

1 **New evidence for a long Rhaetian from a Panthalassan succession (Wrangell**
2 **Mountains, Alaska) and regional differences in carbon cycle perturbations at the**
3 **Triassic-Jurassic transition**

4

5 *¹Caruthers, A.H., ²Marroquín, S.M., ³Gröcke, D.R., ⁴Golding, M., ⁵Aberhan, M., ⁶Them,
6 T.R., II, ⁷Veenma, Y.P., ⁸Owens, J.D., ⁹McRoberts, C.A., ¹⁰Friedman, R.M., ¹¹Trop, J.M.,
7 ¹²Szűcs, D., ^{13, 14}Pálffy, J., ¹⁵Rioux, M., ⁷Trabucho-Alexandre, J.P., and ²Gill, B.C.

8

9 **Affiliations**

10 *¹*Department of Geological and Environmental Sciences, Western Michigan University,*
11 *Kalamazoo, MI 49006, USA (andrew.caruthers@wmich.edu)*

12 ²*Department of Geosciences, Virginia Tech, Blacksburg, VA 24061, USA*

13 ³*Department of Earth Sciences, Durham University, South Road, Durham, County*
14 *Durham, DH1 3LE, UK*

15 ⁴*Geological Survey of Canada, Pacific Division, Vancouver, BC V6B 5J3, Canada*

16 ⁵*Museum für Naturkunde Berlin, Leibniz Institute for Evolution and Biodiversity Science,*
17 *10115 Berlin, Invalidenstraße 43, Germany*

18 ⁶*Department of Geology and Environmental Geosciences, College of Charleston,*
19 *Charleston, SC 29424, USA*

20 ⁷*Department of Earth Sciences, Universiteit Utrecht, P.O. Box 80115, 3508 TC Utrecht, the*
21 *Netherlands*

22 ⁸*Department of Earth, Ocean and Atmospheric Science, National High Magnetic Field*
23 *Laboratory, Florida State University, Tallahassee, Florida 32310-3706, USA*

24 ⁹*Geology Department, State University of New York, Bowers Hall Rm 37, Cortland, NY*
25 *13045, USA*

26 ¹⁰*Pacific Centre for Isotopic and Geochemical Research, University of British Columbia,*
27 *Vancouver BC V6T 1Z4, Canada*

28 ¹¹*Department of Geology and Environmental Geosciences, Bucknell University, Lewisburg,*
29 *PA 17837, USA*

30 ¹²*Camborne School of Mines, University of Exeter, Penryn Campus, Cornwall, TR10 9FE,*
31 *UK*

32 ¹³*Department of Geology, Eötvös Loránd University, Pázmány Péter sétány 1/C, Budapest,*
33 *H-1117, Hungary*

34 ¹⁴*MTA-MTM-ELTE Research Group for Paleontology, Ludovika tér 2, Budapest, H-1083,*
35 *Hungary*

36 ¹⁵*Department of Earth Science, 1006 Webb Hall, University of California, Santa Barbara,*
37 *CA 93106, USA*

38 *Corresponding author

39

40 **Abstract**

41

42 The end-Triassic mass extinction is one of the *big five* extinction events in Phanerozoic
43 Earth history. It is linked with the emplacement of the Central Atlantic Magmatic

44 Province and a host of interconnected environmental and climatic responses that
45 caused profound deterioration of terrestrial and marine biospheres. Current
46 understanding, however, is hampered by (i) a geographically limited set of localities and
47 data; (ii) incomplete stratigraphic records caused by low relative sea-level in European
48 sections during the Late Triassic and earliest Jurassic; and (iii) major discrepancies in the
49 estimated duration of the latest Triassic Rhaetian that limit spatiotemporal evaluation
50 of climatic and biotic responses locally and globally. Here, we investigate the Late
51 Triassic–Early Jurassic time interval from a stratigraphically well-preserved sedimentary
52 succession deposited in tropical oceanic Panthalassa. We present diverse new data
53 from the lower McCarthy Formation exposed at Grotto Creek (Wrangell Mountains,
54 southern Alaska), including ammonoid, bivalve, hydrozoan, and conodont
55 biostratigraphy; organic carbon isotope ($\delta^{13}\text{C}_{\text{org}}$) stratigraphy; and CA-ID TIMS zircon U-
56 Pb dates. These data are consistent with a Norian-Rhaetian Boundary (NRB) of
57 ~209 Ma, providing new evidence to support a long duration of the Rhaetian. They also
58 constrain the Triassic-Jurassic boundary (TJB) to a ~6 m interval in the section. Our TJB
59 $\delta^{13}\text{C}_{\text{org}}$ record from Grotto Creek, in conjunction with previous data, demonstrates
60 consistent features that not only appear correlative on a global scale but also shows
61 local heterogeneities compared to some Tethyan records. Notably, smaller excursions
62 within a large negative carbon isotope excursion [NCIE] known from Tethyan localities
63 are absent in Panthalassan records. This new comparative isotopic record becomes
64 useful for (i) distinguishing regional overprinting of the global signal; (ii) raising
65 questions about the ubiquity of smaller-scale NCIEs across the TJB; and (iii) highlighting

66 the largely unresolved regional vs. global scale of some presumed carbon cycle
67 perturbations. These paleontological and geochemical data establish the Grotto Creek
68 section as an important Upper Triassic to Lower Jurassic succession due to its
69 paleogeographic position and complete marine record. Our record represents the best
70 documentation of the NRB and TJB intervals from Wrangellia, and likely the entire
71 North American Cordillera.

72

73 **Key Words**

74 Norian-Rhaetian boundary, Triassic-Jurassic boundary, stable carbon-isotopes,
75 Wrangellia, Panthalassa, CAMP large igneous province

76

77 **1. Introduction**

78

79 The Late Triassic to Early Jurassic was a dynamic interval of Earth history when the
80 biosphere was severely disrupted by climatic and environmental changes that
81 culminated in a major mass extinction (i.e., the end-Triassic mass extinction or ETE)
82 across the Triassic-Jurassic Boundary (TJB; e.g., Alroy et al., 2008). It is considered one
83 of the largest extinction events in Earth history and may be associated with rapid
84 volcanogenic outgassing during the emplacement of the Central Atlantic Magmatic
85 Province (CAMP; Fig. 1A; Wignall, 2001).

86

87 One of the most significant problems in understanding the timing of events around the
88 ETE is the mass extinction itself. The removal of a large number of organisms from the
89 global biosphere drastically decreased the number of taxa available for relative age
90 assignments and, by consequence, our collective confidence in global stratigraphic
91 correlation. The severity of climatic and environmental disruption at this time,
92 however, significantly impacted global geochemical records, thus allowing alternative
93 techniques (e.g., carbon isotope chemostratigraphy) to correlate strata and assign
94 relative ages.

95

96 Considerable effort has been invested into identifying the global extent of biological
97 turnover and environmental change during the latest Triassic and Early Jurassic using a
98 diverse set of paleontological and geochemical data from the terrestrial and marine
99 records (e.g., McElwain et al., 1999; Pálffy et al., 2000; Hesselbo et al., 2002; Whiteside
100 et al., 2010; Schoene et al., 2010; Schaller et al., 2011; Steinhorsdottir et al., 2011).
101 Detangling the local, regional, and global environmental signals from these datasets,
102 however, remains an outstanding and important challenge that (given the available
103 records) is exacerbated by (i) a geographically biased set of data, with the majority of
104 published records from successions that represent deposition in the western part of the
105 ancient Tethys Ocean and epeiric seaways (i.e., Europe, Fig. 1A); (ii) a low relative sea-
106 level in the Tethys during the Late Triassic and earliest Jurassic which caused shallow-
107 marine sites to be more susceptible to erosion and the development of significant
108 hiatuses (e.g., Schoene et al., 2010); (iii) major discrepancies in current Late Triassic

109 (Rhaetian) timescale models (e.g., Wotzlaw et al., 2014; Li et al., 2017). The latter has
110 complicated the temporal correlation of geochemical datasets commonly used to
111 interpret environmental change and the driving mechanisms of the ETE.

112

113 Here, we seek to address this gap by investigating the Upper Triassic to Lower Jurassic
114 record from a well-preserved and largely unstudied sedimentary succession exposed in
115 the Wrangellia terrane of North America (Fig. 1; Wrangell Mountains, USA). The
116 Triassic to Jurassic rocks of this terrane accumulated in a tropical oceanic environment
117 situated upon a subsiding oceanic plateau (e.g., Greene et al., 2010) in the Panthalassan
118 Ocean. New data generated from the Grotto Creek section represent an important
119 addition to existing end-Triassic records with implications toward a greater
120 understanding of event timing and global carbon cycle perturbations.

121

122 **2. Background**

123

124 *2.1 Trigger and driving mechanisms of the end-Triassic extinction*

125

126 To date, both terrestrial and extraterrestrial causal mechanisms have been proposed
127 for the ETE. As reviewed by Pálffy and Kocsis (2014) and Korte et al. (2019), the timing
128 and magnitude of a bolide impact as the sole extinction mechanism lack significant
129 evidence. The more widely accepted hypothesis links CAMP volcanism with a cascade
130 of climatic and environmental feedbacks, which ultimately led to global mass extinction

131 (e.g., Wignall, 2001; Carter and Hori, 2005; Korte et al., 2019) and is well supported by
132 coeval peak extinction rates in siliceous (i.e., radiolarians) and calcifying organisms
133 during the late Rhaetian (Kocsis et al., 2014). This hypothesis, known as the Volcanic
134 Greenhouse Scenario or VGS (Wignall, 2001), has also been applied to explain several
135 other mass extinctions linked to the emplacement of other large igneous provinces
136 (e.g., Wignall, 2001).

137

138 The VGS proposes that perturbations to the global carbon cycle are one of the most
139 ubiquitous underlying phenomena that accompany mass extinctions (e.g., Wignall,
140 2001). In this scenario, negative carbon isotope excursions (NCIEs) are caused by the
141 input of ^{12}C -enriched carbon into the oceans and atmosphere by CO_2 from volcanic
142 degassing, metamorphism of organic carbon-rich sediments by volcanic intrusions,
143 and/or biogenic CH_4 . Elevated atmospheric $p\text{CO}_2$ during the ETE is supported stomatal
144 index and paleosol data (McElwain et al., 1999; Schaller et al., 2011; Steinhorsdottir et
145 al., 2011). Regardless of carbon source, all scenarios lead to atmospheric and oceanic
146 warming and associated environmental feedbacks such as deoxygenation (and many
147 others).

148

149 The organic carbon isotope ($\delta^{13}\text{C}_{\text{org}}$) records from the former Tethys Ocean and a
150 handful of localities from Panthalassa show brief, large-amplitude NCIEs of $\sim 2\text{--}6\text{‰}$
151 across coeval TJB successions (Ward et al., 2001; Guex et al., 2004; Hesselbo et al.,
152 2002; Pálffy et al., 2007; Korte et al., 2019; and others). These records include what has

153 been termed an *initial* NCIE before the TJB, which appears coeval with the main mass
154 extinction interval (e.g., Korte et al., 2019). In many records, the *initial* NCIE is followed
155 by a transient increase in $\delta^{13}\text{C}_{\text{org}}$ and then a second or *main* NCIE that extends well into
156 the early Hettangian (e.g., Korte et al., 2019). Similar general trends have also been
157 observed in the $\delta^{13}\text{C}$ of fossil wood (Hesselbo et al., 2002) and compound-specific $\delta^{13}\text{C}$
158 (e.g., Whiteside et al., 2010; Williford et al., 2014) at several locations, supporting their
159 global nature.

160

161 Counter to this interpretation, some $\delta^{13}\text{C}_{\text{org}}$ records lack two clear NCIEs from the TJB
162 interval (Pálffy et al., 2007), and other potentially correlatable NCIEs are identified in
163 uppermost Triassic at some European locations with varied interpretations for their
164 correlation (e.g., Lindström et al., 2017). Whether these NCIEs recorded from Tethyan
165 successions exist in Panthalassa remains outstanding (e.g., Du et al., 2020). Until more
166 data are generated that may resolve these smaller NCIEs (e.g., Heimdal et al., 2020),
167 there is insufficient evidence to support a global driver for their occurrence.

168

169 **2.2 The Triassic-Jurassic Boundary Interval**

170

171 Although the Kuhjoch section in Austria was ratified as the GSSP for the base of the
172 Jurassic (Hillebrandt et al., 2013), the choice of this section has drawn criticism (e.g.,
173 Palotai et al., 2017). The formal base of the Jurassic is defined by the lowest occurrence
174 of *Psiloceras spelae tirolicum* (Hillebrandt et al., 2013) and several other variably utilized

175 stratigraphic markers which typically include a combination of paleontological and
176 geochemical data. For example, carbon isotope stratigraphy has been utilized with the
177 TJB demarcated between the initial and main NCIEs (e.g., Hesselbo et al., 2002; Korte
178 et al., 2019). In terms of paleontological markers, the TJB is defined by the
179 disappearance and/or appearance datums of organisms in three taxonomic groups (see
180 Fig. 2): (i) ammonoids, lowest occurrence of *Psiloceras spelae* and *P. tilmanni* above
181 species of *Rhabdoceras*, *Placites*, *Arcestes*, *Vandaites*, *Cycloceltites* and *Megaphyllites*; (ii)
182 conodonts, the total extinction of the class; and (iii) radiolarians, by the disappearance
183 of *Betraccium*, *Risella*, *Globolaxtorum tozeri*, *Livarella valida*, and *Pseudohagiastrum*
184 *giganteum*, and the appearance of low-diversity spumellarians along with genera
185 *Charlottea*, *Udalia*, and *Parahsuum* s.l. (Carter and Hori, 2005). Radiolarians represent a
186 prominent example showing a temporal relationship between the onset of CAMP
187 volcanism (as marked by geochemical anomalies) and rapid species-level turnover at
188 the ETE / TJB transition (Carter and Hori, 2005; Kocsis et al., 2014).

189

190 Although *Aegerchlamys boellingi* was previously suggested as a marker for the basal
191 Hettangian (e.g., McRoberts et al., 2007), recent correlations of the lower Fernie
192 Formation at Williston Lake, British Columbia Canada (Larina et al., 2019) confirm
193 several levels bearing *Aegerchlamys boellingi* (McRoberts unpublished collections)
194 above the last occurrence of *Monotis subcircularis*. Also concerning the extinction of
195 Class Conodonta at the TJB, reports indicate that *Neohindeodella detrei* occurs in the
196 lowermost Hettangian overlapping with *Psiloceras* and Jurassic radiolarians in Csővár,

197 Hungary (Pálfy et al., 2007; Du et al., 2020). Having additional data with which to assess
198 and/or reinforce these stratigraphic relationships with other Rhaetian fauna is
199 imperative for an improved understanding of the TJB interval and the ETE.

200

201 Absolute calibration of the latest Triassic to TJB interval has been the subject of
202 numerous contributions (e.g., Pálfy et al., 2000; Guex et al., 2012) using a wide variety
203 of radiometric dating techniques in terrestrial and marine sedimentary sequences, but
204 with variable results. Recent U-Pb TIMS dating of two ash layers between the last
205 occurrence of *Choristoceras* and the first occurrence of *Psiloceras* within a TJB section
206 from Peru yielded single-grain U–Pb zircon dates of 201.51 ± 0.15 and 201.39 ± 0.14 Ma
207 (Schoene et al., 2010; Guex et al., 2012; recalculated by Wotzlaw et al., 2014 based on
208 revised tracer calibration). These recalculated dates provide robust age constraints on
209 the TJB.

210

211 In addition, magneto- and cyclo-stratigraphic analyses have been applied in an attempt
212 to provide higher-resolution absolute age constraint(s) on this interval (e.g., Kent et al.,
213 2017; Li et al., 2017; Galbrun et al., 2020). Most prominently, data from the fluvial-
214 lacustrine succession in the Newark Basin have been used to develop a Newark
215 astrochronostratigraphic polarity timescale (or Newark APTS; e.g., Kent et al. 2017).
216 While correlations of some marine successions to the Newark APTS have been
217 proposed (e.g., Maron et al. 2019), most studies of marine successions rely on a

218 combination of biostratigraphic and chemostratigraphic data for temporal constraint
219 and correlation.

220

221 **2.3** *A short vs. long Rhaetian*

222 In contrast to the TJB, there is no consensus on the age of the Norian-Rhaetian
223 Boundary (NRB) and the duration of the Rhaetian (i.e., the youngest age of the Late
224 Triassic). At present, there are divergent age models based on a combination of
225 biostratigraphic, geochemical, and magnetostratigraphic datasets and
226 astrochronologic models that suggest conflicting durations (e.g., Wotzlav et al., 2014;
227 Golding et al., 2016; Li et al., 2017; Kent et al., 2017; Rigo et al., 2020; Galbrun et al.,
228 2020). Models suggest either a *short* or *long* Rhaetian where the lower boundary with
229 the Norian is constrained at 205.7 or 209.5 Ma, respectively, corresponding to a total
230 duration (of the Rhaetian) that could have lasted approximately 4 to 8 Ma (see Li et al.,
231 2017).

232

233 The currently accepted definition of the NRB in marine successions is the first
234 appearance of the conodont *Misikella posthernsteini* (Krystyn, 2010). There is, however,
235 disagreement regarding at what point this species can be considered a distinct taxon
236 from its predecessor *Misikella hernsteini* (e.g., Galbrun et al., 2020), a problem
237 exacerbated by recognition of two distinct morphotypes of *M. posthernsteini*. By using
238 the first occurrence of *M. posthernsteini* in a broader sense (*sensu lato, s.l.*), as in the
239 Steinbergkogel Section near Hallstatt, Austria, the NRB occurs just above a change

240 from a normal to a reverse polarity magnetozone in the 207–210 Ma interval,
241 suggesting a *long* ~8–9 Ma Rhaetian (Krystyn et al., 2007; Muttoni et al., 2010; Li et al.,
242 2017). By using the first occurrence of a more developed form (i.e., *sensu stricto*, s.s.),
243 the duration becomes much shorter (Rigo et al., 2016; Wotzlav et al., 2014). The s.s.
244 case is proposed as the marker for the base of the Rhaetian at the Pignola-Abriola
245 section in Italy, where the NRB is very high within a reversed polarity magnetozone
246 (viz., 205.7 Ma), suggesting a ~4 Ma duration (Maron et al., 2015; Kent et al., 2017). An
247 additional problem is the rare occurrence of *M. posthernsteini* (both s.l. and s.s.) outside
248 the Tethys region, which hampers their use for global correlation.

249

250 Interestingly, interpretations from the terrestrial Newark Supergroup (eastern North
251 America) and the astrochronology and geomagnetic polarity timescale (APTS) derived
252 from it have been used to support both *short* and *long* durations for the Rhaetian.
253 Correlations of marine strata to the Newark APTS 2017 (Kent et al. 2017) indicate that
254 the NRB may occur in either the E17 chron (near the *normal* to reverse polarity flip, at
255 ~209.5 Ma) or the E20 chron (*reversed* polarity at ~206–205 Ma) (as summarized by Li
256 et al., 2017, Fig. 1). A short duration for the Rhaetian requires a ~2–5 Ma hiatus in
257 Newark-APTS (Newark Gap; Tanner and Lucas 2015), but whether such a hiatus exists
258 remains highly contentious (e.g., Kent et al., 2017). These discrepancies in the age
259 models for the Rhaetian help reinforce the importance and need for more studies with
260 diverse sets of chronological data focused on the temporal correlation of this critical
261 interval of time.

262

263 Data presented here from an oceanic Panthalassan locality with abundant fossils and
264 radioisotopically datable bentonite beds crucially offer a new opportunity to assess the
265 timing and duration of the NRB and TJB intervals in a conformable succession with a
266 complete record of those intervals. This is critical for refining timescale calibration and
267 assessing the global timing of carbon cycle perturbations and biotic crises during the
268 ETE.

269

270 **3. Geological setting**

271

272 The Triassic to Lower Jurassic portion of the Wrangellia terrane is conformable and
273 rests nonconformably on a thick succession of flood basalts in the Western Cordillera of
274 North America (Greene et al., 2010). The terrane contains several tectonostratigraphic
275 units across nearly 2000 km throughout westernmost British Columbia and Alaska (Fig.
276 1B). The type section, or northern block, is located in the Wrangell Mountains of
277 Southcentral Alaska, whereas the southern block is best documented on Vancouver
278 Island and Haida Gwaii in western British Columbia, Canada. Although its position in
279 Panthalassa and accretionary history have been debated, paleomagnetic,
280 geochronologic, and paleontologic datasets indicate that Wrangellia was located at
281 tropical latitudes in eastern Panthalassa during the Late Triassic (e.g., Caruthers and
282 Stanley, 2008) before colliding with the continental margin of North America during the

283 Middle Jurassic (southern block) and Cretaceous (northern block; e.g., Trop et al.,
284 2020).

285

286 The Upper Triassic portion of Wrangellia represents an extensive carbonate platform
287 and reef system inhabited by abundant and locally diverse marine biota (e.g., Caruthers
288 and Stanley, 2008). In the Wrangell Mountains this section is represented by two
289 calcareous units: the supratidal/intertidal to shallow subtidal, thick- to very thick-
290 bedded, Chitistone Formation and the deeper water, medium- to thick-bedded, Nizina
291 Formation which together form a ~100 m-thick succession deposited during Carnian
292 to late Norian times (Armstrong et al., 1969). During the Norian, thermal subsidence of
293 Wrangellia's northern block is thought to have initiated the drowning of the carbonate
294 platform, resulting in deposition of ~540 m of calcareous and siliceous mudstones
295 comprising the McCarthy Formation (Greene et al., 2010). The uppermost Triassic and
296 lowermost Jurassic strata of the lower McCarthy Formation are the focus of this study.

297

298 **4. Materials and methods**

299

300 We studied the upper Norian to middle Hettangian lower McCarthy Formation along an
301 unnamed tributary of Grotto Creek, located near its headwaters (base of the section:
302 61°30' 13.23"N, 142°26' 31.51"W; Fig. 1C), ~25 km east-northeast of McCarthy, Alaska
303 (Fig. 1C). This section (Grotto Creek section) was originally described by Witmer (2007),
304 who presented a preliminary stratigraphic log and carbon isotope stratigraphy (~20 m

305 sample spacing) along with sparse paleontological samples and preliminary U-Pb zircon
306 dates of ~214 and 209 Ma from two bentonites within and stratigraphically below our
307 measured section. To constrain the age of our measured section, we report final high-
308 precision CA-ID TIMS U-Pb zircon dates herein from the bentonite samples studied by
309 Witmer (2007; see SI Table 1.2).

310

311 We measured and described 96 m of conformable stratigraphy consisting mostly of
312 buff-weathering, black, carbonaceous, siliceous mudstones and calcareous cherts with
313 textures that alternate between fine mudstones, sandy mudstones, and muddy
314 sandstones. Bentonites occur frequently throughout the middle portion of the section.
315 We placed the 0 m datum of the section (i.e., Fig. 3) at the base of an easily
316 recognizable 5 cm-thick bentonite just below the biostratigraphically defined Norian-
317 Rhaetian boundary. The lower ~26 m are more resistant and cliff-forming due to the
318 presence of medium-thick beds of sandy mudstone with fine mudstone partings. These
319 alternate with more recessive intervals of fine mudstones. Several beds within this
320 lower interval are laminated. At ~3 m there is a ~12 m-high asymmetric fold within an
321 otherwise normally bedded stratigraphic succession (Fig. 4A). We interpret this
322 structure as syndimentary soft-sediment deformation related to the depositional
323 slope. The upper ~70 m of the section is a slope-forming succession where thin-bedded
324 fine mudstones are more prevalent than in the lower ~26 m of the section. The more
325 prominent strata are thin to medium-thick beds of calcareous and siliceous sandy

326 mudstones and fine calcareous cherts. In this upper interval, sedimentary structures
327 have mostly been destroyed by bioturbation.

328

329 We collected 70 samples of carbonaceous, siliceous mudstones for $\delta^{13}\text{C}_{\text{org}}$ and whole-
330 rock total organic carbon (TOC_{wr}) analyses using continuous-flow isotope ratio mass
331 spectrometry (SI Text 1), and four bentonite samples for zircon U-Pb CA-ID TIMS
332 analysis (SI Text 1-3). Additionally, we collected 30 samples for conodont analysis and
333 103 *in situ* and float macrofossil specimens (ammonoids, bivalves, and hydrozoans)
334 from 51 fossiliferous horizons. Fossils are preserved as whole-body specimens and as
335 internal and external molds.

336

337 Ammonoid zonation follows Tozer (1994) for the Upper Triassic and Taylor et al. (2001)
338 for the Lower Jurassic, applicable to assemblage zones. Paleontological data are
339 presented in Figures 3–6, geochemical data in Figures 3, 7, and 8, and supplementary
340 files contain expanded methodologies, expanded results, and interpretation of
341 geochronology analytical details (SI Text 1–4; SI Fig. 1; SI Tables 1–5). Collected
342 paleontological specimens are curated at the Wrangell-St. Elias National Park and
343 Preserve, with corresponding collections permit numbers (see acknowledgements and
344 SI Table 1.1).

345

346 Magnetostratigraphy was not attempted on the Grotto Creek Section. Previous studies
347 by Coe et al. (1985) and Hillhouse and Coe (1994) have shown generally that while

348 Mesozoic volcanic rocks of northern Wrangellia most likely preserved their primary
349 signal, the interbedded and overlying sediments (viz., Cretaceous and Tertiary) have
350 most likely been re-magnetized. Stamatakos et al. (2001) also reinforced these findings
351 by showing that while Cretaceous strata exposed ~20 km south of Grotto Creek at
352 MacColl Ridge are not remagnetized, the sediments in the Grotto Creek section (i.e.,
353 those lying within the outcrop belt of Neogene volcanics/intrusions known as the
354 Wrangell arc) have likely had their paleomagnetic record reset. This is further bolstered
355 by preliminary Rock-Eval pyrolysis data from the McCarthy Formation by Witmer (2007,
356 p. 29, Appendix C) showing high maturity and T_{\max} values from 461 to 482 °C.
357 Altogether, this evidence suggests that the McCarthy Formation may not be a suitable
358 candidate for magnetostratigraphic analysis.

359

360 5. Results

361

362 Paleontological data from the base of the section, below reported carbon isotope
363 values, show that the bivalve *Monotis* (*M. cf. alaskana*, *M. subcircularis*, and *M. sp.*)
364 occurs in abundance from -30 m to ~-19 m, with the highest occurrence as float at
365 -18.85 m (Fig. 3). At -18.65 m, the conodonts *Mockina* sp., *Norigondolella*
366 *steinbergensis*, and *Misikella hernsteini* were recovered along with float ammonoids
367 *Rhacophylites debilis* (-20.6 to ~-4 m). At -15.23 m, there is a narrow ~0.5 m- thick
368 interval with abundant *in situ* species of the hydrozoan *Heterastridium*, the spheroidal
369 form *H. conglobatum* (Fig. 4B), and the flattened discoidal form *H. disciforme* (Fig. 4D–

370 H, J). Species identification of this group is based on revised systematic descriptions in
371 Senowbari-Daryan and Link (2019). The conodont *Mockina* sp. was recovered at
372 -10.1 m and float specimens of the bivalve ?*Leptochondria* sp. and the ammonoid
373 *Rhacophylites debilis* at -4 m.

374

375 From -2.1 to 6.95 m, the conodont *Mockina bidentata* was recovered close to a float
376 ammonoid *Sagenites* sp. 1 (~-2.1 m; Fig. 3), with *in situ* and float specimens of the
377 bivalve *Agerchlamys boellingi* overlapping with ammonoids *Rhacophylites debilis* and
378 *Sagenites* sp. 2 (2.95 to 4.1 m). At 4.15 m the ammonoid *Vandaites* cf. *suttonensis* was
379 found *in situ* along with the ammonoids ?*Paracochloceras* cf. *amoenum* and *Placites*
380 *polydactylus* and *Agerchlamys* cf. *boellingi* (4.95 to 6.95 m). This is followed by a ~20 m-
381 thick interval with several *in situ* and float taxa including: *Agerchlamys boellingi*, *Mockina*
382 *bidentata*, *Mockina englandi*, *Mockina mosheri* morphotype B, *Norigondolella* sp.,
383 *Sagenites* cf. *minaensis*, and *Choristoceras rhaeticum*.

384

385 At 29.42 m, the ammonoid ?*Psiloceras* sp. was recovered *in situ* along with the
386 conodont *Neohindeodella* sp. followed by float and *in situ* occurrences of the ammonoid
387 *Psiloceras tilmanni* (~33.95 to 35.45 m), *Agerchlamys* cf. *boellingi* (~37.95 to 38.95 m), and
388 float specimens of the ammonoid *Psiloceras polymorphum* (~40.95 to 45.95 m). Near
389 the top of the section, the ammonoids *Transipsiloceras* sp., *Nevadaphyllites* aff.
390 *compressus*, and *Pleuroacanthites* cf. *biformis* were recovered along with *Agerchlamys*
391 cf. *boellingi* (spanning ~45.95 to 64.75 m; Fig. 3).

392

393 The four sampled bentonites were collected from (i) 50 m above the base of the
394 McCarthy Formation (i.e., Grot-1, Fig. 7, occurring below the base of our measured
395 section); (ii) approximately -6 to 0 m in our section (i.e., Grot-124, position
396 approximated based on correlation with Witmer, 2007, discussed below in section 6.1);
397 (iii) 0 m (i.e., 2017GC3.8); (iv) 11.07 m (i.e., 2017GC14.9) (Figs. 3, 7). Bentonites (i) and (ii)
398 are finalized data originally collected by Witmer (2007) and (iii) and (iv) are new to this
399 study. We interpret the bentonites as four separate volcanic events and associated
400 settling of volcanic ash through the water column with no sedimentary evidence for
401 reworking or abrasion of the grains. The bentonites form yellow-weathering thin
402 (<10 cm) recessive beds and contain elongate euhedral to subhedral crystals with minor
403 inclusions. Well-developed zoning patterns are present in imaged grains (sample
404 2017GC3.8, SI Fig. 1), and tight clusters of dates occur from analyzed grains within each
405 respective sample (see SI text 2, 3 for an expanded justification for our interpretation of
406 the bentonites).

407

408 U-Pb chemical abrasion-isotope dilution (CA-ID) TIMS analysis were carried out at the
409 University of British Columbia (UBC) and the Massachusetts Institute of Technology
410 (MIT). All samples were run using the EARTHTIME 535 tracer (calibration v. 3), thus
411 minimizing interlaboratory biases. Complete results, photomicrographs and/or
412 cathodoluminescence images of zircon grains, and laser ablation-derived trace element

413 concentration data are presented as Supplemental Information (SI Text 2; Fig. 1; Tables
414 2-5).

415

416 Eleven single-grain analyses from sample Grot-124 yielded overlapping Th-corrected
417 $^{206}\text{Pb}/^{238}\text{U}$ dates from 210.10 ± 0.16 to 209.73 ± 0.25 Ma (Fig. 7A), with a weighted mean
418 of 209.92 ± 0.043 Ma (MSWD = 1.6), which we interpret as the eruption age of the
419 sample (reported uncertainties are 2-sigma internal). Ten single-grain analyses from
420 sample Grot-1 yielded a range of Th-corrected $^{206}\text{Pb}/^{238}\text{U}$ dates from 245.8 ± 2.0 to
421 213.2 ± 1.6 Ma (excluding a single low precision analysis, z27). Eight of the 10 analyses
422 shown on Fig. 7A overlap within uncertainty with a Th-corrected weighted mean
423 $^{206}\text{Pb}/^{238}\text{U}$ date of 214.36 ± 0.19 Ma (MSWD = 1.2), which we interpret as the eruption
424 age of this sample—the two older zircon grains (246–221 Ma) are likely inherited (not
425 shown on Fig. 7). Six dated grains from sample 2017GC3.8 (0 m, Fig. 3) yielded dates of
426 210.60 ± 0.31 to 209.73 ± 0.25 Ma. The data comprises distinct younger (3 results) and
427 older (2 results) groupings, and a relatively imprecise result (not plotted, Fig. 7A) that
428 spans the two clusters. A weighted mean $^{206}\text{Pb}/^{238}\text{U}$ date of 209.86 ± 0.16 Ma for the
429 younger cluster is interpreted as the best estimate age, with older grains interpreted as
430 antecrysts or xenocrysts. For sample 2017GC14.9 (11.07 m, Fig. 3), two younger grains
431 yield a weighted mean $^{206}\text{Pb}/^{238}\text{U}$ date of 208.25 ± 0.25 Ma, and a single older grain is
432 likely a xenocryst (Fig. 7).

433

434 TOC_{wr} values range ~0.5–3 wt%, with an average of 1.5 wt% (Fig. 3). TOC_{wr} is variable
435 through the upper Norian (up to ~4.15 m) in the section, followed by a trend towards
436 lower values in the Rhaetian (~19.95 m) before gradually increasing across the TJB,
437 peaking at 2.7 wt% (~31.95 m; Fig. 3). Values stabilize through the Spelae-Pacificum
438 zones and remain below 2 wt% (apart from one value of 2.6 wt% at 51.97 m) to the top
439 of the section. $\delta^{13}\text{C}_{\text{org}}$ values become gradually less negative from -29‰ to -28‰
440 through the Rhaetian with two decreases occurring in close proximity to the TJB: the
441 first from -27.56‰ to -29.22‰ (26.42 to 30.03 m), and a second from -27.92‰ to
442 -29.26‰ (32.46 to 35.97 m). Above this, $\delta^{13}\text{C}_{\text{org}}$ values gradually increase from ~-29‰
443 to -27.5‰ at the top of the measured section (Fig. 3).

444

445 6. Discussion

446

447 Our data from the Wrangellia terrane represent an important addition to the global
448 database of Upper Triassic to Lower Jurassic successions. Biostratigraphy shows a
449 complete (i.e., *Cordilleranus* to *Mulleri*) ammonite zonation in the Grotto Creek section
450 with no obvious long breaks in sedimentation, suggesting a complete record from
451 upper Norian to lower-middle Hettangian. These data not only improve the resolution
452 of timescale calibrations, but also provide a more holistic understanding of
453 biogeochemical dynamics associated with the ETE from Panthalassa. Here, we
454 establish the Grotto Creek section as an important succession with respect to the (i)
455 debated long vs. short duration of the Rhaetian, (ii) paleontological and geochemical

456 trends across the TJB, and (iii) implications of the VGS and controlling mechanisms of
457 the ETE.

458

459 **6.1** *A case for a long Rhaetian*

460

461 Precise quantification of the duration of the Rhaetian Stage is pivotal for understanding
462 the timing of the events surrounding the ETE. At present, various lines of indirect
463 evidence are used to argue for the initiation of CAMP magmatism prior to the oldest
464 dated igneous bodies (e.g., Davies et al., 2017). These include seismites, basalt-derived
465 sediments directly below CAMP basalts, and eustatic sea-level fall during the Rhaetian,
466 as evidence of short-term climatic cooling (induced by volcanic SO₂) and the VGS (e.g.,
467 Schoene et al., 2010). Importantly, this early initiation is invoked to explain possible
468 diachroneity between mass extinction in the marine and terrestrial records (e.g., Pálffy
469 et al., 2000), and therefore it is essential to better constrain the duration of the
470 Rhaetian.

471

472 In the Grotto Creek section the NRB (Fig. 4A, yellow line) occurs at 4.15 m, just above
473 the ~12 m-high soft-sediment deformation fold (Fig. 4A at right), temporally
474 constrained through biostratigraphic data and the ~209 Ma U-Pb zircon CA-ID-TIMS
475 dates from bentonites in the lower McCarthy Formation (Figs. 3, 7; SI Text 2, SI Fig. 1, SI
476 Tables 1-5).

477

478 From the section base to 4.15 m, a late Norian Cordilleranus Zone age is indicated by
479 occurrences of *Monotis*, *Heterastridium*, ammonoids, and age-specific conodonts (Figs.
480 2-6). The last *in situ Monotis* occurs at -24.87 m, uppermost float *M. subcircularis* at
481 -18.85 m, and lowest *in situ Heterastridium* at -15.23 m. According to Senowbari-
482 Daryan and Link (2019), previous accounts of *Heterastridium* from the Carnian and
483 Rhaetian stages are doubtful, and this genus is restricted to the Norian Stage. From
484 3.24 to 4.15 m, *in situ Rhacophyllites debilis* overlaps with the lowest *in situ Agerchlamys*
485 *boellingi* and the strictly Rhaetian ammonoid *Vandaites cf. suttonensis* (at 4.15 m),
486 marking the NRB at Grotto Creek (~4 m, Fig. 3).

487

488 The abundance of bentonite beds (orange lines in Fig. 3) in this part of the section
489 hampers the exact placement of the dated bentonite bed collected by Witmer (2007;
490 i.e., Grot-124, Figs. 3, 7, 209.92 ± 0.043 Ma) within our measured section. Witmer (2007)
491 noted that Grot-124 occurs 19 m above the last occurrence of *Monotis*. This is estimated
492 at ~-6 to 0 m in our section, bounded by our uppermost measured *in situ Monotis* (at
493 -24.87 m) and the uppermost float *M. subcircularis* (-18.83 m); this is demarcated by a
494 dashed, red-lined box of uncertainty in Fig. 3. Stratigraphically, this interval is just
495 below our new dates of 209.86 ± 0.16 Ma and 208.25 ± 0.25 Ma from 0 and 11.07 m,
496 respectively, which span the NRB (~4 m, Fig. 3). The characteristics of the zircons (SI
497 Text 2, 3; SI Fig. 1) and the tight clusters of dates (Fig. 7) indicate a primary magmatic
498 age. Overall, this is consistent with a long duration (~8 Ma) for the Rhaetian from ~209–
499 201.4 Ma.

500

501 The interpretation presented here of a long duration Rhaetian Stage is similar to that
502 derived from the Steinbergkogel Austria section (e.g., Li et al., 2017; Fig. 1), which uses
503 *M. posthernsteini* s.l. for the NRB datum, but in the Grotto Creek section we use the first
504 occurrence of the ammonoid *Vandaïtes suttonensis* as the NRB indicator (which has
505 been shown to be restricted to the Rhaetian; Tozer, 1994; e.g., Fig. 2). In the Grotto
506 Creek section, samples collected for conodont analysis from this interval were barren
507 and no specimens of *Misikella posthernsteini* (s.s. or s.l.) were recovered. A dominance
508 of late Norian taxa low in the section followed directly by *in situ* *Agerchlamys boellingi*
509 and *Vandaïtes* cf. *suttonensis* at ~3.9 m, with a variety of Rhaetian-restricted taxa
510 above, however, strongly support the placement of NRB.

511

512 Our duration for the Rhaetian appears at odds with the record from Levanto in Peru
513 where similar lines of evidence are used in support of a short-duration Rhaetian (i.e.,
514 last occurrence of *Monotis* below *Vandaïtes* with no reported occurrence of NRB-
515 defining conodont *M. posthernsteini* s.s. or s.l.; Wotzlaw et al., 2014). An important
516 detail concerning the Levanto succession, however, is that Wotzlaw et al. (2014; fig. 2)
517 report primary magmatic dates of ~205 Ma from bentonites that occur ~5 meters above
518 the last occurrence of *M. subcircularis* and ~50 meters below the first occurrence of
519 *Vandaïtes*. At Grotto Creek, primary magmatic dates of ca. 209 to 208 Ma were derived
520 from bentonites that occur above the last occurrence of *M. subcircularis* and bracket the
521 first occurrence of *Vandaïtes* cf. *suttonensis* (i.e., Figs. 3, 7B). Per Wotzlaw et al. (2014)

522 and using a similar argument as Galbrun et al. (2020), if the extinction of *Monotis* was
523 relatively globally synchronous, then the discrepancy between the Grotto Creek and
524 Levanto stratigraphies and our probable primary magmatic dates suggest that the
525 Levanto section contains unidentified hiatus(es) and/or is condensed over the Norian-
526 Rhaetian transition.

527

528 In summary, it becomes apparent that given the wide array of complicating factors
529 surrounding the NRB (i.e., current definition and potential stratigraphic complexities
530 with the existing records), the definition should be revised to include multiple lines of
531 data that can be applied globally. As previously noted, various correlations of marine
532 strata to the Newark-APTS have been used to argue for both a long and short Rhaetian.
533 The new U-Pb dates from Grotto Creek place the NRB in the reverse or normal polarity
534 intervals of the E17 chron of Newark-APTS 2017 (Kent et al. 2017). This correlation
535 supports age models that lack a gap in the Newark succession (e.g., Kent et al. 2017)
536 and also that the first appearance *Misikella posthernsteini s.l.* and not *Misikella*
537 *posthernsteini s.s.* marks the NRB (e.g., Krystyn et al., 2007).

538

539 Carbon isotope stratigraphy has recently been suggested to provide an additional
540 constraint, as recent work has suggested that a NCIE may occur in the NRB interval
541 (Rigo et al., 2020). Although rigorous evaluation of the geographic extent of this CIE is
542 outstanding, the negative values at -2.79 and 0.22 m in the Grotto Creek section may
543 correlate with this NRB NCIE. Since our data do not extend below this interval, we

544 cannot at present confidently identify this trend at Grotto Creek as being correlative
545 with this suspected NRB NCIE. Nevertheless, a new multi-faceted definition of the NRB
546 is needed to provide a means to overcome shortcomings in any one kind of datum and
547 provide a more utilitarian means to correlate strata globally.

548

549 **6.2** *The Triassic-Jurassic boundary Interval at Grotto Creek*

550

551 A TJB transition interval is defined with our combined paleontological and geochemical
552 ($\delta^{13}\text{C}_{\text{org}}$) data from the Grotto Creek section. Overlying the NRB, there is a ~22 m-thick
553 interval (up to 26.65 m) that contains Rhaetian ammonoids and an assortment of
554 Norian-Rhaetian conodonts and bivalves (Figs. 3, 5, 6). While *Choristoceras rhaeticum* is
555 known to be restricted to the Crickmayi Zone (Tozer, 1994), its occurrence at 26.65 m is
556 from float and therefore we cannot currently designate a Crickmayi Zone boundary.
557 Furthermore, the lowest *in situ* *Agerchlamys boellingi* is 0.08 m below the NRB, which
558 places this species within the uppermost Norian, in agreement with previous accounts
559 for a Late Triassic origin (e.g., Larina et al., 2019) and refuting its utility as a defining
560 species of the TJB.

561

562 From 29.42 to 35.46 m, the TJB is defined based on the co-occurrence of the lowest *in*
563 *situ* strictly Jurassic genus *Psiloceras* (i.e., ?*Psiloceras*) and the highest *in situ* conodont
564 (*Neohindeodella* sp.), both at 29.42 m, and the lowest *in situ* *Psiloceras* cf. *tilmanni* at
565 35.46 m (Fig. 3 shaded region; Fig. 4A red line). The poor preservation of ?*Psiloceras* (at

566 29.42 m) above the highest float *Choristoceras rhaeticum* precludes unequivocal
567 delineation of the TJB, which requires a TJB interval of ~6 m in the section. Regardless,
568 the occurrence of *P. cf. tilmanni* is a robust indication of the lower Hettangian (Figs. 2A,
569 5), which marks the upper limit (of the ~6 m TJB interval). This is followed by two *in situ*
570 occurrences of *A. cf. boellingi* and an assortment of float ammonoids from the
571 Pacificum (e.g., *Psiloceras pacificum*), Polymorphum (e.g., *Psiloceras polymorphum* and
572 *Transipsiloceras* sp.), and Mulleri (e.g., *Pleuroacanthites cf. biformis*) zones representing
573 the lower to middle Hettangian (Figs. 2, 5).

574

575 Organic carbon isotopes in the uppermost Rhaetian record a ~1.3‰ positive carbon
576 isotope excursion (PCIE) from 23.69 to 26.42 m (Fig. 3). This is followed by an abrupt
577 NCIE of 1.7‰ that is broad in character (i.e., ~15 m in stratigraphic thickness), which
578 begins at 26.42 m and extends through to the top of the Spelae–Pacificum zones at
579 40.94 m (Figs. 3, 8). Within this broad NCIE, two further NCIEs occur with a magnitude
580 of 1.7‰ and ~1.3‰ at 26.42 and 32.46 m, respectively. Altogether, this broad trend in
581 organic carbon isotope values is consistent with other global TJB records (Fig. 8, see
582 discussion below).

583

584 **6.3 Global vs. regional carbon cycle perturbations and the ETE**

585

586 Available records of the TJB interval show numerous small-magnitude fluctuations in
587 organic carbon isotopes. The stratigraphic and geographic distribution of these CIEs

588 have implications regarding their underlying drivers and utility for regional to global
589 correlation. Here, we briefly review some of the existing carbon isotope records in
590 attempt to reconcile important differences and help develop a more complete
591 understanding of environmental changes enveloping the ETE.

592

593 Most studies of the ETE and TJB $\delta^{13}\text{C}_{\text{org}}$ records are from the westernmost Tethys and
594 have signatures that commonly delineate two NCIEs: the first occurs below the TJB,
595 commonly referred to as the *initial* NCIE ($\sim 2\text{--}5\text{‰}$), and the second, referred to as the
596 *main* isotope excursion ($\sim 5\text{‰}$), occurs just above the base of the Jurassic (Hesselbo et
597 al., 2002). Additionally, the available terrestrial carbon-isotope records across this
598 interval (i.e., East Greenland, Poland, and Denmark) show a similar *initial* NCIE below
599 the TJB with a *main* NCIE above (e.g., Steinhorsdottir et al. 2011; Pieńkowski et al.
600 2012; Korte et al., 2019).

601

602 Recent work by Ruhl and Kürschner (2011), Lindström et al. (2017), and others expand
603 the number of NCIEs to three based on ammonoid and palynoflora occurrences in
604 sections primarily from the westernmost Tethys, identifying them (in stratigraphic
605 order) as the: Precursor (or Marshi; correlative within the last occurrence of the
606 Rhaetian ammonite *Choristoceras marshi*), Spelae (correlative with the *initial* NCIE
607 occurring within the earliest Hettangian), and top-Tilmanni (correlative with the *main*
608 NCIE occurring at a slightly higher position in the early Hettangian). Most recently,

609 Kovács et al. (2020) show many small-scale anomalies in both the $\delta^{13}\text{C}_{\text{carb}}$ and $\delta^{13}\text{C}_{\text{org}}$
610 records across this TJB transition from the western Tethys shelf (Csővár, Hungary).
611
612 To date, however, the three larger-magnitude and multiple higher-frequency, smaller-
613 magnitude NCIEs observed in the Tethyan records have not been clearly identified
614 within Panthalassan successions. Here, we assess features of the TJB organic carbon
615 isotope record that can be delineated and reliably correlated across Panthalassa and
616 then assess potential correlations to records from the Tethys (Fig. 8). This opens the
617 door to a discussion concerning the ubiquity of these smaller NCIEs and helps to
618 delineate regional *versus* global signals across the TJB organic carbon isotope record.
619
620 Compilation TJB data from Wrangellia and Eastern Panthalassa show a PCIE of $\sim 1.5\text{‰}$
621 to $\sim 2.0\text{‰}$ that occurs in the upper Rhaetian (green shading on Fig. 8), which appears of
622 larger ($\sim 5\text{‰}$) magnitude in Central Panthalassa (e.g., deep-water chert deposits in
623 Japan). This is followed by a NCIE that initiates toward the end of the Rhaetian near the
624 top of the Crickmayi/Marshi ammonite zone beginning just at, or before, the extinction
625 interval that precedes the TJB (blue shading on Fig. 8). The overall magnitude of the
626 NCIE varies from 1.66‰ to 4.94‰ and appears to contain higher-order oscillations in
627 most of the Panthalassan successions. In Nevada, however, it should be noted that
628 existing data do not extend low enough in the stratigraphy to confirm a PCIE. Timing of
629 the initiation of the PCIE and NCIE are constrained by the Peruvian Levanto section,

630 where two bentonite beds at these intervals have been dated to 201.87 ± 0.17 Ma and
631 201.51 ± 0.15 Ma, respectively.

632

633 We compare these features of Panthalassa to those recorded in the Tethys and suggest
634 a more simplified global correlation. Here, we use the St. Audrie's Bay (England) and
635 Kuhjoch West (Austria) records as points of reference, as nearly all other Tethyan
636 records are compared to these (e.g., Korte et al., 2019; Kovács et al., 2020). We note,
637 however, that these records are inherently problematic: the TJB transition at St.
638 Audrie's Bay records a transition from continental / marginal marine to fully marine
639 environments, and a shear zone deforms the Kuhjoch West section at the stratigraphic
640 interval that records the onset of the *main* NCIE (Ruhl et al., 2009, Palotai et al., 2017).

641

642 Nevertheless, in comparison to these schemes, the PCIE from Panthalassa corresponds
643 to a $\sim 5.5\%$ PCIE in the upper Rhaetian at St. Audrie's Bay that is just below the *initial* (= *Spelae*
644 CIE) and well below the *main* (= top-Tilmanni CIE). A similar feature occurs
645 broadly at the same level in many other Tethyan $\delta^{13}\text{C}_{\text{org}}$ records (e.g., Lindström et al.,
646 2017; Korte et al., 2019). Specifically, at Kuhjoch West, an *initial* NCIE occurs below 0 m
647 and the *main* NCIE at ~ 2.5 m in section (Fig. 8; Ruhl et al., 2009, Hillebrandt et al., 2013).

648

649 The overlying NCIE spans the uppermost Rhaetian into the Hettangian, corresponding
650 to (and containing) the *initial* (*Spelae*) and *main* (top-Tilmanni) CIEs. These events are
651 likely higher-frequency oscillations contained within a temporally broader NCIE. To this

652 point, the St. Audrie's Bay and Kuhjoch West records also contain other higher-
653 frequency $\delta^{13}\text{C}$ oscillations (or NCIEs) of similar magnitude (up to 3‰) stratigraphically
654 above and below the previously described *initial* and *main* NCIEs.

655

656 Given that these higher-order features observed in the Tethys either do not appear or
657 are subdued in the open ocean records of Panthalassa, there exists at present a need for
658 a more conservative definition of the global $\delta^{13}\text{C}_{\text{org}}$ record of the TJB interval. This new
659 definition should be centered on open ocean records and account for local dynamics
660 that either magnify $\delta^{13}\text{C}_{\text{org}}$ in regional records of individual sedimentary basins or
661 dampen global signals.

662

663 Deciphering such global *versus* regional signals across the TJB has important
664 implications for environmental changes and carbon cycle dynamics controlling the ETE.
665 The driving mechanisms at the onset of the broader NCIE are coincident (within error)
666 with the first major evidence of CAMP volcanism dated to 201.566 ± 0.031 Ma
667 (Blackburn et al., 2013). Alternatively, Davies et al. (2017) emphasized the role of
668 subvolcanic intrusions whose emplacement preceded the first eruptive phase and may
669 have contributed degassing of greenhouse gases through contact with organic-rich
670 sedimentary rocks. Regardless, input of ^{12}C -enriched carbon to the ocean-atmosphere
671 from CAMP has long been invoked as the driver of these NCIEs.

672

673 The finer-scale NCIEs, if global, could reflect inputs of ^{12}C -enriched carbon to the ocean
674 and atmosphere from discrete eruptive phases of CAMP or other carbon cycle
675 feedbacks (e.g., methane releases, global declines in productivity, response of
676 terrestrial carbon cycling; e.g., Heimdal et al., 2020.). This is substantiated by a second
677 known eruptive phase at 201.274 ± 0.032 Ma (Blackburn et al., 2013), which potentially
678 correlates in time to the initiation of a second negative shift in $\delta^{13}\text{C}_{\text{org}}$ at Levanto (e.g.,
679 ~65 m in that section; Fig. 8). Alternatively, if higher-order NCIEs are only regionally
680 correlative (i.e., do not occur in open-ocean Panthalassan environments), this could
681 indicate a dominance of local/regional influences on the $\delta^{13}\text{C}_{\text{org}}$ record, which should
682 not be factored into interpretations and modeling of the global carbon cycle.

683

684 Therefore, it becomes evident that determining the global *versus* regional nature of
685 isotope excursions surrounding the TJB remains an outstanding and important
686 challenge, critical to understand the end-Triassic mass extinction. We posit that new
687 multi-proxy, multi-lithology, and higher-resolution studies are required to fully address
688 the underlying mechanisms, magnitudes, and outstanding uncertainties of the carbon
689 isotope record around the ETE.

690

691 **7. Conclusions**

692

693 Paleontological and geochemical data were collected from the Grotto Creek section
694 (Wrangell Mountains, Alaska) representing undisturbed deposition on the oceanic

695 plateau of Wrangellia in open Panthalassa during Late Triassic to Early Jurassic time.
696 Data suggest (i) an upper Norian (Cordilleranus Zone) succession spanning the lower
697 ~34 m of the section, well constrained by abundant occurrences of *Monotis*,
698 *Heterastridium*, and age-specific conodonts; (ii) the NRB at 4.15 m marked by the
699 appearance of the Rhaetian heteromorph ammonoid *Vandaites* cf. *suttonensis*,
700 supported by overlying Rhaetian-restricted ammonoids and assorted Norian–Rhaetian
701 conodonts and bivalves; (iii) three new primary magmatic U-Pb CA-ID TIMS dates of
702 209.92 ± 0.043 , 209.86 ± 0.16 and 208.25 ± 0.25 Ma from bentonites that straddle the
703 NRB, suggesting a boundary age of ~209 Ma (in line with a longer, ~8 Ma, Rhaetian); (iv)
704 a stratigraphically continuous TJB transition interval from 29.42 to 35.46 m marked by
705 *?Psiloceras* sp., *Neohindeodella* sp., and *P.* cf. *tilmanni*, and followed by an assortment of
706 float ammonoids from the early to middle Hettangian Polymorphum to Mulleri zones;
707 and (v) a new, simplified, interpretation of the $\delta^{13}\text{C}_{\text{org}}$ record across the TJB, whereby a
708 PCIE of variable magnitude is directly followed by an NCIE that is subdued in open-
709 ocean Panthalassa but contains many second-order features in the Tethys and marginal
710 Panthalassa, potentially highlighting regional carbon cycle dynamics during a time of
711 global carbon cycle perturbation. This combined biostratigraphic and geochemical
712 record of the Upper Triassic to Lower Jurassic succession at Grotto Creek (Alaska) is the
713 best-known record of the NRB and TJB intervals from not only Wrangellia, but from all
714 the other terranes in western North America.

715

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747

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1027

1028 **Figure Captions**

1029

1030 **Figure 1: A.** Global Late Triassic (~220 Ma) paleogeographic reconstruction showing
1031 the approximated location of the Central Atlantic Magmatic Province (CAMP) at the
1032 TJB, the allochthonous terrane Wrangellia, and relevant coeval marine and terrestrial
1033 records (base map after Blakey, 2014; data localities after Hesselbo et al., 2002;
1034 Whiteside et al., 2010; Schoene et al., 2010; Williford et al., 2014 and references
1035 therein). Dashed arrow indicates hypothetical direction of future tectonic displacement
1036 of northern Wrangellia. **B.** Present-day tectonic map of western North America
1037 showing location of the Wrangellia composite terrane, the Wrangell Mountains, and
1038 Haida Gwaii (modified from Colpron and Nelson, 2009). **C.** Photograph of the Grotto
1039 Creek section showing the relevant stratigraphy, approximate location of measured

1040 section (A–A' ; base is below ridge in foreground at 61°30' 13.23"N, 142°26' 31.51"W),
1041 and positions of the Norian-Rhaetian boundary (yellow line) and Triassic-Jurassic
1042 boundary (red line).

1043

1044 **Figure 2:** Taxonomic range chart and zonal schemes for selected Late Triassic to Early
1045 Jurassic faunas of North America. **A.** Hydrozoans (after Senowbari-Daryan and Link,
1046 2019), bivalves (after McRoberts et al., 2007) conodonts (Rigo et al., 2018 and others);
1047 ammonoids (after Tozer, 1994; Taylor et al., 2001; Guex et al., 2004; Longridge et al.,
1048 2007; 2008). **B.** Relevant conodont, ammonoid and radiolarian zones of North America
1049 (after Rigo et al., 2018 and references therein). *Note:* Nor = Norian; Ammonoid
1050 zonations used herein denote a zone name with reference to an assemblage of
1051 taxonomic ranges, rather than the range of a particular species.

1052

1053 **Figure 3:** Compilation data from the Grotto Creek section, Alaska (base at
1054 61°30' 13.23"N, 142°26' 31.51"W) showing combined lithological, paleontological, and
1055 geochemical results. *Note:* Shaded area represents the suspected TJB interval; vertical
1056 hash marks indicate intervals of poor exposure; Dashed red box and corresponding
1057 dashed red arrows represent suspected interval of dated ash by Witmer (2007), solid
1058 black arrows and boxes denote new dates in this study; filled circles are *in situ* fossil
1059 occurrences, open circles are float specimens; TOC_{wr} denotes Total Organic Carbon
1060 measured from whole rock; Sp. = Spelae; Pac. = Pacificum; exp. = exposure.

1061

1062 **Figure 4:** Photographs of selected strata and specimens in the lower McCarthy
1063 Formation, Grotto Creek section. Fossil horizons refer to stratigraphic location in Fig. 3;
1064 all specimens natural size unless indicated (e.g., X2). **A.** Field photograph of the Norian-
1065 Rhaetian Boundary (NRB; yellow line) and Triassic-Jurassic Boundary (TJB; red line)
1066 intervals; asymmetric fold at right is ~12 m high. **B.** Field photograph of the middle to
1067 late Norian spherical hydrozoan *Heterastridium conglobatum* at -17.67 m in the section;
1068 *in situ* between fossil horizons 11 and 12 on Fig. 3 (specimen not collected). **C.** *Monotis*
1069 *subcircularis* (multiple) *in situ* at fossil horizon 6, Cordilleranus Zone, late Norian, natural
1070 size. **D and E.** *Heterastridium disciforme* float at fossil horizon 2, middle to late Norian,
1071 natural size (D, surface view; E, longitudinal view). **F.** Longitudinal view of
1072 *Heterastridium disciforme*, float at fossil horizon 2, middle to late Norian, natural size. **G.**
1073 **and H.** *Heterastridium disciforme* float at fossil horizon 2, middle to late Norian, natural
1074 size (G, surface view; H, longitudinal view). **I.** *Monotis* cf. *alaskana*, float at fossil horizon
1075 4, Cordilleranus Zone, late Norian, natural size. **J.** Field photograph showing many
1076 discoid specimens of *Heterastridium disciforme in situ* at -15.23 m (fossil horizon 12, Fig.
1077 3), middle to late Norian.

1078

1079 **Figure 5:** Selected ammonoids from the McCarthy Formation at Grotto Creek, Alaska.
1080 Fossil horizons refer to stratigraphic position in Fig. 3; all specimens natural size unless
1081 indicated (e.g., X2). **A.** *Sagenites* sp. 1, fossil horizon 18, Cordilleranus Zone, late Norian.
1082 **B.** *Transipsiloceras* sp., fossil horizon 47, Polymorphum Zone, lower Hettangian. **C.**
1083 *Pleuroacanthites* cf. *biformis*, fossil horizon 50, Mulleri to Pleuroacanthitoides zones,

1084 middle Hettangian (X₂). **D**, *Rhacophyllites debilis*, fossil horizon fossil horizon 14,
1085 Columbianus to Crickmayi, late Norian-Rhaetian (X₂). **E**. ?*Psiloceras* sp., fossil horizon
1086 40, Spelae to Pacificum zones, lower Hettangian. **F**. *Psiloceras* cf. *tilmanni*, fossil
1087 horizon 41, Spelae to Pacificum zones, lower Hettangian (X₂). **G**. *Placites polydactylus*,
1088 fossil horizon fossil horizon 31, Amoenum Zone, Rhaetian. **H**. *Vandaite* cf. *suttonensis*,
1089 fossil horizon 27, Amoenum to Crickmayi zones, Rhaetian (moldic impression). **I**.
1090 *Psiloceras polymorphum*, fossil horizon 45, Polymorphum Zone, lower Hettangian.

1091

1092 **Figure 6:** Conodonts from the McCarthy Formation at Grotto Creek, Alaska. Fossil
1093 horizons refer to stratigraphic location in Fig. 3; Scale bar = 200 μm. **A-C**, *Misikella*
1094 *hernsteini*, fossil horizon 11, GSC Type No. 139577, from GSC cur. no. V-016700, late
1095 Norian. **D-F**. *Norigondolella steinbergensis*, fossil horizon 11, GSC Type No. 139578, from
1096 GSC cur. no. V-016700, late Norian. **G-I**. *Mockina englandi*, fossil horizon 32, GSC Type
1097 No. 139579, from GSC cur. no. V-016722, Rhaetian. **J-L**. *Mockina bidentata*, fossil
1098 horizon 34, GSC Type No. 139580, from GSC cur. no. V-016725, Rhaetian. **M-O**. *Mockina*
1099 *mosheri* morphotype B sensu Carter and Orchard, fossil horizon 32, GSC Type No.
1100 139581, from GSC cur. no. V-016722, Rhaetian. **P**. *Neohindeodella* sp., fossil horizon 39,
1101 GSC Type No. 139582, from GSC cur. no. V-016726, Hettangian.

1102

1103 **Figure 7: A.** Th-corrected single grain CA-ID-TIMS zircon data for sampled ash beds in
1104 the Grotto Creek section. Results shown as blue error ellipses are 2σ and provide the
1105 basis for age estimates. Data for older grains inferred to be antecrysts and/or

1106 xenocrysts are plotted as grey error ellipses. Two inherited grains (z21, z23) and a single
1107 low-precision analysis (z27) were excluded from sample Grot-1; as well as a relatively
1108 imprecise result (z18) from 2017GC3.8. Ages along concordia are in Ma, and gray bands
1109 (on concordia) show 2σ uncertainties based on decay-constant uncertainties of $^{238}\text{U} =$
1110 0.107% and $^{235}\text{U} = 0.136\%$ (Jaffey et al., 1971). Reported dates are weighted mean
1111 $^{206}\text{Pb}/^{238}\text{U}$ dates—uncertainties are reported as internal/internal+tracer
1112 calibration/internal+tracer calibration+decay constant uncertainties.
1113 Concordia uncertainties are too small to see for Grot-1. **B and C.** Age distribution data
1114 for all bentonite samples in the Grotto Creek section (**B**) is LA-ICPMS U-Pb data from
1115 2017GC3.8 and (**C**) is CA-ID-TIMS U-Pb data from all four bentonite samples. **B and C**
1116 show $^{206}\text{Pb}/^{238}\text{U}$ distributions that are in-line with crystals from a primary ash bed,
1117 rather than a volcanoclastic sandstone containing population(s) of significantly older
1118 zircon grains.

1119

1120 **Figure 8:** Composite carbon isotope data across the TJB interval from Panthalassa and
1121 northwestern Tethys oceans showing the broadly defined PCIE and NCIE intervals.
1122 Colored $\delta^{13}\text{C}_{\text{org}}$ data curve refers to locality in Figure 1A. Red hash marks denote
1123 position of sampled bentonites that provide new and previously established U-Pb age
1124 constraints. See Korte et al. (2019), Ruhl et al. (2020), and Du et al. (2020) for individual
1125 section citations. Rad = radiolarian, Bv = bivalve, Am = Ammonoid, Cordill. =
1126 Cordilleranus, Sp. = Spelae, Pac. = Pacificum, Ch. = *Choristoceras*, Psi. = *Psiloceras*, FAD

1127 = First Appearance Datum, LAD = Last Appearance Datum, Pol. = *Polymorphum*, and

1128 VPDB = Vienna PeeDee Belemnite.