Figure 7. Modeling of the  $\epsilon\text{Hf}-\epsilon\text{Nd}$  covariation diagram in Fig. 4a. Fig. 7a  
908 summarizes published data for potential subducted materials from the Sites 801 and  
909 1149 (Fig. 1a). Basement composites (subscript C) represent whole-core sections  
910 (lavas plus interstitial materials), analysed by Chauvel et al. (2009). Site 801 MORB  
911 and TAB (Top Alkali Basalt) data were from pure basalt cores using  $\epsilon\text{Hf}$  and  $\epsilon\text{Nd}$   
912 data for the data points (Pearce et al., 1999) with additional data on  $\epsilon\text{Nd}$  only (Hauff  
913 et al., 2003) used to establish their dispersion. Fig. 7b-c include further sediment data  
914 from Plank et al. (2007) and Vervoort et al. (2011). Mixing lines between M (residual  
915 FAB mantle) and  $S_1$  (subduction component 1) and between H (hybrid mantle wedge)  
916 and  $S_2$  (subduction component 2) constrain the compositions of bulk subducted  
917 materials.

918

919 Figure 8. Plots of  $\epsilon\text{Hf}$ ,  $\epsilon\text{Nd}$ , and  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $^{207}\text{Pb}/^{204}\text{Pb}$  isotope ratios against  $\Delta 8/4$ ,  
920 where  $\Delta 8/4$  is the displacement from the Northern Hemisphere Reference Line

921 (NHRL) in  $^{208}\text{Pb}/^{204}\text{Pb}$ - $^{206}\text{Pb}/^{204}\text{Pb}$  space. By definition, points on the NHRL have  
922  $\Delta 8/4$  values of 0, while the Indian-Pacific mantle domain boundary on the same  
923 diagram has a value of +20. Philippine Sea and Western Pacific mantle domains (Fig.  
924 5a) are drawn as rectangles either side of this value. Alteration will, however, displace  
925 these domains to lower  $\Delta 8/4$ . The two subduction trends (modeled as described in the  
926 text) are clearly defined as running between 1) a residual FAB, 'Indian' domain,  
927 mantle (M) and a 'Pacific' domain slab melt ( $S_1$ ), and 2) a hybrid (slab melt +  
928 depleted mantle) mantle wedge (H) and a combination of Pacific Pelagic Sediment  
929 (PPS) and Pacific Altered Oceanic Crust (AOC) fluid or melt ( $S_2$ ).  $\epsilon_{\text{Hf}}$  and  $\epsilon_{\text{Nd}}$   
930 values of  $S_1$  and  $S_2$  are constrained by Fig. 5 and used in turn to constrain the Pb and  
931 Sr end-member compositions.

932

933 Figure 9. A conceptual model linking the isotope interpretations to the geodynamics  
934 of subduction initiation. A. The model starts at a hypothetical transform fault at the  
935 Izanagi (IZA)-Pacific (PAC) plate boundary. B) Shallow, slow subduction then  
936 accretes sediment, allowing a sediment-free plate to subduct. C) Once subduction is  
937 sufficiently deep for phase transformation of subducted crust to eclogite facies, initial  
938 rapid rollback leads to extension with formation of FAB with no or little subduction  
939 input. D) As the slab sinks further, it heats up to melting temperatures while still at  
940 shallow depth, and these slab melts then interact with residual mantle to produce low-  
941 Si boninites. E) Further roll-back and perhaps the start of dip-slip subduction causes  
942 accreted sediment to subduct and contribute to off-axis magma genesis, with fluid  
943 interacting with unmelted mixtures of residual mantle and slab melt (hybrid mantle  
944 wedge) to produce high-Si boninites. Alternative, but we believe less likely,  
945 mechanisms for explaining the isotopic features are discussed in the text. Note that

946 the model covers only the c. 52-50Ma period immediately following subduction  
947 initiation and so does not extend to the subsequent (c. 50-44Ma) growth and initial  
948 rifting of the protoarc and transition into normal arc magmatism. Similarly, the model  
949 does not extend to include the c. 50-49 Ma extensional event that formed the forearc  
950 basalts at IODP Site U1438 further to the west (Yogodzinski et al, 2018; see also Fig.  
951 1 and Fig. 5 of this paper) and which significantly post-dates the forearc basalt  
952 spreading event depicted in Fig. 9C (Reagan et al., 2019). IMM and PMM =  
953 Indian/Pacific MORB Mantle.

954

Figure 1

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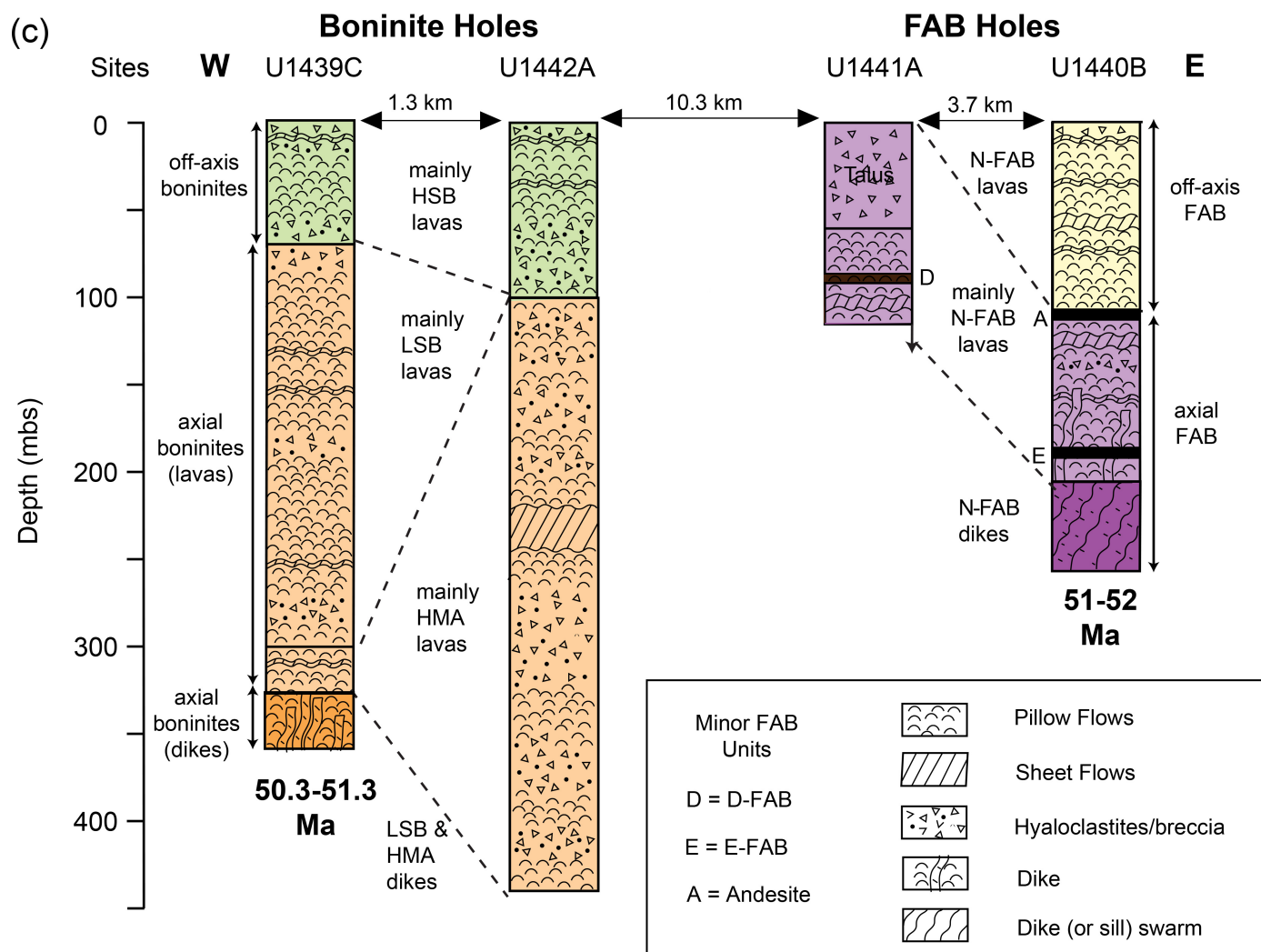
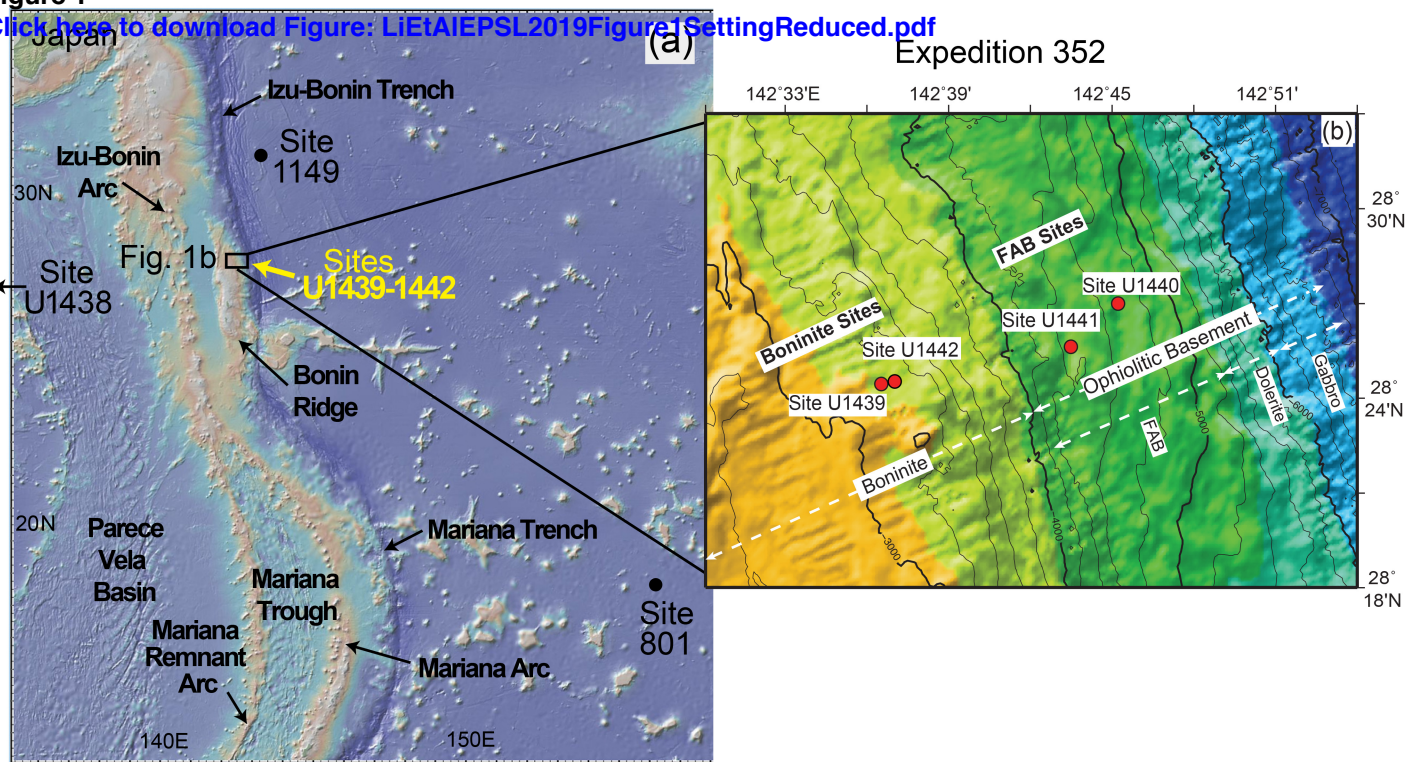


Figure 2

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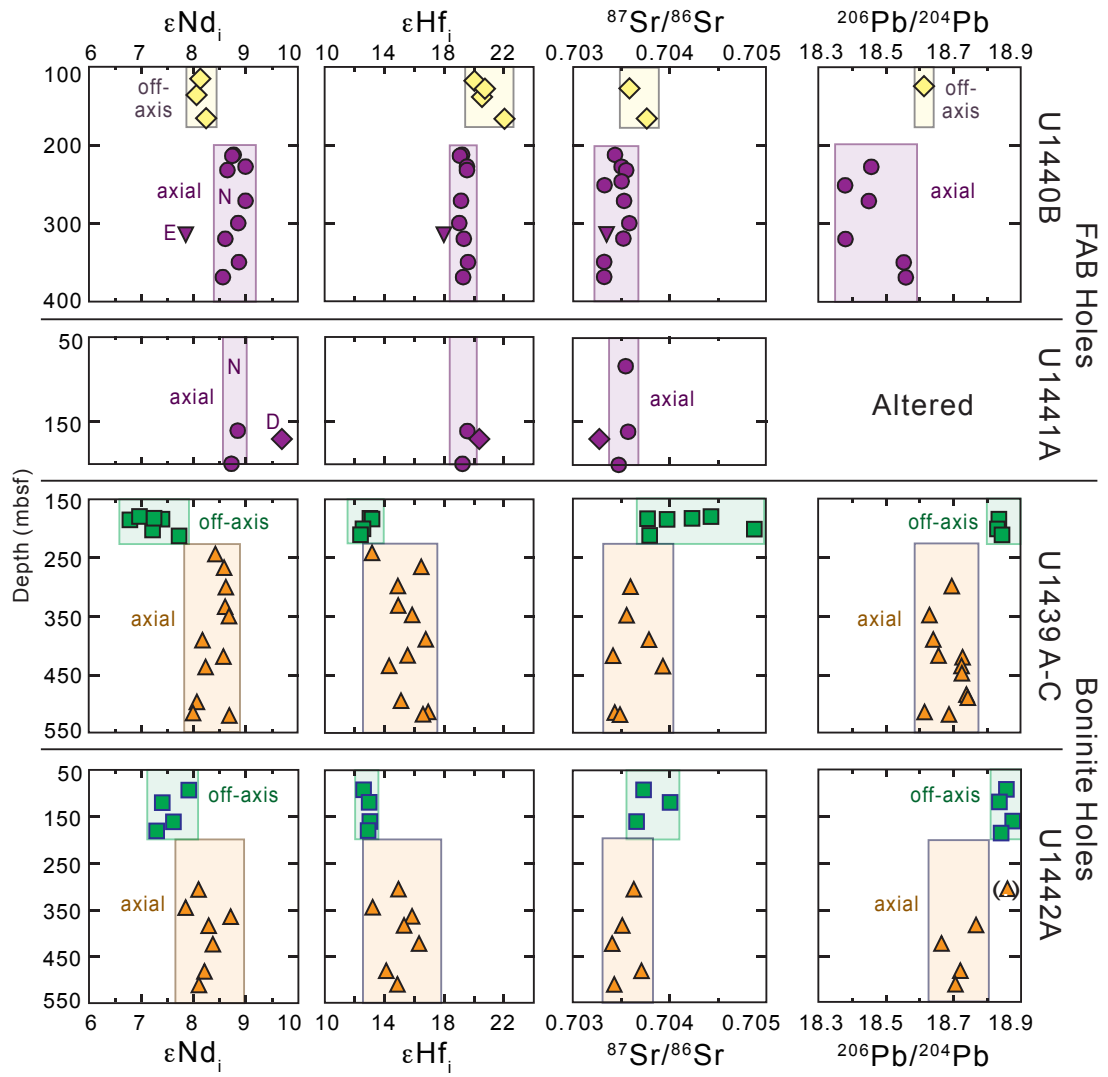


Figure 3

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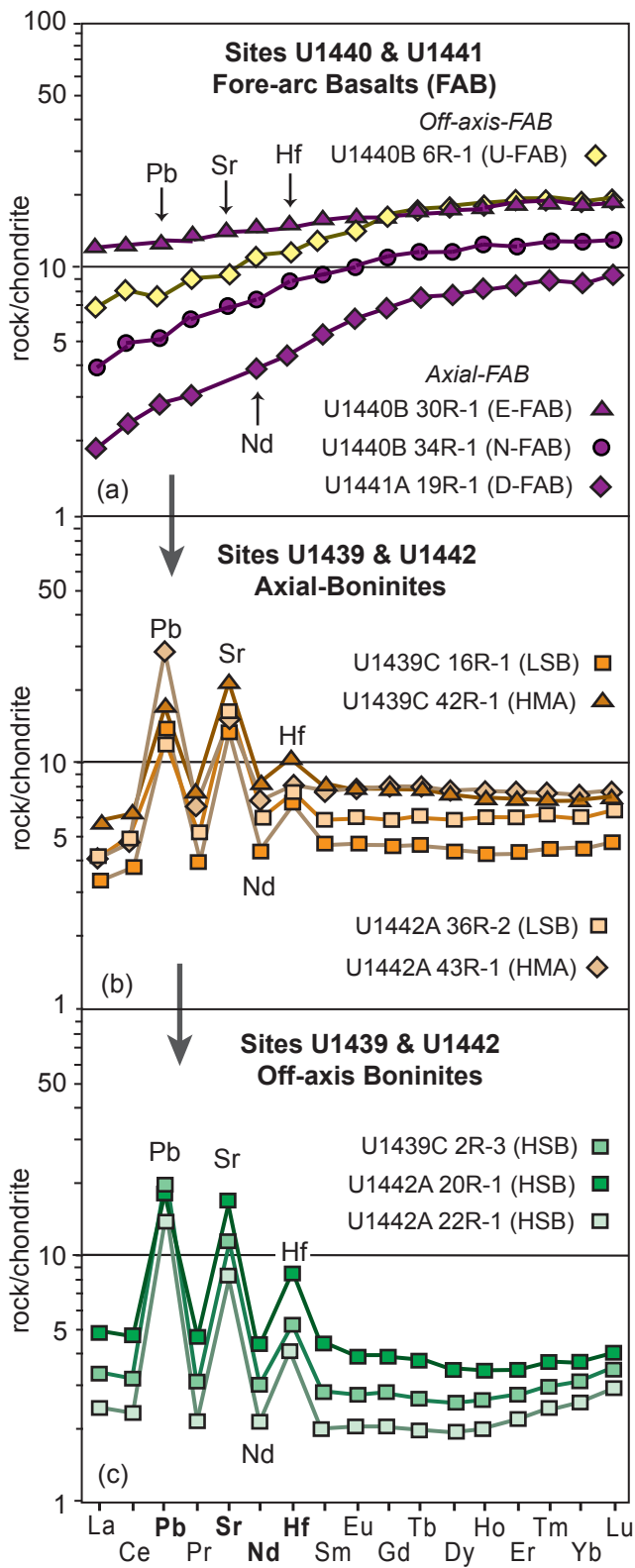
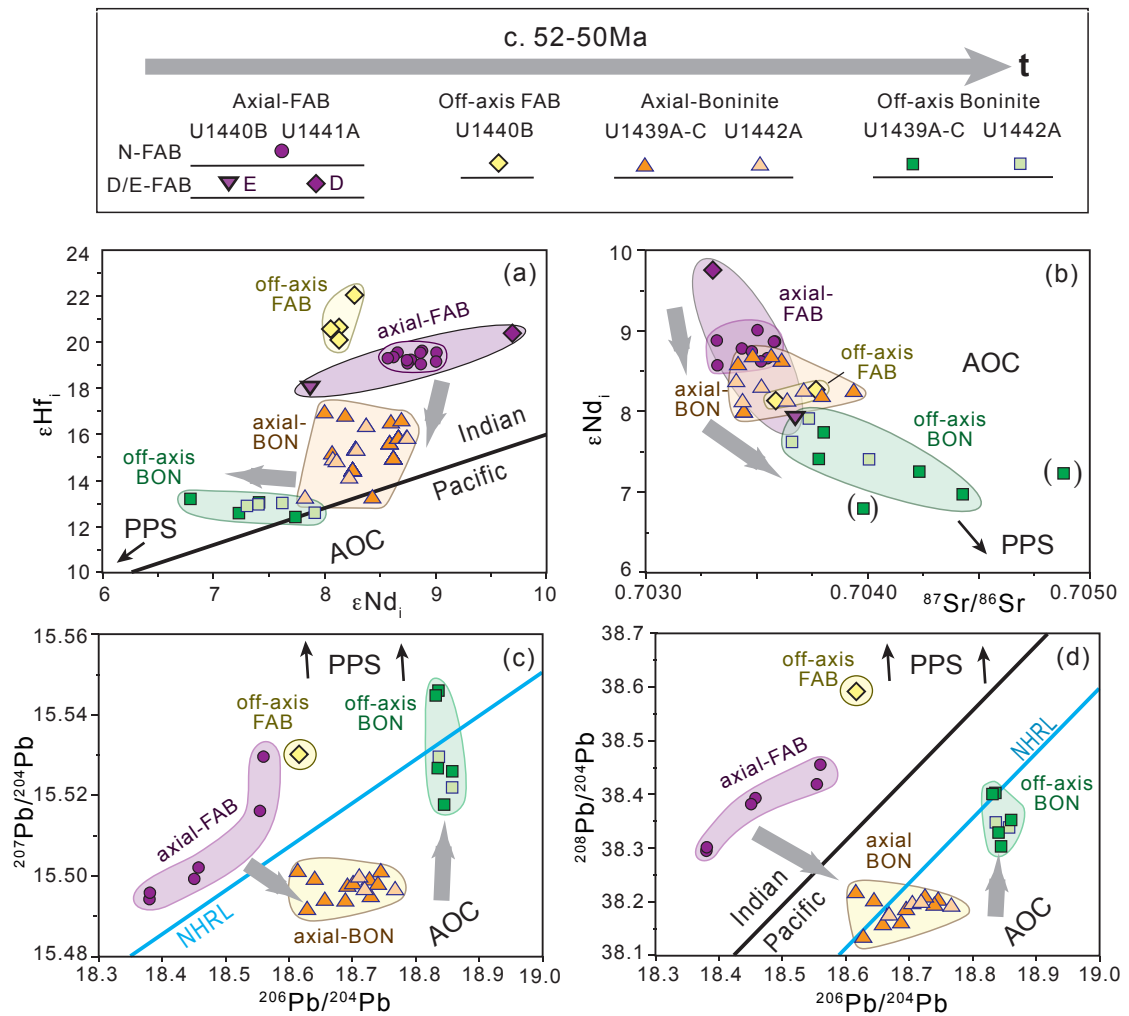




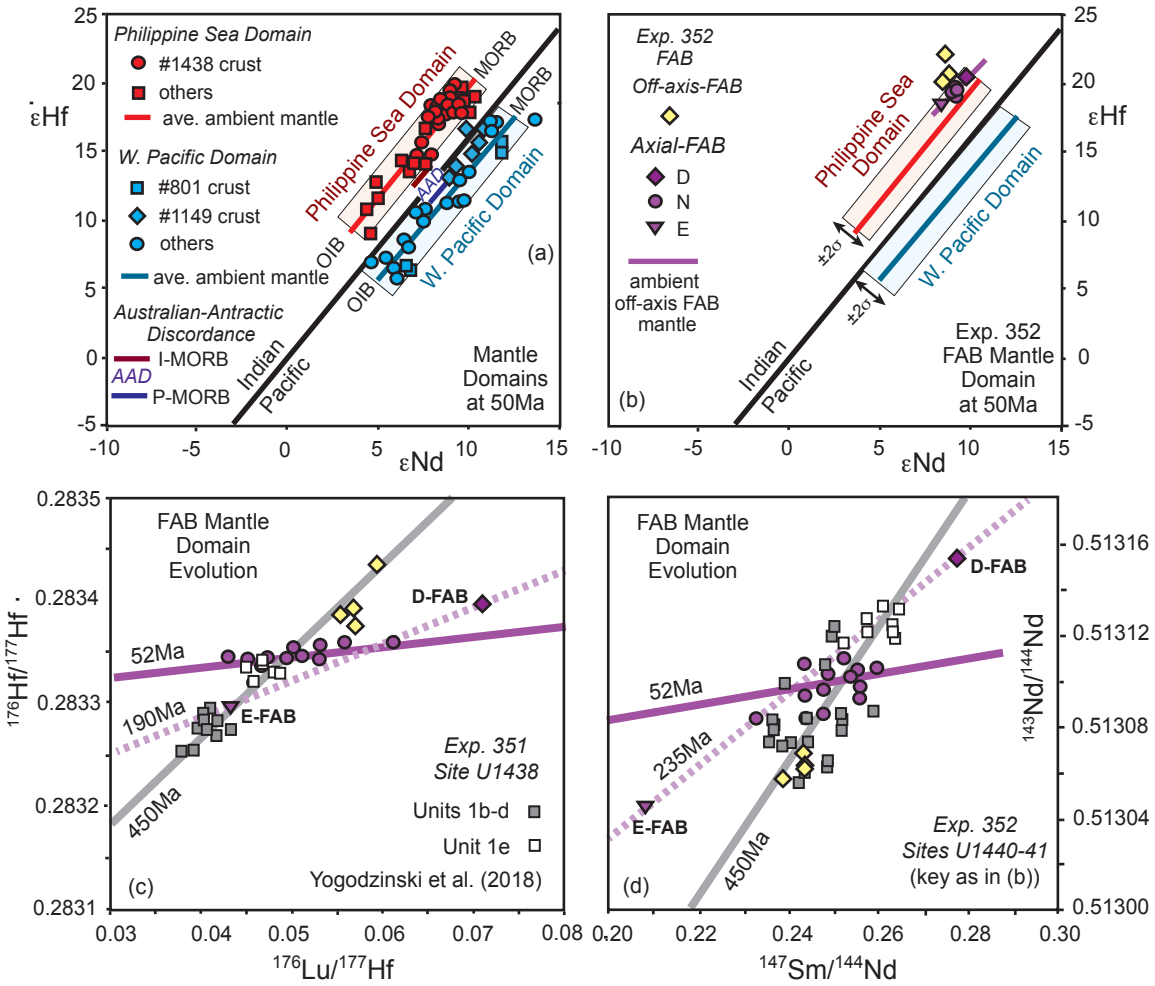
Figure 4

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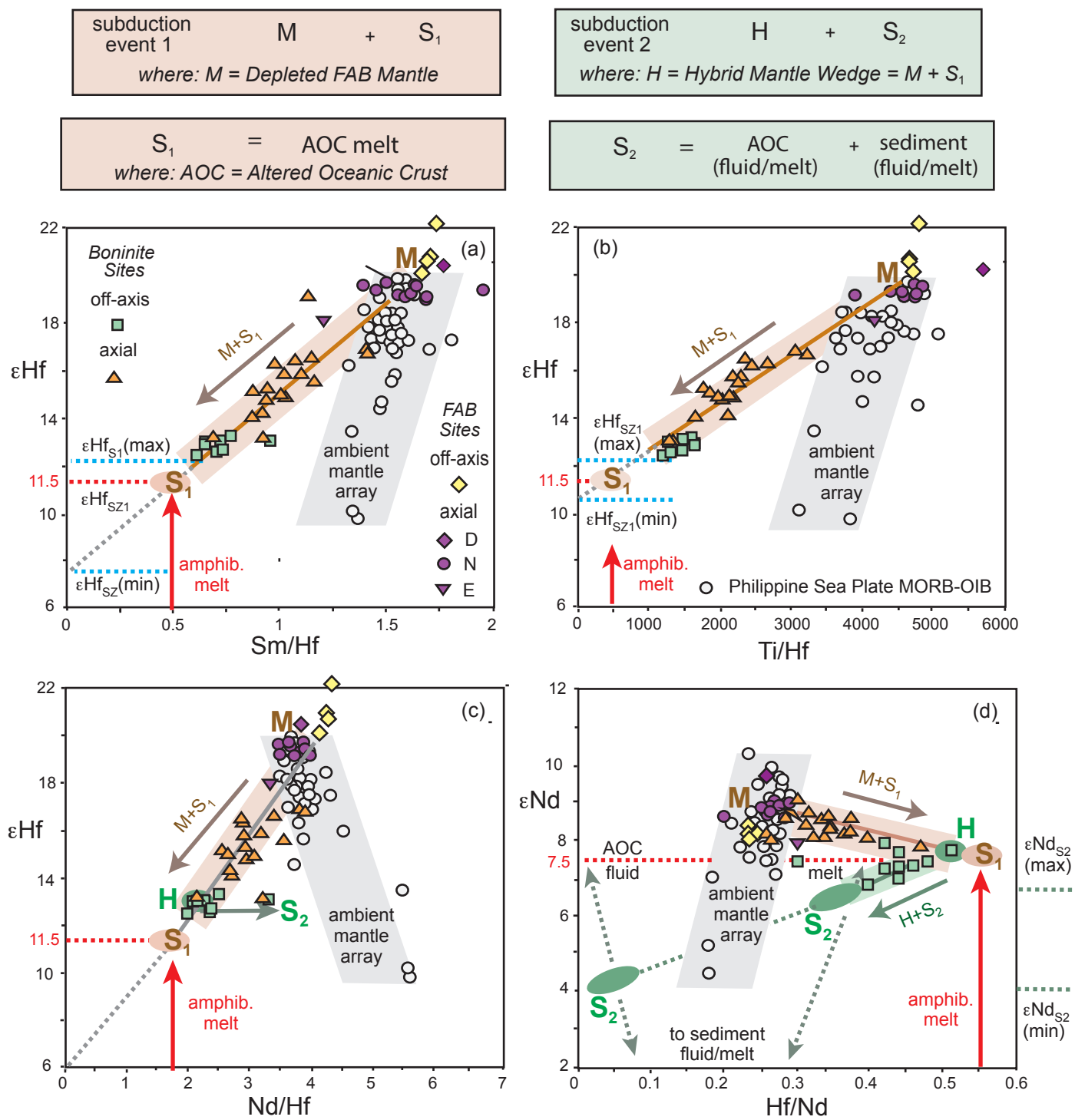


**Figure 5**

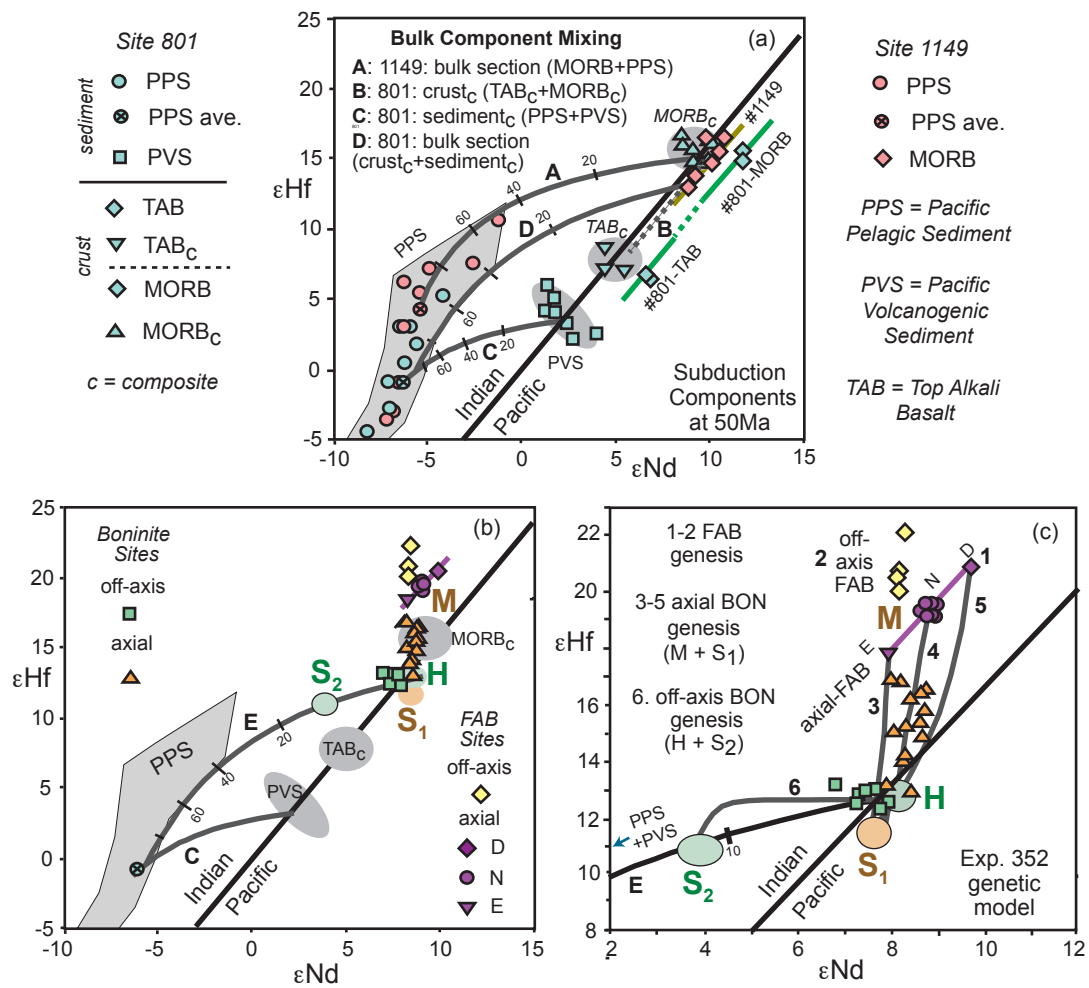
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**Figure 6**  
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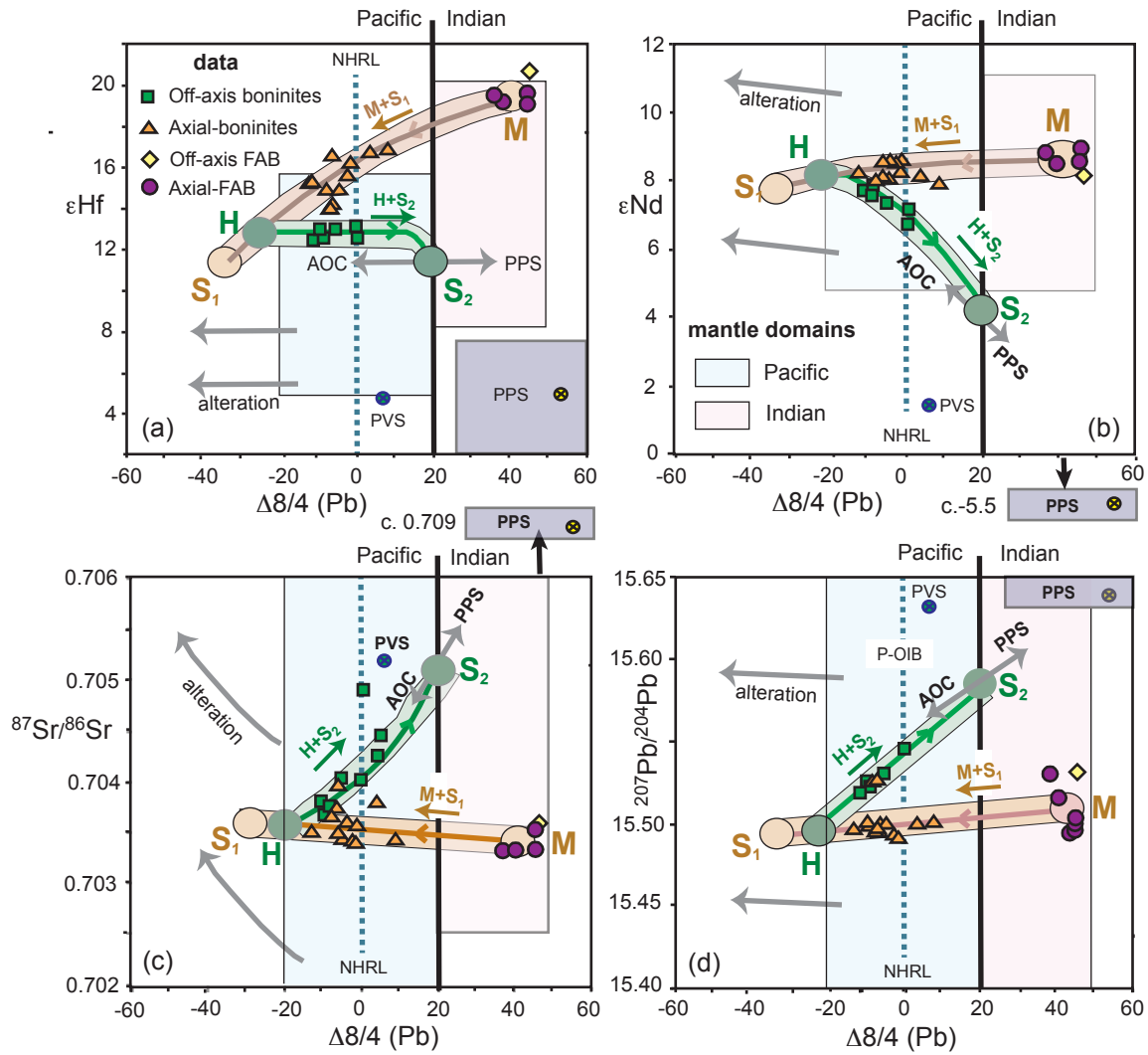


**Figure 7**  
[Click here to download Figure: LIETaIEPSL2019Figure7HfNdSubductionComponents.pdf](#)



**Figure 8**

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**Figure 9**

[Click here to download Figure: LIETaIEPSL2019Figure9IsotopeEvolution.pdf](#)

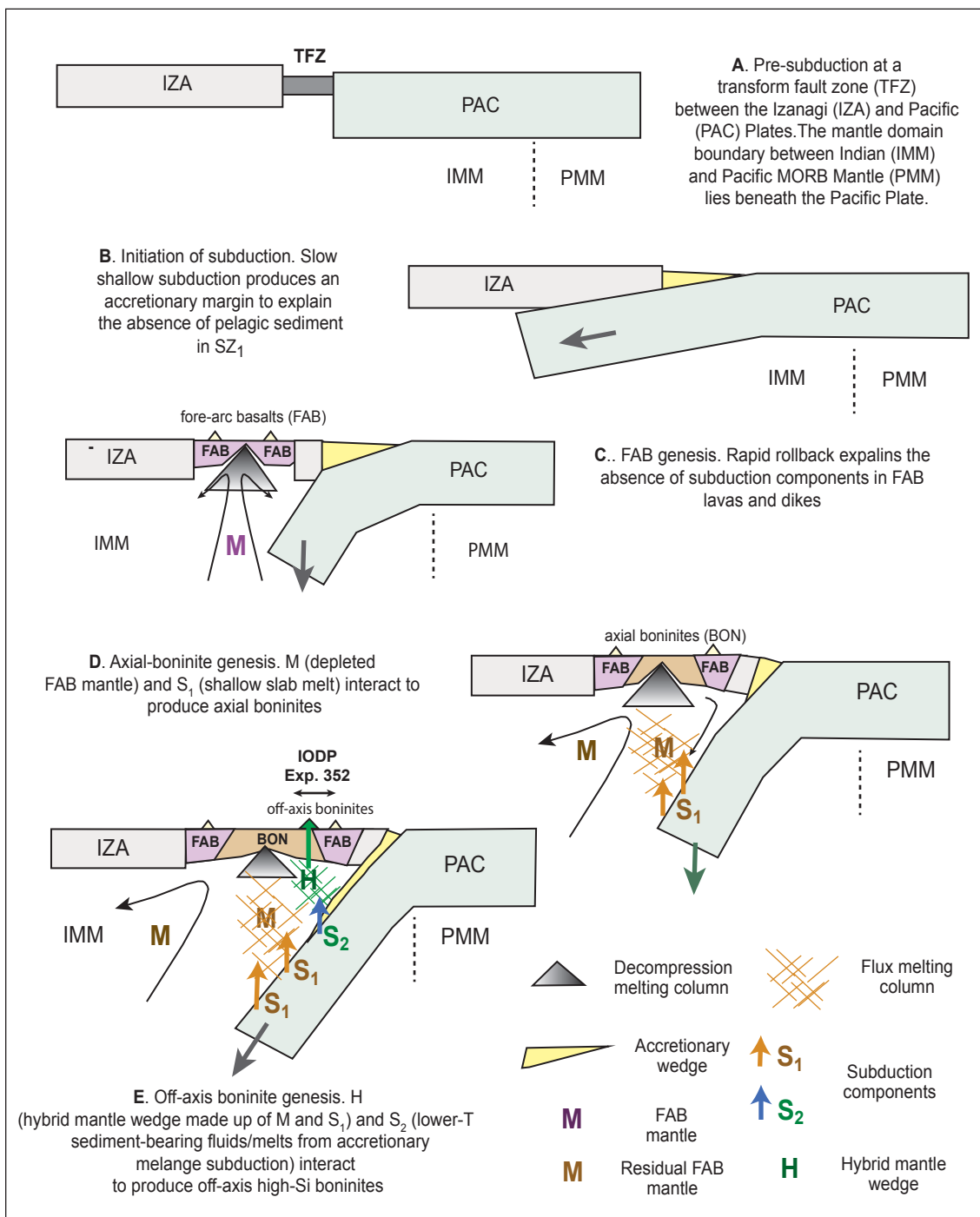


Table 1

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Sample	Series	Type	Unit	Depth	$^{206}\text{Pb}/^{204}\text{Pb}$	$^{207}\text{Pb}/^{204}\text{Pb}$	$^{208}\text{Pb}/^{204}\text{Pb}$	$^{87}\text{Sr}/^{86}\text{Sr}$	$\epsilon\text{Nd}_{50}$	$\epsilon\text{Hf}_{50}$
<b>U1440B</b>										
6R-1	Off-axis	N-FAB	2	127.11	18.6154	15.5301	38.592	0.703581	8.1	20.7
12R-2	Off-axis	N-FAB	3	165.14		altered		0.703768	8.3	22.1
19R-1*	Axial	N-FAB	6	227.40	18.4577	15.5021	38.393	0.703482	9.0	19.6
24R-1*	Axial	N-FAB	7	251.20	18.3803	15.4942	38.295	0.703327		
26R-1*	Axial	N-FAB	8	271.14	18.4508	15.4992	38.382	0.703470	9.0	19.2
29R-1	Axial	N-FAB	11	299.75		altered		0.703584	8.9	19.1
30R-1	Axial	E-FAB	13	309.53		altered		0.703426	8.1	17.8
31R-1*	Axial	N-FAB	14	319.74	18.3809	15.4958	38.302	0.703521	8.6	19.4
34R-1	Axial <sup>D</sup>	N-FAB	15	349.72	18.5544	15.5161	38.419	0.703322	8.9	19.6
36R-1	Axial <sup>D</sup>	N-FAB	15	368.94	18.5600	15.5296	38.455	0.703324	8.6	19.3
<b>U1441A</b>										
18R-1	Axial	N-FAB	2	160.97		altered		0.703578	8.9	19.5
19R-1	Axial	D-FAB	3	171.16		altered		0.703305	9.7	20.4
22R-1	Axial	N-FAB	4	199.85		altered		0.703479	8.7	19.2
<b>U1439C</b>										
2R-3	Off-axis	HSB	2a	185.13	18.8358	15.5460	38.403	0.703981	6.8	13.2
4R-1	Off-Axis	HSB	3a	201.86	18.8341	15.5438	38.404	0.704884	7.2	12.6
16R-1	Axial	LSB	6	299.76	18.6961	15.4979	38.189	0.703604	8.6	14.9
22R-1	Axial	LSB	6	349.19	18.6305	15.4918	38.135	0.703565	8.7	15.9
26R-3	Axial	LSB	7	390.66	18.6420	15.4993	38.204	0.703794	8.2	16.8
29R-2	Axial	LSB	8	418.27	18.6566	15.4938	38.158	0.703423	8.6	15.6
41R-1	Axial <sup>D</sup>	LSB	10	515.09	18.6154	15.5012	38.219	0.703443	8.0	17.0
42R-1	Axial <sup>D</sup>	HMA	10	518.88	18.6883	15.4943	38.161	0.703495	8.7	16.6
<b>U1442A</b>										
11R-1	Off-axis	HSB	1a	92.22	18.8572	15.5219	38.339	0.703736	7.9	12.6
15R-1	Off-axis	HSB	1b	119.40	18.8364	15.5295	38.348	0.704007	7.4	13.0
20R-1	Off-axis	HSB	1c	160.64	18.8751	15.5245	38.356	0.703661	7.6	13.0
35R-1	Axial	LSB	2b	306.01	18.8603	15.5259	38.350	0.703634	8.1	14.9
43R-1	Axial	HMA	2b	383.86	18.7666	15.4966	38.190	0.703515	8.3	15.3
47R-1	Axial	HMA	3	423.10	18.6646	15.4914	38.177	0.703409	8.4	16.3
53R-1	Axial	HMA	4	481.39	18.7205	15.4967	38.197	0.703714	8.2	14.1
56R-1	Axial	LSB	4	510.49	18.7054	15.4993	38.198	0.703433	8.1	14.9

- glass separates used for Pb and Sr isotope analyses (except 31R-1: Pb only)